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DIOPTRICAE
PARS PRIMA
CONTINENS
LIBRVM PRIMVM,
DE
EXPLICATIONE
PRINCIPIORVM,
EX QVIBVS
CONSTRVCTIO TAM TELESCOPIORVM
QVAM
MICROSCOPIORVM
EST PETENDA.

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ACAD. PETROP. PARISIN. ET LOND.

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**PETROPOLI**

Impensis Academiae Imperialis Scientiarum  
1769.









## INDEX CAPITVM

*In Tomo I. contentorum.*

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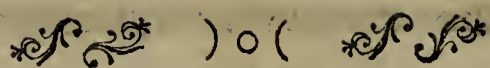
CAPVT I. De diffusione imaginis per vnicam lentem repraesentatae.

CAPVT II. De diffusione imaginis per plures lentes repraesentatae.

CAPVT III. De lentibus compositis seu multiplicatis.

CAPVT IV. De confusione visionis nec non de magnitudine apparente et claritate.





CAPVT V. De campo apparente oculique loco  
maxime idoneo.

CAPVT VI. De confusione a diuersa radiorum  
indole oriunda.

CAPVT VII. De constructione instrumentorum  
Dioptricorum in genere.





LIBER PRIMVS  
CONTINENS  
EXPLICATIONEM  
PRINCIPIORVM,  
EX QVIBVS CONSTRUCTIO  
TAM TELESCOPIORVM  
QVAM  
MICROSCOPIORVM  
EST PETENDA.



LIBER PRIMVS

CONTINENS

EXPLICATIONEM

PRINCIPIORVM

EX QVIBVS CONSTAT

QVAM TELLVRCOPVVM

AVT

INSTRVCTIONEM

AVT





## C A P V T I.

DE

### DIVERSIONE IMAGINIS

### PER VNICAM LENTEM

### REPRESENTATAE.

#### Definitio. I.

**I**mago principalis vocatur ea, quae a radiis axi lentis  
proximis per lentem refractis repraesentatur. Tab. I.  
Fig. I.

Scilicet si lentis PP axis sit EF in eaque tantum spatium minimum aAa sit apertum, per quod radiis transitus concedatur, radii a puncto lucido E emissi in vno puncto F colligentur, quod punctum imago principalis vocatur.

A 2

Coroll.



## Coroll. 1.

2. Cum nempe apertura  $aAa$  sit minima, omnes radii, qui a puncto quocunque in eam incidunt, ita aequabilem patiuntur refractionem, vt omnes iterum in vnum punctum colligantur, id quod non solum de puncto lucido  $E$  in axe lentis fito, sed etiam de quibusuis aliis extra axem positis est intelligendum.

## Coroll. 2.

3. Quodsi ergo lentis apertura fuerit minima, singula cuiusuis obiecti puncta post refractionem iterum singulis punctis referentur, sicque imago principalis erit distincta et confusione carebit: siquidem confusio tum demum oritur, quando radii ex vno puncto emissi non iterum in vno puncto colliguntur.

## Coroll. 3.

4. Et quamdiu apertura lentis  $aAa$  est minima, nihil interest, secundum quamnam figuram facies lentis sint elaboratae; quaecumque enim earum figura fuerit, quoniam tantum portiuncula minima in computum venit, ea semper vt sphaerica spectari poterit.

## Coroll. 4.

5. Pendet ergo locus imaginis principalis primum a loco puncti lucidi  $E$ , siue id in axe siue extra axem fuerit situm; deinde a sphaericitate  
vtrius-



vtriusque faciei  $aAa$  et  $bBb$  refringentis ; tertio ab earum distantia  $AB$  seu lentis crassitie, et quarto a ratione refractionis, quam radii in transitu per lentem patiuntur.

### Scholion I.

6. Refractio radiorum per huiusmodi lentium aperturas minimas transmissorum, ideoque determinatio imaginum principalium satis accurate in elementis Dioptricis tradi solet, quod igitur negotium hic fufius non prosequar; sed potius in eam refractionis rationem hic inquirere constitui, quando lentium apertura est modicae quantitatis, in quo imprimis ad vtriusque faciei figuram est spectandum. Hic autem perpetuo lentium figuras sphaericas assumo, propterea quod haec figura vulgo lentibus induci, vel saltem, nisi forte accurate successerit, intendi solet. In praxi certe nulla adhuc alia figura lentibus commode et accurate tribui poterit, atque adeo a sphaerica figura, etsi ad praxin maxime est accommodata, ab artificibus frequenter aberrari solet. A sollertioribus autem talia vitia iam plerumque satis feliciter evitantur, unde non adeo erit verendum, ne ea, quae per calculum ex hypothese sphaericae figurae elicientur, experientiae non consentanea sint futura. Hanc ob rem hic perpetuo postulo, vt ambae facies lentium exactissime secundum sphaericam figuram sint elaboratae.



## Scholion 2.

7. Sphaerica autem figura hoc laborat incommodo, quod statim ac lenti maior apertura tribuatur, non amplius omnes radii, qui quidem ab vno obiecti puncto sunt profecti, post refractionem ad vnum punctum dirigantur radiique EM longius ab axe lentis transmissi non amplius in puncto F concurrant. Vnde confusionem eo maiorem nasci necesse est, quo magis hi radii remotiores ab iis, qui prope axem transeunt, declinauerint; et quoniam talis declinatio a figura sphaerica originem ducit, eo maior euadet, quo maior lenti apertura tribuatur. Quanta igitur quouis casu futura sit haec confusio, hoc capite definire constitui; quae cum ceteris paribus a quantitate aperturae pendeat, hic in perpetuum moneo, me cuiusuis lentis aperturam circularem assumere, per cuius centrum axis lentis transeat: ita vt semidiameter huius circuli simul mensuram aperturae exhibeat. Ita si in superficie lentis PP spatium MAM apertum relinquatur, reliqua parte MP velamine opaco obducta, punctorum extremorum MM distantia diametrum aperturae eiusque semissis semidiametrum aperturae praebebit.

## Definitio 2.

8. *Imago extrema est ea, quam radii per extremitatem aperturae transmissi exhibent.*

Ita



Ita si  $MM$  sit apertura lentis, radiique  $EM$ ,  $EM$  a puncto lucido  $E$  circa oram aperturæ transmissi conueniant in puncto  $f$ , in hoc ipso puncto  $f$  erit imago extrema.

## C o r o l l. 1.

9. Si punctum lucidum  $E$  est in ipso axe lentis, nullum est dubium, quin radii inde per marginem circulem  $MM$  transeuntes iterum in vno axis puncto  $f$  concurrant, imaginemque distinctam repræsentent, quæ imago extrema vocatur.

## C o r o l l. 2.

10. Verum si punctum lucidum non esset in axe lentis, hoc neutiquam eueniet, radiique per marginem illum circulem transmissi non amplius in vno puncto colligentur; vnde hoc casu imago extrema eo magis erit confusa, quo magis punctum lucidum ab axe fuerit remotum.

## S c h o l i o n.

11. Quomodo se habeat refractione radiorum, quando punctum lucidum extra axem lentis fuerit constitutum, quæstio est non solum difficillima, sed etiam ita prolixis calculis inuoluitur, vt inde vix quicquam concludi possit. Ceterum in vsu, ad quem lentes accommodantur, nunquam objecta ab axe remotiora spectari solent, atque contentos esse nos oportet, dummodo objecta in ipso axe lentis sita distin-

te



cte repraesententur, neque etiam confusio, qua obiecta axi proxima afficiuntur, sensibilis esse potest; nam cum imago extrema puncti  $E$  in ipso lentis axe sit sit punctum  $f$ , nulla confusione inquinatum; etiamsi id parumper esset remotum ab axe, vix sensibilis confusio se immiscere poterit. Quam ob causam investigationes sequentes tantum ad obiecta in ipso lentis axe sita adstringam.

### Definitio 3.

12. *Spatium diffusionis vocatur interuallum, inter imaginem principalem et extremam interceptum.*

Ita si imago principalis sit in  $F$ , extrema vero in  $f$ , interuallum  $Ff$  appellatur spatium diffusionis.

### Coroll. 1.

13. Si ergo apertura lentis  $MM$  euanescat, spatium simul diffusionis euanescit, tum enim tantum radii axi proximi transmittuntur, quibus imago distincta in  $F$  effingitur. Ex quo intelligi licet, quo maior fuerit apertura lentis, eo maius fore spatium diffusionis  $Ff$ .

### Coroll. 2.

14. Cum in  $F$  imago a radiis axi proximis, in  $f$  autem imago a radiis circa marginem circula-rem  $MM$  transmissis formetur; si totam aperturam lentis infinitis circulis concentricis diuisam concipiamus,  
radii



radii per singulos circulos transmissi imagines intermedias exhibebunt, quibus interuallum *Ff* replebitur.

### COROLL. 3.

15. Si enim apertura primum nulla, tum vero continuo increfcens flatuatur, imago extrema primum cum principali congruet; tum vero continuo magis ab ea discedet, sicque cum vsque ad *MM* fuerit aucta, omnes illae imagines etiam nunc subsistent, spatiumque *Ff* implebunt.

### SCHOLIUM I.

16. Spatium hoc diffusionis causam continet confusionis, qua repraesentatio imaginis perturbatur; cum enim eiusdem puncti lucidi *E* infinitae imagines per interuallum *Ff* dispositae exhibeantur; earum commistio confusionem pariat necesse est, quae eo maior erit, quo maius fuerit spatium diffusionis *Ff*. Quemadmodum enim ad repraesentationem distinctam requiritur, vt omnes radii ex eodem obiecti puncto emissi iterum in vnico puncto colligantur; ita si hi radii in plura puncta coeant, pluresque eiusdem puncti imagines referant, prout hae magis minusue inter se discrepant, maior inde minorue confusio nascetur. Quemadmodum autem hanc confusionem aestimari oporteat, deinceps demum explicare licebit, cum ante spatium diffusionis accurate definire docuerimus: quo circa in hoc capite, cum proposita fuerit lens quaecunque vitrea faciebus sphaericis terminata



pro quavis puncti lucidi  $E$  ab ea distantia et quavis apertura spatium diffusionis  $Ff$  inuestigare constitui. Quo facto eandem inuestigationem pro duabus pluribusue lentibus inter se coniunctis suscipi conueniet, vt tandem inde confusionem in quibusuis instrumentis dioptricis assignare valeamus.

### Scholion 2.

17. Quaestio igitur principalis huius capituli in hoc versatur, vt proposita lente  $PP$  in eiusque axe puncto lucido  $E$ , radius quicunque incidens  $EM$  consideretur, eiusque per lentem refractione definiatur, vnde punctum  $f$ , vbi iterum in axem incidat, assignari queat. Cum enim in eodem puncto  $f$  omnes radii per totam peripheriam circularem  $MM$  transmissi concurrant; ab his imago quaeque puncti  $E$  in  $f$  exprimetur, quae erit extrema, si circulus  $MM$  in lente eius aperturam determinet; sin autem capiat minor, imago quaedam intermedia habebitur. Cum autem hic duplex refractione eueniat, altera in ingressu radii  $EM$  in vitrum, altera in egressu eiusdem e vitro, quemadmodum eius directio in vtraque inflectatur, seorsim est inuestigandum; vnde duo nascuntur problemata quasi praeliminaria, ex quorum combinatione deinceps negotium conficietur. Verum quo haec problemata commodius calculo expediri queant, quaedam lemmata ex doctrina angulorum petita praemittere oportet.

Lem-



## Lemma I.

18. Si angulus  $\Phi$  triginta gradus non excedat eius sinus satis accurate erit  $\sin. \Phi = \Phi - \frac{1}{6} \Phi^3$ , si quidem in circulo, cuius radius est  $= 1$ , arcus  $\Phi$ , qui pro illius anguli mensura habetur, in partibus radii exprimatur.

## Demonstratio.

Quantuscunque fuerit angulus  $\Phi$ , notum est, eius finum hac serie infinita exprimi:

$$\sin. \Phi = \Phi - \frac{1}{6} \Phi^3 + \frac{1}{120} \Phi^5 - \frac{1}{5040} \Phi^7 + \text{etc.}$$

sumtis igitur tantum binis primis terminis, error committitur reliquis neglectis aequalis, si ergo statuamus  $\sin. \Phi = \Phi - \frac{1}{6} \Phi^3$ , hincque pro variis angulis  $\Phi$  finus colligamus, eorum comparatio cum tabulis finuum errores manifestabit. Ita si sumatur  $\Phi = 30^\circ$ , quia

arcus  $180^\circ$  valet 3, 14159265, erit in partibus radii

$\Phi = 0,52359877$ , et  $\frac{1}{6} \Phi^3 = 0,0239246$ , hincque

$$\Phi - \frac{1}{6} \Phi^3 = 0,4996741. \text{ at est reuera}$$

$$\sin. \Phi = 0,5000000 \text{ vnde habetur}$$

error  $= 0,0003259$ , qui ergo ne ad  $\frac{1}{3000}$  partem radii quidem affurgit. At si angulus  $\Phi$  caperetur duplo minor, scilicet  $\Phi = 15^\circ$ , reperiretur

$$\Phi - \frac{1}{6} \Phi^3 = 0,2588088$$

$$\text{cum tamen sit } \sin. \Phi = 0,2588190$$

$$\text{errore existente } = 0,0000102$$

qui tricies bis minor est quam casu praecedente. Cum ergo in praxi error partem adeo termillesimam radii adaequans facile tolerari possit, multo magis, si an-



gulus  $\Phi$  fuerit  $30^\circ$  minor, expressio  $\Phi - \frac{1}{6}\Phi^3$  verum eius sinum exhibere censenda erit.

### Lemma 2.

19. Vicissim si anguli triginta gradibus minoris detur sinus  $= s$ , ex eo ipse angulus ita proxime definitur, ut sit in circulo, cuius radius  $= 1$ , arcus eum metiens  $= s + \frac{1}{6}s^3$ .

### Demonstratio.

Si enim  $\Phi$  designet istum angulum, cuius sinus proponitur  $= s$ ; modo vidimus, esse satis exacte  $s = \Phi - \frac{1}{6}\Phi^3$ , hinc autem per conuersionem oritur proxime  $\Phi = s + \frac{1}{6}s^3$ , quae expressio quantum peccet in angulo triginta graduum, ut videamus, fit  $s = \frac{1}{2}$  eritque  $s + \frac{1}{6}s^3 = 0,5208333$ . Cum autem numerus 3, 14159265 respondeat angulo  $180^\circ$ , hic numerus 0, 5208333 praebet angulum  $29^\circ 50' 30''$ ; verus autem angulus est  $30^\circ$ , ita ut error sit  $9' 30''$ . Sit autem  $s = \frac{1}{4}$  cui sinui respondet angulus  $= 14^\circ 28' 39''$ , erit  $s + \frac{1}{6}s^3 = 0,2526041 = 14^\circ 28' 23''$ , ita ut hoc casu error non excedat  $16''$ . Patet ergo dum angulus minor sit  $30^\circ$ , hoc modo satis exacte ex dato sinu reperiri angulum.

### Problema I.

Tab. I. 20. Si ex puncto lucido E in superficiem sphaer-  
Fig. 2. ricam reuolutione arcus circularis AM circa axem EC  
genitam incidat radius EM, definire eius, postquam  
fuerit refractus, concursum cum axe EC.

Solutio.



## Solutio.

Sit  $C$  centrum superficiei sphaericae, eiusque radius  $CA = CM = f$ , et distantia puncti lucidi  $E$  a superficie refringente  $EA = a$ , tum vero sit ratio refractionis radiorum ex aere in vitrum  $= n:1$ . Vocetur pro radio incidente  $CM$  angulus  $ACM = \Phi$ , eritque huius anguli sinus  $= \Phi - \frac{1}{6}\Phi^3$ . Ponatur iam breuitatis gratia distantia  $EC = a + f = c$ , et quia in triangulo  $ECM$  dantur latera  $CM = f$ ,  $EC = c$  cum angulo  $CEM = \Phi$ , fiet producta  $CM$  in  $c$

$CM (f) : \sin. CEM = CE (c) : \sin. EMc$   
ideoque

$$\sin. EMc = \frac{c}{f} \sin. \Phi = \frac{c}{f} \left( \Phi - \frac{1}{6}\Phi^3 \right)$$

vnde ipse angulus  $EMc$  elicitur, neglectis tertia altioribus potestatibus ipsius  $\Phi$ :

$$EMc = \frac{c}{f} \Phi - \frac{c}{6f} \Phi^3 + \frac{c^3}{6f^3} \Phi^3 = \frac{c}{f} \Phi + \frac{c(cc - ff)}{6f^3} \Phi^3$$

qui cum sit  $= ECM + CEM$ , erit

$$ECM = \frac{c-f}{f} \Phi + \frac{c(cc - ff)}{6f^3} \Phi^3$$

Verum angulus  $EMc$  est angulus incidentiae, ac si  $MO$  referat radium refractum, erit  $CMO$  angulus refractionis, sicque per hypothesein

$$\sin. EMc : \sin. CMO = n:1$$

vnde colligitur

$$\sin. CMO = \frac{c}{nf} \left( \Phi - \frac{1}{6}\Phi^3 \right)$$

hincque ipse angulus

$$CMO = \frac{c}{nf} \Phi + \frac{c(cc - nnff)}{6n^3f^3} \Phi^3$$



quo ablato ab angulo ECM relinquitur angulus

$$\text{COM} = \frac{(n-1)c-nf}{nf} \Phi + \frac{c((n^3-1)cc-nn(n-1)ff)}{6n^3j^3} \Phi^3$$

cuius finus propterea erit

$$\text{fin. COM} = \frac{(n-1)c-nf}{nf} \Phi + \frac{3(n-1)c^3 + 3(n-1)^2ccf - 4n(n-1)cff + nnf^3}{6nnj^3} \Phi^3$$

Iam uero ob

fin. COM : CM (f) = fin. CMO : CO, obtinebitur

$$\text{CO} = \frac{\frac{c}{n} - \frac{c}{6n} \Phi \Phi}{\frac{(n-1)c-nf}{nf} + \frac{3(n-1)c^3 + 3(n-1)^2ccf - 4n(n-1)cff + nnf^3}{6nnj^3}} \Phi^2$$

cuius formulae euolutio praebet :

$$\text{CO} = \frac{cf}{(n-1)c-nf} - \frac{cf \Phi \Phi}{6((n-1)c-nf)} - \frac{c(3(n-1)c^3 + 3(n-1)^2ccf - 4n(n-1)cff + nnf^3)}{6nj((n-1)c-nf)^2} \Phi \Phi$$

quae porro reducitur ad hanc formam :

$$\text{CO} = \frac{cf}{(n-1)c-af} - \frac{(n-1)cc(c-f)(c+nf)}{2nj((n-1)c-nf)^2} \Phi \Phi$$

cui si addamus CA = f prodibit

$$\text{AO} = \frac{nf(c-f)}{(n-1)c-nf} - \frac{(n-1)cc(c-f)(c+nf)}{2nj((n-1)c-nf)^2} \Phi \Phi$$

atque hinc positio radii refracti MO primo per interuallum AO modo inuentum definitur, tum vero insuper ex angulo AOM, qui erit :

$$\text{AOM} = \frac{(n-1)c-nf}{nf} \Phi + \frac{c((n^3-1)cc-nn(n-1)ff)}{6n^3j^3} \Phi^3$$

siue

$$\text{AOM} = \frac{(n-1)c-nf}{nf} \Phi + \frac{(n-1)c(nn+n-1)cc-nnf^3}{6n^3j^3} \Phi^3$$

### Coroll. I.

21. Cum posuerimus  $c=a+f$  erit

$c-f=a$ , et  $(n-1)c-nf=(n-1)a-f$ ,

atque



atque

$(nn+n+1)cc-nmff=(nn+n+1)aa+2(nn+n+1)af+(n+1)ff$   
quibus valoribus substitutis habebimus

$$AO = \frac{naf}{(n-1)a-f} - \frac{(n-1)a(a+f)^2(a+(n+1)f)}{2nf((n-1)a-f)^2} \Phi \Phi \text{ et}$$

$$AOM = \frac{(n-1)a-f}{nfa} \Phi + \frac{(n-1)(a+f)(nn+n+1)(a+2f)+(n+1)ff}{6n^3f^3} \Phi^3$$

### COROLL. 2.

22. Si M sit in extremitate aperturæ, erit  
semidiameter aperturæ = f sin. ECM = aΦ satis exacte:  
neque enim necesse est, aperturam tam exacte nosse.  
Vnde si semidiameter aperturæ ponatur = x, erit  
prope x = aΦ, ideoque  $\Phi = \frac{x}{a}$ .

### COROLL. 3.

23. Si accuratius rem definire velimus, quia est

$$\text{sin. ECM} = \frac{c-f}{f} \Phi + \frac{(c-f)(3c-f)}{6Jf} \Phi^3 = \frac{a}{f} \Phi + \frac{a(3a+2f)}{6Jf} \Phi^3$$

$$\text{erit } x = a\Phi + \frac{a(3a+2f)}{6J} \Phi^3$$

$$\text{ideoque } \Phi = \frac{x}{a} - \frac{(3a+2f)x^3}{6a^3f}$$

sed quia in valore ipsius AO non ultra secundam dimensionem ipsius Φ ascendimus, hac expressione non indigemus.

### COROLL. 4.

24. Etsi hae expressiones tantum sunt prope  
verae; tamen in praxi sine errore adhiberi poterunt,  
dummodo anguli, qui in calculum sunt ingressi, infra  
30°



30° subsistant. Non solum ergo necesse est, vt angulus  $AE M = \Phi$  sed etiam angulus  $EMc$  seu  $\frac{c}{f}\Phi = \Phi + \frac{a}{f}\Phi$  minor sit 30 gradibus.

### Scholion.

25. Summo quidem rigore geometrico distantiam  $EO$  definire potuissimus, neque opus fuisset ad approximationes confugere: scilicet si posuissimus angulum  $ECM = \omega$ , et distantiam  $CO = u$ ; deuenissemus ad hanc determinationem.

$\frac{cf}{u} = -c \cos. \omega + \sqrt{(nn(c \cos. \omega - f)^2 + (nn - 1)cc \sin. \omega^2)}$   
foretque

$$\cos. \omega = \frac{c}{f} \sin. \Phi^2 + \cos. \Phi \sqrt{(1 - \frac{cc}{ff} \sin. \Phi^2)} \text{ et}$$

$$\sin. \omega = \frac{c}{f} \sin. \Phi \cos. \Phi - \sin. \Phi \sqrt{(1 - \frac{cc}{ff} \sin. \Phi^2)}$$

hisque valoribus substitutis

$$\frac{cf}{u} = \frac{-cc \sin. \Phi^2}{f} - c \cos. \Phi \sqrt{(1 - \frac{cc}{ff} \sin. \Phi^2)} \\ + (c \cos. \Phi - \sqrt{(ff - cc \sin. \Phi^2)}) \sqrt{(nn - \frac{cc}{ff} \sin. \Phi^2)}$$

vnde ponendo angulo  $\Phi$  valde paruo per approximationem superior expressio eliceretur. Verum praece-dens analysis magis ad praesens institutum videtur accommodata. Interim si quis propius ad veritatem accedere voluerit, ex formula vera hic exhibita adipiscetur.

$$\frac{cf}{u} = (n-1)c \cos. \Phi - nf + \frac{(n-1)cc}{2nff} ((n-1)f + c \cos. \Phi) \sin. \Phi^2 \\ + \frac{c^4}{8n^3f^4} ((nn-1)^2 f + (n^3-1)c \cos. \Phi) \sin. \Phi^4 \\ \text{seu}$$

feu

$$\frac{cf}{n} = (n-1)c - nf + \frac{(n-1)c(c-1)(c+nf)}{2nff} \Phi\Phi$$

$$+ \frac{(n-1)(c-1)(nn+1+c^2+2nn(n+1)cf+nn(n-1)cf^2-2^2f^3)}{24n^3f^4} \Phi^3.$$

## Problema 2.

26. Si prout in problemate praecedente inuenimus, radius MO in vitrum missus per superficiem sphaericam BN iterum in aerem erumpat, definire punctum V, vbi is cum axe concurret.

## Solutio.

Ponatur interuallum  $AB=d$ ; sitque superficiei sphaericae BN centrum in D eiusque radius  $DB=DN=g$ . Quoniam igitur positio radii incidentis MNO datur, ponamus  $BO=b$ , et angulum  $BON=\psi$ , et brevitatis ergo interuallum  $DO=b+g=e$ . Cum nunc in triangulo DON dentur latera  $DN=g$ ,  $DO=e$  cum angulo  $BON=\psi$ , reperitur

$$\sin. DNM = \frac{e}{g} \sin. \psi = \frac{e}{g} \psi - \frac{e}{6g} \psi^3, \text{ hincque}$$

$$DNM = \frac{e}{g} \psi + \frac{e(ee-gg)}{6g^3} \psi^3 \text{ et}$$

$$ODN = \frac{e-g}{g} \psi + \frac{e(ee-gg)}{6g^3} \psi^3$$

Sed hic  $\sin. DNM$  est sinus incidentiae, cui respondet angulus refractionis  $VNd$ , cuius sinus propterea est ad illum vt  $n:1$ , vnde fit

$$\sin. VNd = \frac{ne}{g} \psi - \frac{ne}{6g} \psi^3$$

Tom. I.

C

hinc-



hincque

$$VNd = \frac{ne}{g} \psi + \frac{ne(nne - gg)}{6g^3} \psi^3$$

a quo ablato angulo ODN relinquitur angulus

$$DVN = \frac{(n-1)e+g}{g} \psi + \frac{e((n^3-1)ee - (n-1)gg)}{6g^3} \psi^3$$

ergo

$$\text{fin. DVN} = \frac{(n-1)e+g}{g} \psi + \frac{3n(n-1)e^3 - 3(n-1)^2 eeg - 4(n-1)egg - g^3}{6g^3} \psi^3$$

Cum nunc fit

$$\text{fin. DVN} : g = \text{fin. VNd} : DV \text{ erit}$$

$$\frac{g}{DV} = \frac{(n-1)e+g + \frac{1}{6gg}(3n(n-1)e^3 - 3(n-1)^2 eeg - 4(n-1)egg - g^3) \psi^2}{ne - \frac{1}{6} ne \psi^2}$$

quae expressio reducitur ad hanc formam :

$$\frac{g}{DV} = \frac{(n-1)e+g}{ne} + \frac{(n-1)(e-g)(ne+g) \psi^2}{2ngg}$$

vnde reciproce oritur

$$DV = \frac{neg}{(n-1)e+g} - \frac{n(n-1)ee(e-g)(ne+g)}{2g((n-1)e+g)^2} \psi^2$$

et

$$BV = \frac{g(e-g)}{(n-1)e+g} - \frac{n(n-1)ee(e-g)(ne+g)}{2g((n-1)e+g)^2} \psi^2$$

Puncto autem V inuento notandum est esse angulum

$$BVN = \frac{(n-1)e+g}{g} \psi + \frac{e((n^3-1)ee - (n-1)gg)}{6g^3} \psi^3$$

### Coroll. I.

27. Cum fit

$$e = b + g \text{ erit } (n-1)e + g = (n-1)b + ng, \\ \text{quo}$$

quo valore restituto habebimus.

$$BV = \frac{bg}{(n-1)b+ng} - \frac{n(n-1)b(b+g)^2(nb+(n+1)g)}{2g((n-1)b+ng)^2} \psi^2 \text{ et}$$

$$BVN = \frac{(n-1)b+ng}{g} \psi + \frac{(b+g)((n^3-1)(b+g)^2-(n-1)gg)}{6g^3} \psi^3$$

### Coroll. 2.

28. Hae formulae etiam ex praecedentibus erui possunt, si pro litteris  $n, a, f$  scribantur  $\frac{1}{n}, -b$  et  $-g$ , quoniam hoc modo casus praecedentis problematis ad hunc reducitur. Praeterea autem, qui angulus ibi erat  $\Phi$  hic est  $\psi$ .

### Coroll. 3.

29. Quia in praecedente problemate inuenimus tam lineam AO quam angulum AOM erit hoc problemate ad radium illum refractum MO accommodato

$$BO = b = \frac{naf}{(n-1)a-f} - d - \frac{(n-1)a(a+f)^2(a+(n+1)f)}{2nf((n-1)a-f)^2} \Phi \Phi$$

et

$$\psi = \frac{(n-1)a-f}{nf} \Phi + \frac{(n-1)(a+f)((nn+n+1)a(a+2f)+(n+1)ff)}{6n^3f^3} \Phi^3$$

### Problema 3.

30. Proposita lente vitrea MABN faciebus Tab. I. sphaericis AM et BN terminata, si a puncto quo- Fig. 2. cunque E in eius axe posito incidat in eam radius EM, definire punctum V, in quo is post geminam refractionem iterum cum axe lentis fit concursurus.

C 2

Solutio.



## Solutio.

Consideremus lentem vt vtrunque conuexam, sitque faciei anterioris AM radius  $AC = f$ , posterioris vero BN radius  $BD = g$ ; ipsius lentis autem crassities  $AB = d$ . Lens porro sit vitrea, ita vt si in eam radius lucis ex aere incidat, sit sinus incidentiae ad sinum refractionis vt  $n$  ad 1. Iam recta iungens centra vtriusque faciei C et D erit axis lentis, in quo reperiatur punctum lucidum E ante lentem in distantia  $AE = a$ , vnde sub angulo  $AEM = \Phi$  in lentem incidat radius EM, qui prima refractione ita inflectatur, vt productus cum axe concurrat in O. Quodsi iam ex iis, quae problemate primo sunt inuenta, ponamus

$$BO = \frac{naf}{(n-1)a-f} - d - \frac{(n-1)a(a+f)^2(a+(n+1)f)}{2nf((n-1)a-f)^2} \Phi \Phi = b$$

$$\text{et ang. BON} = \frac{(n-1)a-f}{nf} \Phi = \psi$$

in valore enim anguli  $\psi$  negligere licet terminum  $\Phi^3$  inuoluentem, quoniam calculum tantum ad secundam potestatem ipsius  $\psi$  extendimus; his positis in problemate secundo inuenimus fore.

$$BV = \frac{bg}{(n-1)b+ng} - \frac{n(n-1)b(b+g)^2(nb+(n+1)g)}{2g((n-1)b+ng)^2} \psi \psi$$

et

$$BVN = \frac{(n-1)b+ng}{g} \psi.$$

Totum ergo negotium huc redit, vt isthic pro  $b$  et  $\psi$  valo-

valores assignatos substituamus, quod quo facilius fieri possit, statuamus

$$b = P - Q\Phi\Phi \text{ et } \psi = R\Phi \text{ vt fit}$$

$$P = \frac{naf}{(n-1)a-f} - d = \frac{naf - (n-1)ad + df}{(n-1)a-f}$$

$$Q = \frac{(n-1)a(a+f)^2(a+(n+1)f)}{2nf((n-1)a-f)^2} \text{ et } R = \frac{(n-1)a-f}{nf}$$

Hinc erit:

$$\frac{bg}{(n-1)+ng} = \frac{Pg - Qg\Phi\Phi}{(n-1)P + ng - (n-1)Q\Phi\Phi} = \frac{Pg}{(n-1)+ng} - \frac{nQgg\Phi\Phi}{((n-1)P + ng)^2}$$

at in altero membro sufficit pro  $b$  scribere  $P$ : ex quo obtinebimus

$$BV = \frac{Pg}{(n-1)+ng} - \frac{nQgg\Phi\Phi}{((n-1)P + ng)^2} - \frac{n(n-1)PRR(P+g)^2(nP+(n+1)g)}{2g((n-1)P + ng)^2} \Phi\Phi$$

$$\text{et } BVN = \frac{(n-1)P + ng}{g} R\Phi$$

Pro his autem substitutionibus notandum est fore:

$$(n-1)P + ng = \frac{n(n-1)a(f+g) - nfg - (n-1)^2ad + (n-1)df}{(n-1)a-f}$$

$$P + g = \frac{naf + (n-1)ag - fg - (n-1)ad + df}{(n-1)a-f}$$

$$nP + (n+1)g = \frac{nna + (nn-1)ag - (n+1)fg - 1(n-1)ad + ndf}{(n-1)a-f}$$

unde concluditur:

$$BV = \frac{nafg - (n-1)adg + dfg}{n(n-1)a(f+g) - nfg - (n-1)^2ad + (n-1)df}$$

$$- \frac{(n-1)agg(a+f)^2(a+(n+1)f)}{2f(n(n-1)a(f+g) - nfg - (n-1)^2ad + (n-1)df)^2} \Phi\Phi$$

$$- \frac{(n-1)(naf - (n-1)ad + df)(naf + (n-1)ag - fg - (n-1)ad + df)^2(nna + (nn-1)ag - (n+1)fg - n(n-1)ad + ndf)}{2nffg(n(n-1)a(f+g) - nfg - (n-1)^2ad + (n-1)df)^2} \Phi\Phi$$

$$\text{et } BVN = \frac{n(n-1)a(f+g) - nfg + (n-1)df - (n-1)^2ad}{nfg} \Phi$$



## Coroll. II.

31. Si angulus  $\Phi$  prorsus euanescat, punctum V cadet in imaginem principalem, cuius si a lente distantia dicatur  $= a$  erit

$$a = \frac{nafg - (n-1)adg + df g}{n(n-1)a(f+g) - nfg - (n-1)^2ad + (n-1)df}$$

quae si vt data spectetur, hac aequatione relatio inter  $f$  et  $g$  definitur, vt haec distantia principalis oriatur.

## Coroll. 2.

32. Quare si distantia obiecti ante lentem fit  $= a$ , eiusque imago principalis post lentem ad distantiam  $= a$  proici debeat huic aequationi satisfieri oportet.

$$n(n-1)aa(f+g) - nafg + (n-1)adg - (n-1)^2aad = 0 \\ -nafg + (n-1)adf - df g$$

## Coroll. 3.

33. Hanc autem imaginis principalis distantiam  $a$  in calculum introducendo expressiones nostrae inuentae haud mediocriter contrahentur. Cum enim sit  $a = \frac{Pg}{(n-1)P + ng}$ , erit hinc  $P = \frac{n\alpha g}{g - (n-1)\alpha}$

et porro

$$(n-1)P + ng = \frac{ngg}{g - (n-1)\alpha}; P + g = \frac{gg + \alpha g}{g - (n-1)\alpha} = \frac{g(\alpha + g)}{g - (n-1)\alpha} \\ nP + (n+1)g = \frac{g(\alpha + (n+1)g)}{g - (n-1)\alpha}$$

vnde

vnde facta substitutione sequentes determinaciones simpliciores assequemur.

$$BV = \alpha - \frac{(n-1)a(a+f)^2(g-(n-1)\alpha)^2(a+(n-1)f)}{2nfgg((n-1)a-f)^2} \Phi \Phi$$

$$- \frac{(n-1)\alpha(\alpha+g)^2((n-1)a-f)^2(\alpha+(n-1)g)}{2nfgg(g-(n-1)\alpha)^2} \Phi \Phi$$

et

$$BVN = \frac{((n-1)a-f)g}{(g-(n-1)\alpha)f} \Phi.$$

### Scholion.

34. Non solum hae formulae multo sunt breui-  
ores et concinniores, quam primo inuentae, sed etiam  
perspicuus ordo in iis obseruatur quo litterae  $\alpha$  et  $g$   
cum litteris  $a$  et  $f$  ita connexae sunt, vt permutatio-  
nem admittant. Nullum igitur est dubium, quin  
si statim distantiam  $\alpha$  in calculum introduxissemus,  
via breuiori ad eas peruenire licuisset. Ceterum quia  
hae formulae posteriores non amplius crassitiem len-  
tis  $AB = d$  inuoluunt, euidens est eas ex primo in-  
uentis nasci, si introducta distantia  $\alpha$  crassities  $d$  elimi-  
netur, seu eius loco hic valor surrogetur.

$$d = \frac{n(n-1)a\alpha(f+g) - n(a+\alpha)fg}{(n-1)^2 a\alpha - (n-1)(ag + \alpha f) + fg} = n \frac{(n-1)a\alpha(f+g) - (a+\alpha)fg}{((n-1)a-f)((n-1)\alpha-g)}$$

qui labor autem ne suscipi quidem mereretur, nisi  
iam ante de eximio eius vsu certiores essemus facti.  
In sequentibus igitur his elegantioribus formulis uta-  
mur, quoad negotium adhuc succinctius expedire di-  
dicerimus.

Proble.



## Problema 4.

35. Proposita lente quacunque faciebus sphaericis terminata, si obiectum in data ab ea distantia sit constitutum, pro data lentis apertura spatium diffusionis assignare.

## Solutio.

Tab. I.

Fig. I.

Concipiamus lentem vtrinque conuexam, sitque faciei anterioris  $MAM$  radius  $=f$ , posterioris  $NBN=g$ , lentisque crassities  $AB=d$ . Sit porro  $MM$  lentis huius apertura, cuius semidiameter sit  $=x$ ; atque in axe lentis expositum sit obiectum vel saltem punctum lucidum  $E$ , cuius a lente ponatur distantia  $AE=a$ . Iam primo quaeramus eius imaginem principalem, quae cadat in  $F$  atque supra (31) inuenimus fore

$$BF = \frac{nafg - (n-1)adg + dfg}{n(n-1)a(f+g) - nfg - (n-1)^2ad + (n-1)df}$$

Vocetur ergo haec distantia  $=\alpha$ , et si radii  $EM$  circa oram aperturae transeant, erit angulus  $AEM = \Phi = \frac{x}{a}$ , quo valore in superioribus formulis substituto prodibit distantia imaginis extremae  $f$  a lente scilicet:

$$Bf = \alpha - \frac{(n-1)a(a+f)^2(g-(n-1)\alpha)^2(a+(n+1)f)}{2nnfgg((n-1)a-f)^2} \cdot \frac{xx}{a^2}$$

$$- \frac{(n-1)\alpha(\alpha+g)^2((n-1)a-f)^2(\alpha+(n+1)g)}{2nnffg(g-(n-1)\alpha)^2} \cdot \frac{xx}{a^2}$$

ex quo colligitur spatium diffusionis quaesitum:

$$Ef = + \frac{(n-1)a(a+f)^2(g-(n-1)\alpha)^2(a+(n+1)f)}{2nnfgg((n-1)a-f)^2} \cdot \frac{xx}{a^2}$$

$$+ \frac{(n-1)\alpha(\alpha+g)^2((n-1)a-f)^2(\alpha+(n+1)g)}{2nnffg(g-(n-1)\alpha)^2} \cdot \frac{xx}{a^2}$$

ac

ac praeterea angulus  $BfN$  erit

$$BfN = \frac{((n-1)a-f)g}{(g-(n-1)\alpha)f} \cdot \frac{x}{a}.$$

### Coroll. I.

36. Quoties ergo spatium diffusionis hoc modo expressum est positium, imago extrema propius ad lentem cadit quam principalis seu est  $Bf < BF$ . Contra autem si ista expressio valorem obtineat negatiuum, imago extrema a lente longius erit remota principali.

### Coroll. 2.

37. Patet hinc etiam spatium diffusionis cum apertura ita crescere, vt sit quadrato semidiametri aperturae proportionale, sequetur ergo ipsam aperturae rationem.

### Coroll. 3.

38. Spatium diffusionis etiam hoc modo exprimi potest

$$Ff = \left\{ \begin{array}{l} \frac{+(n-1)a(1+\frac{a}{f})^2(1-\frac{(n-1)\alpha}{g})^2(n+1+\frac{a}{f})}{2nn(\frac{(n-1)a}{f}-1)^2} \cdot \frac{xx}{aa} \\ \frac{+(n-1)\alpha(1+\frac{\alpha}{g})^2(\frac{(n-1)a}{f}-1)^2(n+1+\frac{\alpha}{g})}{2nn(1-\frac{(n-1)\alpha}{g})^2} \cdot \frac{xx}{aa} \end{array} \right.$$

et angulus

$$BfN = \frac{\frac{(n-1)a-f}{f}}{1-\frac{(n-1)\alpha}{g}} \cdot \frac{x}{a}.$$

Tom. I.

D

Coroll.



## Coroll. 4.

39. Quia hae formulae introducendis litterarum valoribus reciprocis redditae sunt simpliciores, in hunc modum etiam aequatio (§. 32) exhibita tractetur quae per  $aadf g$  diuisa abit in hanc formam.

$$n(n-1)\frac{1}{d}\left(\frac{1}{f}+\frac{1}{g}\right)-n\frac{1}{d}\left(\frac{1}{a}+\frac{1}{\alpha}\right)+(n-1)\left(\frac{1}{af}+\frac{1}{ag}\right)-(n-1)^2\frac{1}{fg}-\frac{1}{a\alpha}=0$$

seu

$$\frac{n}{d}\left((n-1)a\alpha\left(\frac{1}{f}+\frac{1}{g}\right)-a-\alpha\right)=\left(\frac{(n-1)a}{f}-1\right)\left(\frac{(n-1)\alpha}{g}-1\right)$$

quae commodior erit tam ad relationem inter  $a$  et  $\alpha$ , quam inter  $f$  et  $g$  definiendam.

## Scholion.

40. Scilicet si proposita lente obiectum in variis distantiiis exponatur ac pro singulis distantiam imaginis principalis definire velimus, aequatio hoc modo referatur:

$$\frac{1}{a\alpha}-(n-1)\left(\frac{1}{af}+\frac{1}{ag}\right)+n\left(\frac{1}{da}+\frac{1}{d\alpha}\right)-n(n-1)\left(\frac{1}{df}+\frac{1}{dg}\right)+(n-1)^2\frac{1}{fg}=0$$

quae per factores ita adornari poterit

$$\left(\frac{1}{a}-\frac{(n-1)}{f}+\frac{n}{d}\right)\left(\frac{1}{\alpha}-\frac{(n-1)}{g}+\frac{n}{d}\right)=\frac{nn}{dd}$$

Vnde patet productum ex his duobus factoribus semper esse idem. Cuius resolutio quo facilius instituat, capiantur numeri  $\mu$  et  $\nu$  ita, vt eorum summa sit  $= 1$ , scilicet  $\mu + \nu = 1$  ac statuatur

$$\frac{1}{a}-\frac{(n-1)}{f}+\frac{n}{d}=\frac{\mu n}{vd} \text{ erit } \frac{1}{a}=\frac{n-1}{f}-\frac{n}{\nu d}$$

$$\frac{1}{\alpha}-\frac{(n-1)}{g}+\frac{n}{d}=\frac{\nu n}{\mu d} \text{ erit } \frac{1}{\alpha}=\frac{n-1}{g}-\frac{n}{\mu d}$$

Quare

Quare si distantia obiecti  $AE = a$  ita capiatur, ut sit  $\frac{1}{a} = \frac{n-1}{f} - \frac{n}{v d}$ , distantia imaginis principalis  $BF = \alpha$  ita se habebit, ut sit  $\frac{1}{\alpha} = \frac{n-1}{g} - \frac{n}{\mu d}$ .

Eodem modo si dentur binae distantiae  $AE = a$  et  $BF = \alpha$  cum crassitie lentis  $AB = d$ , radii facierum  $f$  et  $g$  ita debent esse comparati, ut sit

$$\frac{1}{f} = \frac{1}{(n-1)a} + \frac{n}{v(n-1)d} \quad \text{et} \quad \frac{1}{g} = \frac{1}{(n-1)\alpha} + \frac{n}{\mu(n-1)d}$$

id quod infinitis modis praestari potest, cum numeri  $\mu$  et  $v$  pro arbitrio accipi queant, dummodo eorum summa  $\mu + v$  aequetur unitati. Si hoc modo etiam in formulis pro spatio diffusionis inuentis omnes litterae in denominatores detrudantur, reperietur:

$$Ff = \frac{(n-1)\alpha\alpha x x}{2nn} \left\{ \frac{\left(\frac{1}{a} + \frac{1}{f}\right)^2 \left(\frac{1}{\alpha} - \frac{(n-1)}{g}\right)^2 \left(\frac{n+1}{a} + \frac{1}{f}\right)}{\left(\frac{n-1}{f} - \frac{1}{a}\right)^2} + \frac{\left(\frac{1}{\alpha} + \frac{1}{g}\right)^2 \left(\frac{n-1}{f} - \frac{1}{a}\right)^2 \left(\frac{n+1}{\alpha} + \frac{1}{g}\right)}{\left(\frac{1}{\alpha} - \frac{(n-1)}{g}\right)^2} \right\}$$

et

$$\text{angulus } BfN = \frac{\frac{n-1}{f} - \frac{1}{a}}{\frac{1}{\alpha} - \frac{(n-1)}{g}} \cdot \frac{x}{\alpha} = \frac{\frac{n}{v d} x}{\frac{n}{\mu d} \alpha} = -\frac{\mu}{v} \cdot \frac{x}{\alpha}$$

Cum autem sit

$$\frac{n-1}{f} - \frac{1}{a} = \frac{n}{v d} \quad \text{et} \quad \frac{1}{\alpha} - \frac{(n-1)}{g} = \frac{n}{\mu d} \quad \text{erit}$$

$$Ff = \frac{(n-1)\alpha\alpha x x}{2nn} \left( \frac{v v}{\mu \mu} \left(\frac{1}{a} + \frac{1}{f}\right)^2 \left(\frac{n+1}{a} + \frac{1}{f}\right) + \frac{\mu \mu}{v v} \left(\frac{1}{\alpha} + \frac{1}{g}\right)^2 \left(\frac{n+1}{\alpha} + \frac{1}{g}\right) \right)$$

et

$$\text{angulus } BfN = -\frac{\mu x}{v \alpha} = \frac{\mu x}{(\mu-1)\alpha}$$

D 2

Proble-



## Problema 5.

Tab. I.

Fig. I.

41. Datis distantiiis obiecti ante lentem  $AE = a$  et imaginis principalis post lentem  $BF = \alpha$  vna cum lentis crassitie  $AB = d$ , definire omnes lentes satisfaci-  
entes, simulque pro singulis spatium diffusionis  $Ff$ .

## Solutio.

Modo vidimus si lentis faciei anterioris  $AM$  radius ponatur  $= f$ , posterioris  $BN = g$ , lente vt conuexa vtrunque spectata, hos duos radios ita comparatos esse debere, vt sit

$$\frac{n-1}{f} = \frac{1}{a} + \frac{n}{vd} \text{ et } \frac{n-1}{g} = \frac{1}{\alpha} + \frac{n}{\mu d}$$

sumtis pro  $\mu$  et  $\nu$  numeris quibuscunque, vt sit  $\mu + \nu = 1$ ; vnde infinitae lentes quaesito satisfaci-  
entes obtinentur. Deinde si huius lentis semidiameter aperturæ ponatur  $= x$ , spatium diffusionis  $Ff$  ita exprimi potest vt sit

$$Ff = \frac{\alpha \alpha x x}{2n(n-1)^2} \left( \frac{\nu \nu}{\mu \mu} \left( \frac{n-1}{a} + \frac{n-1}{f} \right)^2 \left( \frac{n-1}{a} + \frac{n-1}{f} \right) + \frac{\mu \mu}{\nu \nu} \left( \frac{n-1}{\alpha} + \frac{n-1}{g} \right)^2 \left( \frac{n-1}{\alpha} + \frac{n-1}{g} \right) \right)$$

et angulus  $BfN = \frac{\mu x}{(\mu-1)\alpha}$ .

Quod si iam hic pro  $\frac{n-1}{f}$  et  $\frac{n-1}{g}$  valores assignati substituantur spatium diffusionis erit

$$Ff = \frac{n \alpha \alpha x x}{2(n-1)^2} \left( \frac{\nu \nu}{\mu \mu} \left( \frac{1}{a} + \frac{1}{vd} \right)^2 \left( \frac{n}{a} + \frac{1}{vd} \right) + \frac{\mu \mu}{\nu \nu} \left( \frac{1}{\alpha} + \frac{1}{\mu d} \right)^2 \left( \frac{n}{\alpha} + \frac{1}{\mu d} \right) \right)$$

feu

$$Ff = \frac{n \alpha \alpha x x}{2(n-1)^2} \left\{ + \frac{\nu \nu}{\mu \mu} \left( \frac{n}{a^3} + \frac{2n+1}{v a a d} + \frac{n+2}{v v a d d} + \frac{1}{v^3 d^3} \right) \right\}$$

$$\left\{ + \frac{\mu \mu}{\nu \nu} \left( \frac{n}{\alpha^3} + \frac{2n+1}{\mu \alpha \alpha d} + \frac{n+2}{\mu \mu \alpha d d} + \frac{1}{\mu^3 d^3} \right) \right\}$$

Si

Si crassities lentis fuerit valde parua, ne confusio fiat enormis, necesse est pro  $\mu$  et  $\nu$  sumi numeros vehementer magnos, alterum scilicet positium alterum negatiuum.

Cum igitur sit  $\mu d + \nu d = d$ , statuatur  $\mu d = \frac{d-k}{2}$  et  $\nu d = \frac{d+k}{2}$ , vt sit

$$\frac{n-1}{f} = \frac{1}{a} + \frac{2n}{k+d} \text{ et } \frac{n-1}{g} = \frac{1}{a} - \frac{2n}{k-d} \text{ seu}$$

$$f = \frac{(n-1)a(k+d)}{2na+k+d} \text{ et } g = \frac{(n-1)a(k-d)}{k-d-2n\alpha}$$

hincque obtinetur spatium diffusionis

$$Ff = \frac{n\alpha x x}{2(n-1)^2} \left\{ + \left( \frac{k+d}{k-d} \right)^2 \left( \frac{n}{a} + \frac{2}{k+d} \right) \left( \frac{1}{a} + \frac{2}{k+d} \right)^2 \right. \\ \left. + \left( \frac{k-d}{k+d} \right)^2 \left( \frac{n}{a} - \frac{2}{k-d} \right) \left( \frac{1}{a} - \frac{2}{k-d} \right)^2 \right\}$$

in quibus formulis paruitas crassitiei lentis  $AB=d$  nullum negotium facessit: tum vero est

$$\text{angulus } BfN = \frac{k-d}{k+d} \cdot \frac{x}{a}.$$

### C O R O L L. I.

42. Propositis ergo binis distantis  $AE=\alpha$  et  $BF=\alpha$  vna cum crassitie lentis  $AB=d$ , infinitis modis lentes idoneae parari possunt cum pro  $k$  quantitates pro lubitu siue positivae siue negativae assumi queant.

### C O R O L L. 2.

43. Cum semidiameter aperturæ  $x$  faciem anteriorem respiciat, et angulus  $BfN$  seu  $BFN$  sit  $= \frac{k-d}{k+d} \cdot \frac{x}{a}$ , existente distantia  $BF=\alpha$ , manifestum est semidiametrum



trum aperturæ posterioris faciei esse debere  $= \frac{k-d}{k+d} \cdot x$   
vel saltem non minorem.

### Coroll. 3.

44. Quodsi lentis crassities tam sit parua, vt ea  
prae  $k$  contemni queat, formulae nostrae fient simpli-  
ciores

$$f = \frac{(n-1)ak}{k+2na} \text{ et } g = \frac{(n-1)ak}{k-2na}$$

et spatium diffusionis

$$Ff = \frac{n\alpha x}{2(n-1)^2} \left( \left( \frac{n}{a} + \frac{2}{k} \right) \left( \frac{1}{a} + \frac{2}{k} \right)^2 + \left( \frac{n}{a} - \frac{2}{k} \right) \left( \frac{1}{a} - \frac{2}{k} \right)^2 \right)$$

angulusque  $BfN = \frac{x}{a}$ .

### Scholion.

45. Sufficiat haec de spatio diffusionis in genere  
quaecunque fuerit lentis crassities, tradidisse, cum non  
obstante his satis concinnis transformationibus calculus  
nimium fieret molestus, si in determinatione spatii  
diffusionis rationem crassitiei lentis habere vellemus.  
Licebit autem, vt modo vidimus, crassitiem lentis  
negligere non solum, quando ipsa est per se valde  
exigua verum etiam dummodo prae quantitate  $k$  fuerit  
perparua. Atque hinc etiam in sequentibus facile  
iudicare poterimus, vtrum crassitiem lentis recte in  
calculo contemserimus, nec ne? quouis enim casu  
consideretur quantitas pro  $k$  assumpta, quae si multo-  
ties fuerit maior quam  $d$  error nullus erit pertime-  
scendus;

scendus; contra vero si  $k$  non multum superet  $d$  multum aberrabitur, quantumvis exigua fuerit crassities ipsa per se.

Vt fiat  $Ff$  minimum, definiri potest conueniens valor ipsius  $k$  hoc modo. Ponatur  $\frac{k-d}{k+d} = z$  et pro minimo peruenitur ad hanc aequationem:

$$\begin{aligned} 0 = & 2z^4 \left( \frac{n}{a} + \frac{1}{d} \right) \left( \frac{1}{a} + \frac{1}{d} \right)^2 \\ & - \frac{2nz^3}{a} \left( \frac{1}{a} + \frac{1}{d} \right) - \frac{z^3}{d} \left( \frac{1}{a} + \frac{2}{d} \right)^2 \\ & + \frac{2nz}{a} \left( \frac{1}{a} + \frac{1}{d} \right) + \frac{z}{d} \left( \frac{1}{a} + \frac{2}{d} \right)^2 \\ & - 2 \left( \frac{n}{a} + \frac{1}{d} \right) \left( \frac{1}{a} + \frac{1}{d} \right)^2 \end{aligned}$$

unde valor ipsius  $z$  erui debet.

### Problema 6.

46. Neglecta lentis crassitie, si detur cum obiecti ante lentem distantia  $AE = a$ , tum imaginis principalis post lentem distantia  $BF = \alpha$ , eam definire lentem quae pro data apertura minimam pariat diffusionem.

### Solutio.

Positis radiis faciei anterioris  $AM = f$  et posterioris  $BN = g$  vtraque vt conuexa spectata, vidimus omnes lentes datis distantis  $a$  et  $\alpha$  conuenientes his formulis determinari: si scilicet pro  $k$  scribamus  $2k$ .

$$\frac{n-1}{f} = \frac{1}{a} + \frac{n}{k} \quad \text{et} \quad \frac{n-1}{g} = \frac{1}{\alpha} - \frac{n}{k}$$

deno-



denotante  $k$  quantitatem quamcunque. Tum autem spatium diffusionis ita exprimi est repertum

$$Ff = \frac{n\alpha\alpha\alpha\alpha}{2(n-1)^2} \left( \left( \frac{n}{a} + \frac{1}{k} \right) \left( \frac{1}{a} + \frac{1}{k} \right)^2 + \left( \frac{n}{\alpha} - \frac{1}{k} \right) \left( \frac{1}{\alpha} - \frac{1}{k} \right)^2 \right)$$

quae expressio ad hanc reducitur formam :

$$Ff = \frac{n\alpha\alpha\alpha\alpha}{2(n-1)^2} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( n \left( \frac{1}{a\alpha} - \frac{1}{a\alpha} + \frac{1}{\alpha\alpha} \right) + \frac{2n+1}{k} \left( \frac{1}{a} - \frac{1}{\alpha} \right) + \frac{n+2}{k k} \right)$$

Quaestio igitur huc reducitur, vt definiatur quantitas  $k$  qua huic expressioni valor minimus concilietur; cui quidem requisito satisfacit valor

$$\frac{1}{k} = \frac{-(2n+1)}{2(n+2)} \left( \frac{1}{a} - \frac{1}{\alpha} \right) \text{ feu } \frac{1}{k} = \frac{2n+1}{2(n+2)} \left( \frac{1}{\alpha} - \frac{1}{a} \right)$$

quo substituto habebitur :

$$\frac{n-1}{f} = \frac{4+n-2nn}{2(n+2)a} + \frac{n(2n+1)}{2(n+2)\alpha} \text{ et } \frac{n-1}{g} = \frac{4+n-2nn}{2(n+2)\alpha} + \frac{n(2n+1)}{2(n+2)a}$$

et spatium diffusionis quod est minimum :

$$Ff = \frac{n\alpha\alpha\alpha\alpha}{2(n-1)^2} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( n \left( \frac{1}{a\alpha} - \frac{1}{a\alpha} + \frac{1}{\alpha\alpha} \right) - \frac{(2n+1)^2}{4(n+2)} \left( \frac{1}{a} - \frac{1}{\alpha} \right)^2 \right)$$

At est

$$n \left( \frac{1}{a\alpha} - \frac{1}{a\alpha} + \frac{1}{\alpha\alpha} \right) = n \left( \frac{1}{a} - \frac{1}{\alpha} \right)^2 + \frac{n}{a\alpha} \text{ ideoque}$$

$$Ff = \frac{n\alpha\alpha\alpha\alpha}{2(n-1)^2} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \frac{4n-1}{4(n+2)} \left( \frac{1}{a} - \frac{1}{\alpha} \right)^2 + \frac{n}{a\alpha} \right)$$

quae expressio etiam in hanc formam transfunditur

$$Ff = \frac{n\alpha\alpha\alpha\alpha}{2(n-1)^2} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \frac{4n-1}{4(n+2)} \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{(n-1)^2}{(n+2)a\alpha} \right)$$

feu

$$Ff = \frac{n(4n-1)\alpha\alpha\alpha\alpha}{8(n-1)^2(n+2)} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{4(n-1)^2}{(4n-1)a\alpha} \right)$$

Coroll.

## Coroll. I.

47. Vt igitur lens minimum spatium diffusionis producat, eam ita formari necesse est vt sit

$$f = \frac{2(n-1)(n+2)a\alpha}{n(2n+1)a + (4+n-2nn)\alpha} \text{ et}$$

$$g = \frac{2(n-1)(n+2)a\alpha}{n(2n+1)\alpha + (4+n-2nn)a}$$

neglecta scilicet lentis crassitie, quae quidem recte negligitur, si modo fuerit vehementer parua prae quantitate  $k = \frac{2(n+2)a\alpha}{(2n+1)(a-\alpha)}$ .

## Coroll. 2.

48. Si igitur sit  $a = \alpha$ , seu  $BF = AE$  crassities lentis, quantacunque fuerit, nihil turbat in spatio diffusionis. Pro hoc autem casu erit  $f = g = (n-1)a$  et spatium diffusionis ipsum  $Ff = \frac{nnxx}{(n-1)^2a}$ . At quo magis distantiae  $a$  et  $\alpha$  a se inuicem discrepant, seu minor fuerit quantitas  $\frac{a\alpha}{a-\alpha}$ , eo magis haec determinatio ob lentis crassitiem fiet erronea.

## Coroll. 3.

49. Spatium autem hoc confusionis minimum  $Ff$  pluribus modis exhiberi potest, inter quos commodissimum eligi conuenit; sunt autem praecipui:

$$\text{I. } Ff = \frac{n(4n-1)\alpha\alpha xx}{8(n-1)^2(n+2)} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{4(n-1)^2}{(4n-1)\alpha} \right)$$

$$\text{II. } Ff = \frac{n(4n-1)\alpha\alpha xx}{8(n-1)^2(n+2)} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \left( \frac{1}{a} - \frac{1}{\alpha} \right)^2 + \frac{4n(n+2)}{(4n-1)a\alpha} \right)$$

$$\text{III. } Ff = \frac{n(4n-1)\alpha\alpha xx}{8(n-1)^2(n+2)} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \left( \frac{1}{aa} + \frac{1}{\alpha\alpha} \right) + \frac{2(2nn+1)}{(4n-1)a\alpha} \right)$$

$$\text{IV. } Ff = \frac{n(4n-1)\alpha\alpha xx}{8(n-1)^2(n+2)} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \left( \frac{1}{aa} - \frac{1}{a\alpha} + \frac{1}{\alpha\alpha} \right) + \frac{(2n+1)^2}{(4n-1)\alpha} \right)$$

Tom. I.

E

Coroll.



## Coroll. 4.

50. Cum ergo hoc spatium diffusionis sit minimum, si lenti alia quaecunque figura tribuatur, ita tamen, vt obiecti ad distantiam  $AE = a$  remoti imaginem principalem in distantia  $BF = a$  exhibeat, spatium diffusionis erit maius, quam hic inuenimus.

## Scholion.

51. Quia  $n : 1$  denotat rationem refractionis ex aere in vitrum, haec autem pro radiorum natura est variabilis, conueniet pro  $n$  medium valorem assumi, qui est  $n = \frac{31}{20}$ ; hinc ergo erit

$$n - 1 = \frac{11}{20}; n + 2 = \frac{71}{20}; 2n + 1 = \frac{41}{10}; 4 + n - 2nn = \frac{149}{200},$$

hincque

$$\frac{n(2n+1)}{2(n-1)(n+2)} = \frac{1271}{781} = 1,627401$$

$$\frac{4+n-2nn}{2(n-1)(n+2)} = \frac{149}{781} = 0,190781$$

vnde lens minimam confusionem pariens ita definietur

$$\frac{1}{f} = \frac{149}{781a} + \frac{1271}{781a} = \frac{0,190781}{a} + \frac{1,627401}{a}$$

$$\frac{1}{g} = \frac{149}{781a} + \frac{1271}{781a} = \frac{0,190781}{a} + \frac{1,627401}{a}$$

feu

$$f = \frac{781a\alpha}{149\alpha + 1271\alpha} = \frac{a\alpha}{0,190781\alpha + 1,627401\alpha}$$

et

$$g = \frac{781a\alpha}{149\alpha + 1271\alpha} = \frac{a\alpha}{0,190781\alpha + 1,627401\alpha}$$

Pro

Pro spatio autem diffusionis ipso definiendo ob  $4n-1=\frac{26}{6}$

$$\text{erit } \frac{n(4n-1)}{8(n-1)^2(n+2)} = \frac{8060}{8591} = 0,938191$$

$$\text{et } \frac{4(n-1)^2}{4n-1} = \frac{121}{520} = 0,232692$$

hincque

$$Ff = 0,938191 \alpha \alpha x x \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{0,232692}{a\alpha} \right)$$

vnde reliquae formulæ facile deducuntur. Ceterum hic monendum est, quoniam valor  $n=\frac{31}{20}$  ex experimentis est conclusus, neque ipse pro omnibus radiorum generibus valet, superfluum fore in praxi hos numeros inuentos nimis studiose obseruare; quin etiam ipsa natura minimi aliquam aberrationem permittit. Nihilo vero minus has fractiones decimales ad tot figuras producere visum est, quo facilius quantum ab hac hypothesis aberretur, dignosci queat.

### Problema 7.

52. Neglecta lentis crassitie, si detur cum obiecti ante lentem distantia  $AE=a$ , tum imaginis principalis post lentem distantia  $BF=\alpha$ , eam definire lentem, quae pro data apertura non minimum, sed datum pariat spatium diffusionis  $Ff$ .

### Solutio.

Lente vtrunque vt conuexa spectata, sit faciei anterioris radius  $=f$ , posterioris  $=g$ ; atque vt ex data

E 2

distan-



distantia obiecti  $AE = a$ , data oriatur imaginis principalis distantia  $= \alpha$  necesse est fit in genere

$$\frac{n-1}{f} = \frac{1}{a} + \frac{n}{k} \quad \text{et} \quad \frac{n-1}{g} = \frac{1}{\alpha} - \frac{n}{k}$$

vnde spatium diffusionis fit

$$Ff = \frac{n\alpha\alpha x x}{2(n-1)^2} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( n \left( \frac{1}{a} - \frac{1}{\alpha} + \frac{1}{\alpha\alpha} \right) + \frac{2n+1}{k} \left( \frac{1}{a} - \frac{1}{\alpha} \right) + \frac{n+2}{kk} \right)$$

Cum autem spatium diffusionis minimum repertum fit

$$\frac{n(n-1)\alpha\alpha x x}{8(n-1)^2(n+2)} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{4(n-1)^2}{(4n-1)\alpha} \right)$$

necesse est vt illud fit maius: statuaturs ergo:

$$Ff = \frac{n(4n-1)\alpha\alpha x x}{8(n-1)^2(n+2)} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{4(n-1)^2}{(4n-1)\alpha} + S \right)$$

et habebitur ista aequatio:

$$4n(n+2) \left( \frac{1}{a} - \frac{1}{\alpha} + \frac{1}{\alpha\alpha} \right) + \frac{4(n+2)(2n+1)}{k} \left( \frac{1}{a} - \frac{1}{\alpha} \right) + \frac{4(n+2)^2}{kk} \\ = (4n-1) \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{4(n-1)^2}{\alpha\alpha} + (4n-1)S$$

quae in hanc formam redigitur.

$$(2n+1)^2 \left( \frac{1}{a} - \frac{1}{\alpha} \right)^2 + \frac{4(n+2)(2n+1)}{k} \left( \frac{1}{a} - \frac{1}{\alpha} \right) + \frac{4(n+2)^2}{kk} = (4n-1)S$$

vnde radice extracta fit

$$(2n+1) \left( \frac{1}{a} - \frac{1}{\alpha} \right) + \frac{2(n+2)}{k} = \sqrt{(4n-1)S}$$

$$\text{et} \quad \frac{1}{k} = \frac{(2n+1)}{2(n+2)} \left( \frac{1}{a} - \frac{1}{\alpha} \right) + \frac{\sqrt{(4n-1)S}}{2(n+2)}$$

Quare erit

$$\frac{n-1}{f} = \frac{1}{a} - \frac{n(2n+1)}{2(n+2)} \left( \frac{1}{a} - \frac{1}{\alpha} \right) + \frac{n}{2(n+2)} \sqrt{(4n-1)S}$$

$$\frac{n-1}{g} = \frac{1}{\alpha} + \frac{n(2n+1)}{2(n+2)} \left( \frac{1}{a} - \frac{1}{\alpha} \right) - \frac{n}{2(n+2)} \sqrt{(4n-1)S}$$

fiue

fiue

$$f = \frac{2(n-1)(n+1)a\alpha}{(1+n-2nn)\alpha + 1(2n+1)a + 1a\alpha\sqrt{(4n-1)S}}$$

$$g = \frac{2(n-1)(n+1)a\alpha}{(1+n-2nn)\alpha + 1(2n+1)a - n\alpha\alpha\sqrt{(4n-1)S}}$$

Oportet ergo pro S sumi quantitatem positiuam, atque vt cum reliqua parte, cui jungitur, sit homogenea, ponamus

$$S = (\lambda - 1) \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2$$

vbi  $\lambda$  denotat numerum vnitatem maiorem; atque effici poterit, vt spatium diffusionis ita exprimatur

$$Ff = \frac{n(4n-1)\alpha\alpha\alpha\alpha/x}{8(n-1)^2(n+1)} + \frac{1}{\alpha} \left( \lambda \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{4(n-1)^2}{(4n-1)a\alpha} \right)$$

quod cum vt datum spectetur, numerus  $\lambda$  pro dato assumi poterit, ac lens hunc effectum produciens ita determinabitur.

$$f = \frac{2(n-1)(n+1)a\alpha}{(1+n-2nn)\alpha + 1(2n+1)a + 1(a+\alpha)\sqrt{(4n-1)(\lambda-1)}}$$

$$g = \frac{2(n-1)(n+1)a\alpha}{(1+n-2nn)\alpha + 1(2n+1)a - n(a+\alpha)\sqrt{(4n-1)(\lambda-1)}}$$

quod ergo cum signum radicale ambiguo illatum inuoluat, duplici modo fieri poterit.

### Coroll. I.

53. Omnes igitur lentes, quae obiecti ad distantiam  $AE = a$  remoti, imaginem principalem in distantia  $BF = \alpha$  repraesentant, hanc habent proprietatem vt sit:

$$\frac{n-1}{f} + \frac{n-1}{g} = \frac{1}{a} + \frac{1}{\alpha} \text{ seu } \frac{fg}{f+g} = \frac{(n-1)a\alpha}{a+\alpha}.$$

E 3

et



et cum fit

$$n = \frac{31}{20} \text{ erit } \frac{fg}{f+g} = \frac{11a\alpha}{20(a+\alpha)} = \frac{0,55a\alpha}{a+\alpha}.$$

### Coroll. 2.

54. Si ergo lens fit vtrunque aequae conuexa seu  $f=g$ , oportet esse  $f=g = \frac{2(n-1)a\alpha}{a+\alpha} = \frac{11a\alpha}{10(a+\alpha)}.$

sin autem lens desideretur plano-conuexa vt fit  $f=\infty$ , capi debet

$$g = \frac{(n-1)a\alpha}{a+\alpha} = \frac{11a\alpha}{20(a+\alpha)} = \frac{0,55a\alpha}{a+\alpha}$$

At si lens debeat esse conuexo-plana seu  $g=\infty$ , capi oportet

$$f = \frac{(n-1)a\alpha}{a+\alpha} = \frac{11a\alpha}{20(a+\alpha)} = \frac{0,55a\alpha}{a+\alpha}$$

### Scholion. I.

55. Substituamus pro  $n$  valorem ipsi conuenientem  $\frac{31}{20}$ , et iam vidimus fore:

$$\frac{n(4n-1)}{8(n-1)^2(n+2)} = \frac{8060}{8591} = 0,938191$$

$$\frac{4(n-1)^2}{4n-1} = \frac{121}{520} = 0,232692$$

$$\frac{4+7-2nn}{2(n-1)(n+2)} = \frac{149}{781} = 0,190781$$

$$\frac{n(2n+1)}{2(n-1)(n+2)} = \frac{1271}{781} = 1,627401$$

nunc vero notari oportet esse:

$$\frac{n\sqrt{4n-1}}{2(n-1)(n+2)} = \frac{62\sqrt{130}}{781} = 0,905133$$

Quare

Quare si lens ita construatur vt fit

$$f = \frac{a\alpha}{0,190781\alpha + 1,627401\alpha \pm 0,905133(a+\alpha)\sqrt{\lambda-1}}$$

$$g = \frac{a\alpha}{0,190781\alpha + 1,627401\alpha \pm 0,905133(a+\alpha)\sqrt{\lambda-1}}$$

erit pro eius apertura, cuius semidiameter  $=x$ ,  
spatium diffusionis

$$Ff = 0,938191\alpha\alpha xx\left(\frac{1}{a} + \frac{1}{\alpha}\right)\left(\lambda\left(\frac{1}{a} + \frac{1}{\alpha}\right)^2 + \frac{0,232692}{a\alpha}\right)$$

Cum autem in posterum hi numeri frequentissime  
occurrant eorum loco ad abbreviandum certis caracte-  
ribus vtamur, ponamus ergo

$$\frac{n(4n-1)}{8(n-1)^2(n+2)} = 0,938191 = \mu$$

$$\frac{4(n-1)^2}{4n-1} = 0,232692 = \nu$$

$$\frac{4+n-2nn}{2(n-1)(n+2)} = 0,190781 = \varrho$$

$$\frac{n(2n+1)}{2(n-1)(n+2)} = 1,627401 = \sigma$$

$$\frac{n\sqrt{4n-1}}{2(n-1)(n+2)} = 0,905133 = \tau$$

feu

$$\mu = \frac{1}{4(n+2)} + \frac{1}{4(n-1)} + \frac{1}{8(n+1)^2}$$

$$\varrho = \frac{1}{2(n-1)} + \frac{1}{n+2} - 1$$

$$\sigma = 1 + \frac{1}{2(n-1)} - \frac{1}{n+2}$$

$$\tau = \frac{1}{3}\left(\frac{1}{2(n-1)} + \frac{1}{n+2}\right)\sqrt{4n-1}$$

vnde pro quouis alio medio pellucido hi valores  
supputari possunt

ficque



ficque vt prodeat spatium diffusionis

$$Ff = \mu \alpha \alpha x x \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{v}{a\alpha} \right)$$

lentis constructio erit

$$f = \frac{a\alpha}{\varrho a + \sigma a \pm \tau(a+\alpha)\sqrt{\lambda-1}}$$

$$g = \frac{a\alpha}{\varrho a + \sigma a \pm \tau(a+\alpha)\sqrt{\lambda-1}}$$

Dummodo igitur  $\lambda$  fuerit numerus positivus unitate non minor talis lens duplici modo effici poterit. Casu autem  $\lambda = 1$ , quo spatium diffusionis est minimum, vnica lens proposito satisfaciens construi potest.

### Coroll. 3.

56. Si igitur lens fit vtrunque aequaliter convexa, ideoque  $f = g = \frac{2(n-1)a\alpha}{a+\alpha} = \frac{11}{10} \cdot \frac{a\alpha}{a+\alpha}$ , in expressio-  
ne nostra pro spatium diffusionis inventa valor ipsius  $\lambda$   
ex aequalitate inter  $f$  et  $g$  statuta definietur, vnde fit

$$(\sigma - \varrho)(a - \alpha) = 2\tau(a + \alpha)\sqrt{\lambda - 1} \text{ hincque}$$

$$\lambda = 1 + \frac{0,629795\alpha\alpha - 1,259589a\alpha + 0,629795aa}{(a + \alpha)^2}.$$

### Coroll. 4.

57. Si lens capiatur plano convexa vt fit:

$$f = \infty \text{ et } g = \frac{(n-1)a\alpha}{a+\alpha} = \frac{11}{20} \cdot \frac{a\alpha}{a+\alpha}$$

pro spatium diffusionis habebitur

$$\varrho a + \sigma a = \pm \tau(a + \alpha)\sqrt{\lambda - 1}$$

vnde in numeris colligitur

$$\lambda = 1 + \frac{0,044427\alpha\alpha + 0,757940a\alpha + 3,232692aa}{(a + \alpha)^2}$$

Coroll.

## Coroll. 5.

58. Si denique lens adhibeatur conuexo-plana, ut sit:

$$g = \infty \text{ et } f = \frac{(n-1)a\alpha}{a^2 + \alpha^2} = \frac{11}{20} \cdot \frac{a\alpha}{a + \alpha}$$

pro spatio diffusionis inueniendo poni oportet

$$ga + \sigma\alpha = \pm \tau(a + \alpha)V(\lambda - 1)$$

vnde in numeris colligimus:

$$\lambda = 1 + \frac{3,232692\alpha\alpha + 0,757940\alpha\alpha + 0,044427\alpha\alpha}{(a + \alpha)^2}$$

## Scholion 2.

59. Quod ad aperturam attinet iam initio au- Tab. I.  
mauertimus, in ea maiores arcus comprehendi non Fig. 2.  
debere, quam qui sint principiis stabilitis conformes.  
Scilicet ut nullus angulus supra 30° gradus occurrat,  
anguli ACM et BDN certe minores 30 gradibus esse  
debent, cum anguli EMc et VNd ipsis sint maiores,  
quorum alter cum quandoque ad duplum affurgere  
possit, poterimus hanc regulam statuere, ut aperturæ  
femidiameter  $x$  neque  $\frac{1}{4}f$  neque  $\frac{1}{4}g$  superet. Verum  
quouis casu ad ipsos angulos EMc et VNd, qui sunt  
maximi, attendi conueniet, atque tanta apertura ad-  
mitti poterit, vnde neuter horum angulorum 30 gra-  
dus superans oriatur, si cautius procedere velimus,  
etiam angulos 20 gradibus minores euitare poterimus,  
quo pacto apertura magis restringetur.



## Problema 8.

Tab. I. 60. Non neglecta lentis crassitie AB si pro  
Fig. 3. distantia obiekti AE =  $a$  detur distantia imaginis prin-  
cipalis BF =  $\alpha$ , obiektum autem parumper longius  
in  $e$  remouetur, definire locum imaginis principalis  $f$ .

## Solutio.

Posita lentis crassitie AB =  $d$ , radiisque faciei  
anterioris AM =  $f$  et posterioris BN =  $g$ , supra  
vidimus hos radios ita a binis distantis  $a$  et  $\alpha$  atque  
crassitie lentis  $d$  pendere debere, vt fit

$$\frac{n-1}{f} = \frac{1}{a} + \frac{2n}{k+d} \text{ et } \frac{n-1}{g} = \frac{1}{\alpha} - \frac{2n}{k-d}$$

denotante  $k$  quantitatem quamcunque. Hinc ergo  
cum fit

$$\frac{k+d}{2n} = \frac{af}{(n-1)a-f} \text{ et } \frac{k-d}{2n} = \frac{\alpha g}{g-(n-1)\alpha}$$

erit eliminando  $k$ .

$$\frac{d}{n} = \frac{af}{(n-1)a-f} - \frac{\alpha g}{g-(n-1)\alpha}$$

Ponamus iam distantiam AE =  $a$  crescere particula  
Ee =  $da$  ac per differentiationem inueniemus, quantum  
inde distantia imaginis BF =  $\alpha$  immutetur: habebimus  
scilicet:

$$\frac{-ffda}{((n-1)a-f)^2} - \frac{ggd\alpha}{(g-(n-1)\alpha)^2} = 0$$

unde elicimus

$$d\alpha = \frac{-ff(g-(n-1)\alpha)^2}{gg((n-1)a-f)^2} da = \frac{-\alpha\alpha'}{aa} da \left(\frac{k+d}{k-d}\right)^2$$

Quare

Quare obiecto E per spatium minimum Ee longius a lente remoto, imago principalis ex F propius ad lentem admovebitur per spatium minimum Ff, ita ut sit

$$Ff = \frac{\alpha \alpha}{a a} \left( \frac{k+d}{k-d} \right)^2. Ee = \frac{ff(g-(n-1)\alpha)^2}{gg((n-1)\alpha-f)^2}. Ee.$$

### COROLL. I.

61. Quia quantitas  $\frac{\alpha \alpha}{a a} \left( \frac{k+d}{k-d} \right)^2$  necessario est positiva, euidens est si obiectum longius a lente remoueat, imaginem semper propius ad lentem admoueri. Contra ergo etiam si obiectum propius ad lentem accedat, imago longius ab ea recedet.

### COROLL. 2.

62. Si crassities lentis  $d$  euanescat erit  $Ff = \frac{\alpha \alpha}{a a} Ee$ , fin autem ea non euanescat fieri potest ut sit vel  $Ff > \frac{\alpha \alpha}{a a} Ee$  vel  $Ff < \frac{\alpha \alpha}{a a} Ee$ , prius eueniet si  $k$  sit quantitas positiua, posterius si negatiua. At si sit vel  $k = \infty$  vel  $k = 0$ , utroque casu erit  $Ff = \frac{\alpha \alpha}{a a} Ee$  etiam si crassities lentis non euanescat.

### COROLL. 3.

63. Cum in distantis  $a$  et  $\alpha$  mutatio minima fieri concipiatur, spatium diffusionis nullam inde variationem subire censendum est: siue ergo obiectum in E siue e reperiatur, ac semidiameter aperturæ



lentic in facie anteriore fuerit  $=x$ , erit spatium diffusionis :

$$\frac{n\alpha\alpha x x}{2(n-1)} \left\{ + \left( \frac{k+d}{k-d} \right)^2 \left( \frac{n}{\alpha} + \frac{z}{k+d} \right) \left( \frac{1}{\alpha} + \frac{z}{k+d} \right)^2 \right\}$$

$$\left\{ + \left( \frac{k-d}{k+d} \right)^2 \left( \frac{n}{\alpha} - \frac{z}{k-d} \right) \left( \frac{1}{\alpha} - \frac{z}{k-d} \right)^2 \right\}$$

in facie autem posteriori semidiameter aperturæ debet esse  $= \frac{k-d}{k+d} \cdot x$ .

### Problema 9.

Tab. I. 64. Rationem definire, quam habet magnitudo  
Fig. 3. imaginis ad magnitudinem obiecti non neglecta lentis  
crassitie.

### Solutio.

Sit obiecti ante lentem distantia  $AE=a$ , imaginis vero  $BF=\alpha$  existente lentis crassitie  $=d$ ; ubi quidem tantum imaginem principalem spectamus neglecto spatio diffusionis. Sit iam  $E\epsilon$  obiectum, cui tribuatur magnitudo quam minima  $E\epsilon=z$ , normaliter axi lentis insistens; eiusque imago in  $F\zeta$  exhibebitur, cuius magnitudo  $F\zeta$  quaeritur. Cum igitur punctum  $\zeta$  a puncto  $\epsilon$  oriatur, ita ut radii ab  $\epsilon$  emissi in  $\zeta$  colligantur, consideretur radius quicumque  $\epsilon M$ , qui productus cum axe in  $e$  occurrat: et perinde est, ac si hic radius ex axis puncto  $e$  emanaret. Quare is post refractionem cum axe in  $f$  concurret, ut sit  $Ff = \frac{\alpha\alpha}{a} \left( \frac{k+d}{k-d} \right)^2$ .  $Ee$ , indeque ad  $\zeta$  perget: ex quo magnitudinem  $F\zeta$  definire licebit. Statuatur in

in hunc finem  $AM = x$ , erit  $BN = \frac{k-d}{k+d} \cdot x$ ; hincque colligemus has proportiones:

$$Ee : E\varepsilon = eA : AM = a : x$$

$$Ff : F\zeta = fB : BN = \alpha : \frac{k-d}{k+d} x$$

ob spatiola  $Ee$  et  $Ff$  quam minima: vnde habebimus

$$\frac{Ff}{Ee} : \frac{F\zeta}{E\varepsilon} = \frac{\alpha}{a} : \frac{k-d}{k+d} = \frac{\alpha\alpha}{aa} \left( \frac{k+d}{k-d} \right)^2 : \frac{F\zeta}{E\varepsilon}$$

Concluditur ergo  $\frac{F\zeta}{E\varepsilon} = \frac{\alpha}{a} \cdot \frac{k+d}{k-d}$ , ex quo cum magnitudo obiecti  $E\varepsilon$  posita sit  $= z$ , erit magnitudo imaginis  $F\zeta = \frac{\alpha(k+d)}{a(k-d)} z$ .

### COROLL. I.

65. Secundum hanc ergo rationem diameter obiecti immutatur. Vbi notandum est, si expressio  $\frac{\alpha(k+d)}{a(k-d)}$  habeat positium valorem obiecti imaginem situ inuerso repraesentari, contra autem situ erecto, si  $\frac{\alpha(k+d)}{a(k-d)}$  negatiuum valorem adipiscatur.

### COROLL. 2.

66. Si crassities lentis euanescat, fit  $F\zeta = \frac{\alpha}{a} z$ . quo ergo casu recta iungens puncta extrema  $\varepsilon$  et  $\zeta$  transit per centrum lentis. At si crassities in computum ducatur, recta  $\varepsilon\zeta$  modo intra lentem modo extra, axem lentis secare poterit.

### SCHOLIUM.

67. Sic igitur omnia, quae circa vnam lentem quameunque nosse oportet, expediimus, vt etiam



crassitiei rationem habuerimus. Ac primo quidem ad binas distantias lentis quasi determinatrices spectari conuenit, quae sunt obiecti distantia ante lentem  $AE = a$  et imaginis principalis post lentem distantia  $BF = \alpha$ , quibus addi debet lentis crassities  $AB = d$ . His autem conditionibus infinitae lentes satisfaciunt; si enim radius faciei anterioris  $AM$  dicatur  $= f$ , et posterioris  $BN = g$ , vtraque tanquam conuexa considerata, constructio lentis his continetur formulis:

$$\frac{n-1}{f} = \frac{1}{a} + \frac{2n}{k+d} \quad \text{et} \quad \frac{n-1}{g} = \frac{1}{\alpha} - \frac{2n}{k-d}$$

feu

$$f = \frac{(n-1)a(k+d)}{k+d+2na} \quad \text{et} \quad g = \frac{(n-1)\alpha(k-d)}{k-d-2n\alpha}$$

existente  $n = \frac{31}{20}$ ; vbi  $k$  est quantitas arbitraria, hincque lentes innumerabiles quaesito satisfaciētes obtinentur.

Quod si iam obiecti diameter ponatur  $= z$ , erit imaginis principalis repraesentatae diameter  $= \frac{\alpha(k+d)}{a(k-d)} z$ , quatenus imago situ inuerso exhibita consideratur.

Deinde si aperturæ in facie anterioris lentis semidiameter fit  $= x$ , spatium diffusionis, quatenus ab imagine principali ad lentem porrigitur, ita exprimetur ut fit:

$$\frac{n\alpha\alpha xx}{2(n-1)^2} \cdot \left\{ + \left( \frac{k+d}{k-d} \right)^2 \left( \frac{n}{a} + \frac{2}{k+d} \right) \left( \frac{1}{a} + \frac{2}{k+d} \right)^2 \right. \\ \left. + \left( \frac{k-d}{k+d} \right)^2 \left( \frac{n}{\alpha} - \frac{2}{k-d} \right) \left( \frac{1}{\alpha} - \frac{2}{k-d} \right)^2 \right\}$$

in facie autem posteriori aperturæ semidiameter debet esse  $= \frac{k-d}{k+d} \cdot x$  vel saltem non minor. Quia spatium diffu-

diffusionis factorem habet  $\alpha\alpha x x$ , breuitatis gratia id ita  $P\alpha\alpha x x$  indicabimus. Haecque in genere teneantur etiam crassitiei lentis ratione habita. At si crassities lentis euanescat, formulas magis euoluere licuit, scilicet si breuitatis gratia ponatur:

$$\mu = 0,938191; \rho = 0,190781; \tau = 0,905133$$

$$\nu = 0,232692; \sigma = 1,627401; \lambda > 1 \text{ arbitr.}$$

sumaturque pro formatione lentis:

$$f = \frac{a'\alpha}{\rho\alpha + \sigma a \pm \tau \frac{a'\alpha}{(a+\alpha)\sqrt{(\lambda-1)}}$$

$$g = \frac{a\alpha}{\rho a + \sigma\alpha + \tau \frac{a\alpha}{(a+\alpha)\sqrt{(\lambda-1)}}$$

spatium diffusionis erit pro aperturae semidiametro  $x$

$$\mu\alpha\alpha x x \left(\frac{1}{a} + \frac{1}{\alpha}\right) \left(\lambda \left(\frac{1}{a} + \frac{1}{\alpha}\right)^2 + \frac{\nu}{a\alpha}\right)$$

et obiecti diametro existente  $= z$  imaginis diameter erit  $= \frac{\alpha}{a} z$ . His ergo pro vna lente determinatis, videamus quomodo in combinatione duarum pluriumve lentium spatium diffusionis definiatur: vt deinceps in omnis generis instrumentis dioptricis siue Telescopiis siue microscopiis confusionem assignare, modumque eam diminuendi inuestigare possimus.





# CAPVT II.

## DE DIFFVSIONE IMAGINIS PER PLVRES LENTES REPRÆSENTATÆ.

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### Problema I.

68.

Tab. I.  
Fig. 4.

**S**i loco obiecti adsit imago per spatium  $Ee$  diffusa, indeque radii per lentem  $PP$  aperture indefinite transmittantur, determinare spatium diffusionis  $Ff$  per hanc lentem productum.

### Solutio.

Loco obiecti veri hic consideramus imaginem iam per aliam lentem repræsentatam, quæ sit diffusa per spatium  $Ee$ , ita ut in  $E$  sit imago principalis per radios axi proximos formata in  $e$  autem imago extrema, radiis scilicet per marginem aperture lentis præcedentis transmissis formata; qui radii cum axe constituent angulum  $=\Phi$ . Quare a puncto  $E$  nonnisi radii axi proximi in lentem  $PP$  emittuntur, a puncto  $e$  autem eiusmodi tantum radii, qui ad axem  $eA$  sub angulo  $MeA = \Phi$  sint inclinati. Ponatur  
iam

iam distantia  $EA = a$ , prae qua spatium diffusionis  $Ee$  ut valde paruum spectetur, lens autem  $PP$  sit eiusmodi ut obiecti in  $E$  existentis imaginem principalem referat in distantia  $BF = a$ , existente lentis crassitie  $AB = d$ . Hanc ob rem si ponatur faciei anterioris  $AM$  radius  $= f$ , posterioris  $BN = g$ , lente ut conuexa utrinque spectata oportet esse:

$$f = \frac{(n-1)a(k+d)}{k+d+2na} \text{ et } g = \frac{(n-1)a(k-d)}{k-d-2na}$$

vbi est  $n = \frac{31}{20}$  et  $k$  quantitas arbitraria. Hinc autem vti demonstrauiamus, pro aperturae semidiametro  $= x$  nasceretur spatium diffusionis

$$\frac{n\alpha\alpha x x}{2(n-1)^2} \cdot \left\{ + \left( \frac{k+d}{k-d} \right)^2 \left( \frac{n}{a} + \frac{2}{k+d} \right) \left( \frac{1}{a} + \frac{2}{k+d} \right)^2 \right\} \\ \left\{ + \left( \frac{k-d}{k+d} \right)^2 \left( \frac{n}{a} - \frac{2}{k-d} \right) \left( \frac{1}{a} - \frac{2}{k-d} \right)^2 \right\}$$

pro quo scribamus breuitatis gratia  $P\alpha\alpha x x$ . Iam puncti  $E$ , quia inde nonnisi radii axi proximi ad lentem emittuntur imago exhibebitur in  $F$ , ut sit distantia  $BF = a$ : videamus ergo quorsum imago puncti  $e$  debeat cadere. Ac si radii ex puncto  $e$  emissi essent axi proximi, quia id a lente magis est remotum quam  $E$ , eius imago propius ad lentem caderet puta in  $\Phi$ , ut esset, sicut §. 60 definiuimus

$$F\Phi = \frac{\alpha\alpha}{a} \left( \frac{k+d}{k-d} \right)^2 \cdot Ee.$$

in  $\Phi$  scilicet caderet imago principalis, si obiectum esset in  $e$ . Sed quia ab  $e$  tantum radii  $eM$  ad axem angulo  $AeM = \Phi$  inclinati emittuntur, hi lenti in

Tom. I.

G

punctis



punctis M ab A interuallo  $AO = eA$ .  $\Phi = a\Phi$  remotis occurrent, quandoquidem interuallum  $Ee$  prae distantia  $Ae = a$  contemnimus, perinde igitur est ac si lenti apertura tribueretur, cuius semidiameter esset  $= a\Phi$ , et obiecti in  $e$  existentis imago extrema definiri deberet, quae cadat in  $f$  ita vt  $\Phi f$  sit spatium diffusionis obiecto in  $e$  esistenti, et aperturæ lentis, cuius semidiameter  $= a\Phi$ , conueniens. Hinc ergo erit interuallum  $\Phi f = Pa\alpha\alpha\alpha\Phi\Phi$ ; et quia puncti E imago in F, puncti  $e$  autem imago in  $f$  exhibetur, erit spatium diffusionis per lentem PP productum

$$Ff = \frac{\alpha\alpha}{a\alpha} \left( \frac{k+d}{k-d} \right)^2 \cdot Ee + Pa\alpha\alpha\alpha\Phi\Phi.$$

In imaginē autem extrema  $f$  radii  $Nf$  ita cum axe concurrent, vt sit

$$\text{angulus } BfN = \frac{k-d}{k+d} \cdot \frac{a\Phi}{\alpha}.$$

### C o r o l l. I.

69. Si ergo imago iam diffusa per spatium  $Ee$  locum obiecti teneat respectu lentis PP, per hanc spatium diffusionis nouum producitur  $Ff$ , ita vt imago principalis cadat in F extrema vero in  $f$ : fierique poterit, vt hoc nouum spatium diffusionis  $Ff$  maius sit vel minus proposito  $Ee$ .

### C o r o l l. 2.

70. Aperturam lentis PP vt indefinitam assumfi, manifestum autem est sufficere, dum eius semidiameter non sit minor quam  $a\Phi$ . Si enim esset  
minor

minor, radii ex puncto  $e$  emissi plane non per lentem transitum inuenirent, neque imago puncti  $e$  exprimeretur, sed imaginis  $Ff$  punctum extremum responderet imagini cuiusdam intermediae spatii  $Ee$ .

### COROLL. 3.

71. Si diameter obiecti seu potius imaginis in  $E$  sitae sit  $=z$ , tum imaginis inde per lentem  $PP$  in  $F$  formatae diameter erit  $=\frac{\alpha(k+d)}{\alpha(k-d)}z$  quae expressio si sit positiua simul declarat situm imaginis in  $F$  esse inuersum respectu eius, quae est in  $E$ .

### SCHOLION.

72. Cum in hoc capite plures lentes sim consideraturus, pro cuiuslibet determinatione, spectandae sunt primo binae distantiae determinatrices, altera obiecti seu imaginis, a qua lens radios accipit ante lentem, altera imaginis inde per lentem repraesentatae post lentem: quae distantiae ex imaginibus principalibus aestimentur. Deinde cuiusque lentis crassities in computum est ducenda. Tertio cum his lens nondum prorsus determinetur, insuper pro quaque lente accedit distantia quaedam arbitraria hactenus per litteram  $k$  indicata. Quarto vero imprimis ratio haberi deberet, spatii diffusionis, quod cuique lenti pro data apertura conueniat. Quemadmodum autem ex binis distantibus determinatricibus, crassitie lentis et quantitate illa arbitraria  $k$ , cum binae lentis facies tum



etiam spatium diffusionis pro apertura, cuius semidiameter ponitur  $= x$  definiatur, in praecedente capite fusius est expositum. Hic igitur cum plures lentes sint considerandae, pro singulis haec elementa sequentibus litteris indicabo.

| Pro Lente | Distantiae determinatrices |               | Crassities lentis | Quantitas arbitraria | Spatium diffusionis pro aperturae semidiametro $x$ |
|-----------|----------------------------|---------------|-------------------|----------------------|----------------------------------------------------|
|           | obiecti                    | imaginis      |                   |                      |                                                    |
| Prima     | $a$                        | $\alpha$      | $v$               | $k$                  | $P\alpha\alpha xx$                                 |
| Secunda   | $b$                        | $\beta$       | $v^I$             | $k^I$                | $Q\beta\beta xx$                                   |
| Tertia    | $c$                        | $\gamma$      | $v^{II}$          | $k^{II}$             | $R\gamma\gamma xx$                                 |
| Quarta    | $d$                        | $\delta$      | $v^{III}$         | $k^{III}$            | $S\delta\delta xx$                                 |
| Quinta    | $e$                        | $\varepsilon$ | $v^{IV}$          | $k^{IV}$             | $T\varepsilon\varepsilon xx$                       |

Sin autem ratio refractionis in singulis lentibus discrepet, pro prima lente eam ponamus  $= n$ ; pro secunda  $= n^I$ ; pro tertia  $= n^{II}$  etc.

Prima autem lens hic mihi perpetuo erit ea, quae obiecto est proxima indeque recedendo reliquas lentes ordine numero: ex quo simul habebuntur distantiae lentium, scilicet secundae a prima  $= \alpha + b$ , tertiae a secunda  $= \beta + c$ ; quartae a tertia  $= \gamma + d$ ; quintae a quarta  $= \delta + e$  etc. quae distantiae ex sua natura debent esse positivae etiamsi singulae distantiae determinatrices praeter primam  $\alpha$ , quippe quae ad ipsum obiectum necessario ante lentem primam constituendum refertur, sint quandoque negativae. Crassitiem lentium  
hic

hic littera  $v$  designo, quia littera  $d$  inter distantias determinatrices, si numerus lentium ultra ternarium assurgat, reperitur. Pro crassitie autem et quantitate arbitraria iisdem litteris vtor, commatibus inscriptis distinguendis ob penuriam litterarum diuersarum. Ceterum obseruandum est me omnes lentes tanquam super communi axe dispositas assumere.

### Problema 2.

73. Si radii ab obiecto  $E\varepsilon$  emissi per duas lentes  $PP$  et  $QQ$  transmittantur definire spatium diffusionis  $Gg$ , a data apertura primae lentis oriundum; vt et magnitudinem imaginis principalis in  $G$  exhibitae. Tab. I.  
Fig. 5.

### Solutio.

Sit obiecti magnitudo  $E\varepsilon = z$ , eiusque imago principalis per primam lentem  $PP$  proiciatur in  $F\zeta$ , at per ambas in  $G\eta$  erunt distantiae determinatrices pro lente prima  $PP$ , obiecti  $EA = a$ , imaginis  $aF = \alpha$  pro lente secunda  $QQ$ , obiecti  $FB = b$ , imaginis  $bG = \varepsilon$  Deinde fit

pro lente  $PP$ , crassities  $A\alpha = v$ , quantitas arbitr:  $= k$   
 pro lente  $QQ$ , crassities  $Bb = v'$ , quantitas arbitr:  $= k'$   
 Denique pro apertura in anteriori facie vtriusque lentis cuius semidiameter fit  $= x$

spatium diffusionis primae lentis  $PP = P\alpha\alpha xx$

- - - - - secundae lentis  $QQ = Q\varepsilon\varepsilon xx$ .

G 3,

His



His positis pro imagine per primam lentem proiecta  $F\zeta$  erit eius magnitudo  $F\zeta = \frac{\alpha(k+v)}{a(k-v)} z$  pro inuersa habenda, si haec expressio fuerit positua. Deinde si lentis primae PP semidiameter aperturæ in anteriori facie ponatur  $=x$ , erit spatium diffusionis  $Ff = P\alpha\alpha x x$  et inclinatio radiorum in  $f$  ad axem  $= \frac{k-v}{k+v} \cdot \frac{x}{\alpha}$ . tum vero in facie lentis PP posteriori semidiameter aperturæ non minor esse debet quam  $\frac{k-v}{k+v} x$ . Tota iam haec imago diffusa per spatium  $Ff$  respectu alterius lentis QQ tanquam obiectum spectari debet, cuius proinde repræsentatio Gg per propositionem præcedentem determinabitur. Erit autem hic  $\Phi = \frac{k-v}{k+v} \cdot \frac{x}{\alpha}$ , et spatium ibi expressum  $Ee = P\alpha\alpha x x$ ; tum vero pro  $a, \alpha, k, d$  et P hic scribi oportet  $b, \xi, k', v'$  et Q, unde fiet spatium diffusionis quaesitum:

$$Gg = \frac{\xi\xi}{bb} \left( \frac{k'+v'}{k'-v'} \right)^2 P\alpha\alpha x x + \frac{bb\xi\xi}{\alpha\alpha} \left( \frac{k-v}{k+v} \right)^2 Qxx$$

siue

$$Gg = \frac{\xi\xi}{bb} \left( \frac{k'+v'}{k'-v'} \right)^2 P\alpha\alpha x x + \frac{bb}{\alpha\alpha} \left( \frac{k-v}{k+v} \right)^2 Q\xi\xi x x$$

Radiorum porro in g incidentium inclinatio ad axem est  $\left( \frac{k-v}{k+v} \right) \left( \frac{k'+v'}{k'-v'} \right) \frac{bx}{\alpha\xi}$ . Denique cum sit  $F\zeta = \frac{\alpha(k+v)}{a(k-v)} z$  erit magnitudo imaginis in G ut erutae consideratae:

$$G\eta = \frac{\alpha\xi}{ab} \left( \frac{k+v}{k-v} \right) \left( \frac{k'+v'}{k'-v'} \right) z.$$

### COROLL. I.

74. Quod ad aperturam facierum lentis QQ attinet, ea maior esse debet spatio transitus radiorum; hinc ergo erit semidiameter aperturæ

faciei

$$\text{faciei anterioris} > \left(\frac{k-v}{k+v}\right) \frac{bx}{a}$$

$$\text{faciei posterioris} > \left(\frac{k-v}{k+v}\right) \left(\frac{k'-v'}{k'+v'}\right) \frac{bx}{a}$$

## Coroll. 2.

75. Si ponatur pro lente prima PP radius faciei anterioris  $=f$ , et posterioris  $=g$  erit posito  $n = \frac{31}{23}$

$$f = \frac{(n-1)a(k+v)}{k+v+2na} \text{ et } g = \frac{(n-1)a(k-v)}{k-v+2na}$$

Similique modo si pro lente altera QQ radius faciei anterioris ponatur  $=f'$ , et posterioris  $=g'$  erit

$$f' = \frac{(n-1)b(k'+v')}{k'+v'+2nb} \text{ et } g' = \frac{(n-1)b(k'-v')}{k'-v'+2nb}$$

omnibus scilicet faciebus ut conuexis consideratis.

## Coroll. 3.

76. Pro spatio autem diffusionis ab vtraque lente productae erit quemadmodum inuenimus:

$$P = \frac{n}{2(n-1)^2} \cdot \left\{ + \left(\frac{k+v}{k-v}\right)^2 \left(\frac{n}{a} + \frac{2}{k+v}\right) \left(\frac{1}{a} + \frac{2}{k+v}\right)^2 \right\} \\ \left\{ + \left(\frac{k-v}{k+v}\right)^2 \left(\frac{n}{a} - \frac{2}{k-v}\right) \left(\frac{1}{a} - \frac{2}{k-v}\right)^2 \right\}$$

similique modo

$$Q = \frac{n}{2(n-1)^2} \cdot \left\{ + \left(\frac{k'+v'}{k'-v'}\right)^2 \left(\frac{n}{b} + \frac{2}{k'+v'}\right) \left(\frac{1}{b} + \frac{2}{k'+v'}\right)^2 \right\} \\ \left\{ + \left(\frac{k'-v'}{k'+v'}\right)^2 \left(\frac{n}{b} - \frac{2}{k'-v'}\right) \left(\frac{1}{b} - \frac{2}{k'-v'}\right)^2 \right\}$$

Non opus est ut omnibus lentibus eadem refractionis ratio  $n$  tribuatur, sed simili modo pro pluribus lentibus poni potest  $n$ ,  $n^I$ ,  $n^{II}$ ,  $n^{III}$ , etc. ut iam supra monuimus unde nullum aliud discrimen nascitur, nisi



nisi quod in formulis pro  $f^l$  et  $g^l$  inuentis loco  $n$  scribi debeat  $n^l$ , et  $Q$  statui debeat.

$$= \frac{n^l}{2(n^l - 1)^2} \left\{ \begin{aligned} &+ \left( \frac{k' + v'}{k' - v'} \right)^2 \left( \frac{n'}{b} + \frac{2}{k' + v'} \right) \left( \frac{1}{b} + \frac{2}{k' + v'} \right)^2 \\ &+ \left( \frac{k' - v'}{k' + v'} \right)^2 \left( \frac{n'}{b} - \frac{2}{k' - v'} \right) \left( \frac{1}{b} - \frac{2}{k' - v'} \right)^2 \end{aligned} \right\}$$

#### COROLL. 4.

77. Sin autem crassities lentium euanescat, et pro quantitate arbitraria  $k, k^l$  introducamus numerum  $\lambda, \lambda^l$ , erit

pro lente PP

$$f = \frac{a\alpha}{\rho a + \sigma a \pm \tau(a + \alpha)\sqrt{(\lambda - 1)}}; \quad g = \frac{a\alpha}{\rho a + \sigma a \pm \tau(a + \alpha)\sqrt{(\lambda - 1)}}$$

et

$$P = \mu \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{\nu}{a\alpha} \right)$$

at pro lente QQ

$$f^l = \frac{b\beta}{\rho b + \sigma b \pm \tau(b + \beta)\sqrt{(\lambda' - 1)}}; \quad g^l = \frac{b\beta}{\rho b + \sigma b \pm \tau(b + \beta)\sqrt{(\lambda' - 1)}}$$

$$Q = \mu \left( \frac{1}{b} + \frac{1}{\beta} \right) \left( \lambda' \left( \frac{1}{b} + \frac{1}{\beta} \right)^2 + \frac{\nu}{b\beta} \right).$$

In expressionibus autem inuentis formulae  $\left( \frac{k-v}{k+v} \right)$  et  $\left( \frac{k'-v'}{k'+v'} \right)$  abeunt in unitatem.

Numerorum vero  $\mu, \nu, \rho, \sigma, \tau$  indoles in §. 55 expofita est.

Si refractionis lentium differat etiam litterae  $\mu, \nu, \rho, \sigma, \tau$  diuerfos valores pro singulis lentibus fortientur, quae  
litte-

litterae si etiam commatibus a prioribus distinguantur, ut sit

$$\xi' = \frac{1}{2(n' - 1)} + \frac{1}{n' + 2} - 1$$

$$\sigma' = 1 + \frac{1}{2(n' - 1)} - \frac{1}{n' + 2}$$

$$\tau' = \frac{1}{3} \left( \frac{1}{2(n' - 1)} + \frac{1}{n' + 2} \right) \sqrt{(4n' - 1)}$$

pro secunda lente erit

$$f' = \frac{b\xi}{\xi'b + \sigma'b + \tau(b + \xi)\sqrt{(\lambda' - 1)}}$$

$$g' = \frac{b\xi}{\xi'b + \sigma'b + \tau(b + \xi)\sqrt{(\lambda' - 1)}}$$

et ob eandem rationem erit

$$\mu' = \frac{1}{4(n' + 2)} + \frac{1}{4(n' - 1)} + \frac{1}{8(n' + 1)^2}$$

$$\text{et } \nu' = \frac{4(n' - 1)^2}{4n' - 1}$$

quod et de sequentibus lentibus est intelligendum, si forte diuersa refractionis lege fuerint praeditae.

### Scholion.

78. Quo formulas in problemate inuentas magis contrahamus, ut cum ad plures lentes processerimus, succinctiores euadant, ponamus

$$\frac{k-v}{k+v} = i \text{ et } \frac{k'-v'}{k'+v'} = i'$$

ita, ut hi numeri  $i$  et  $i'$  abeant in unitatem euanescente lentium crassitie  $v$  et  $v'$ . Tum autem erit spatium diffusionis

$$Gg = \frac{1}{i'v'} \cdot \frac{\xi\xi}{bb} \cdot Pa\alpha x x + ii \cdot \frac{bb}{\alpha\alpha} \cdot Q\xi\xi x x$$

Tom. I.

H

et



et radiorum in  $g$  incidentium inclinatio ad axem  
 $= i i' \cdot \frac{bx}{\alpha^2}$  : porro magnitudo imaginis  $G \eta = \frac{1}{i' r} \cdot \frac{\alpha^2}{ab} z$

Ac pro apertura singularum facierum erit vt sequitur

| Semidiam. aperturæ<br>lentis | faciei<br>anterioris        | faciei<br>posterioris          |
|------------------------------|-----------------------------|--------------------------------|
| primæ PP                     | $x$                         | $i x$                          |
| secundæ QQ                   | $i \cdot \frac{bx}{\alpha}$ | $i i' \cdot \frac{bx}{\alpha}$ |

aperturæ autem hæ præter primam maiores esse debent assignatis : valores enim assignati sufficerent , si obiectum esset merum punctum in axe positum ; quando autem habet magnitudinem , radii ab eius terminis per primam faciem ingressi latius diuagantur , et ad sequentes facies maiorem aperturam exigunt.

### Problema 3.

Tab. II.  
Fig. 6.

79. Si radii ab obiecto  $E\epsilon$  emissi per tres lentes PP, QQ et RR refringantur definire spatium diffusionis  $Hb$  ob datam aperturam lentis primæ oriundum , vt et magnitudinem imaginis principalis in  $H$  exhibitæ.

### Solutio.

Posita obiecti magnitudine  $E\epsilon = z$ , cadat eius , imago principalis per lentem primam PP in  $F\zeta$ , dehinc per secundam QQ in  $G\eta$ , tum vero per tertiam RR in  $H\theta$ . Erunt ergo distantiae determinatrices.

pro lente PP, obiecti  $EA = a$ , imaginis  $aF = \alpha$

pro lente QQ, obiecti  $FB = b$ , imaginis  $bG = \beta$

pro lente RR, obiecti  $GC = c$ ; imaginis  $cH = \gamma$

Deinde

Deinde fit

pro lente PP, crassities  $Aa=v$ , quantitas arbitr:  $=k$   
 pro lente QQ, crassities  $Bb=v'$ , quantitas arbitr:  $=k'$   
 pro lente RR crassities  $Cc=v''$ , quantitas arbitr:  $=k''$   
 ac ponatur breuitatis gratia:

$$\frac{k-v}{k+v} = i; \frac{k'-v'}{k'+v'} = i'; \frac{k''-v''}{k''+v''} = i''$$

Denique pro quavis lente, si esset singularis, eiusque  
 aperturae semidiameter esset  $=x$ , fit

$$\begin{aligned} \text{spatium diffusionis primae lentis PP} &= Paaxx \\ \text{secundae lentis QQ} &= Q\epsilon\epsilon xx \\ \text{tertiaе lentis RR} &= R\gamma\gamma xx \end{aligned}$$

Iam in praecedente problemate inuenimus fore spatium  
 diffusionis per duas lentes priores productum

$$Gg = \frac{1}{i'}. \frac{\epsilon\epsilon}{bb} Paaxx + ii. \frac{bb}{aa}. Q\epsilon\epsilon xx$$

et radiorum in g concurrentium inclinationem ad axem  
 $= ii'. \frac{bx}{a\epsilon}$  imaginisque in G magnitudinem  $G\eta = \frac{1}{i'}. \frac{a\epsilon}{ab} z$ .

Haec igitur si ad problema I. transferantur, ibi loco

$$Ee, \frac{aa}{aq}, \left(\frac{k+d}{k-d}\right)^2, Paaxx \text{ et } \Phi \text{ scribi debet}$$

$$Gg; \frac{\gamma\gamma}{cc}; \frac{1}{i''i'}, Rcc\gamma\gamma \text{ et } ii'. \frac{bx}{a\epsilon},$$

hincque spatium diffusionis per tres lentes productum  
 orietur.

$$\begin{aligned} Hh &= \frac{1}{i'i'.i''i'}. \frac{\epsilon\epsilon\gamma\gamma}{bbcc} Paaxx + \frac{ii}{i''i'}. \frac{bb\gamma\gamma}{aacc}. Q\epsilon\epsilon xx \\ &\quad + ii. i' i'. \frac{bbcc}{aa\epsilon\epsilon}. R\gamma\gamma xx \end{aligned}$$

at radiorum in h concurrentium inclinatio ad axem

H 2

erit



erit  $= i i' i'' \cdot \frac{b c x}{\alpha \beta \gamma}$ . Denique imaginis in H formatae magnitudo erit  $H \theta = \frac{1}{i i' i''} \cdot \frac{\alpha \beta \gamma}{a b c} x$  ad situm inuersum relata.

### C o r o l l. 1.

80. Quod ad aperturas singularum facierum attinet, eas sequenti modo comparatas esse conuenit.

| Semid. aperturæ | faciei<br>anterioris                    | faciei<br>posterioris                              |
|-----------------|-----------------------------------------|----------------------------------------------------|
| Lentis          |                                         |                                                    |
| Primæ P P       | $x$                                     | $i x$                                              |
| Secundæ Q Q     | $i \cdot \frac{b x}{\alpha}$            | $i i' \cdot \frac{b x}{\alpha}$                    |
| Tertiæ R R      | $i i' \cdot \frac{b c x}{\alpha \beta}$ | $i i' i'' \cdot \frac{b c x}{\alpha \beta \gamma}$ |

his scilicet valoribus non debent esse minores.

### C o r o l l. 2.

81. Si inclinatio radiorum in h concurrentium ad axem vocetur  $= \Phi$ . Cum sit  $\Phi = i i' i'' \cdot \frac{b c x}{\alpha \beta \gamma}$  et  $H \theta = \frac{1}{i i' i''} \cdot \frac{\alpha \beta \gamma}{a b c} x$  erit  $\Phi \cdot H \theta = \frac{x x}{a}$ : quae proprietas pro lentium numero quantumuis magno locum habet.

### C o r o l l. 3.

82. Quemadmodum radii singularum facierum determinandi sint, ex praecedentibus facile liquet:

| Erit nempe<br>pro | Radius faciei<br>anterioris           | posterioris                                     |
|-------------------|---------------------------------------|-------------------------------------------------|
| Lente prima P P   | $\frac{(n-1)a(k+v)}{k+v+2na}$         | $\frac{(n-1)\alpha(k-v)}{k-v-2n\alpha}$         |
| Lente secunda Q Q | $\frac{(n-1)b(k'+v')}{k'+v'+2nb}$     | $\frac{(n-1)\beta(k'-v')}{k'-v'-2n\beta}$       |
| Lente tertia R R  | $\frac{(n-1)c(k''+v'')}{k''+v''+2nc}$ | $\frac{(n-1)\gamma(k''-v'')}{k''-v''-2n\gamma}$ |

existente pro vitro  $n = \frac{31}{20}$ .

Coroll.

## COROLL. 4.

83. Tum vero valores literarum P, Q, R ita se habebunt

$$P = \frac{n}{2(n-1)^2} \left( \frac{1}{ii} \left( \frac{n}{a} + \frac{2}{k+v} \right) \left( \frac{1}{a} + \frac{2}{k+v} \right)^2 + ii \left( \frac{n}{\alpha} - \frac{2}{k-v} \right) \left( \frac{1}{\alpha} - \frac{2}{k-v} \right)^2 \right)$$

$$Q = \frac{n}{2(n-1)^2} \left( \frac{1}{i'i'} \left( \frac{n}{b} + \frac{2}{k'+v'} \right) \left( \frac{1}{b} + \frac{2}{k'+v'} \right)^2 + i'i' \left( \frac{n}{\beta} - \frac{2}{k'-v'} \right) \left( \frac{1}{\beta} - \frac{2}{k'-v'} \right)^2 \right)$$

$$R = \frac{n}{2(n-1)^2} \left( \frac{1}{i''i'''} \left( \frac{n}{c} + \frac{2}{k''+v''} \right) \left( \frac{1}{c} + \frac{2}{k''+v''} \right)^2 + i''i''' \left( \frac{n}{\gamma} - \frac{2}{k''-v''} \right) \left( \frac{1}{\gamma} - \frac{2}{k''-v''} \right)^2 \right)$$

quibus valoribus spatia diffusionis definiuntur.

Si refractio differat :

| litteris | tribuatur refr. |
|----------|-----------------|
| P        | $n$             |
| Q        | $n'$            |
| R        | $n''$           |

## COROLL. 5.

84. Iuvabit etiam spatia diffusionis, prout per unam, duas ac tres lentes producuntur, inter se comparasse, quod ita commodissime fieri videtur.

$$Ff = \alpha \alpha x x. P$$

$$Gg = \beta \beta x x \left( \frac{1}{i'i'} \cdot \frac{\alpha \alpha}{\beta \beta} P + ii \cdot \frac{\beta \beta}{\alpha \alpha} Q \right)$$

$$Hh = \gamma \gamma x x \left( \frac{1}{i''i'''} \cdot \frac{\alpha \alpha \beta \beta}{\beta \beta \gamma \gamma} P + \frac{ii}{i''i'''} \cdot \frac{\beta \beta \gamma \gamma}{\alpha \alpha \gamma \gamma} Q + ii \cdot i'i' \cdot \frac{\beta \beta \gamma \gamma}{\alpha \alpha \gamma \gamma} R \right)$$

## SCHOLIUM.

85. Hinc facile perspicitur, quomodo determinatio spatii diffusionis, ad plures lentes extendi debeat;

H. 3.

vnde



vnde problema generale statim ad numerum lentium quemcunque accommodabo, idque pro quacunque lentium crassitie.

### Problema 4.

86. Si radii ab obiecto  $E\epsilon$  emissi per lentes quocunque PP, QQ, RR, SS etc. super communi axe dispositas refringantur, definire spatium diffusionis a data apertura lentis primae oriundum vt et magnitudinem imaginis repraesentatae.

### Solutio.

Posita obiecti magnitudine  $E\epsilon = z$  imagines eius principales contemplemur: cadat igitur eius imago principalis per lentem primam PP in  $F\zeta$  deinde per secundam QQ in  $G\eta$ , tum per tertiam RR in  $H\theta$ , porro per quartam SS in  $I\iota$ , per quintam TT in  $K\kappa$  etc. Hinc pro singulis lentibus habebimus distantias determinatrices quas ita litteris indicemus.

crassitiem

Pro lente PP obiecti  $EA = a$ , imaginis  $aF = \alpha$ ;  $Aa = v$

Pro lente QQ obiecti  $FB = b$ ; imaginis  $bG = \beta$ ;  $Bb = v^I$

Pro lente RR obiecti  $GC = c$ ; imaginis  $cH = \gamma$ ;  $Cc = v^{II}$

Pro lente SS obiecti  $HD = d$ ; imaginis  $dI = \delta$ ;  $Dd = v^{III}$

Pro lente TT obiecti  $IE = e$ ; imaginis  $eK = \epsilon$ ;  $Ee = v^{IV}$

etc.

Deinde

Deinde cum determinatio cuiusvis lentis non solum has distantias determinatrices cum crassitie cuiusque sed etiam quantitatem quandam arbitrariam inuoluat, a qua spatium diffusionis cuiusque penderet, ponamus si quaelibet lens esset solitaria eiusque aperturæ semidiameter  $= x$ :

| Pro lente  | quant :<br>arbitr : | Spatium<br>Diffusionis       |
|------------|---------------------|------------------------------|
| Prima PP   | $k$                 | $P\alpha\alpha xx$           |
| Secunda QQ | $k^I$               | $Q\beta\beta xx$             |
| Tertia RR  | $k^{II}$            | $R\gamma\gamma xx$           |
| Quarta SS  | $k^{III}$           | $S\delta\delta xx$           |
| Quinta TT  | $k^{IIII}$          | $T\varepsilon\varepsilon xx$ |
|            | etc.                |                              |

Hinc ergo constructio cuiusque lentis ita se habebit posito  $n = \frac{31}{20}$ : vel si refractio differat, cuique lenti suus tribuatur valor primæ  $n$ , secundæ  $n^I$ , tertiæ  $n^{II}$  etc.

| Erit nempe       | Radius faciei                                                   |                                                                                     |
|------------------|-----------------------------------------------------------------|-------------------------------------------------------------------------------------|
| pro              | anterioris                                                      | posterioris                                                                         |
| Lente prima PP   | $\frac{(n-1)a(k+v)}{k+v+2na}$                                   | $\frac{(n-1)\alpha(k-v)}{k-v-2n\alpha}$                                             |
| Lente secunda QQ | $\frac{(n-1)b(k'+v')}{k'+v'+2nb}$                               | $\frac{(n-1)\beta(k'-v')}{k'-v'-2n\beta}$                                           |
| Lente tertia RR  | $\frac{(n-1)c(k''+v'')}{k''+v''+2nc}$                           | $\frac{(n-1)\gamma(k''-v'')}{k''-v''-2n\gamma}$                                     |
| Lente quarta SS  | $\frac{(n-1)d(k''' + v''')}{k''' + v''' + 2nd}$                 | $\frac{(n-1)\delta(k''' - v''')}{k''' - v''' - 2n\delta}$                           |
| Lente quinta TT  | $\frac{(n-1)e(k^{IIII} + v^{IIII})}{k^{IIII} + v^{IIII} + 2ne}$ | $\frac{(n-1)\varepsilon(k^{IIII} - v^{IIII})}{k^{IIII} - v^{IIII} - 2n\varepsilon}$ |
|                  | etc.                                                            |                                                                                     |

at



at si ponamus breuitatis gratia :

$$\frac{k_1 - v}{k + v} = i; \frac{k' - v'}{k' + v'} = i'; \frac{k'' - v''}{k'' + v''} = i''; \frac{k''' - v'''}{k''' + v'''} = i'''; \frac{k'''' - v''''}{k'''' + v''''} = i'''' \text{ etc.}$$

pro spatiis diffusionis habebimus hos valores :

$$P = \frac{n}{2(n-1)^2} \left( \frac{1}{i i'} \left( \frac{n}{a} + \frac{2}{k + v} \right) \left( \frac{1}{a} + \frac{2}{k + v} \right)^2 + i i' \left( \frac{n}{a} - \frac{2}{k - v} \right) \left( \frac{1}{a} - \frac{2}{k - v} \right)^2 \right)$$

$$Q = \frac{n}{2(n-1)^2} \left( \frac{1}{i' i''} \left( \frac{n}{b} + \frac{2}{k' + v'} \right) \left( \frac{1}{b} + \frac{2}{k' + v'} \right)^2 + i' i'' \left( \frac{n}{b} - \frac{2}{k' - v'} \right) \left( \frac{1}{b} - \frac{2}{k' - v'} \right)^2 \right)$$

$$R = \frac{n}{2(n-1)^2} \left( \frac{1}{i'' i'''} \left( \frac{n}{c} + \frac{2}{k'' + v''} \right) \left( \frac{1}{c} + \frac{2}{k'' + v''} \right)^2 + i'' i''' \left( \frac{n}{c} - \frac{2}{k'' - v''} \right) \left( \frac{1}{c} - \frac{2}{k'' - v''} \right)^2 \right)$$

$$S = \frac{n}{2(n-1)^2} \left( \frac{1}{i''' i''''} \left( \frac{n}{d} + \frac{2}{k''' + v'''} \right) \left( \frac{1}{d} + \frac{2}{k''' + v'''} \right)^2 + i''' i'''' \left( \frac{n}{d} - \frac{2}{k''' - v'''} \right) \left( \frac{1}{d} - \frac{2}{k''' - v'''} \right)^2 \right)$$

$$T = \frac{n}{2(n-1)^2} \left( \frac{1}{i'''' i'''''} \left( \frac{n}{e} + \frac{2}{k'''' + v''''} \right) \left( \frac{1}{e} + \frac{2}{k'''' + v''''} \right)^2 + i'''' i''''' \left( \frac{n}{e} - \frac{2}{k'''' - v''''} \right) \left( \frac{1}{e} - \frac{2}{k'''' - v''''} \right)^2 \right)$$

etc.

His constitutis pro magnitudine singularum imaginum habebimus :

ad situm

pro vna lente imaginem  $F \zeta = \frac{1}{i} \cdot \frac{\alpha}{a} z$  inuersum

pro duabus lentibus . . . .  $G \eta = \frac{1}{i i'} \cdot \frac{\alpha \beta}{ab} z$  erectum

pro tribus lentibus . . . .  $H \theta = \frac{1}{i' i''} \cdot \frac{\alpha \beta \gamma}{a b c} z$  inuersum

pro quatuor lentibus . . . .  $I \iota = \frac{1}{i'' i''' i''''} \cdot \frac{\alpha \beta \gamma \delta}{a b c d} z$  erectum

pro quinque lentibus . . . .  $K \kappa = \frac{1}{i''' i'''' i''''' i'''''} \cdot \frac{\alpha \beta \gamma \delta \epsilon}{a b c d e} z$  inuersum

etc.

Deni-

Denique dum aperturæ lentium non sint minores, quam sequentes formulæ exhibent:

| Semid. aperturæ<br>Lentis | faciei<br>anterioris                                         | faciei<br>posterioris                                              |
|---------------------------|--------------------------------------------------------------|--------------------------------------------------------------------|
| primæ PP                  | $x$                                                          | $i x$                                                              |
| secundæ QQ                | $i. \frac{bx}{\alpha}$                                       | $i i'. \frac{bx}{\alpha}$                                          |
| tertiæ RR                 | $i i'. \frac{bcx}{\alpha \epsilon}$                          | $i i' i''. \frac{bcx}{\alpha \epsilon}$                            |
| quartæ SS                 | $i i' i''. \frac{bcdx}{\alpha \epsilon \gamma}$              | $i i' i'' i'''. \frac{bcdx}{\alpha \epsilon \gamma}$               |
| quintæ TT                 | $i i' i'' i'''. \frac{bcdex}{\alpha \epsilon \gamma \delta}$ | $i i' i'' i''' i''''. \frac{bcdex}{\alpha \epsilon \gamma \delta}$ |
|                           | etc.                                                         |                                                                    |

erit vt sequitur pro quolibet lentium numero:

#### I. Pro vna lente

Spatium diffusionis  $Ff = \alpha \alpha x x$ . P  
inclinatio radiorum in  $f$  concurrentium ad axem  $= i. \frac{x}{\alpha}$

#### II. Pro duabus lentibus

Spatium diffusionis

$Gg = \epsilon \epsilon x x \left( \frac{1}{i' i''} \cdot \frac{\alpha \alpha}{b b} P + i i'. \frac{b b}{\alpha \alpha} Q \right)$   
et radiorum in  $g$  concurrentium  
inclinatio ad axem  $= i i'. \frac{bx}{\alpha \epsilon}$

#### III. Pro tribus lentibus

Spatium diffusionis:

$Hh = \gamma \gamma x x \left( \frac{1}{i' i'' i'''} \cdot \frac{\alpha \alpha \epsilon \epsilon}{b b c c} P + \frac{i i'}{i'' i'''} \cdot \frac{b b \epsilon \epsilon}{\alpha \alpha c c} Q + i i'. i' i''. \frac{b b c c}{\alpha \alpha \epsilon \epsilon} R \right)$   
et radiorum in  $h$  concurrentium  
inclinatio ad axem  $= i i' i''. \frac{bcx}{\alpha \epsilon \gamma}$

Tom. I.

I

IV.



## IV. Pro quatuor lentibus

Spatium diffusionis:

$$Ii = \delta\delta xx \left\{ \begin{array}{l} \frac{1}{i' i'' i''' i'''} \cdot \frac{\alpha\alpha\epsilon\epsilon\gamma\gamma}{bbccdd} P + \frac{ii}{i'' i''' i'''} \cdot \frac{bb\epsilon\epsilon\gamma\gamma}{\alpha\alpha ccdd} Q \\ + \frac{ii \cdot i' i''}{i''' i'''} \cdot \frac{bbcc\gamma\gamma}{\alpha\alpha\epsilon\epsilon ad} R + ii \cdot i' i'' \cdot i''' i'''' \cdot \frac{bbccdd}{\alpha\alpha\epsilon\epsilon\gamma\gamma} S \end{array} \right\}$$

et radiorum in  $i$  concurrentium.

$$\text{inclinatio ad axem} = ii' i'' i''' \cdot \frac{bcdx}{\alpha\epsilon\gamma\delta}$$

## V. Pro quinque lentibus

Spatium diffusionis:

$$Kk = \epsilon\epsilon xx \left\{ \begin{array}{l} + \frac{1}{i' i'' i''' i'''} \cdot \frac{\alpha\alpha\epsilon\epsilon\gamma\gamma\delta\delta}{bbccdde\epsilon} P \\ + \frac{ii}{i'' i''' i'''} \cdot \frac{bb\epsilon\epsilon\gamma\gamma\delta\delta}{\alpha\alpha ccdde\epsilon} Q \\ + \frac{ii \cdot i' i''}{i''' i'''} \cdot \frac{bbcc\gamma\gamma\delta\delta}{\alpha\alpha\epsilon\epsilon dde\epsilon} R \\ + \frac{ii \cdot i' i'' \cdot i''' i'''}{i'''' i''''} \cdot \frac{bbccdd\delta\delta}{\alpha\alpha\epsilon\epsilon\gamma\gamma ee} S \\ + ii \cdot i' i'' \cdot i''' i'''' \cdot \frac{bbccdde\epsilon}{\alpha\alpha\epsilon\epsilon\gamma\gamma\delta\delta} T \end{array} \right\}$$

et radiorum in  $k$  concurrentium

$$\text{inclinatio ad axem} = ii' i'' i''' i'''' \cdot \frac{bcde x}{\alpha\epsilon\gamma\delta\epsilon}$$

Vnde progressio harum formularum ad plures adhuc lentes satis est manifesta. Si in lentibus ratio refractionis sit diuersa atque ad singulas lentes ordine his litteris indicetur  $n, n', n'', n''',$  etc.; haec diuersitas facile ad formulas hic inuentas accommodabitur. Primum enim haec correctio occurrit in formulis pro radiis lentium,

lentium, ita, ut quemadmodum formulae  $f$  et  $g$  pro prima lente numerum  $n$  inuolunt, ita pro secunda lente numerus  $n^I$ , pro tertia  $n^{II}$  et ita porro introducatur. Similem correctionem etiam requirunt valores litterarum  $P, Q, R, S$  etc. et loco litterae  $n$ , quae in valore  $P$  occurrit, in valoribus  $Q, R, S$  etc. scribi oportet  $n^I, n^{II}, n^{III}$  etc.

### C o r o l l. 1.

87. Si obiectum sit tantum punctum in axe positum, sufficit ut lentium aperturae sint tantae, quantas assignauimus sin autem obiectum habeat quandam magnitudinem, tum aperturae praeter primam eo magis mensuras assignatas superare debent, quo maior fuerit obiecti magnitudo  $z$ .

### C o r o l l. 2.

88. In expressione spatii diffusionis quadratum semidiametri aperturae primae faciei  $xx$  primo multiplicatur per quadratum distantiae postremae imaginis ab vltima lentium: quae ergo si fuerit infinita, etiam spatium diffusionis fit infinitum.

### C o r o l l. 3.

89. Ceteris ergo paribus, quotcunque fuerint lentes, spatium diffusionis semper est proportionale quadrato diametri aperturae primae faciei, hoc est ipsi huic aperturae. Vnde diametro aperturae primae



faciei ad semiffem redactō spatium diffusionis quadruplo fiet minus.

### Scholiō.

90. Considerauimus hic statim loca singularum imaginum principalium tanquam data ex iisque structuram cuiusque lentis quantitatem arbitrariam introducendo determinauimus. Quod si vero ipsae lentes fuerint datae ita vt tam radii ambarum facierum cuiusque quam crassities, vna cum earum interuallis cognoscantur tum ope formularum exhibitarum, vicissim distantiae determinatrices innotescunt. Sint scilicet radii facierum anterioris et posterioris primae lentis PP,  $f, g$ , secundae lentis QQ,  $f', g'$ ; tertiae lentis RR,  $f'', g''$  etc. crassitie earum existente  $v, v', v''$  etc. tum verodentur distantiae  $aB = F; bC = G; cD = H$  etc. Praeterea autem distantia obiecti ante lentem primam sit  $AE = a$ , ac sequenti modo omnia elementa ad problema superius necessaria elicientur;

$$\left\{ \begin{array}{l} 1. \frac{n-1}{f} = \frac{1}{a} + \frac{2n}{k+v}, \text{ hinc reperitur } k \end{array} \right.$$

$$\left\{ \begin{array}{l} 2. \frac{n-1}{g} = \frac{1}{a} - \frac{2n}{k-v}, \text{ hinc vero } a \end{array} \right.$$

$$\left\{ \begin{array}{l} 3. F = a + b \text{ vnde } b = F - a \end{array} \right.$$

$$\left\{ \begin{array}{l} 4. \frac{n-1}{f'} = \frac{1}{b} + \frac{2n}{k'+v'}, \text{ hinc reperitur } k' \end{array} \right.$$

$$\left\{ \begin{array}{l} 5. \frac{n-1}{g'} = \frac{1}{b} - \frac{2n}{k'-v'}, \text{ hinc vero porro } b \end{array} \right.$$

$$\left\{ \begin{array}{l} 6. G = b + c \text{ vnde } c = G - b \end{array} \right.$$

(7.

$$\left\{ \begin{array}{l} 7. \frac{n-1}{f''} = \frac{1}{c} + \frac{2n}{k''+v''}, \text{ hinc reperitur } k'' \\ 8. \frac{n-1}{g''} = \frac{1}{\gamma} - \frac{2n}{k''-v''}, \text{ hinc vero } \gamma \\ 9. H = \gamma + d \text{ unde } d = H - \gamma \end{array} \right.$$

etc.

Vtcunque ergo lentes datae fuerint dispositae super axe communi, si ante eas constituatur obiectum in data distantia  $AE = a$ , inde singulae distantiae determinatrices  $\alpha, b, \xi, c, \gamma$  etc. cum arbitrariis  $k, k', k'', k'''$  etc. facile definiuntur ex iisque porro spatium diffusionis cum reliquis phaenomenis in solutione problematis commemoratis assignabitur. Operae pretium autem erit casum, quo crassities lentium vt euanescens spectatur, accuratius euoluiffe.

### Problema 5.

91. Si crassities lentium euanescat, et quotcunque huiusmodi lentes super communi axe fuerint dispositae ante quas existat obiectum  $E\varepsilon$ , definire spatium diffusionis, per quod imago erit dissipata, vt et magnitudinem imaginis.

### Solutio.

Si obiecti magnitudo  $E\varepsilon = z$ , cuius imagines principales successiue cadant in  $F\zeta, G\eta, H\theta, I\iota, K\kappa$  etc. hincque pro singulis lentibus sequentes habebimus distantias determinatrices, cum imago per quamuis lentem repraesentata respectu lentis sequentis vicem obiecti gerat.

I 3

distan-



|              | distantia          |          | distantia          |
|--------------|--------------------|----------|--------------------|
| Pro lente PP | obiecti EA = $a$ ; | imaginis | aF = $\alpha$      |
| Pro lente QQ | obiecti FB = $b$ ; | imaginis | bG = $\beta$       |
| Pro lente RR | obiecti GC = $c$ ; | imaginis | cH = $\gamma$      |
| Pro lente SS | obiecti HD = $d$ ; | imaginis | dI = $\delta$      |
| Pro lente TT | obiecti IE = $e$ ; | imaginis | eK = $\varepsilon$ |
|              | etc.               |          |                    |

Porro autem sint numeri arbitrarii unitate maiores cuiusque lentis figuram determinantes,  $\lambda$  pro lente PP,  $\lambda'$  pro QQ,  $\lambda''$  pro RR,  $\lambda'''$  pro SS,  $\lambda''''$  pro TT, etc. ita ut ponendo breuitatis gratia

$$\rho = 0,190781, \sigma = 1,627401, \tau = 0,905133$$

$$\begin{aligned} \text{Pro lente PP radius faciei} & \begin{cases} \text{anter.} = \frac{a\alpha}{\rho\alpha + \sigma\alpha + \tau(a+\alpha)\sqrt{(\lambda-1)}} \\ \text{poster.} = \frac{a\alpha}{\rho\alpha + \sigma\alpha + \tau(a+\alpha)\sqrt{(\lambda-1)}} \end{cases} \\ \text{Pro lente QQ radius faciei} & \begin{cases} \text{anter.} = \frac{b\beta}{\rho b + \sigma b + \tau(b+\beta)\sqrt{(\lambda'-1)}} \\ \text{poster.} = \frac{b\beta}{\rho b + \sigma b + \tau(b+\beta)\sqrt{(\lambda'-1)}} \end{cases} \\ \text{Pro lente RR radius faciei} & \begin{cases} \text{anter.} = \frac{c\gamma}{\rho c + \sigma c + \tau(c+\gamma)\sqrt{(\lambda''-1)}} \\ \text{poster.} = \frac{c\gamma}{\rho c + \sigma c + \tau(c+\gamma)\sqrt{(\lambda''-1)}} \end{cases} \\ \text{Pro lente SS radius faciei} & \begin{cases} \text{anter.} = \frac{d\delta}{\rho d + \sigma d + \tau(d+\delta)\sqrt{(\lambda'''-1)}} \\ \text{poster.} = \frac{d\delta}{\rho d + \sigma d + \tau(d+\delta)\sqrt{(\lambda'''-1)}} \end{cases} \\ \text{Pro lente TT radius faciei} & \begin{cases} \text{anter.} = \frac{e\varepsilon}{\rho e + \sigma e + \tau(e+\varepsilon)\sqrt{(\lambda''''-1)}} \\ \text{poster.} = \frac{e\varepsilon}{\rho e + \sigma e + \tau(e+\varepsilon)\sqrt{(\lambda''''-1)}} \end{cases} \\ & \text{etc.} \end{aligned}$$

Deinde

Deinde si quaelibet lens cum binis suis distantis determinatricibus seorsim consideretur, eiusque aperturæ semidiameter foret  $= x$ , posito  $\mu = 0,938191$  et  $\nu = 0,232692$  esset spatium diffusionis

$$\text{Lentis PP} = \mu \alpha \alpha x x \left( \frac{1}{a} + \frac{1}{a} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{a} \right)^2 + \frac{\nu}{a \alpha} \right)$$

$$\text{Lentis QQ} = \mu \beta \beta x x \left( \frac{1}{b} + \frac{1}{b} \right) \left( \lambda' \left( \frac{1}{b} + \frac{1}{b} \right)^2 + \frac{\nu}{b \beta} \right)$$

$$\text{Lentis RR} = \mu \gamma \gamma x x \left( \frac{1}{c} + \frac{1}{c} \right) \left( \lambda'' \left( \frac{1}{c} + \frac{1}{c} \right)^2 + \frac{\nu}{c \gamma} \right)$$

$$\text{Lentis SS} = \mu \delta \delta x x \left( \frac{1}{d} + \frac{1}{d} \right) \left( \lambda''' \left( \frac{1}{d} + \frac{1}{d} \right)^2 + \frac{\nu}{d \delta} \right)$$

$$\text{Lentis TT} = \mu \epsilon \epsilon x x \left( \frac{1}{e} + \frac{1}{e} \right) \left( \lambda'''' \left( \frac{1}{e} + \frac{1}{e} \right)^2 + \frac{\nu}{e \epsilon} \right)$$

etc.

His constitutis pro magnitudine singularum imaginum habebimus

$$\text{Pro vna lente F} \zeta = \frac{\alpha}{a} z \text{ situ inuerso}$$

$$\text{Pro duabus lentibus G} \eta = \frac{\alpha \beta}{a b} z \text{ situ erecto}$$

$$\text{Pro tribus lentibus H} \theta = \frac{\alpha \beta \gamma}{a b c} z \text{ situ inuerso}$$

$$\text{Pro quatuor lentibus I} \iota = \frac{\alpha \beta \gamma \delta}{a b c d} z \text{ situ erecto}$$

$$\text{Pro quinque lentibus K} \kappa = \frac{\alpha \beta \gamma \delta \epsilon}{a b c d e} z \text{ situ inuerso}$$

etc.

At si semidiameter aperturæ primæ lentis PP ponatur  $= x$ , necesse est, vt reliquarum lentium aperturæ superent sequentes valores

Semi-



Semidiameter aperturæ

Lentis secundæ  $QQ > \frac{b}{a} x$

Lentis tertiæ  $RR > \frac{bc}{a\epsilon} x$

Lentis quartæ  $SS > \frac{bcd}{a\epsilon\gamma} x$

Lentis quintæ  $TT > \frac{bcde}{a\epsilon\gamma\delta} x$

etc.

Hinc spatium diffusionis pro quolibet lentium numero ita se habebit.

### I. Pro vna lente

spatium diffusionis

$$Ff = \mu a a x x \left( \frac{1}{a} + \frac{1}{a} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{a} \right)^2 + \frac{v}{a a} \right)$$

radiatorum in  $f$  concurrentium

inclinatio ad axem  $= \frac{x}{a}$

### II. Pro duabus lentibus

spatium diffusionis

$$Gg = \mu \epsilon \epsilon x x \left\{ \begin{aligned} &+ \frac{a a}{b b} \left( \frac{1}{a} + \frac{1}{a} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{a} \right)^2 + \frac{v}{a a} \right) \\ &+ \frac{b b}{a a} \left( \frac{1}{b} + \frac{1}{\epsilon} \right) \left( \lambda' \left( \frac{1}{b} + \frac{1}{\epsilon} \right)^2 + \frac{v}{b \epsilon} \right) \end{aligned} \right.$$

et radiatorum in  $g$  concurrentium

inclinatio ad axem  $= \frac{b x}{a \epsilon}$

### III. Pro tribus lentibus

spatium diffusionis

$$Hh = \mu \gamma \gamma x x \left\{ \begin{aligned} &+ \frac{a a \epsilon \epsilon}{b b c c} \left( \frac{1}{a} + \frac{1}{a} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{a} \right)^2 + \frac{v}{a a} \right) \\ &+ \frac{b b \epsilon \epsilon}{a a c c} \left( \frac{1}{b} + \frac{1}{\epsilon} \right) \left( \lambda' \left( \frac{1}{b} + \frac{1}{\epsilon} \right)^2 + \frac{v}{b \epsilon} \right) \\ &+ \frac{b b c c}{a a \epsilon \epsilon} \left( \frac{1}{c} + \frac{1}{\gamma} \right) \left( \lambda'' \left( \frac{1}{c} + \frac{1}{\gamma} \right)^2 + \frac{v}{c \gamma} \right) \end{aligned} \right.$$

et

et radiorum in  $b$  concurrentium

inclinatio ad axem  $\frac{bcx}{a\epsilon\gamma}$

#### IV. Pro quatuor lentibus

spatium diffusionis

$$Ii = \mu \delta \delta x x \left\{ \begin{array}{l} + \frac{a\alpha\epsilon\epsilon\gamma\gamma}{bbccdd} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{v}{a\alpha} \right) \\ + \frac{bb\epsilon\epsilon\gamma\gamma}{a\alpha ccdd} \left( \frac{1}{b} + \frac{1}{\epsilon} \right) \left( \lambda' \left( \frac{1}{b} + \frac{1}{\epsilon} \right)^2 + \frac{v}{b\epsilon} \right) \\ + \frac{bbcc\gamma\gamma}{a\alpha\epsilon\epsilon dd} \left( \frac{1}{c} + \frac{1}{\gamma} \right) \left( \lambda'' \left( \frac{1}{c} + \frac{1}{\gamma} \right)^2 + \frac{v}{c\gamma} \right) \\ + \frac{bbccdd}{a\alpha\epsilon\epsilon\gamma\gamma} \left( \frac{1}{d} + \frac{1}{\delta} \right) \left( \lambda''' \left( \frac{1}{d} + \frac{1}{\delta} \right)^2 + \frac{v}{d\delta} \right) \end{array} \right.$$

et radiorum in  $i$  concurrentium

inclinatio ad axem  $= \frac{bcdx}{a\epsilon\gamma\delta}$

#### V. Pro quinque lentibus

spatium diffusionis

$$Kk = \mu \epsilon \epsilon x x \left\{ \begin{array}{l} + \frac{a\alpha\epsilon\epsilon\gamma\gamma\delta\delta}{ouccdde\epsilon} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{v}{a\alpha} \right) \\ + \frac{bb\epsilon\epsilon\gamma\gamma\delta\delta}{a\alpha ccade\epsilon} \left( \frac{1}{b} + \frac{1}{\epsilon} \right) \left( \lambda' \left( \frac{1}{b} + \frac{1}{\epsilon} \right)^2 + \frac{v}{b\epsilon} \right) \\ + \frac{bbcc\gamma\gamma\delta\delta}{a\alpha\epsilon\epsilon dde\epsilon} \left( \frac{1}{c} + \frac{1}{\gamma} \right) \left( \lambda'' \left( \frac{1}{c} + \frac{1}{\gamma} \right)^2 + \frac{v}{c\gamma} \right) \\ + \frac{bbccdd\delta\delta}{a\alpha\epsilon\epsilon\gamma\gamma ee} \left( \frac{1}{d} + \frac{1}{\delta} \right) \left( \lambda''' \left( \frac{1}{d} + \frac{1}{\delta} \right)^2 + \frac{v}{d\delta} \right) \\ + \frac{bbccdde\epsilon}{a\alpha\epsilon\epsilon\gamma\gamma\delta\delta} \left( \frac{1}{e} + \frac{1}{\epsilon} \right) \left( \lambda'''' \left( \frac{1}{e} + \frac{1}{\epsilon} \right)^2 + \frac{v}{e\epsilon} \right) \end{array} \right.$$

et radiorum in  $k$  concurrentium

inclinatio ad axem  $= \frac{bcdex}{a\epsilon\gamma\delta\epsilon}$

neque ergo casus, quibus plures occurrunt lentes, vlla amplius laborant difficultate.

Si lentes ratione refractionis discrepent, ad easque referendae sint litterae  $n, n', n'', n'''$ , etc. formu-

Tom. I.

K

lae



lae in hoc problemate inuentae sequenti modo facile ad hunc casum latius patentem adcommo-  
dabuntur. Primo scilicet in formulis pro radiis facierum inuen-  
tis litterae  $\varrho$ ,  $\sigma$ , et  $\tau$  tantum ad primam lentem  
pertinent, earumque loco pro secunda lente scribi  
oportet  $\varrho'$ ,  $\sigma'$ , et  $\tau'$ ; pro tertia autem  $\varrho''$ ,  $\sigma''$  et  $\tau''$  et ita  
porro. Praeterea vero spatia diffusionis hinc aliquam  
mutationem requirunt, eritque spatium diffusionis.

## I. Pro vna lente

$$\alpha\alpha xx\left(\frac{1}{a} + \frac{1}{\alpha}\right)\mu\left(\lambda\left(\frac{1}{a} + \frac{1}{\alpha}\right)^2 + \frac{\nu}{a\alpha}\right).$$

## II. Pro duabus lentibus

$$\epsilon\epsilon xx \left\{ \begin{aligned} &+ \frac{\mu\alpha\alpha}{bb}\left(\frac{1}{a} + \frac{1}{\alpha}\right)\left(\lambda\left(\frac{1}{a} + \frac{1}{\alpha}\right)^2 + \frac{\nu}{a\alpha}\right) \\ &+ \frac{\mu'\epsilon\epsilon}{\alpha\alpha}\left(\frac{1}{b} + \frac{1}{\epsilon}\right)\left(\lambda'\left(\frac{1}{b} + \frac{1}{\epsilon}\right)^2 + \frac{\nu'}{b\epsilon}\right) \end{aligned} \right\}.$$

## III. Pro tribus lentibus

$$\gamma\gamma xx \left\{ \begin{aligned} &+ \frac{\mu\alpha\alpha\epsilon\epsilon}{bbcc}\left(\frac{1}{a} + \frac{1}{\alpha}\right)\left(\lambda\left(\frac{1}{a} + \frac{1}{\alpha}\right)^2 + \frac{\nu}{a\alpha}\right) \\ &+ \frac{\mu'\epsilon\epsilon\gamma\gamma}{\alpha\alpha cc}\left(\frac{1}{b} + \frac{1}{\epsilon}\right)\left(\lambda'\left(\frac{1}{b} + \frac{1}{\epsilon}\right)^2 + \frac{\nu'}{b\epsilon}\right) \\ &+ \frac{\mu''bbcc}{\alpha\alpha\epsilon\epsilon}\left(\frac{1}{c} + \frac{1}{\gamma}\right)\left(\lambda''\left(\frac{1}{c} + \frac{1}{\gamma}\right)^2 + \frac{\nu''}{c\gamma}\right) \end{aligned} \right\}.$$

## IV. Pro quatuor lentibus

$$\delta\delta xx \left\{ \begin{aligned} &+ \frac{\mu\alpha\alpha\epsilon\epsilon\gamma\gamma}{bbccda}\left(\frac{1}{a} + \frac{1}{\alpha}\right)\left(\lambda\left(\frac{1}{a} + \frac{1}{\alpha}\right)^2 + \frac{\nu}{a\alpha}\right) \\ &+ \frac{\mu'\epsilon\epsilon\gamma\gamma}{\alpha\alpha ccdd}\left(\frac{1}{b} + \frac{1}{\epsilon}\right)\left(\lambda'\left(\frac{1}{b} + \frac{1}{\epsilon}\right)^2 + \frac{\nu'}{b\epsilon}\right) \\ &+ \frac{\mu''bbcc\gamma\gamma}{\alpha\alpha\epsilon\epsilon dd}\left(\frac{1}{c} + \frac{1}{\gamma}\right)\left(\lambda''\left(\frac{1}{c} + \frac{1}{\gamma}\right)^2 + \frac{\nu''}{c\gamma}\right) \\ &+ \frac{\mu'''\epsilon\epsilon\gamma\gamma}{\alpha\alpha\epsilon\epsilon\gamma\gamma}\left(\frac{1}{d} + \frac{1}{\delta}\right)\left(\lambda'''\left(\frac{1}{d} + \frac{1}{\delta}\right)^2 + \frac{\nu'''}{d\delta}\right) \end{aligned} \right\}.$$

valores autem harum litterarum  $\mu'$ ,  $\mu''$ ,  $\mu'''$  etc.  
iam supra definiuimus §. 77. Coroll.

## Coroll. I.

92. Si lentis primae PP ponatur radius faciei anterioris  $=f$  et posterioris  $=g$  erit

$$\frac{1}{f} = \frac{\rho}{a} + \frac{\sigma}{a} \pm \tau \left( \frac{1}{a} + \frac{1}{a} \right) \sqrt{\lambda - 1}$$

$$\frac{1}{g} = \frac{\rho}{a} + \frac{\sigma}{a} \mp \tau \left( \frac{1}{a} + \frac{1}{a} \right) \sqrt{\lambda - 1}$$

unde si detur distantia obiecti  $EA = a$ , primo intuen-  
tur  $\alpha$  ex hac aequatione

$$\frac{1}{f} + \frac{1}{g} = (\rho + \sigma) \left( \frac{1}{a} + \frac{1}{a} \right) = 1, \text{ 818182 } \left( \frac{1}{a} + \frac{1}{a} \right) = \frac{20}{11} \left( \frac{1}{a} + \frac{1}{a} \right)$$

Inuenta autem distantia  $\alpha$  numerus  $\lambda$  reperitur ex  
hac aequatione.

$$\frac{1}{f} - \frac{1}{g} + (\sigma - \rho) \left( \frac{1}{a} - \frac{1}{a} \right) = 2 \tau \left( \frac{1}{a} + \frac{1}{a} \right) \sqrt{\lambda - 1}.$$

## Coroll. 2.

93. Deinde si distantia secundae lentis a prima sit  
 $=F$ , ob  $F = a + b$  habetur  $b = F - a$ ; qua distantia  $b$   
cognita, si pro lente secunda datus sit radius faciei anterioris  
 $=f'$ , et posterioris  $=g'$  habebuntur iterum duae aequationes

$$\frac{1}{f'} = \frac{\rho}{b} + \frac{\sigma}{b} \pm \tau \left( \frac{1}{b} + \frac{1}{b} \right) \sqrt{\lambda' - 1}$$

$$\frac{1}{g'} = \frac{\rho}{b} + \frac{\sigma}{b} \mp \tau \left( \frac{1}{b} + \frac{1}{b} \right) \sqrt{\lambda' - 1}$$

ex quibus cum distantiam  $b$  tum numerum  $\lambda'$  definire  
licet. Similique modo ex forma sequentium lentium  
earumque distantia reliqua elementa innotescunt.

## Coroll. 3.

94. Si singulae lentes ad minimum spatium diffu-  
sionis fuerint accommodatae, erit  $\lambda = 1, \lambda' = 1, \lambda'' = 1, \lambda''' = 1$  etc.  
sin autem hae lentes alia forma fuerint praeditae, isti  
numeri erunt unitate maiores.



## Scholion.

95. Quo plures fuerint lentes eo pluribus constabit membris spatium diffusionis ab iis productum; neque tamen propterea aucto lentium numero spatium diffusionis necessario augetur. Cum enim quantitates  $\alpha, b, \epsilon, c, \gamma, d, \delta$ , etc. valores quoque negativos recipere queant, dummodo binarum summae  $\alpha + b$ ;  $\epsilon + c$ ;  $\gamma + d$ ;  $\delta + e$  etc. vtpote lentium distantiae maneant positivae, fieri potest, ut vnum vel aliquot membra fiant negatiua, hincque spatium diffusionis diminuatur, quin etiam interdum prorsus euanescat, quo casu repraesentatio sine dubio erit perfectissima. Verum in instrumentis dioptricis ad visionem instructis veluti Telescopiis ac Microscopiis non tam hoc, quod definiuimus, spatium diffusionis quam confusio in ipsa visione orta spectari debet; quae autem etsi a spatio diffusionis plurimum differt, tamen ex eo definiri potest vti mox explicabimus. Ante autem conueniet lentes compositas seu multiplicatas considerare, cuiusmodi oriuntur, si duae pluresue lentes, quarum crassities tam est parua vt negligi queat, immediate iungantur quo quidem pacto instar lentium simplicium spectari possunt; verum tali coniunctione effici potest, ut spatium diffusionis multo fiat minus quam si lens simplex adhiberetur, atque adeo euanescat valorque numeri  $\lambda$  istiusmodi lenti compositae conueniens vnitae minor sit proditurus vnde maxima commoda ad confusionem diminuendam obtinebuntur.



## CAPUT III.

DE

LENTIBVS COMPOSITIS  
SEV MVLTIPPLICATIS.

## Definitio I.

**L**<sup>96.</sup>ens duplicata oritur, si duae lentes super communi  
axe sibi immediate iungantur.

Crassitiem hic vtriusque lentis tanquam nullam assumo  
et quia distantia inter lentes nulla ponitur, crassities  
etiam lentis duplicatae pro nulla haberi poterit.

## Coroll. I.

97. Binae ergo lentes PP et QQ sibi ad con- Tab. II.  
tactum fere coniunctae lentem duplicatam constituunt; Fig. 7.  
de qua tamen notandum est, eius crassitiem minus  
tuto negligi posse quam vtriusque lentis simplicis  
seorsim sumtae. Si enim lentes immediate se in  
puncto contingerent; phaenomena colorum a *Newtono*  
obseruata essent metuenda tum vero etiam ostende-  
mus, quomodo ratio distantiae inter binas lentes ha-  
beri possit.



## Coroll. 2.

98. Si lentis anterioris PP distantiae determinatrices sint  $a$  et  $\alpha$ , lentis posterioris vero QQ,  $b$  et  $\beta$ , necesse est ut sit  $a+b=0$  seu  $\alpha=-b$ . Tum vero obiecti ante lentem ad distantiam  $AE=a$  positi imago principalis repraesentabitur post lentem ad distantiam  $G=\beta$ .

## Coroll. 3.

99. Erunt ergo  $a$  et  $\beta$  quasi distantiae determinatrices lentis duplicatae; ac sumendo  $a$  vel  $b$  ad libitum infinita paria lentium pro his distantis exhiberi possunt. Cum deinde utraque lens praeterea numerum indefinitum  $\lambda$  recipiat insuper infinita varietas locum habet.

## Coroll. 4.

100. Quia crassities pro nihilo reputatur aperturae in singulis faciebus eadem est ratio; scilicet si in prima facie semidiameter aperturae sit  $=x$ , in reliquis quoque faciebus apertura eadem vel saltem non minor esse debet.

## Scholion.

101. Non opus est, ut lentes plane ad contactum coniungantur quoniam forte refractionis lex turbari posset: hoc autem vel minima interposita distantia evitabitur id quod ad institutum nostrum sufficit, cum etiam crassities utriusque non omnino sit nulla.

Proble-

## Problema I.

102. Omnes lentes duplicatas describere, quibus obiectum  $E\varepsilon$  in data ante lentem distantia  $AE$  propositum post lentem in data distantia  $bG$  repraesentetur, simulque diffusionem imaginis  $Gg$  pro data lentis apertura definire.

## Solutio.

Sit distantia obiecti  $AE = a$ , imaginis principalis  $bG = \varepsilon$ , tum vero lentis primae  $PP$  distantiae determinatrices  $a$ , et  $\alpha$ ; lentis posterioris vero  $QQ$ ,  $b$  et  $\varepsilon$ ; iam quia distantia lentium est nulla, erit  $\alpha + b = 0$ , seu  $\alpha = -b$ ; Hinc si obiecti magnitudo sit  $E\varepsilon = z$ , erit imaginis principalis magnitudo  $G\eta = \frac{\varepsilon}{a}z$  pro situ inuerso. Cum porro vtraque lens infinitis modis formari possit, sit  $\lambda$  numerus arbitarius pro prima  $PP$  et  $\lambda'$  pro secunda  $QQ$ , quarum ergo constructio ita se habebit:

Pro Lente

$$\begin{aligned} PP \text{ radius faciei} & \begin{cases} \text{anterioris} = \frac{a\alpha}{\varrho\alpha + \sigma\alpha \pm \tau(a+\alpha)\sqrt{(\lambda-1)}} \\ \text{posterioris} = \frac{a\alpha}{\varrho\alpha + \sigma\alpha \pm \tau(a+\alpha)\sqrt{(\lambda-1)}} \end{cases} \\ QQ \text{ radius faciei} & \begin{cases} \text{anterioris} = \frac{b\varepsilon}{\varrho\varepsilon + \sigma b \pm \tau(b+\varepsilon)\sqrt{(\lambda'-1)}} \\ \text{posterioris} = \frac{b\varepsilon}{\varrho\varepsilon + \sigma b \pm \tau(b+\varepsilon)\sqrt{(\lambda'-1)}} \end{cases} \end{aligned}$$

Si ratio refractionis sit diuersa; pro secunda lente scribi debet  $\varrho'$ ,  $\sigma'$  et  $\tau'$ , loco  $\varrho$ ,  $\sigma$  et  $\tau$ .

Pro



Pro spatio autem diffusionis  $Gg$  inueniendo, sit semidiameter aperturae lentis duplicatae  $=x$ , eritque

$$Gg = \mu \xi \xi x x \cdot \left\{ \begin{array}{l} + (\frac{1}{a} + \frac{1}{\alpha})(\lambda(\frac{1}{a} + \frac{1}{\alpha})^2 + \frac{v}{a\alpha}) \\ + (\frac{1}{b} + \frac{1}{\beta})(\lambda'(\frac{1}{b} + \frac{1}{\beta})^2 + \frac{v}{b\beta}) \end{array} \right.$$

et radiorum in  $g$  concurrentium inclinatio ad axem  $= \frac{x}{\xi}$ .

Si ratio refractionis discrepet; in parte ex lente secunda orta scribatur  $\mu'$ ,  $v'$ , loco  $\mu$ ,  $v$ .

### C O R O L L. I.

103. Huius ergo lentis duplicatae distantiae determinatrices sunt  $a$  et  $\xi$ , praeterea vero duo numeri arbitrarii  $\lambda$  et  $\lambda'$ , vna cum distantia  $a$  vel  $b$  eius perfectam determinationem constituunt, vnde in huiusmodi lentibus multo maior varietas locum habet, quam in lentibus simplicibus.

### C O R O L L. 2.

104. Si ex eisdem distantis determinatricibus  $a$  et  $\xi$  adiungendo numero arbitrario  $\lambda^o$  lens simplex construatur, ea imaginem eadem magnitudine  $G\eta = \frac{\xi}{a}x$  referet; sed pro eadem apertura, cuius semidiameter  $=x$  habebitur spatium diffusionis

$$Gg = \mu \xi \xi x x (\frac{1}{a} + \frac{1}{\xi})(\lambda^o(\frac{1}{a} + \frac{1}{\xi})^2 + \frac{v}{a\xi})$$

### C O R O L L. 3.

105. Fieri ergo poterit vt lens duplicata modo maiorem modo minorem diffusionem gignat. Lens autem

autem simplex minimam parit diffusionem, si  $\lambda^{\circ} = 1$ ; ergo tum demum lentes duplicatae simplicibus erunt praeferendae, cum adhuc minorem diffusionem producent.

### Scholion.

106. Concipi quidem semper poterit lens simplex eandem diffusionem gignens ac lens duplicata, si pro numero  $\lambda^{\circ}$  omnes valores admittamus; quomocunque enim lens duplicata fuerit comparata, si spatium diffusionis inde productum huic ex lente simplici nato aequale statuatur, determinatus valor pro numero  $\lambda^{\circ}$  elicitur: qui si fuerit positivus et unitate maior, realis lens simplex aequivalens exhiberi poterit, sin autem prodeat unitate minor, vel adeo negativus, lens simplex inter imaginaria erit referenda. Quando autem fit  $\lambda^{\circ} > 1$ , evidens est lentem simplicem eundem plane effectum esse edituram ac duplicatam, ideoque semper expediet lente simplici potius uti quam duplicata; quodsi vero prodierit  $\lambda^{\circ} < 1$ , quo casu lens simplex fit imaginaria, tum lentes duplicatae effectum praestabunt a simplicibus non expectandum, qui adeo cum insigni hoc commodo, quod spatium diffusionis futurum sit minus, erit coniunctus. Talibus ergo lentibus duplicatis maximo cum successu uti poterimus, eoque magis eae simplicibus erunt anteponendae, quo minor fuerit valor numeri  $\lambda^{\circ}$  iis respondens.



## Problema 2.

107. Data lente duplicata ad binas distantias determinatrices  $AE=a$  et  $bG=g$  relata pro iisdem distantis definire lentem simplicem, quae pro eadem apertura, eandem diffusionem imaginis producat.

## Solutio.

Totum ergo negotium huc redit, ut spatium diffusionis  $Gg$  a lente simplici productum (104) aequale ponatur spatio diffusionis a lente duplicata orto, cuius expressio in problemate praecedente (101) est inuenta, indeque valor numeri  $\lambda^o$  pro constructione lentis simplicis eliciatur. Quae inuestigatio quo commodius institui possit, ponamus:

$$\frac{1}{\alpha} = \frac{f-1}{a} + \frac{f}{g} \text{ eritque } \frac{1}{b} = \frac{-f+1}{a} - \frac{f}{g}$$

ita ut loco quantitatis  $\alpha$  vel  $b$  numerum  $f$  introduamus eritque

$$\frac{1}{a} + \frac{1}{\alpha} = f\left(\frac{1}{a} + \frac{1}{g}\right) \text{ et } \frac{1}{b} + \frac{1}{g} = (1-f)\left(\frac{1}{a} + \frac{1}{g}\right)$$

vnde spatium diffusionis a lente duplicata ortum prodit:

$$Gg = \mu g g x x \left\{ \begin{aligned} &+ f\left(\frac{1}{a} + \frac{1}{g}\right) \left( \lambda f f \left(\frac{1}{a} + \frac{1}{g}\right)^2 + \nu \left( \frac{f-1}{a a} + \frac{f}{a g} \right) \right) \\ &+ (1-f)\left(\frac{1}{a} + \frac{1}{g}\right) \left( \lambda (1-f) \left(\frac{1}{a} + \frac{1}{g}\right)^2 + \nu \left( \frac{1-f}{a g} - \frac{f}{g g} \right) \right) \end{aligned} \right.$$

quae expressio reducitur ad hanc formam ::

$$Gg = \mu g g x x \left( \frac{1}{a} + \frac{1}{g} \right) \left( \left( \lambda f^3 + \lambda (1-f)^3 \right) \left( \frac{1}{a} + \frac{1}{g} \right)^2 + \nu \left( \frac{f(f-1)}{a a} + \frac{1-2f+2ff}{a g} + \frac{f(f-1)}{g g} \right) \right)$$

Verum postremum membrum  $\frac{f(f-1)}{a a} + \frac{1-2f+2ff}{a g} + \frac{f(f-1)}{g g}$  mutatur

mutatur in  $f(f-1)(\frac{1}{a}+\frac{1}{e})^2+\frac{1}{ae}$ , ficque habebimus pro lente duplicata:

$$Gg = \mu \epsilon \epsilon x x (\frac{1}{a}+\frac{1}{e})((\lambda f^3 + \lambda'(1-f)^3 - \nu f(1-f))(\frac{1}{a}+\frac{1}{e})^2 + \frac{\nu}{ae})$$

quae forma iam facillime cum spatio diffusionis lentis simplicis comparatur indeque manifesto colligitur:

$$\lambda^0 = \lambda f^3 + \lambda'(1-f)^3 - \nu f(1-f)$$

Cum autem loco quantitatum  $a$  et  $b$  numerum  $f$  introduxerimus constructio lentis duplicatae ita se habebit

Pro lente

radius faciei

$$\text{prima PP} \left\{ \begin{array}{l} \text{anterioris} = \frac{ae}{(\rho - \sigma(1-f))e + \sigma f a \pm \tau f(a+e)\sqrt{(\lambda-1)}} \\ \text{posterioris} = \frac{ae}{(\sigma - \rho(1-f))e + \rho f a \pm \tau f(a+e)\sqrt{(\lambda-1)}} \end{array} \right.$$

$$\text{secunda QQ} \left\{ \begin{array}{l} \text{anterioris} = \frac{ae}{(\sigma - \rho f)a + \rho(1-f)e \pm \tau(1-f)(a+e)\sqrt{(\lambda-1)}} \\ \text{posterioris} = \frac{ae}{(\rho - \sigma f)a + \sigma(1-f)e \pm \tau(1-f)(a+e)\sqrt{(\lambda-1)}} \end{array} \right.$$

Tum vero inuento numero  $\lambda^0$  constructio lentis simplicis aequivalentis erit

$$\text{Radius faciei} \left\{ \begin{array}{l} \text{anterioris} = \frac{ae}{\rho e + \sigma a \pm \tau(a+e)\sqrt{(\lambda^0-1)}} \\ \text{posterioris} = \frac{ae}{\rho a + \sigma e \pm \tau(a+e)\sqrt{(\lambda^0-1)}} \end{array} \right.$$

### COROLL. I.

108. Quoties ergo fuerit

$$\lambda f^3 + \lambda'(1-f)^3 - \nu f(1-f) > 1$$

semper lens simplex parari potest duplicatae aequiva-

L 2

lens



lens iisque ergo casibus praestabit lente simplici uti potius quam lente duplicata.

### Coroll. 2.

109. Verum si fuerit  
 $\lambda f^3 + \lambda'(1-f)^3 - \nu f(1-f) < 1$   
 ob  $\lambda^\circ < 1$ , constructio lentis simplicis fit impossibilis, ac lens duplicata minorem pariet diffusionem quam per ullam lentem simplicem obtineri potest.

### Coroll. 3.

110. Si esset  $f=0$ , prodiret  $\lambda^\circ=\lambda'$ , et in lente duplicata anteriori nullam refractionem produceret ob facies parallelas, resque eodem rediret, ac si posterior sola adesset. Sin autem sumatur  $f=1$ , fit  $\lambda^\circ=\lambda$ , et lens posterior superflua; utroque ergo casu nullum lucrum impetratur.

### Coroll. 4.

111. Cum autem numeri  $\lambda$  et  $\lambda'$  unitate nequeant esse minores, sitque  $\nu=0,232692$ , patet pro  $f$  nullum valorem inter limites 0 et 1 assumi posse, unde fiat  $\lambda^\circ=0$ . At si  $f$  extra hos limites capiatur, utique pro  $\lambda$  et  $\lambda'$  eiusmodi numeri unitate maiores assignari poterunt, ut fiat  $\lambda^\circ=0$ .

### Scholion I.

112. Ratione ergo numeri  $f$  tres casus lentium duplicatarum considerari conueniet, prout vel  $f$  intra limi-

limites 0 et 1 continetur, vel fuerit  $f > 1$  vel  $f < 0$ . Primo casu euenire non potest, vt fiat  $\lambda^{\circ} = 0$ ; sed plurimum intererit eam determinasse lentem duplicatam, pro qua  $\lambda^{\circ}$  minimum obtineat valorem qui quo magis infra vnitatem cadat, eo perfectior lens erit censenda, maioreque iure simplicibus anteferenda. Binis reliquis vero casibus  $f > 1$  et  $f < 0$  eiusmodi adeo lentes duplicatae parari poterunt, quae praebeant  $\lambda^{\circ} = 0$ , quae ergo pro perfectissimis essent habendae. Verum hic quoque ad praxin est respiciendum, quae cum semper a praeceptis Theoriae aberrare soleat, euenire potest, vt leui errore commissio numerus  $\lambda^{\circ}$  non solum non euanescat, sed adeo vnitatem excedat quo casu vtique expediret lente vti simplici.

### Scholion 2.

113. Cum tanti sit momenti rationem aberrationis a qua praxis vix liberari potest habere, eam etiam in lentibus simplicibus perpendi conueniet. Postulauimus autem pro datis distantis determinatricibus lentem construere posse in qua numerus  $\lambda$  datum obtineat valorem dum ne sit vnitatem minor; hic igitur obseruari oportet, quo maior fuerit  $\lambda$ , eo difficilius fore errorem euitare; si enim in constructione leuissimus error committatur alius eo magis diuersus valor pro  $\lambda$  orietur, quo maior fuerit  $\lambda$ . Verum e contrario cum vnitatem sit minimus valor, quem  $\lambda$  recipere potest ex natura minimi liquet etiam si in praxi a praescripta regula notabiliter recedatur tamen inde vix



fenfibile difcrimen in valorem  $\lambda$  effe redundaturum. Ex quo concludimus feliciffimo cum fucceffu eiusmodi lentes fimples parari poffe, pro quibus futurum fit  $\lambda = 1$ , neque hic errores praxeos, nifi fuerint enormes, admodum effe pertimescendos. Deinde quo minus numerus  $\lambda$  vnitatem fuperare debeat, eo certiores effe poterimus de fucceffu, fed non eo gradu, quo cafu,  $\lambda = 1$ ; at fi opus fit eiusmodi lente, pro qua valor ipfius  $\lambda$  debeat effe numerus fatis magnus, difficillime per praxin fatisfiet ac fortaffe ingentem lentium numerum parare oportebit, antequam vna obtineatur fcopo fatisfaciens. Quamobrem fi praxi confulere velimus, vix alias lentes exigere debemus, nifi pro quibus numerus  $\lambda$  vel fit vnitatis ipfa, vel parumper maior. Sin autem ad infigne aliquod commodum aliae lentes requirantur, labori non erit parcendum, cum fortaffe non nifi poft plurimos conatus irritos voti tandem compotes reddi queamus.

### Problema 3.

114. Definire eam lentem duplicatam, pro qua, fi numerus  $f$  intra limites 0 et 1 accipiatur numerus  $\lambda^0$  minimum adipiscatur valorem.

### Solutio.

Positis iisdem, quae in praecedentibus problematibus funt constituta, inuenimus effe

$$\lambda^0 = \lambda f^3 + \lambda' (1 - f)^3 - \nu f (1 - f)$$

vbi cum  $f$  intra limites 0 et 1 assumi debeat, ambo termini

termini  $\lambda f^3$  et  $\lambda'(1-f)^3$  erunt positiui. Quare vt  $\lambda^\circ$  omnium minimum valorem nanciscatur, necesse est vtrique numero  $\lambda$  et  $\lambda'$  minimum valorem cuius est capax tribui.

Sit ergo  $\lambda = 1$  et  $\lambda' = 1$ , ac habebimus:

$$\lambda^\circ = 1 - 3f + 3ff - \nu f + \nu ff = 1 - (3 + \nu)f(1 - f)$$

quae expressio vt minima reddatur, oportet fieri  $f(1-f)$  maximum, id quod fit sumendo  $f = \frac{1}{2}$ ; hincque oritur:

$$\lambda^\circ = 1 - \frac{1}{4}(3 + \nu) = \frac{1-\nu}{4} = 0,191827$$

Quare constructio huius lentis duplicatae ita se habebit

$$\text{Pro lente PP radius faciei} \begin{cases} \text{anterioris} = \frac{2a\mathfrak{E}}{(2\mathfrak{E} - \sigma)\mathfrak{E} + \sigma a} \\ \text{posterioris} = \frac{2a\mathfrak{E}}{(2\sigma - \mathfrak{E})\mathfrak{E} + \mathfrak{E}a} \end{cases}$$

$$\text{Pro lente QQ radius faciei} \begin{cases} \text{anterioris} = \frac{2a\mathfrak{E}}{(2\sigma - \mathfrak{E})a + \mathfrak{E}\mathfrak{E}} \\ \text{posterioris} = \frac{2a\mathfrak{E}}{(2\mathfrak{E} - \sigma)a + \sigma\mathfrak{E}} \end{cases}$$

et si aperturae semidiameter sit  $= x$ , erit spatium diffusionis

$$Gg = \mu \mathfrak{E} \mathfrak{E} x x \left( \frac{1}{a} + \frac{1}{\mathfrak{E}} \right) (0,191827 \left( \frac{1}{a} + \frac{1}{\mathfrak{E}} \right)^2 + \frac{\nu}{a\mathfrak{E}})$$

Si eiusmodi vitro utamur, pro quo est  $n = 1,60 = \frac{8}{5}$  tum ob  $\nu = \frac{4}{15}$ , prodiret  $\lambda^\circ = 0,183333$  ideoque haec vitri species adhuc minorem confusionem pareret.

Coroll.



## Coroll. 1.

115. Si pro iisdem distantis determinatricibus  $a$  et  $b$  lens simplex minimam diffusionem pariens construatur quod fit sumendo

$$\text{radius faciei} \begin{cases} \text{anterioris} = \frac{ab}{b^2 + a^2} \\ \text{posterioris} = \frac{ab}{a^2 + b^2} \end{cases}$$

spatium diffusionis foret

$$\mu \cdot b^2 x x \left( \frac{1}{a} + \frac{1}{b} \right) \left( \left( \frac{1}{a} + \frac{1}{b} \right)^2 + \frac{v}{a^2 b^2} \right).$$

## Coroll. 2.

116. Apparet ergo a lente duplicata descripta multo minorem oriri diffusionem, quam a lente simplici, etiam si haec iam ad minimam diffusionem sit instructa: Cum enim ceterae partes sint pares coefficientis membri  $\left( \frac{1}{a} + \frac{1}{b} \right)^2$  plus quam quintuplo minor est in duplicata quam in simplici.

## Coroll. 3.

117. Si ponamus  $\lambda = 1$  et  $\lambda' = 1$ , vel saltem  $\lambda' = \lambda$ , minor valor pro  $\lambda^\circ$  obtineri nequit, quam inuenimus, etiam si pro  $f$  alios valores admittere velimus. Vnde si vtraque lens per se iam minimam diffusionem pariat, pro lente duplicata valor ipsius  $\lambda^\circ$  minor quam 0,191827 fieri nequit.

## Scholion.

118. Huiusmodi ergo lentes duplicatae maxime sunt notatu dignae cum loco simplicium adhibitae multo

multo minorem diffusionem pariant, ex quo in constructione Telescopiorum et Microscopiorum earum amplissimus erit usus. Neque vero hac insigni proprietate sunt praeditae, sed etiam earum constructio in praxi minimis difficultatibus est obnoxia: propterea quod etsi a praescriptis regulis parumper aberretur, effectus tamen inde vix ullam mutationem patiatur. Siue enim in constructione vtriusque seorsim levis error committatur, valores numerorum  $\lambda$  et  $\lambda'$  unitatem haud sensibilibiter excedent, siue in quantitate  $f$  valor iustus  $f = \frac{1}{2}$  non exacte obseruetur, error vix sentietur, quoniam hi numeri ex natura minimi sunt eruti. Quomocunque autem a regulis praescriptis aberretur, valor ipsius  $\lambda^0$  inde paulisper maior prodibit. Veluti si eiusmodi errores committantur, ut sit  $\lambda = 1 + \frac{1}{10}$ ;  $\lambda' = 1 + \frac{1}{10}$  et  $f = \frac{1}{2} + \frac{1}{20}$ , prodibit  $\lambda^0 = 0,191827 + 0,02558$  seu  $\lambda^0 = 0,2174$ , ita ut discrimen partem tantum quadragesimam unitatis conficiat. Facile autem intelligitur, dummodo  $\lambda^0$  prodeat minus quam  $\frac{1}{4}$ , quod nunquam non, facile praestari posse videtur, ab his lentibus duplicatis insignem utilitatem expectandam esse. Etsi ergo eiusmodi lentes duplicatae confici possunt, pro quibus numerus  $\lambda^0$  plane euanescat, ob harum lubricam constructionem illae istis anteferendae videntur, ut mox clarius patebit.

#### Problema 4.

119. Pro datis distantis determinatricibus  $AE = a$  et  $bG = b$  eas lentes duplicatas inuenire, in quibus sit  $\lambda^0 = 0$ .

Tom. I.

M

Solutio.



## Solutio.

Cum fieri nequeat  $\lambda^0 = 0$ , nisi numerus  $f$  extra limites 0 et 1 accipiatur, simulque numeri  $\lambda$  et  $\lambda'$  fuerint inaequales, ita ut alterutra saltem lens non debeat seorsim minimam diffusionem parere; ponamus esse vel  $f > 1$  vel  $f < 0$ . Sit ergo primo  $f = 1 + \xi$ , et cum sit

$$\lambda^0 = \lambda(1 + \xi)^3 - \lambda'\xi^3 + \nu\xi(1 + \xi)$$

ut fiat  $\lambda^0 = 0$  oportet esse:

$$\lambda' = \lambda\left(1 + \frac{1}{\xi}\right)^3 + \frac{\nu(1 + \xi)}{\xi^2}$$

vnde  $\lambda'$  necessario unitatem superabit, cuius valor ne prodeat nimis magnus, sumi conueniet  $\lambda = 1$ , ita ut sit

$$\lambda' = \left(1 + \frac{1}{\xi}\right)^3 + \frac{0,232602(1 + \xi)}{\xi^2} \text{ et } \lambda = 1.$$

Quicumque ergo valor ipsi  $\xi$  tribuatur, lens duplicata habetur, pro qua sit  $\lambda^0 = 0$ , ac propterea spatium diffusionis  $= \mu \sqrt[3]{xx} \left(\frac{1}{a} + \frac{1}{b}\right) \frac{\nu}{a^2 b}$ . Videamus nonnullos casus speciales.

$$f = \frac{3}{2}; \xi = \frac{1}{2}; \lambda = 1 \text{ et } \lambda' = 28,396152$$

$$f = 2; \xi = 1; \lambda = 1 \text{ et } \lambda' = 8,465384$$

$$f = 3; \xi = 2; \lambda = 1 \text{ et } \lambda' = 3,549519$$

$$f = 4; \xi = 3; \lambda = 1 \text{ et } \lambda' = 2,473789$$

$$f = 5; \xi = 4; \lambda = 1 \text{ et } \lambda' = 2,025841$$

$$f = 6; \xi = 5; \lambda = 1 \text{ et } \lambda' = 1,783846$$

etc.

Pro

Pro altero casu fit  $f = -\xi$ , ideoque

$$\lambda^0 = -\lambda \xi^3 + \lambda' (1 + \xi)^3 + \nu \xi (1 + \xi)$$

vnde facto  $\lambda^0 = 0$  prodit

$$\lambda = \lambda' (1 + \frac{1}{\xi})^3 + \frac{\nu(1 + \xi)}{\xi^2}.$$

Statui ergo conueniet  $\lambda' = 1$ , ac pro  $\lambda$  notentur casus sequentes.

$$f = -\frac{1}{2}; \xi = \frac{1}{2}; \lambda = 28,396152 \text{ et } \lambda' = 1$$

$$f = -1; \xi = 1; \lambda = 8,465384 \text{ et } \lambda' = 1$$

$$f = -2; \xi = 2; \lambda = 3,549519 \text{ et } \lambda' = 1$$

$$f = -3; \xi = 3; \lambda = 2,473789 \text{ et } \lambda' = 1$$

$$f = -4; \xi = 4; \lambda = 2,025841 \text{ et } \lambda' = 1$$

$$f = -5; \xi = 5; \lambda = 1,783846 \text{ et } \lambda' = 1$$

ficque patet infinitis modis huiusmodi lentes duplicatas parari posse, pro quibus fit  $\lambda^0 = 0$ , et spatium diffusionis

$$Gg = \mu \xi \xi x x (\frac{1}{a} + \frac{1}{\xi}) \frac{\nu}{a \xi}.$$

### Scholion I.

120. Si huiusmodi lentes accuratissime parari possent, nullum est dubium, quin praecedentibus essent anteferendae propterea quod diffusio iis adhuc magis diminuitur. Verum dolendum est, quod minimus error in earum constructione commissus omnem



fere vsum destruat. Quo hoc facilius diiudicare queamus, examinemus eum casum, quo est

$$f=5, \lambda=1 \text{ et } \lambda'=2,025841$$

hincque

$$\lambda^{\circ} = \lambda f^3 - \lambda' (f-1)^3 + \nu f (f-1) = 0.$$

Ponamus autem in constructione errorem esse commissum vt reuera non sit  $f=5$  sed  $f=5\frac{1}{10}$ , dum numeri  $\lambda$  et  $\lambda'$  suos iustos valores obtineant: ob hunc autem vix vitandum errorem non sit  $\lambda^{\circ}=0$ , sed adeo  $\lambda^{\circ}=-2,011$ ; sicque haec lens duplicata simplicibus longe est postponenda, simili modo si  $f$  esset  $=5$ , sed vel  $\lambda$  vel  $\lambda'$  tantillum a praescripto valore aberraret, enorme statim discrimen in valorem ipsius  $\lambda^{\circ}$  redundaret. Minus quidem error metuendus videtur in ea specie, qua  $\lambda=1$ ,  $\lambda'=28,396152$  et  $f=\frac{3}{2}$ ; sed praeterquam, quod lens simplex posterior difficillime parari queat, pro qua  $\lambda'$  praecise valorem assignatum consequatur, talis lens ad vsum dioptricum plane est inepta, ob ingentem alterius faciei curvaturam. Quae cum ita sint, quia tam exiguus error in paratione huiusmodi lentium commissus facit, vt  $\lambda^{\circ}$  adeo supra vnitatem excrescat, vix sperare poterimus, vt vnquam talis lens duplicata perficiatur, pro qua  $\lambda^{\circ}$  vsque ad  $\frac{1}{2}$  diminuatur. In lentibus autem praecedentis generis successus vix fallere poterit, nisi in praxi enormiter a praescripta regula aberretur: ex quo his solis lentibus duplicatis cum fructu vti licebit, dum  
contra

contra eae, quas in praesente problemate descripsimus, penitus profligandae videntur.

### Scholion 2.

121. Pro datis ergo binis distantis determinatricibus  $a$  et  $b$  semper eiusmodi lens duplicata parari potest, ex qua nascatur spatium diffusionis

$$Gg = \mu b b x x \left( \frac{1}{a} + \frac{1}{b} \right) \left( \lambda^0 \left( \frac{1}{a} + \frac{1}{b} \right)^2 + \frac{v}{a b} \right)$$

ita ut  $\lambda^0$  numerum quemcunque denotare possit. Ad hoc enim satisfieri oportet huic aequationi:

$$\lambda^0 = \lambda f^3 + \lambda' (1 - f)^3 - v f (1 - f)$$

id quod semper fieri potest, cum  $f$  plane ab arbitrio nostro pendeat, et numeri  $\lambda$  et  $\lambda'$  tantum non unitate minores accipi debeant. Definitis autem his tribus numeris  $\lambda$ ,  $\lambda'$  et  $f$  ita, ut  $\lambda^0$  datum valorem obtineat, binae lentes simplices, ex quibus duplicata est componenda, secundum formulas §. 107 datas confici debent. Vbi quidem tenenda sunt ea, quae modo observauimus, in praxi eas lentes facillime obtineri, quando numeri  $\lambda$  et  $\lambda'$  unitatem parum superant,  $f$  vero propemodum  $\frac{1}{2}$  denotat, cum contra quo magis hi numeri ab istis terminis recedant, eo maius sit periculum ne effectus enormiter fallat. Ceterum cum diffusionem a lente duplicata oriundam ad eandem formam reduxerimus, qua diffusio lentis simplicis exprimitur, inde id commodi consequimur, ut simili modo diffusionem a lentibus magis multiplicatis ortum definire valeamus.



## Supplementum I.

De lentibus duplicatis.

Si pro lente anteriori ratio refractionis sit  $n : 1$ , pro lente posteriori vero alia ratio  $n' : 1$  locum habeat; problemata hic tractata sequenti modo resolui poterunt

## Pro Problemate I.

Si pro numeris  $\varrho, \sigma, \tau$ , qui ex  $n$  oriuntur, quaerantur simili modo ex  $n'$  valores  $\varrho', \sigma'$ , et  $\tau'$  erunt

Pro Lente

$$P P \text{ radius faciei} \begin{cases} \text{anter.} = \frac{a\alpha}{\varrho\alpha + \sigma\alpha + \tau(a+\alpha)\sqrt{(\lambda-1)}} \\ \text{poster.} = \frac{a\alpha}{\varrho\alpha + \sigma\alpha + \tau(a+\alpha)\sqrt{(\lambda-1)}} \end{cases}$$

$$Q Q \text{ radius faciei} \begin{cases} \text{anter.} = \frac{b\beta}{\varrho'\beta + \sigma'\beta + \tau'(b+\beta)\sqrt{(\lambda'-1)}} \\ \text{poster.} = \frac{b\beta}{\varrho'\beta + \sigma'\beta + \tau'(b+\beta)\sqrt{(\lambda'-1)}} \end{cases}$$

et definitis simili modo valoribus  $\mu', \nu'$  ex ratione  $n' : 1$  reperietur spatium diffusionis

$$\xi \xi x x \begin{cases} + \mu \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{\nu}{a\alpha} \right) \\ + \mu' \left( \frac{1}{b} + \frac{1}{\beta} \right) \left( \lambda' \left( \frac{1}{b} + \frac{1}{\beta} \right)^2 + \frac{\nu'}{b\beta} \right) \end{cases}$$

reliqua manent vt in problemate.

## Pro Problemate II.

Quoniam hic duae vitri species occurrunt, ponamus lentem simplicem quaesitam ex alio quocunque vitri genere parari, cuius ratio refractionis sit  $n^\circ : 1$ , vnde prodeant numeri  $\mu^\circ$  et  $\nu^\circ$ ; superfluum autem foret, numeros  $\varrho^\circ, \sigma^\circ$  et  $\tau^\circ$  computari, quoniam ra-

dii

dii facierum fiunt imaginarii, ita, vt eos exprimere non sit opus, et quoniam pro hac lente simplici aequivalente numerus  $\lambda^{\circ}$  est introductus, erit huius lentis spatium diffusionis

$$\frac{1}{a} \frac{1}{b} x x \cdot \mu^{\circ} \left( \frac{1}{a} + \frac{1}{b} \right) \left( \lambda^{\circ} \left( \frac{1}{a} + \frac{1}{b} \right)^2 + \frac{\nu^{\circ}}{a b} \right)$$

quod vt cum spatio diffusionis lentis duplicatae aequale fiat; ponatur, vti in problemate est factum,

$$\frac{1}{a} = \frac{f-1}{a} + \frac{f}{b} \quad \text{et} \quad \frac{1}{b} = \frac{1-f}{a} - \frac{f}{b}$$

vt prodeat.

$$\frac{1}{a} + \frac{1}{a} = f \left( \frac{1}{a} + \frac{1}{b} \right) \quad \text{et} \quad \frac{1}{b} + \frac{1}{b} = (1-f) \left( \frac{1}{a} + \frac{1}{b} \right)$$

Vnde peruenietur ad hanc aequationem

$$\begin{aligned} & \mu f \left( \lambda f^2 \left( \frac{1}{a} + \frac{1}{b} \right)^2 + \nu \left( \frac{f-1}{a^2} + \frac{f}{a b} \right) \right) \\ & + \mu' (1-f) \left( \lambda' (1-f)^2 \left( \frac{1}{a} + \frac{1}{b} \right)^2 + \nu' \left( \frac{1-f}{a b} - \frac{f}{b^2} \right) \right) \\ & = \mu^{\circ} \left( \lambda^{\circ} \left( \frac{1}{a} + \frac{1}{b} \right)^2 + \frac{\nu^{\circ}}{a b} \right) \end{aligned}$$

quae ita repraesentari potest

$$\begin{aligned} & \left( \frac{1}{a} + \frac{1}{b} \right)^2 \left( \mu \lambda f^3 + \mu' \lambda' (1-f)^3 \right) \\ & + \left( \frac{f-1}{a} + \frac{f}{b} \right) \left( \frac{\mu \nu f}{a} - \frac{\mu' \nu' (1-f)}{b} \right) \\ & = \mu^{\circ} \lambda^{\circ} \left( \frac{1}{a} + \frac{1}{b} \right)^2 + \frac{\nu^{\circ} \mu^{\circ}}{a b} \end{aligned}$$

Vnde

$$\lambda^{\circ} = \frac{\mu \lambda f^3 + \mu' \lambda' (1-f)^3}{\mu^{\circ}}$$

$$\frac{a a b b}{\mu^{\circ} (a + b)^2} \left( \left( \frac{f-1}{a} + \frac{f}{b} \right) \left( \frac{\mu \nu f}{a} - \frac{\mu' \nu' (1-f)}{b} \right) - \frac{\nu^{\circ} \mu^{\circ}}{a b} \right)$$

Deni-



Denique si ex ratione refractionis  $n' : n$  computentur numeri  $\rho'$ ,  $\sigma'$  et  $\tau'$ , radii facierum lentis duplicatae erunt

Pro Lente

$$\begin{aligned} \text{P P radius faciei} & \begin{cases} \text{anterioris} = \frac{a \epsilon}{(\rho - \sigma(1-f))\epsilon + \sigma f a + \tau f(a + \epsilon)\sqrt{\lambda - 1}} \\ \text{posterioris} = \frac{a \epsilon}{(\sigma - \rho(1-f))\epsilon + \rho f a + \tau f(a + \epsilon)\sqrt{\lambda - 1}} \end{cases} \\ \text{Q Q radius faciei} & \begin{cases} \text{anter.} = \frac{a \epsilon}{(\sigma' - \rho' f)a + \rho'(1-f)\epsilon + \tau'(1-f)(a + \epsilon)\sqrt{\lambda' - 1}} \\ \text{poster.} = \frac{a \epsilon}{(\rho' - \sigma' f)a + \sigma'(1-f)\epsilon + \tau'(1-f)(a + \epsilon)\sqrt{\lambda' - 1}} \end{cases} \end{aligned}$$

### Ad Problema III.

Hoc problema non solum pro diuersa refractione  $n$  et  $n'$  hic generalius pertractabo, sed etiam rationem distantiae inter binas lentes habebo. Primum igitur vtramque lentem ad distantias determinatrices lentis duplicatae,  $a$  et  $\epsilon$ , reuocabo, ponendo

$$\frac{1}{a} + \frac{1}{\alpha} = f\left(\frac{1}{a} + \frac{1}{\epsilon}\right) \text{ et } \frac{1}{b} + \frac{1}{\epsilon} = g\left(\frac{1}{a} + \frac{1}{\epsilon}\right)$$

vt fit

$$\frac{1}{\alpha} = \frac{f-1}{a} + \frac{f}{\epsilon} \text{ et } \frac{1}{b} = \frac{g}{a} + \frac{g-1}{\epsilon}$$

hincque

$$a = \frac{a \epsilon}{f a + (f-1)\epsilon} \text{ et } b = \frac{a \epsilon}{(g-1)a + g \epsilon}$$

ideoque distantia lentium

$$a + b = \frac{a \epsilon (f + g - 1)(a + \epsilon)}{(f a + (f-1)\epsilon)((g-1)a + g \epsilon)}$$

quae si deberet esse  $= 0$ , capi oporteret  $g = 1 - f$  sed si distantiam aliquam inter lentes admittamus, statuamus

tuamus  $f+g-1=\omega$ , denotante  $\omega$  fractionem quandam minimam, siue positiuam siue negatiuam vt distantia lentium prodeat positua et valde parua. Cum hinc igitur fit  $g=1+\omega-f$ ; erit lentium distantia

$$a+b=\frac{a\epsilon(a+\epsilon)\omega}{(fa+(f-1)\epsilon)((\omega-f)a+(1+\omega-f)\epsilon)}$$

Spatium autem diffusionis nunc ita exprimetur:

$$\epsilon\epsilon xx.(\frac{1}{a}+\frac{1}{\epsilon})\left\{\begin{array}{l} +\lambda\mu f^3(\frac{1}{a}+\frac{1}{\epsilon})^2+\lambda'\mu'g^3(\frac{1}{a}+\frac{1}{\epsilon})^2 \\ +\frac{\mu v f}{a}(\frac{f-1}{a}+\frac{f}{\epsilon}) \\ +\frac{\mu'v'g}{\epsilon}(\frac{g}{a}+\frac{g-1}{\epsilon}) \end{array}\right.$$

quae cum in hoc problemate ita tractari debeat, vt tam  $f$ , quam  $g$  positue sumantur, et haec formula minima reddatur; euident est, litteris  $\lambda$  et  $\lambda'$  minimos valores tribui debere, scilicet  $\lambda=1$  et  $\lambda'=1$ ; deinde pro hoc casu minimi  $f$  et  $g$  conuenienter definiantur, vbi quia  $f+g-1=\omega$  ideoque constans in differentiatione, habebimus  $dg=-df$ ; vnde obtinebimus hanc aequationem

$$(3\mu f^2-3\mu'g^2)(\frac{1}{a}+\frac{1}{\epsilon})^2+\frac{\mu v}{a}(\frac{2f-1}{a}+\frac{2f}{\epsilon})-\frac{\mu'v'}{\epsilon}(\frac{2g}{a}+\frac{2g-1}{\epsilon})=0$$

cui proxime satisfit, ponendo  $f=g=\frac{1+\omega}{2}$  quibus valoribus substitutis spatium diffusionis ipsum minimum, erit proxime

$$\frac{\epsilon^2 x^2 (1+\omega)}{8}(\frac{1}{a}+\frac{1}{\epsilon})\left\{\begin{array}{l} (\frac{1}{a}+\frac{1}{\epsilon})^2(1+\omega)^2(\mu+\mu') \\ +\frac{2\mu v}{a}(\frac{\omega-1}{a}+\frac{1+\omega}{\epsilon}) \\ +\frac{2\mu'v'}{\epsilon}(\frac{1+\omega}{a}+\frac{\omega-1}{\epsilon}) \end{array}\right.$$

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Tum



Tum vero distantia lentium erit

$$a+b = \frac{4ab(a+g)\omega}{((1+\omega)a + (\omega-1)g)((\omega-1)a + (1+\omega)g)}$$

$$= \frac{4ab(a+g)\omega}{-(1-\omega^2)(a^2+g^2) + 2(1+\omega)^2 ab}$$

cuius denominator cum sit negatiuus ob  $\omega$  minimum, necesse est, fractionem  $\omega$  sumi debere negatiuam; hincque adeo spatium diffusionis minus reddetur.

### COROLL.

Si ergo distantia obiecti  $a$  fuerit infinita seu  $a = \infty$ , habebitur primo distantia lentium  $= \frac{+g\omega}{1-\omega^2}$  et secundo spatium diffusionis

$$\frac{x^2(1+\omega)}{8g} \left\{ \begin{array}{l} (1+\omega)^2(\mu+\mu') \\ - 2\mu'\nu'(1-\omega) \end{array} \right\}$$

quod cum  $\omega$  debeat esse negatiuum, non mediocriter minus erit, quam si distantia lentium esset nulla.

### Ad Problema IV.

In hoc problemate etiam distantiam lentium non negligamus; factaque reductione, ut ante, statuamus spatium diffusionis plane euanesceat; id quod fieri nequit, nisi altera litterarum  $f$  et  $g$  sit negatiua, quod cum etiam fiat, quando confusio a diuersa radiorum refrangibilitate oriunda ad nihilum redigi debet, ut infra videbimus; hic casus multo magis euolutionem meretur. Ponatur igitur  $g = -\zeta f$ , vbi  $\zeta$  ex illa

illa conditione determinatur, vt primo pro distantia lentium fit

$$\alpha + b = \frac{-a\epsilon(a+\epsilon)(f-\zeta f-1)}{(fa+(f-1)\epsilon)((\zeta f+1)a+\zeta f\epsilon)}$$

et posito  $f+g-1=\omega$ , prodeat  $f-\zeta f=1+\omega$  ideoque

$$f=\frac{1+\omega}{1-\zeta} \text{ et } g=\frac{-\zeta(1+\omega)}{1-\zeta} \text{ ficque}$$

$$\alpha + b = \frac{-a\epsilon(a+\epsilon)\omega(1-\zeta)^2}{((1+\omega)a+(\omega+\zeta)\epsilon)((1+\omega\zeta)a+\zeta(1+\omega)\epsilon)}$$

ac si in denominatore  $\omega$  reiciatur, erit

$$\alpha + b = \frac{-a\epsilon(a+\epsilon)(1-\zeta)^2\omega}{(a+\zeta\epsilon)}$$

ficque patet  $\omega$  negative capi debere.

Posito autem  $g=-\zeta f$ , fit spatium diffusionis

$$\epsilon^2 x^2 \left(\frac{1}{a} + \frac{1}{\epsilon}\right) \left\{ \begin{array}{l} (\lambda\mu f^3 - \lambda'\mu'\zeta^3 f^3) \left(\frac{1}{a} + \frac{1}{\epsilon}\right)^2 \\ + \frac{\mu v f}{a} \left(\frac{f-1}{a} + \frac{f}{\epsilon}\right) \\ - \frac{\mu' v \zeta f}{\epsilon} \left(\frac{-\zeta f}{a} - \frac{(\zeta f+1)}{\epsilon}\right) \end{array} \right.$$

ad nihilum redigendum; vnde sequitur fore

$$\lambda' = \frac{\lambda\mu}{\mu'\zeta^3} + \frac{\mu v a \epsilon^2}{\mu'\zeta^3 f^2 (a+\epsilon)^2} \left(\frac{f-1}{a} + \frac{f}{\epsilon}\right) + \frac{v a^2 \epsilon}{\zeta^2 f^2 (a+\epsilon)^2} \left(+ \frac{\zeta f}{a} + \frac{\zeta f+1}{\epsilon}\right)$$

qui valor cum debeat esse maior vnitatem, si forte eueniat, vt minor prodeat, tunc non  $\lambda'$  sed  $\lambda$  definiri conueniet, vbi notandum est esse  $f=\frac{1+\omega}{1-\zeta}$ . Hincque per formulas ante datas facile eruuntur radii facierum vtriusque lentis.

Etsi formula superius data pro spatium diffusionis iam ad casum, quo distantia lentium est nulla, est



adcommodata ; tamen quia hic distantiam minimam assumimus , nullus inde error est metuendus.

### Definitio 2.

122. *Lens triplicata est , quae consistit ex tribus lentibus simplicibus sibi immediate iunctis ad communem axem.*

Hic quidem etiam crassitiem negligo , etiam si necessario maior sit quam in lentibus duplicatis. In supplemento autem ostendetur , quomodo etiam distantiarum inter lentes ratio sit habenda.

### Corollarium.

123. Potest ergo lens triplicata considerari , quasi composita ex lente duplicata et lente simplici , hocque duplici modo , prout vel binae anteriores , vel binae posteriores lentem duplicatam constituere concipiuntur.

### Scholion.

124. Lentibus triplicatis tum demum usum in praxi concedi conueniet , cum eadem commoda per lentes simplices vel duplicatas consequi non licet. Haec autem commoda in paruitate numeri  $\lambda$  consistunt , qui quamdiu unitate fuerit maior , semper lente simplici uti praestat , cum vero circumstantiae in expressione diffusionis minorem numerum  $\lambda$  requirunt , ad lentes multiplicatas erit confugiendum. Quoniam igitur eiusmodi lentes duplicatas conficere docuimus ,  
in

in quibus valor ipsius  $\lambda$  non solum ad nihilum vsque, sed adeo ad negatiua diminui queat, vsus lentium triplicatarum superfluous videtur. Verum iam obseruauimus in praxi aegre eiusmodi duplicatas lentes parari posse, pro quibus valor ipsius  $\lambda$  minor sit quam 0, 191827, propterea quod si leuissimus error committatur, omnis labor irritus reddatur. His igitur casibus imprimis, quando minori valore numeri  $\lambda$  opus est, lentes triplicatae in usum erunt vocandae: et quia respectu ad praxin habito non omnes aequo successu construere licet, errores ineuitabiles hic quoque imprimis spectari oportet, vt pateat quousque numerus  $\lambda$  cum successu diminui queat, ac si adhuc minore valore opus fuerit lentes adeo quadruplicate erunt inducendae.

### Pr o b l e m a 5.

125. Datis binis distantis determinatricibus omnes Tab. II. lentes triplicatas definire, simulque spatium diffusionis Fig. 8. ab iis ortum.

### S o l u t i o.

Sit  $E\epsilon$  obiectum, cuius magnitudo  $=z$ , et distantia a lente  $AE=a$ : tum primae lentis  $PP$  distantiae determinatrices sint  $a$  et  $\alpha$ , secundae lentis  $QQ$   $b$  et  $\beta$ ; ac tertiae  $RR$ ,  $c$  et  $\gamma$ . His positis quia lentes immediate iunctae sumuntur, erit  $a+b=0$ ,  $\beta+c=0$ , et  $a$  et  $\gamma$  distantiae determinatrices lentis triplicatae, ita vt  $\alpha$  et  $\beta$  arbitrio nostro relinquantur.

N. 3

Si



Si igitur lens prima sola PP adestet, imago repraesentaretur in F $\zeta$ , vt effet AF =  $\alpha$  et F $\zeta$  =  $\frac{\alpha}{a}z$ : si binae priores PP et QQ solae adestent imago exhiberetur in G $\eta$ , vt effet AG =  $\epsilon$  et G $\eta$  =  $\frac{\epsilon}{a}z$ ; at per lentem triplicatam referetur in H $\theta$  vt sit cH =  $\gamma$  et H $\theta$  =  $\frac{\gamma}{a}z$ , pro situ inuerso ob  $\frac{\alpha\epsilon}{b\gamma} = 1$ . Haftenus scilicet res perinde se habet, ac si in A haberetur lens simplex ad distantias determinatrices  $a$  et  $\gamma$  accommodata.

At si ad spatium diffusionis Hb respiciamus figuram singularum lentium in computum ducere debemus, quatenus praeter distantias determinatrices numeri arbitrarii  $\lambda, \lambda', \lambda''$  inuoluuntur; ex quibus facies singularum lentium supra §. 91 sunt definitae. Indidem autem colligitur pro apertura cuius semidiameter =  $x$ , spatium diffusionis ob  $\frac{\alpha}{b} = -1$  et  $\frac{\epsilon}{c} = -1$  fore:

$$Hb = \mu \gamma \gamma' x x \left\{ \begin{array}{l} + (\frac{1}{a} + \frac{1}{\alpha})(\lambda(\frac{1}{a} + \frac{1}{\alpha})^2 + \frac{\nu}{\alpha a}) \\ + (\frac{1}{b} + \frac{1}{\epsilon})(\lambda'(\frac{1}{b} + \frac{1}{\epsilon})^2 + \frac{\nu}{b \epsilon}) \\ + (\frac{1}{c} + \frac{1}{\gamma})(\lambda''(\frac{1}{c} + \frac{1}{\gamma})^2 + \frac{\nu}{c \gamma}) \end{array} \right.$$

quae expressio, vt ad formam vni lenti respondentem reducatur statuamus:

$$\frac{1}{a} + \frac{1}{\alpha} = f(\frac{1}{a} + \frac{1}{\gamma})$$

$$\frac{1}{b} + \frac{1}{\epsilon} = g(\frac{1}{a} + \frac{1}{\gamma})$$

$$\frac{1}{c} + \frac{1}{\gamma} = h(\frac{1}{a} + \frac{1}{\gamma})$$

et

et quia  $\frac{1}{a} + \frac{1}{b} = 0$ , et  $\frac{1}{b} + \frac{1}{c} = 0$ , his aequationibus addendis adipiscimur:  $f + g + b = 1$ . Porro vero erit

$$\frac{1}{a} = f\left(\frac{1}{a} + \frac{1}{\gamma}\right) - \frac{1}{a}; \quad \frac{1}{b} = \frac{1}{a} - f\left(\frac{1}{a} + \frac{1}{\gamma}\right)$$

$$\frac{1}{b} = (f + g)\left(\frac{1}{a} + \frac{1}{\gamma}\right) - \frac{1}{a}; \quad \frac{1}{c} = \frac{1}{a} - (f + g)\left(\frac{1}{a} + \frac{1}{\gamma}\right)$$

sive ob  $1 = f + g + b$

$$\frac{1}{a} = \frac{f + g + b}{a}; \quad \frac{1}{a} = -\frac{g + b}{a} + \frac{f}{\gamma}$$

$$\frac{1}{b} = \frac{g + b}{a} - \frac{f}{\gamma}; \quad \frac{1}{b} = -\frac{b}{a} + \frac{f + g}{\gamma}$$

$$\frac{1}{c} = \frac{b}{a} - \frac{f + g}{\gamma}; \quad \frac{1}{\gamma} = \frac{f + g + b}{\gamma}$$

Vnde cum spatium diffusionis fiat.

$$Hb = \mu \gamma \gamma x a \left(\frac{1}{a} + \frac{1}{\gamma}\right) (\lambda f^3 + \lambda' g^3 + \lambda'' b^3) \left(\frac{1}{a} + \frac{1}{\gamma}\right)^2 + \nu \left(\frac{f}{a a} + \frac{g}{b b} + \frac{b}{c c}\right)$$

reducetur id ad hanc formam

$$Hb = \mu \gamma \gamma x a \left(\frac{1}{a} + \frac{1}{\gamma}\right) (\lambda f^3 + \lambda' g^3 + \lambda'' b^3 - \nu(1 - f)(1 - g)(1 - b)) \left(\frac{1}{a} + \frac{1}{\gamma}\right)^2 + \frac{\nu}{a \gamma}$$

Radiatorum autem in  $b$  concurrentium inclinatio ad axem erit  $= \frac{\infty}{\gamma}$ .

### COROLL. I.

126. Haec igitur lens triplicata idem producit spatium diffusionis quod produceret lens simplex ad easdem distantias determinatrices instructa, numero eius arbitrario (per litteram  $\lambda$  indicato) existente

$$\lambda f^3 + \lambda' g^3 + \lambda'' b^3 - \nu(1 - f)(1 - g)(1 - b)$$

vbi quidem est  $f + g + b = 1$ .

Coroll.



## Coroll. 2.

127. Quatenus ergo haec quantitas reddi potest minor non solum unitate ; sed etiam fractione 0,191827 ita scilicet , vt praxis non enormi aberrationi sit exposita , eatenus lentibus triplicatis vsus erit concedendus.

## Coroll. 3.

128. Si sit vel  $f=0$  , vel  $g=0$  , vel  $b=0$  , vna lentium habebit facies parallelas , et lens triplicata aequiualebit duplicatae ac si duae litterarum  $f, g, b$  simul euanescant , tertia in unitatem abeunte , casus habebitur lentis simplicis.

## Coroll. 4.

129. Si sit  $f=1$  , ideoque  $b=-g$  , valor numeri  $\lambda$  pro lente triplicata erit  $\lambda + (\lambda' - \lambda'')g^3$  , ideoque si  $\lambda'' = \lambda'$  lens triplicata simplici aequiualebit , quod idem euenit si fuerit vel  $g=1$  vel  $b=1$ .

## Coroll. 5.

130. Sumtis autem pro  $f, g, b$  numeris idoneis ob  $f+g+b=1$  , constructio lentis triplicatae ex formulis §. 91 exhibitis est petenda sumendo.

$$\frac{1}{a} = \frac{-1+f}{a} + \frac{f}{\gamma} ; \quad \frac{1}{b} = \frac{1-f}{a} - \frac{f}{\gamma}$$

$$\frac{1}{c} = \frac{-b}{a} + \frac{1-b}{\gamma} ; \quad \frac{1}{c} = \frac{b}{a} - \frac{1+b}{\gamma}$$

Scholion

## Scholion 1.

131. Quemadmodum hinc expressio inuenta prodeat, notandum est fore  $\frac{f}{a\alpha} + \frac{g}{b\beta} + \frac{b}{c\gamma} =$

$$+ \frac{1}{a\alpha}(-f(1-f) - gb(1-f))$$

$$+ \frac{1}{a\gamma}(+ff + g(1-f)(1-b) + fgb + hb)$$

$$+ \frac{1}{\gamma\gamma}(-fg(1-b) - b(1-b))$$

sed  $-f(1-f) - gb(1-f) = -(1-f)(f+gb) = -(1-f)(1-g)(1-b)$

ob  $f = 1-g-b$  ideoque  $f+gb = (1-g)(1-b)$ . Simili modo pro

$\frac{1}{\gamma\gamma}$  est  $-fg(1-b) - b(1-b) = -(1-b)(b+fg) = -(1-f)(1-g)(1-b)$

ob  $b = 1-f-g$ . Denique pro  $\frac{1}{a\gamma}$ , quia est  $ff + hb = (f+b)^2 - 2fb = (1-g)^2 - 2fb = 1 - 2g + gg - 2fb$ ,

hoc valore substituto coefficiens ipsius  $\frac{1}{a\gamma}$  erit

$$1 - 2g - 2fb + gg + fgb + g(1-f)(1-b) =$$

$$1 + (g+fb)(g-2) - g(1-f)(1-b) = 1 - 2(1-f)(1-g)(1-b)$$

ob  $g+fb = (1-f)(1-b)$ . Consequenter colligitur

$$\frac{f}{a\alpha} + \frac{g}{b\beta} + \frac{b}{c\gamma} = -(1-f)(1-g)(1-b)\left(\frac{1}{a} + \frac{1}{\gamma}\right)^2 + \frac{1}{a\gamma}.$$

## Scholion 2.

132. Hoc problema etiam ope praecedentium facilius sequenti modo resolui potest. Considerentur scilicet binae lentes PP et QQ iunctim sumtae tanquam lens duplicata ad distantias determinatrices  $a$  et  $\beta$  per numeros arbitrarios  $\lambda$ ,  $\lambda'$  et  $f$  instructa, ac posito  $\lambda f^3 + \lambda'(1-f)^3 - \nu f(1-f) = \lambda^{(2)}$ , spatium diffusionis ex ea sola ortum erit

$$\mu \beta \beta x x \left(\frac{1}{a} + \frac{1}{\beta}\right) (\lambda^{(2)} \left(\frac{1}{a} + \frac{1}{\beta}\right)^2 + \frac{\nu}{a\beta})$$

Tom. I.

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quae



quae lens si iam in compositione cum tertia R R  
tanquam simplex tractetur exinde elicitur spatium  
diffusionis perinde atque ex coniunctione duarum  
simplicium

$$Hb = \mu \gamma \gamma x x \begin{cases} + (\frac{1}{a} + \frac{1}{g})(\lambda^{(2)}(\frac{1}{a} + \frac{1}{g})^2 + \frac{v}{ag}) \\ + (\frac{1}{c} + \frac{1}{\gamma})(\lambda''(\frac{1}{c} + \frac{1}{\gamma})^2 + \frac{v}{c\gamma}) \end{cases}$$

vbi notandum est esse  $g + c = 0$ , et constructio lentis  
duplicatae ex §. 107 erit petenda lentis vero simpli-  
cis R R ex distantis determinatricibus  $c = -g$  et  $\gamma$   
vna cum numero arbitrario  $\lambda''$ . Ponatur iam  $\frac{1}{g} = \frac{g-1}{a} + \frac{g}{\gamma}$   
vt fit  $\frac{1}{c} = \frac{1-g}{a} - \frac{g}{\gamma}$ , eritque spatium diffusionis huius  
lentis triplicatae,

$$Hb = \mu \gamma \gamma x x (\frac{1}{a} + \frac{1}{\gamma}) ((\lambda^{(2)} g^3 + \lambda''(1-g)^3 - \nu g(1-g))(\frac{1}{a} + \frac{1}{\gamma})^2 + \frac{v}{a\gamma})$$

Quare si pro lente triplicata ponatur

$$\lambda^{(2)} g^3 + \lambda''(1-g)^3 - \nu g(1-g) = \lambda^{(3)}$$

ita vt iam numerus  $g$  insuper arbitrio nostro relin-  
quatur, habebitur spatium diffusionis more hactenus  
recepto expressum

$$Hb = \mu \gamma \gamma x x (\frac{1}{a} + \frac{1}{\gamma}) (\lambda^{(3)}(\frac{1}{a} + \frac{1}{\gamma})^2 + \frac{v}{a\gamma})$$

Hinc iam id intelligitur, quod ex praecedente solutione  
minus patet; si numerus  $\lambda^{(3)}$  fuerit vnitatem maior loco  
binarum priorum lentium PP et QQ commodius  
unicam simplicem adhiberi; ex quo eatenus tantum  
lentes triplicatae resultare censendae sunt, quatenus  
nume-

numerus  $\lambda^{(2)}$  unitate est minor. Vidimus autem successum tuto sperari non posse, nisi  $\lambda^{(1)}$  aequalis sit fractioni  $0,191827$ , vel ea non multo maior; unde si praxi consulere velimus, ipsi  $\lambda^{(2)}$  minorem valorem tribui non couenit atque ob eandem rationem numerus  $g$  intra terminos  $0$  et  $1$  accipi debet; cuius valor imprimis ad praxin erit accommodatus si reddat numerum  $\lambda^{(3)}$  minimum, quia tum leues errores negotium minime turbant. Quodsi vero pro  $\lambda^{(2)}$  valorem assignatum substituamus, obtinebimus

$$\lambda^{(3)} = \lambda f^3 g^3 + \lambda' g^3 (1-f)^3 + \lambda'' (1-g)^3 - \nu g^3 f(1-f) - \nu g(1-g)$$

Conducat autem vtramque expressionem pro numero  $\lambda^{(3)}$ , quo spatium diffusionis a lente triplicata ortum definitur, hic exposuisse cum aliae conclusiones ex altera facilius deducantur. Etsi autem hinc omnes valores pro  $\lambda^{(3)}$  obtineri possunt tamen eos tantum, qui prope minimum subsistunt ad praxin adhiberi conueniet.

### Problema 6.

133. Datis distantis determinatricibus  $AE = a$  et  $cH = \gamma$  definire eam lentem triplicatam, quae minimum spatium diffusionis producat.

### Solutio I.

Duplici modo hoc problema solui potest, prout spatium diffusionis vel ita exprimitur, vti in solutione problematis praecedentis, vel in Scholio 2. Pri-



ori modo numeros  $f, g, h$  ita determinari oportet, ut minima reddatur haec expressio:

$\lambda f^3 + \lambda' g^3 + \lambda'' h^3 - v(1-f)(1-g)(1-h)$   
vbi notandum est esse  $f+g+h=1$ . Eius ergo differentiali nihilo aequali posito habebimus:

$$3\lambda f f df + 3\lambda' g g dg + 3\lambda'' h h dh + v df(1-g)(1-h) + v dg(1-f)(1-h) + v dh(1-f)(1-g) = 0.$$

Cum autem fit  $dh = -df - dg$  erit

$$\left. \begin{aligned} &+ 3\lambda f f df - 3\lambda'' h h df + v df(1-g)(f-h) \\ &+ 3\lambda' g g dg - 3\lambda'' h h dg + v dg(1-f)(g-h) \end{aligned} \right\} = 0$$

quia vero bina differentialia  $df$  et  $dg$  a se inuicem non pendent ambo membra huius aequationis seorsim euanescere debent, unde ob  $1-g=f+h$  et  $1-f=g+h$  has duas nanciscimur aequationes:

$$3\lambda f f - 3\lambda'' h h + v f f - v h h = 0$$

$$3\lambda' g g - 3\lambda'' h h + v g g - v h h = 0$$

ex quibus elicimus:

$$f = h \sqrt{\frac{3\lambda'' + v}{3\lambda + v}} \text{ et } g = h \sqrt{\frac{3\lambda'' + v}{3\lambda' + v}}$$

Cum autem sit

$$f+g+h=1 \text{ seu } \frac{1}{h} = 1 + \frac{f}{h} + \frac{g}{h} \text{ erit}$$

$$\frac{1}{h} = 1 + \sqrt{\frac{3\lambda'' + v}{3\lambda + v}} + \sqrt{\frac{3\lambda'' + v}{3\lambda' + v}}$$

$$\frac{1}{g} = 1 + \sqrt{\frac{3\lambda'' + v}{3\lambda + v}} + \sqrt{\frac{3\lambda'' + v}{3\lambda' + v}}$$

$$\frac{1}{f} = 1 + \sqrt{\frac{3\lambda'' + v}{3\lambda' + v}} + \sqrt{\frac{3\lambda'' + v}{3\lambda + v}}$$

Quicun-

Quicumque ergo numeri  $\lambda$ ,  $\lambda'$ ,  $\lambda''$  in constructione singularum lentium fuerint usurpati, hinc numeri  $f$ ,  $g$ , et  $h$  determinantur, ex quibus spatium diffusionis minimum resultet. Pro lentium autem constructione hinc distantiae  $a$ ,  $b$ ,  $\xi$ ,  $c$  ita definiuntur, ut sit

$$\frac{1}{a} = \frac{1-f}{a} + \frac{f}{\gamma}; \quad \frac{1}{b} = \frac{1-f}{a} - \frac{f}{\gamma}$$

$$\frac{1}{\xi} = \frac{-b}{a} + \frac{1-b}{\gamma}; \quad \frac{1}{c} = \frac{b}{a} - \frac{1+b}{\gamma}$$

est vero :

$$f = \frac{\sqrt{(3\lambda' + v)(3\lambda'' + v)}}{\sqrt{(3\lambda + v)(3\lambda' + v)} + \sqrt{(3\lambda + v)(3\lambda'' + v)} + \sqrt{(3\lambda' + v)(3\lambda'' + v)}}$$

$$g = \frac{\sqrt{(3\lambda + v)(3\lambda'' + v)}}{\sqrt{(3\lambda + v)(3\lambda' + v)} + \sqrt{(3\lambda + v)(3\lambda'' + v)} + \sqrt{(3\lambda' + v)(3\lambda'' + v)}}$$

$$h = \frac{\sqrt{(3\lambda + v)(3\lambda' + v)}}{\sqrt{(3\lambda + v)(3\lambda' + v)} + \sqrt{(3\lambda + v)(3\lambda'' + v)} + \sqrt{(3\lambda' + v)(3\lambda'' + v)}}$$

Ex distantis autem  $a$ ,  $\alpha$ ,  $b$ ,  $\xi$ ,  $c$ ,  $\gamma$  cum numeris  $\lambda$ ,  $\lambda'$ ,  $\lambda''$  lentes ipsae per formulas §. 91 exhibitas construuntur.

### COROLL. I.

134. Si pro hac lente triplicata ponatur

$$\lambda f^3 + \lambda' g^3 + \lambda'' h^3 - v(1-f)(1-g)(1-h) = \lambda^{(3)}$$

substituendis his valoribus pro  $f$ ,  $g$ , et  $h$  reperietur

$$\lambda^{(3)} = \frac{1}{3 \left( \frac{1}{\sqrt{(3\lambda + v)}} + \frac{1}{\sqrt{(3\lambda' + v)}} + \frac{1}{\sqrt{(3\lambda'' + v)}} \right)^2 - \frac{1}{3} v}$$

unde spatium diffusionis fit

$$Hb = \mu \gamma \gamma x x \left( \frac{1}{a} + \frac{1}{\gamma} \right) \left( \lambda^{(3)} \left( \frac{1}{a} + \frac{1}{\gamma} \right)^2 + \frac{v}{a\gamma} \right)$$



## Coroll. 2.

135. Hoc autem spatium diffusionis omnium fiet minimum si numeris  $\lambda$ ,  $\lambda'$ ,  $\lambda''$  minimi valores, quos accipere possunt, tribuantur. Sit ergo  $\lambda = 1$ ,  $\lambda' = 1$ ,  $\lambda'' = 1$ , eritque

$$f = \frac{1}{3}; \quad g = \frac{1}{3}, \quad h = \frac{1}{3}$$

$$\text{et } \lambda^{(3)} = \frac{\gamma + 3}{27} - \frac{\gamma}{3} = \frac{3 - 3\gamma}{27} = 0,042165$$

ob 0,232692; qui ergo valor multo est minor, quam casu lentium duplicatarum.

## Coroll. 3.

136. Hoc porro casu erit  $\frac{1}{a} = -\frac{2}{3a} + \frac{1}{3\gamma}$ ;  $\frac{1}{b} = \frac{2}{3a} - \frac{1}{3\gamma}$ ;  $\frac{1}{c} = \frac{1}{3a} + \frac{2}{3\gamma}$  et  $\frac{1}{d} = \frac{1}{3a} - \frac{2}{3\gamma}$ ; vnde constructio lentium ternarum simplicium ita se habebit.

## Pro Lente

$$\text{Prima radius faciei} \begin{cases} \text{anterioris} = \frac{3a\gamma}{(3\rho - 2\sigma)\gamma + \sigma a} \\ \text{posterioris} = \frac{3a\gamma}{(3\sigma - 2\rho)\gamma + \rho a} \end{cases}$$

$$\text{Secunda radius faciei} \begin{cases} \text{anterioris} = \frac{3a\gamma}{(2\rho - \sigma)\gamma + (2\sigma - \rho)a} \\ \text{posterioris} = \frac{3a\gamma}{(2\sigma - \rho)\gamma + (2\rho - \sigma)a} \end{cases}$$

$$\text{Tertia radius faciei} \begin{cases} \text{anterioris} = \frac{3a\gamma}{\rho\gamma + (3\sigma - 2\rho)a} \\ \text{posterioris} = \frac{3a\gamma}{\sigma\gamma + (3\rho - 2\sigma)a} \end{cases}$$

## Solutio altera Problematis.

137. Consideremus binas lentes priores PP et QQ vt lentem duplicatam ad distantias determinatrices  $a$  et  $b$  ita

ita instructam ut posito  $\lambda f^3 + \lambda'(1-f)^3 - \nu f(1-f) = \lambda^{(2)}$  spatium diffusionis inde oriundum sit

$$= \mu \vartheta \vartheta x x \left( \frac{1}{a} + \frac{1}{e} \right) \left( \lambda^{(2)} \left( \frac{1}{a} + \frac{1}{e} \right)^2 + \frac{\nu}{a e} \right)$$

sed pro constructione binarum lentium simplicium, ex quibus haec lens est composita, recordandum est esse

$$\frac{1}{a} = -\frac{1}{a} + f + \frac{f}{e} \quad \text{et} \quad \frac{1}{b} = \frac{1}{a} - f - \frac{f}{e}$$

Adiuncta iam tertia lente RR ad distantias determinatrices  $c = -\vartheta$  et  $\gamma$  per numerum arbitrium  $\lambda''$  instructa si ponamus  $\frac{1}{e} = -\frac{1}{a} + \frac{\vartheta}{\gamma} + \frac{\vartheta}{\gamma} = -\frac{1}{c}$ , et

$$\lambda^{(2)} g^3 + \lambda''(1-g)^3 - \nu g(1-g) = \lambda^{(3)}$$

erit spatium diffusionis

$$Hh = \mu \gamma \gamma x x \left( \frac{1}{a} + \frac{1}{\gamma} \right) \left( \lambda^{(3)} \left( \frac{1}{a} + \frac{1}{\gamma} \right)^2 + \frac{\nu}{a \gamma} \right)$$

quod ut fiat minimum, valor ipsius  $\lambda^{(3)}$  minimus reddi debet, quaeratur ergo primo  $g$  dum  $\lambda^{(2)}$  ut numerus datus spectatur et habebimus

$$3\lambda^{(2)} g g - 3\lambda''(1-g)^2 - \nu + 2\nu g = 0$$

vnde elicitur

$$g = \frac{-3\lambda'' - \nu + \sqrt{(3\lambda^{(2)} + \nu)(3\lambda'' + \nu)}}{3\lambda^{(2)} - 3\lambda''} \quad \text{siue}$$

$$g = \frac{\sqrt{(3\lambda^{(2)} + \nu)}}{\sqrt{(3\lambda^{(2)} + \nu)} + \sqrt{(3\lambda'' + \nu)}}$$

et hinc valor ipsius  $\lambda^{(3)}$  erit

$$\lambda^{(3)} = \frac{1}{3 \left( \frac{1}{\sqrt{(3\lambda^{(2)} + \nu)}} + \frac{1}{\sqrt{(3\lambda'' + \nu)}} \right)^2} - \frac{1}{3} \nu$$

vel



vel

$$\frac{1}{3\lambda^{(3)} + \nu} = \left( \frac{1}{V(3\lambda^{(2)} + \nu)} + \frac{1}{V(3\lambda'' + \nu)} \right)$$

Sicque fit:

$$\frac{1}{V(3\lambda^{(3)} + \nu)} = \frac{1}{V(3\lambda^{(2)} + \nu)} + \frac{1}{V(3\lambda'' + \nu)}$$

Simili modo si  $f$  ita definiatur, ut  $\lambda^{(2)}$  fiat minimum reperietur:

$$f = \frac{\nu(3\lambda' + \nu)}{\nu(3\lambda + \nu) + \nu(3\lambda' + \nu)}$$

hocque valore substituto

$$\frac{1}{V(3\lambda^{(2)} + \nu)} = \frac{1}{V(3\lambda + \nu)} + \frac{1}{V(3\lambda' + \nu)}$$

Quare si numeris  $\lambda, \lambda', \lambda''$ , arbitrio nostro relictis bini numeri  $f$  et  $g$  ita definiantur, ut  $\lambda^{(3)}$  consequatur valorem minimum erit

$$\frac{1}{V(3\lambda^{(3)} + \nu)} = \frac{1}{V(3\lambda + \nu)} + \frac{1}{V(3\lambda' + \nu)} + \frac{1}{V(3\lambda'' + \nu)}$$

vnde idem valor pro  $\lambda^{(3)}$  reperitur, quem ante inuenimus.

### COROLL. I.

138. Ex hac ergo solutione numeri  $f$  et  $g$  ita definiuntur ut sit

$$\frac{1}{fV(3\lambda + \nu)} = \frac{1}{V(3\lambda + \nu)} + \frac{1}{V(3\lambda' + \nu)} \quad \text{et}$$

$$\frac{1}{gV(3\lambda^{(2)} + \nu)} = \frac{1}{V(3\lambda^{(2)} + \nu)} + \frac{1}{V(3\lambda'' + \nu)}$$

siue

sive hoc modo

$$\left(\frac{1}{f} - 1\right) \frac{1}{\sqrt{(3\lambda + \nu)}} = \frac{1}{\sqrt{(3\lambda' + \nu)}} \text{ et}$$

$$\left(\frac{1}{g} - 1\right) \left( \frac{1}{\sqrt{(3\lambda + \nu)}} + \frac{1}{\sqrt{(3\lambda' + \nu)}} \right) = \frac{1}{\sqrt{(3\lambda'' + \nu)}}$$

### Coroll. 2.

139. Ex inuentis minimis valoribus numerorum  $\lambda^{(2)}$  et  $\lambda^{(3)}$  numeri  $f$  et  $g$  etiam ita definiuntur ut sit

$$\frac{1}{f\sqrt{(3\lambda + \nu)}} = \frac{1}{\sqrt{(3\lambda^{(2)} + \nu)}} \text{ seu } f = \sqrt{\left(\frac{3\lambda^{(2)} + \nu}{3\lambda + \nu}\right)} \text{ et}$$

$$\frac{1}{g\sqrt{(3\lambda^{(2)} + \nu)}} = \frac{1}{\sqrt{(3\lambda^{(3)} + \nu)}} \text{ seu } g = \sqrt{\left(\frac{3\lambda^{(3)} + \nu}{3\lambda^{(2)} + \nu}\right)}$$

### Coroll. 3

140. Distantiae autem determinatrices singulorum lentium ita per  $a$  et  $\gamma$  prodibunt expressae:

$$\frac{1}{a} = -\frac{1}{b} = \frac{-1 + fg}{a} + \frac{fg}{\gamma}$$

$$\frac{1}{c} = -\frac{1}{c} = \frac{-1 + g}{a} + \frac{g}{\gamma}$$

vbi cum sit

$$g = \sqrt{\frac{3\lambda^{(3)} + \nu}{3\lambda^{(2)} + \nu}} \text{ notandum est esse } fg = \sqrt{\frac{3\lambda^{(3)} + \nu}{3\lambda + \nu}}$$

### Coroll. 4.

141. Eliminando autem numero  $\lambda^{(2)}$  erit

$$\frac{1}{a} = -\frac{1}{b} = -\frac{1}{a} \sqrt{\frac{3\lambda^{(3)} + \nu}{3\lambda' + \nu}} - \frac{1}{a} \sqrt{\frac{3\lambda^{(3)} + \nu}{3\lambda'' + \nu}} + \frac{1}{\gamma} \sqrt{\frac{3\lambda^{(3)} + \nu}{3\lambda + \nu}}$$

$$\frac{1}{c} = -\frac{1}{c} = -\frac{1}{a} \sqrt{\frac{3\lambda^{(3)} + \nu}{3\lambda'' + \nu}} + \frac{1}{\gamma} \sqrt{\frac{3\lambda^{(3)} + \nu}{3\lambda + \nu}} + \frac{1}{\gamma} \sqrt{\frac{3\lambda^{(3)} + \nu}{3\lambda' + \nu}}$$

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ex quibus formulis, inuento iam valore minimo  $\lambda^{(3)}$  singulae lentes commodissime determinantur.

### Coroll. 5.

142. Si praeterea singulae lentes ita fuerint comparatae vt per se minimam confusionem pariant quod fit si  $\lambda = \lambda' = \lambda'' = 1$ , erit

$$\frac{1}{V(3\lambda^{(2)} + \nu)} = \frac{2}{V(3 + \nu)} \text{ hincque } 3\lambda^{(2)} + \nu = \frac{3 + \nu}{4}$$

vnde fit  $\lambda^{(2)} = \frac{1 - \nu}{4}$ . Deinde vero habebitur:

$$\frac{1}{V(3\lambda^{(3)} + \nu)} = \frac{3}{V(3 + \nu)} \text{ hincque } 3\lambda^{(3)} + \nu = \frac{3 + \nu}{9}$$

ac propterea

$\lambda^{(3)} = \frac{3 - \nu}{27}$ . Sin autem fuerit tantum  $\lambda = \lambda' = \lambda''$  reperietur:

$$\lambda^{(2)} = \frac{3\lambda - \nu}{12} \text{ et } \lambda^{(3)} = \frac{3\lambda - \nu}{27}$$

### Coroll. 6.

143. Eodem autem casu, quo  $\lambda = \lambda' = \lambda''$ , ob  $V(3\lambda^{(3)} + \nu) = \frac{1}{3}V(3\lambda + \nu)$ , distantiae determinatrices pro lentibus simplicibus erunt:

$$\frac{1}{a} = \frac{-2}{3a} + \frac{1}{3\gamma}; \quad \frac{1}{b} = \frac{2}{3a} - \frac{1}{3\gamma}$$

$$\frac{1}{c} = \frac{-1}{3a} + \frac{2}{3\gamma}; \quad \frac{1}{c} = \frac{1}{3a} - \frac{2}{3\gamma}$$

vnde eadem formulae pro earum constructione nascuntur quae supra (136) sunt allatae nisi quod iam in denominatoribus membra  $\pm \tau(a + \gamma)V(\lambda - 1)$  adiungi debeant.

Scholi-

## Scholion.

144. Ternarum ergo lentium simplicium idonea coniunctione effici potest, vt in expressione spatii diffusionis  $Hb = \mu \gamma \gamma x x \left(\frac{1}{a} + \frac{1}{\gamma}\right) \left(\lambda^{(2)} \left(\frac{1}{a} + \frac{1}{\gamma}\right)^2 + \frac{v}{a\gamma}\right)$  numerus  $\lambda^{(3)}$  fiat  $= 0,042165$ . In lentibus autem duplicatis vidimus minimum valorem numeri  $\lambda^{(2)}$  esse  $= 0,191827$ : sicque in triplicatis hic numerus fere quinquies minor reddi potest at is fere vices quater est minor, quam per lentes simplices obtineri potest. Loquor hic autem de lentibus ex principio minimi petitis, quippe quae ad praxin maxime sunt accommodatae, dum constructio leuibus erroribus non admodum turbatur. Quanquam enim lentes triplicatae perinde ac duplicatae parari possent, pro quibus numerus conueniens  $\lambda$  non solum nihilo aequalis, sed etiam negatiuus resultaret tamen earum constructio tam est lubrica, vt minimus error totum laborem irritum reddat. Interim tamen periculum in triplicatis non tantum est quam duplicatis, vnde sequens problema soluisse operae erit pretium.

## Problema 7.

145. Pro datis distantis determinatricibus  $a$  et  $\gamma$  eas definire lentes triplicatas, pro quibus valor ipsius  $\lambda^{(3)}$  prorsus in nihilum abeat.

## Solutio.

Consideretur numerus  $\lambda^{(2)}$  ex duabus prioribus

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lentibus natus ut datus, et cum sit secundum solutionem posteriorem praecedentis problematis

$$\lambda^{(3)} = \lambda^{(1)} g^3 + \lambda'' (1 - g)^3 - \nu g (1 - g)$$

definiri debet  $g$  ita ut ista quantitas evanescat. Verum cum  $\lambda^{(1)}$  commode nequeat minor effici quam  $\frac{1-\nu}{4}$ , statuamus  $\lambda^{(1)} = \frac{1-\nu}{4}$ , fierique oportet:

$$0 = \frac{1}{4} g^3 + \lambda'' (1 - g)^3 - \nu g (1 - \frac{1}{2} g)^2$$

sed quia  $\lambda''$  unitate minor esse nequit, necesse est ut  $g$  capiatur unitate maior euoluantur ergo quidam casus.

$$\text{I. } g = \frac{5}{4}; 0 = \frac{125}{4 \cdot 64} - \frac{\lambda''}{64} - \frac{5 \cdot 0}{4 \cdot 64} \nu \text{ et } \lambda'' = \frac{125 - 45 \nu}{4}$$

$$\text{II. } g = \frac{6}{4}; 0 = \frac{216}{4 \cdot 64} - \frac{8 \lambda''}{64} - \frac{6 \cdot 4}{4 \cdot 64} \nu \text{ et } \lambda'' = \frac{216 - 24 \nu}{4 \cdot 8} = \frac{27 - 3 \nu}{4}$$

$$\text{III. } g = \frac{7}{4}; 0 = \frac{343}{4 \cdot 64} - \frac{27 \lambda''}{64} - \frac{7 \cdot 1}{4 \cdot 64} \nu \text{ et } \lambda'' = \frac{343 - 7 \nu}{4 \cdot 27}$$

$$\text{IV. } g = \frac{8}{4}; 0 = \frac{512}{4 \cdot 64} - \frac{64 \lambda''}{64} \text{ et } \lambda'' = 2$$

$$\text{V. } g = \frac{9}{4}; 0 = \frac{729}{4 \cdot 64} - \frac{125 \lambda''}{64} - \frac{9 \cdot 1}{4 \cdot 64} \nu \text{ et } \lambda'' = \frac{729 - 9 \nu}{4 \cdot 125}$$

$$\text{VI. } g = \frac{10}{4}; 0 = \frac{1000}{4 \cdot 64} - \frac{216 \lambda''}{64} - \frac{10 \cdot 4}{4 \cdot 64} \nu \text{ et } \lambda'' = \frac{1000 - 40 \nu}{4 \cdot 216} = \frac{125 - 5 \nu}{4 \cdot 27}$$

In primo et secundo casu fit valor ipsius  $\lambda''$  nimis magnus, quam ut ista lens commode in praxin recipi queat; ac si ipsi  $g$  multo maior tribuatur valor, levis error ingentem effectum producit. Ponamus enim loco  $g$  per errorem sumi  $g + \omega$ , et cum  $\lambda''$  ex  $g$  rite fuerit definitum, fiet

$$\lambda^{(3)} = \omega \left( \frac{3}{4} g g - 3 \lambda'' (1 - g)^2 - \nu (1 - \frac{1}{2} g) (1 - \frac{3}{4} g) \right)$$

et

et pro  $\lambda''$  substituto valore:

$$\lambda^{(3)} = \frac{\omega}{4(1-g)} (3gg - \nu(4 - gg))$$

vnde patet, quo minus  $g$  unitatem excedat, ac simul quo maius fuerit  $g$ , valorem ipsius  $\lambda^{(3)}$  ob errorem  $\omega$  eo fieri maiorem. Intelligitur autem hunc errorem fieri minimum si capiatur

$g = 1 + \sqrt{\frac{3-\nu}{\nu+3}}$ , ex hoc autem valore elicitur,

$$\lambda'' = \frac{(3+\nu)\sqrt{3(1-\nu)(3+\nu)} + 3(3+\nu\nu)}{5(1-\nu)}$$

Capi ergo debet  $g = 1, 84384$  vnde colligitur:

$$\lambda'' = \frac{g^3 - \nu g(2-g)^2}{4(g-1)^3} = 2, 60372$$

et si hae mensurae exacte observentur, fiet  $\lambda^{(3)} = 0$ .

At si in valore ipsius  $g$  particula  $\omega$  aberretur, ut sit  $g = 1, 84384 + \omega$ , prodibit ob hunc errorem:

$$\lambda^{(3)} = -2, 981 \omega,$$

ita si effet error  $\omega = \pm \frac{1}{10}$ , loco  $\lambda^{(3)} = 0$ , prodiret:

$\lambda^{(3)} = \pm 0, 2981$  ideoque lens triplicata postponenda duplicatae, longe tamen praeferenda foret simplici.

In genere igitur pro quouis valore alio ipsius  $\lambda^{(2)}$  idem commodum inuestigemus: ac primo cum sit

$$\lambda'' = \frac{\lambda^{(2)}g^3 + \nu g(g-1)}{(g-1)^3}$$

si loco iusti valoris  $g$  capiatur  $g + \omega$  fiet

$$\lambda^{(3)} = \omega(3\lambda^{(2)}g^2 - 3\lambda''(g-1)^2 - \nu + 2\nu g)$$

vbi si pro  $\lambda''$  valor substituat, erit

$$\lambda^{(3)} = \frac{\omega(\nu - (3\lambda^{(2)} + \nu)gg)}{g-1}$$



qui error vt minimus reddatur capi debet

$$g = 1 + \sqrt{\frac{3\lambda^{(2)}}{3\lambda^{(2)} + \nu}}$$

qui est valor maxime idoneus pro  $g$  fumendus, ex quo elicitur

$$\lambda'' = \frac{1}{3\lambda^{(2)}} (\sqrt{(3\lambda^{(2)} + \nu)} + \sqrt{3\lambda^{(2)}}) \left( \frac{1}{3} (6\lambda^{(2)} + \nu) \sqrt{3\lambda^{(2)}} + (2\lambda^{(2)} + \nu) \sqrt{(3\lambda^{(2)} + \nu)} \right)$$

et sumto per errorem  $g + \omega$  pro  $g$  erit

$$\lambda^{(3)} = -2\omega (3\lambda^{(2)} + \nu + \sqrt{3\lambda^{(2)}} (3\lambda^{(2)} + \nu)) \text{ seu}$$

$$\lambda^{(3)} = -2\omega (\sqrt{3\lambda^{(2)}} + \sqrt{(3\lambda^{(2)} + \nu)}) \sqrt{(3\lambda^{(2)} + \nu)}$$

Quo minor igitur iam fuerit valor ipsius  $\lambda^{(2)}$  eo minus erit error metuendus, vnde solutio ante ex valore  $\lambda^{(2)} = \frac{1-\nu}{4}$  eruta prae ceteris est commendanda.

Tum autem erit  $f = \frac{1}{2}$ ;  $\lambda = 1$ ,  $\lambda' = 1$ ,  $\lambda'' = 2$ , 60372 et  $g = 1$ , 84384, vnde pro lentium constructione habemus,

$$\frac{1}{a} = -\frac{1}{b} = \frac{-2+g}{2a} + \frac{g}{2\gamma}$$

$$\frac{1}{c} = -\frac{1}{d} = \frac{-1+g}{a} + \frac{g}{\gamma}$$

vnde singulae lentes per formulas §. 91 construentur.

### COROLL. I.

146. Lens ergo prima PP construi debet ex distantiiis determinatricibus  $a$  et  $\frac{2a\gamma}{ga - (2-g)\gamma}$  cum numero  $\lambda = 1$

Lens vero secunda QQ ex distantiiis determinatricibus

$$\frac{-2a\gamma}{ga - (2-g)\gamma} \text{ cum numero } \lambda' = 1$$

at

at lens tertia RR ex distantiiis determinatricibus

$\frac{-a\gamma}{ga-(1-g)\gamma}$  et  $\gamma$  cum numero  $\lambda''=2, 60372$   
existente  $g=1, 84384$ .

### C O R O L L. 2.

147. Quoniam in formula spatium diffusionis exprimēte quae ob  $\lambda^{(3)}=0$  est  $\mu\gamma\gamma xx(\frac{1}{a}+\frac{1}{\gamma})\frac{\gamma}{a\gamma}$ , reiecto factore primo  $\mu\gamma\gamma xx$ , cuius ratio in his investigationibus non est habita, distantiae  $a$  et  $\gamma$  inter se permutari possunt; hinc etiam alia lens triplicata quaesito aeque satisfaciens exhiberi poterit.

### C O R O L L. 3.

148. Nempe pro hac altera lente triplicata lens prima PP construi debet ex distantiiis determinatricibus

$a$  et  $\frac{+a\gamma}{(1-g)a-g\gamma}$  cum numero  $\lambda=2, 60372$

Lens secunda QQ ex distantiiis determinatricibus

$\frac{-a\gamma}{(1-g)a-g\gamma}$  et  $\frac{+2a\gamma}{(2-g)a-g\gamma}$  cum numero  $\lambda'=1$

at lens tertia ex distantiiis determinatricibus

$\frac{-2a\gamma}{(2-g)a-g\gamma}$  et  $\gamma$  cum numero  $\lambda''=1$

existente ut ante  $g=1, 84384$ .

### C O R O L L. 4.

149. Hinc ergo duas nacti sumus lentes triplicatas pro distantiiis  $a$  et  $\gamma$ , quae producant spatium  
diffu-



diffusionis  $Hb = \mu \gamma \gamma x x (\frac{1}{a} + \frac{1}{\gamma}) \frac{v}{a \gamma}$ . Atque hae inter infinitas alias eundem effectum praestantes hac gaudent praerogatiua, vt leuis error in constructione commissus scopum minime perturbet.

### COROLL. 5.

150. Si in constructione harum lentium per errorem numerus  $g$  parumper maior accipiatur, quam 1, 84384, tum pro lente triplicata numerus  $\lambda^{(3)}$  prodit nihilo minor seu negatiuus. Sin autem numerus  $g$  in praxi aliquantillum maior sumatur numerus  $\lambda^{(3)}$  fit nihilo maior, sicque lens triplicata ad naturam duplicatarum accedet.

### COROLL. 6.

151. Si ergo opus fuerit lente, pro qua numerus  $\lambda$  valorem habeat negatiuum, huic scopo satisfieri commode poterit per lentes descriptas triplicatas, dummodo pro  $g$  numerus aliquanto maior quam 1, 84384 assumatur. Scilicet si sumatur

$$g = 1, 84384 + \omega \text{ fiet } \lambda^{(3)} = -2, 981 \omega.$$

## Supplementum II.

### De lentibus triplicatis

Si pro singulis lentibus refractione fit diuersa, pro prima  $n : 1$  pro secunda  $n' : 1$ , pro tertia  $n''$

$n^{\mu} : 1$ , radiique facierum lentium sequenti modo definiantur

|      | Dist. determinatrices | Refractio et litterae inde pendentes                         |
|------|-----------------------|--------------------------------------------------------------|
| I.   | $a$ et $\alpha$       | $n : 1, \mu, \nu, \rho, \sigma, \tau. \lambda$               |
| II.  | $b$ et $\beta$        | $n' : 1, \mu', \nu', \rho', \sigma', \tau'. \lambda'$        |
| III. | $c$ et $\gamma$       | $n'' : 1, \mu'', \nu'', \rho'', \sigma'', \tau''. \lambda''$ |

scilicet si pro lente prima vocetur radius faciei anterioris  $= F$ ; posterioris  $= G$ ; erit

$$\frac{1}{F} = \frac{\rho}{a} + \frac{\sigma}{\alpha} + \tau \left( \frac{1}{a} + \frac{1}{\alpha} \right) \sqrt{\lambda - 1}$$

$$\text{et } \frac{1}{G} = \frac{\rho}{\alpha} + \frac{\sigma}{a} \pm \tau \left( \frac{1}{a} + \frac{1}{\alpha} \right) \sqrt{\lambda - 1}$$

similique modo pro reliquis lentibus, nempe pro secunda

$$\frac{1}{F'} = \frac{\rho'}{b} + \frac{\sigma'}{\beta} + \tau' \left( \frac{1}{b} + \frac{1}{\beta} \right) \sqrt{\lambda' - 1}$$

$$\frac{1}{G'} = \frac{\rho'}{\beta} + \frac{\sigma'}{b} \pm \tau' \left( \frac{1}{b} + \frac{1}{\beta} \right) \sqrt{\lambda' - 1}$$

tum vero pro tertia

$$\frac{1}{F''} = \frac{\rho''}{c} + \frac{\sigma''}{\gamma} + \tau'' \left( \frac{1}{c} + \frac{1}{\gamma} \right) \sqrt{\lambda'' - 1}$$

$$\frac{1}{G''} = \frac{\rho''}{\gamma} + \frac{\sigma''}{c} \pm \tau'' \left( \frac{1}{c} + \frac{1}{\gamma} \right) \sqrt{\lambda'' - 1}$$

Deinde quia distantiae lentium pro nihilo habentur scilicet  $a + b = 0$  et  $\beta + c = 0$ , erit spatium diffusionis

$$= \gamma \gamma x x \left\{ \begin{array}{l} + \mu \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{\nu}{a \alpha} \right) \\ + \mu' \left( \frac{1}{b} + \frac{1}{\beta} \right) \left( \lambda' \left( \frac{1}{b} + \frac{1}{\beta} \right)^2 + \frac{\nu'}{b \beta} \right) \\ + \mu'' \left( \frac{1}{c} + \frac{1}{\gamma} \right) \left( \lambda'' \left( \frac{1}{c} + \frac{1}{\gamma} \right)^2 + \frac{\nu''}{c \gamma} \right) \end{array} \right.$$

Tom. I.

Q

statua-



statuatur nunc

$$\frac{x}{a} + \frac{x}{a} = f\left(\frac{x}{a} + \frac{x}{\gamma}\right)$$

$$\frac{x}{b} + \frac{x}{e} = g\left(\frac{x}{a} + \frac{x}{\gamma}\right)$$

$$\frac{x}{c} + \frac{x}{\gamma} = h\left(\frac{x}{a} + \frac{x}{\gamma}\right)$$

ut fiat

$$\frac{x}{a} = \frac{f-1}{a} + \frac{f}{\gamma} = \frac{1}{b}; \quad \text{et}$$

$$\frac{x}{c} = \frac{b}{a} + \frac{b-1}{\gamma} = \frac{1}{e}$$

unde fit

$$\frac{x}{b} + \frac{x}{e} = \frac{1-c}{a} + \frac{1-f}{\gamma} = g\left(\frac{x}{a} + \frac{x}{\gamma}\right)$$

hincque  $f+g+h=1$ ; quibus valoribus substitutis erit  
primo pro radiis facierum

$$\frac{1}{F} = \frac{e + (f-1)\sigma}{a} + \frac{\sigma f}{\gamma} + \tau f\left(\frac{x}{a} + \frac{x}{\gamma}\right) \sqrt{\lambda - 1}$$

$$\frac{1}{G} = \frac{\sigma + (f-1)e}{a} + \frac{ef}{\gamma} + \tau f\left(\frac{x}{a} + \frac{x}{\gamma}\right) \sqrt{\lambda - 1}$$

$$\frac{1}{F'} = \frac{(1-f)e' - b\sigma'}{a} + \frac{(1-b)\sigma' - f e'}{\gamma} + \tau' g\left(\frac{x}{a} + \frac{x}{\gamma}\right) \sqrt{\lambda' - 1}$$

$$\frac{1}{G'} = \frac{(1-f)\sigma' - b e'}{a} + \frac{(1-b)e' - f\sigma'}{\gamma} + \tau' g\left(\frac{x}{a} + \frac{x}{\gamma}\right) \sqrt{\lambda' - 1}$$

et pro lente tertia

$$\frac{1}{F''} = \frac{b \cdot e''}{a} + \frac{(b-1)e'' + \sigma''}{\gamma} + \tau'' \cdot h\left(\frac{x}{a} + \frac{x}{\gamma}\right) \sqrt{\lambda'' - 1}$$

$$\frac{1}{G''} = \frac{b \cdot \sigma''}{a} + \frac{(b-1)\sigma'' + e''}{\gamma} + \tau'' \cdot h\left(\frac{x}{a} + \frac{x}{\gamma}\right) \sqrt{\lambda'' - 1}$$

Spatium

Spatium vero diffusionis iam ita exprimetur

$$\gamma\gamma xx \left(\frac{1}{a} + \frac{1}{\gamma}\right) \left\{ \begin{array}{l} \left(\frac{1}{a} + \frac{1}{\gamma}\right)^2 (\mu\lambda f^3 + \mu'\lambda'g^3 + \mu''\lambda''b^3) \\ + \frac{\nu\mu f}{a} \left(\frac{f-1}{a} + \frac{f}{\gamma}\right) \\ + \nu'\mu'g \left(\frac{1-f}{a} - \frac{f}{\gamma}\right) \left(\frac{1-b}{\gamma} - \frac{b}{a}\right) \\ + \frac{\nu''\mu''b}{\gamma} \left(\frac{b}{a} + \frac{b-1}{\gamma}\right) \end{array} \right.$$

Nunc igitur circa has lentes triplicatas sequentia sunt obseruanda

I. Diuersa media refringentia cum tantum in finem adhiberi solent, vt non solum spatium diffusionis hic determinatum ad nihilum redigatur, sed etiam confusio a diuersa radiorum refrangibilitate oriunda tollatur, quippe quod per lentes eiusdem refractionis obtineri nequit. Infra autem videbimus ad hanc conditionem implendam requiri, vt sit  $\zeta f + \eta g + \vartheta b = 0$ , existente  $\zeta = \frac{dn}{n-1}$ ;  $\eta = \frac{dn'}{n'-1}$  et  $\vartheta = \frac{dn''}{n''-1}$ ; ex quibus formis iam perspicitur, si haec differentialia  $dn$ ,  $dn'$  et  $dn''$  essent ipsis  $n-1$ ;  $n'-1$ ;  $n''-1$  proportionalia, vti *Newtonus* statuerat; tum proditura esse  $\zeta = \eta = \vartheta$  siue  $f + g + b = 0$ ; at iam vidimus, esse debere  $f + g + b = 1$ ; quare si *Newtoni* sententia esset vera; tum ne quidem diuersis refractionibus adhibendis diffusioni a diuersa refrangibilitate oriundae remedium adferri posset. Eatenus igitur tantum hoc incommodum vitari poterit, quatenus litterae  $\zeta$ ,  $\eta$  et  $\vartheta$  sunt diuersae, ita, vt simul esse possit, et  $f + g + b = 1$  et  $\zeta f + \eta g + \vartheta b = 0$  ex quo perspicuum est, quantitatum  $f$ ,  $g$  et  $b$  vnā

Q 2

vel



vel adeo duas esse debere negatiuas sicque hac conditione superaddita casus ille principalis, quo omnes tres litterae  $f$ ,  $g$  et  $b$  positiuae sunt assumtae hic locum inuenire nequit.

II. Vt ergo nostrum spatium diffusionis euanescat, satisfaciendum est huic aequationi

$$\begin{aligned} & \left(\frac{1}{a} + \frac{1}{\gamma}\right)^2 (\mu \lambda f^3 + \mu' \lambda' g^3 + \mu'' \lambda'' b^3) \\ & + \frac{\nu \mu f}{a} \left(\frac{f-1}{a} + \frac{f}{\gamma}\right) \\ & + \nu' \mu' g \left(\frac{1-f}{a} - \frac{f}{\gamma}\right) \left(\frac{1-b}{\gamma} - \frac{b}{a}\right) \\ & + \frac{\nu'' \mu'' b}{\gamma} \left(\frac{b}{a} + \frac{b-1}{\gamma}\right) = 0 \end{aligned}$$

Vnde vel  $\lambda$  vel  $\lambda'$  vel  $\lambda''$  quaeri potest, dummodo caueatur, ne valor vel unitate minor vel nimis magnus prodeat, quia prius naturae repugnat; alterum autem, quia constructio lentium fieret nimis lubrica. Hoc autem praestito circa binas reliquas litteras  $\lambda$  nihil amplius definitur neque etiam circa litteras  $f$ ,  $g$  et  $b$ , praeterquam quod supra allatis conditionibus  $f+g+b=1$  et  $\zeta f + \eta g + \vartheta b = 0$  continetur, vnde ex data vna harum litterarum duae reliquae sponte definiuntur sequenti scilicet modo:

$$\begin{aligned} g &= \frac{(\vartheta - \zeta)f - \vartheta}{\eta - \vartheta}; \text{ et} \\ b &= \frac{(\zeta - \eta)f + \eta}{\eta - \vartheta} \end{aligned}$$

III. In praxi autem huiusmodi lentes triplicatae ideo potissimum quaeruntur, vt loco lentis obiectiuae

ætiuae in telescopiis substitui queant, pro quibus est  $a = \infty$ . Statuamus ergo statim  $a = \infty$  et pro radiis facierum singularum lentium habebimus

$$\frac{1}{F} = \frac{\sigma f}{\gamma} + \frac{\tau f}{\gamma} \sqrt{\lambda - 1}; \quad \frac{1}{G} = \frac{\rho f}{\gamma} + \frac{\tau f}{\gamma} \sqrt{\lambda - 1}$$

pro secunda lente

$$\frac{1}{F'} = \frac{(1-b)\sigma' - f\rho'}{\gamma} + \frac{\tau'g}{\gamma} \sqrt{\lambda' - 1}; \quad \frac{1}{G'} = \frac{(1-b)\rho' - f\sigma'}{\gamma} + \frac{\tau'g}{\gamma} \sqrt{\lambda' - 1}$$

pro tertia lente

$$\frac{1}{F''} = \frac{(b-1)\rho'' + \sigma''}{\gamma} + \frac{\tau''b}{\gamma} \sqrt{\lambda'' - 1}; \quad \frac{1}{G''} = \frac{(b-1)\sigma'' + \rho''}{\gamma} + \frac{\tau''b}{\gamma} \sqrt{\lambda'' - 1}$$

Spatium autem diffusionis tum ita exprimetur

$$\frac{x^2}{\gamma} (\mu \lambda f^3 + \mu' \lambda' g^3 + \mu'' \lambda'' b^3 - \nu' \mu' f g (1-b) - \nu'' \mu'' b (1-b))$$

Ita ut satisfieri oporteat huic aequationi

$$\mu \lambda f^3 + \mu' \lambda' g^3 + \mu'' \lambda'' b^3 - \nu' \mu' f g (1-b) - \nu'' \mu'' b (1-b) = 0.$$

Vnde vnus valorum  $\lambda$ ,  $\lambda'$ ,  $\lambda''$ , qui ad vsum commodissimus videtur, determinari debet.

IV. Si tantum lentibus vitreis vti velimus, sufficiet duas tantum vitri species adhiberi; si igitur statuamus, lentem tertiam et primam ex eadem vitri specie parari, ut  $\mu'' = \mu$ ;  $\nu'' = \nu$ ;  $\rho'' = \rho$ ;  $\sigma'' = \sigma$ ;  $\tau'' = \tau$  et  $\zeta = \vartheta$  ob  $n'' = n$ ; pro litteris autem  $f$ ,  $g$  et  $b$  hae determinationes habebuntur, 1<sup>o</sup>. ex aequatione

$$\zeta f + \eta g + \zeta b = 0 \text{ fit } f + b = -\frac{\eta}{\zeta} g;$$

quo valore substituto fiet

$$g \frac{(\zeta - \eta)}{\zeta} = 1; \quad g = \frac{\zeta}{\zeta - \eta} \text{ ideoque } f + b = -\frac{\eta}{\zeta - \eta};$$

Q 3

vnde



unde patet, prouti littera  $g$  fuerit vel positiua vel negatiua, fore vicissim summam  $f+b$  vel negatiuam vel positiuam et aequatio resoluenda iam erit

$$\mu(\lambda f^3 + \lambda'' b^3) + \mu' \lambda' g^3 - \frac{\mu' v' \zeta}{\zeta - \eta} \cdot f(1-b) - \mu v \cdot b(1-b) = 0$$

Hinc igitur elicitur

$$\mu' \lambda' g^3 = -\mu(\lambda f^3 + \lambda'' b^3) + \frac{\mu' v' \zeta}{\zeta - \eta} \cdot f(1-b) + \mu v \cdot b(1-b)$$

cui facile erit pro casu quouis proposito satisfacere.

IV. Dum lens prima et tertia ex eadem vitri specie parantur, media constat ex aqua vel alia materia fluida, ut lens triplicata intra duas lentes vitreas contineat fluidum, et quia fluidum plerumque minorem refractionem patitur, quam vitrum; erit  $n' < n$  indeque porro  $\eta < \zeta$ . Quia igitur pro hoc casu fit  $g$  positium seu lens aquea conuexa, lentes vitreae vel ambae vel vna saltem debent esse concauae.

Inprimis autem praeter determinationes iam inventas necesse est, ut radius faciei anterioris pro lente media aequalis et contrarius sit radio faciei posterioris lentis primae, eodemque modo radius faciei posterioris aequalis et contrarius radio faciei anterioris lentis tertiae unde hae aequalitates nascentur

$$\frac{1}{F} = -\frac{1}{G}, \text{ seu } F' + G = 0; \text{ et } \frac{1}{G'} = -\frac{1}{F''} \text{ seu } F'' + G' = 0.$$

Ideoque satisfieri oportet istis aequationibus:

$$(1-b)\sigma' - f\varepsilon' + \tau'g \cdot \sqrt{\lambda' - 1} = -\varepsilon f + \tau f \sqrt{\lambda - 1} \text{ et}$$

$$(1-b)\varepsilon' - f\sigma' + \tau'g \cdot \sqrt{\lambda' - 1} = (1-b)\varepsilon - \sigma + \tau''b \sqrt{\lambda'' - 1}$$

En

En igitur duas conditiones, quibus satisfieri oportet, unde vel numeri  $\lambda$  et  $\lambda''$  vel alter eorum cum alterutra litterarum  $f$  et  $b$  definiri debent; quem in finem probe obseruandum est, formulas  $\sqrt{\lambda - 1}$  et  $\sqrt{\lambda'' - 1}$  pro lubitu siue positiuas siue negatiuas assumi posse, neque a se inuicem pendere. Pro formula autem  $\sqrt{\lambda' - 1}$  notandum est, si ea in priore aequatione positue ponatur, in posteriore necessario negatiue sumi debere et vicissim.

### Problema 8.

152. Determinare eas lentes quadruplicatas ad Tab. II.  
 datas distantias determinatrices  $AE = a$  et  $dI = \delta$  ac Fig. 9.  
 commodatas, quae minimum spatium diffusionis  $Ii$   
 producant.

### Solutio.

Prima lens PP ad distantias determinatrices  $AE = a$  et  $AF = \alpha$  cum numero  $\lambda$  construatur, secunda QQ ad distantias  $b = -\alpha$  et  $AG = \beta$  cum numero  $\lambda'$ , tertia RR ad distantias  $c = -\beta$  et  $AH = \gamma$  cum numero  $\lambda''$ ; et quarta SS ad distantias  $d = -\gamma$  et  $dI = \delta$  cum numero  $\lambda'''$  construatur: vbi scilicet crassitiem lentium vt euanescentem spectamus. Posito iam semidiametro aperturae lentis  $x$ , sola prima lens PP produceret spatium diffusionis

$$Ef = \mu \alpha x x \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{\nu}{a\alpha} \right)$$

Adiun-



Adiuncta autem secunda lente QQ, positoque

$$\frac{1}{a} = \frac{-1+f}{a} + \frac{f}{e} \quad \text{et} \quad \frac{1}{b} = \frac{1-f}{a} - \frac{f}{e}$$

si breuitatis gratia statuamus

$$\lambda^{(2)} = \lambda f^3 + \lambda' (1-f)^3 - \nu f (1-f)$$

vidimus fore spatium diffusionis :

$$Gg = \mu \epsilon \epsilon x x \left( \frac{1}{a} + \frac{1}{e} \right) \left( \lambda^{(2)} \left( \frac{1}{a} + \frac{1}{e} \right)^2 + \frac{\nu}{a e} \right)$$

Adiungatur insuper tertia lens RR, numerusque  $g$  ita sumatur vt fit

$$\frac{1}{e} = \frac{-1+g}{a} + \frac{g}{\gamma} \quad \text{et} \quad \frac{1}{c} = \frac{1-g}{a} - \frac{g}{\gamma}$$

ac si breuitatis ergo ponamus :

$$\lambda^{(3)} = \lambda^{(2)} g^3 + \lambda'' (1-g)^3 - \nu g (1-g)$$

erit spatium diffusionis :

$$Hb = \mu \gamma \gamma x x \left( \frac{1}{a} + \frac{1}{\gamma} \right) \left( \lambda^{(3)} \left( \frac{1}{a} + \frac{1}{\gamma} \right)^2 + \frac{\nu}{a \gamma} \right)$$

Nunc denique adiungatur lens quarta SS, et numero  $b$  ita in calculum introducto, vt fit

$$\frac{1}{\gamma} = \frac{-1+b}{a} + \frac{b}{\delta} \quad \text{et} \quad \frac{1}{d} = \frac{1-b}{a} - \frac{b}{\delta}$$

si simili modo ponamus

$$\lambda^{(4)} = \lambda^{(3)} b^3 + \lambda''' (1-b)^3 - \nu b (1-b)$$

erit spatium diffusionis a lente quadruplicata productum :

$$Ii = \mu \delta \delta x x \left( \frac{1}{a} + \frac{1}{\delta} \right) \left( \lambda^{(4)} \left( \frac{1}{a} + \frac{1}{\delta} \right)^2 + \frac{\nu}{a \delta} \right)$$

quod igitur minimum reddi debet. Hunc in finem considerentur numeri  $\lambda$ ,  $\lambda'$ ,  $\lambda''$  et  $\lambda'''$  vt dati, et quaerantur

rantur idonei valores pro numeris  $f$ ,  $g$ , et  $h$ : atque ut valor  $\lambda^{(1)}$  minimus euadat, necesse est quoque valores  $\lambda^{(3)}$  et  $\lambda^{(2)}$  minimos fieri. Incipiamus ergo a valore  $\lambda^{(2)}$  qui minimus redditur sumendo

$$f = \frac{V(3\lambda' + \nu)}{V(3\lambda + \nu) + V(3\lambda' + \nu)} \text{ seu } \frac{1}{f} = 1 + V \frac{3\lambda + \nu}{3\lambda' + \nu}$$

vnde fit

$$\frac{1}{V(3\lambda^{(2)} + \nu)} = \frac{1}{V(3\lambda + \nu)} + \frac{1}{V(3\lambda' + \nu)}.$$

Deinde numerus  $\lambda^{(3)}$  minimum induet valorem capi-  
endo

$$g = \frac{V(3\lambda'' + \nu)}{V(3\lambda^{(2)} + \nu) + V(3\lambda'' + \nu)} \text{ seu } \frac{1}{g} = 1 + V \frac{3\lambda^{(2)} + \nu}{3\lambda'' + \nu}$$

hincque colligitur

$$\frac{1}{V(3\lambda^{(3)} + \nu)} = \frac{1}{V(3\lambda^{(2)} + \nu)} + \frac{1}{V(3\lambda'' + \nu)}$$

Denique numerus  $\lambda^{(4)}$  ideoque et spatium diffusionis  $Ii$  minimum efficietur sumendo

$$h = \frac{V(3\lambda''' + \nu)}{V(3\lambda^{(3)} + \nu) + V(3\lambda''' + \nu)} \text{ seu } \frac{1}{h} = 1 + V \frac{3\lambda^{(3)} + \nu}{3\lambda''' + \nu}$$

vnde obtinetur

$$\frac{1}{V(3\lambda^{(4)} + \nu)} = \frac{1}{V(3\lambda^{(3)} + \nu)} + \frac{1}{V(3\lambda''' + \nu)}.$$



Quod si hic valores ante inuentos substituamus, nanciscemur

$$\frac{1}{V(3\lambda^{(2)}+\nu)} = \frac{1}{V(3\lambda+\nu)} + \frac{1}{V(3\lambda'+\nu)} + \frac{1}{V(3\lambda''+\nu)} + \frac{1}{V(3\lambda'''+\nu)}$$

Pro praecedentibus vero erit

$$\frac{1}{V(3\lambda^{(2)}+\nu)} = \frac{1}{V(3\lambda+\nu)} + \frac{1}{V(3\lambda'+\nu)} + \frac{1}{V(3\lambda''+\nu)}$$

et

$$\frac{1}{V(3\lambda^{(2)}+\nu)} = \frac{1}{V(3\lambda+\nu)} + \frac{1}{V(3\lambda'+\nu)}$$

Tum vero ex his porro consequimur:

$$f = V \frac{3\lambda^{(2)}+\nu}{3\lambda+\nu}, g = V \frac{3\lambda^{(3)}+\nu}{3\lambda^{(2)}+\nu}, h = V \frac{3\lambda^{(4)}+\nu}{3\lambda^{(3)}+\nu}$$

Supereft ergo vt constructionem singularum lentium luculentius exponamus: et earum distantias determinatrices per solas propositas  $a$  et  $\delta$  exprimamus: erit igitur

$$\begin{aligned} \frac{1}{\gamma} &= \frac{1+b}{a} + \frac{h}{\delta}; & \frac{1}{d} &= \frac{1-b}{a} - \frac{h}{\delta} \\ \frac{1}{\delta} &= \frac{1+gb}{a} + \frac{gb}{\delta}; & \frac{1}{c} &= \frac{1-gb}{a} - \frac{gb}{\delta} \\ \frac{1}{\alpha} &= \frac{1+fgh}{a} + \frac{fgh}{\delta}; & \frac{1}{b} &= \frac{1-fgh}{a} - \frac{fgh}{\delta} \end{aligned}$$

Ex superioribus vero formulis colligitur:

$$h = V \frac{3\lambda^{(4)}+\nu}{3\lambda^{(3)}+\nu}; g = V \frac{3\lambda^{(3)}+\nu}{3\lambda^{(2)}+\nu}; f = V \frac{3\lambda^{(2)}+\nu}{3\lambda+\nu}$$

vnde

vnde fit

$$g b = V \frac{3\lambda^{(4)} + \nu}{3\lambda^{(2)} + \nu} \text{ et } f g b = V \frac{3\lambda^{(4)} + \nu}{3\lambda + \nu}.$$

et superiores valores ita exprimi poterunt

$$\begin{aligned} -\frac{1}{b} &= -\frac{V(3\lambda^{(4)} + \nu)}{a} \left( \frac{1}{V(3\lambda' + \nu)} + \frac{1}{V(3\lambda'' + \nu)} + \frac{1}{V(3\lambda''' + \nu)} \right) + \frac{V(3\lambda^{(4)} + \nu)}{\delta} \cdot \frac{1}{V(3\lambda + \nu)} \\ -\frac{1}{c} &= -\frac{V(3\lambda^{(4)} + \nu)}{a} \left( \frac{1}{V(3\lambda'' + \nu)} + \frac{1}{V(3\lambda''' + \nu)} \right) + \frac{V(3\lambda^{(4)} + \nu)}{\delta} \left( \frac{1}{V(3\lambda + \nu)} + \frac{1}{V(3\lambda' + \nu)} \right) \\ -\frac{1}{d} &= -\frac{V(3\lambda^{(4)} + \nu)}{a} \cdot \frac{1}{V(3\lambda''' + \nu)} + \frac{V(3\lambda^{(4)} + \nu)}{\delta} \left( \frac{1}{V(3\lambda + \nu)} + \frac{1}{V(3\lambda' + \nu)} + \frac{1}{V(3\lambda'' + \nu)} \right). \end{aligned}$$

### Coroll. I.

153. Si pro lentibus simplicibus numeri  $\lambda, \lambda', \lambda''$  et  $\lambda'''$  sumantur inter se aequales, fiet pro minimo spatio diffusionis:

$$V(3\lambda^{(2)} + \nu) = \frac{1}{2} V(3\lambda + \nu); \quad \lambda^{(2)} = \frac{3\lambda - 1.3\nu}{3.4}$$

$$V(3\lambda^{(3)} + \nu) = \frac{1}{3} V(3\lambda + \nu); \quad \lambda^{(3)} = \frac{3\lambda - 2.4\nu}{3.9}$$

$$V(3\lambda^{(4)} + \nu) = \frac{1}{4} V(3\lambda + \nu); \quad \lambda^{(4)} = \frac{3\lambda - 3.5\nu}{3.16}$$

hinc  $f = \frac{1}{2}$ ;  $g = \frac{2}{3}$ ;  $b = \frac{3}{4}$ , et pro constructione lentium simplicium:

$$\frac{1}{a} = -\frac{3}{4a} + \frac{1}{4\delta}; \quad \frac{1}{b} = +\frac{3}{4a} - \frac{1}{4\delta}$$

$$\frac{1}{c} = -\frac{2}{4a} + \frac{2}{4\delta}; \quad \frac{1}{d} = +\frac{2}{4a} - \frac{2}{4\delta}$$

$$\frac{1}{e} = -\frac{1}{4a} + \frac{3}{4\delta}; \quad \frac{1}{f} = +\frac{1}{4a} - \frac{3}{4\delta}$$



## Coroll. 2.

154. Hinc ex §. 91 sequens quatuor lentium simplicium constructio obtinetur:

Pro lente radius faciei

$$\begin{aligned}
 \text{Prima PP} \quad & \left\{ \begin{aligned} \text{anterioris} &= \frac{4a\delta}{(4\rho-3\sigma)\delta + \sigma a \pm \tau(a+\delta)\sqrt{(\lambda-1)}} \\ \text{posterioris} &= \frac{4a\delta}{(4\sigma-3\rho)\delta + \rho a \pm \tau(a+\delta)\sqrt{(\lambda-1)}} \end{aligned} \right. \\
 \text{Secunda QQ} \quad & \left\{ \begin{aligned} \text{anterioris} &= \frac{4a\delta}{(3\rho-2\sigma)\delta + (2\sigma-\rho)a \pm \tau(a+\delta)\sqrt{(\lambda-1)}} \\ \text{posterioris} &= \frac{4a\delta}{(3\sigma-2\rho)\delta + (\rho-\sigma)a \pm \tau(a+\delta)\sqrt{(\lambda-1)}} \end{aligned} \right. \\
 \text{Tertia RR} \quad & \left\{ \begin{aligned} \text{anterioris} &= \frac{4a\delta}{(2\rho-\sigma)\delta + (3\sigma-2\rho)a \pm \tau(a+\delta)\sqrt{(\lambda-1)}} \\ \text{posterioris} &= \frac{4a\delta}{(2\sigma-\rho)\delta + (3\rho-2\sigma)a \pm \tau(a+\delta)\sqrt{(\lambda-1)}} \end{aligned} \right. \\
 \text{Quarta SS} \quad & \left\{ \begin{aligned} \text{anterioris} &= \frac{4a\delta}{\rho\delta + (4\sigma-3\rho)a \pm \tau(a+\delta)\sqrt{(\lambda-1)}} \\ \text{posterioris} &= \frac{4a\delta}{\sigma\delta + (4\rho-3\sigma)a \pm \tau(a+\delta)\sqrt{(\lambda-1)}} \end{aligned} \right.
 \end{aligned}$$

## Coroll. 3.

155. Si praeterea numeri  $\lambda$ ,  $\lambda'$ ,  $\lambda''$ ,  $\lambda'''$  unitati aequales statuantur, qui est valor minimus, quem recipere possunt erit: ob  $v=0$ , 232692

$$\lambda^{(2)} = \frac{3-3v}{3.4} = 0,191827$$

$$\lambda^{(3)} = \frac{3-8v}{3.9} = 0,042165$$

$$\lambda^{(1)} = \frac{3-15v}{3.16} = -0,010216$$

ficque pro lente quadruplicata valor numeri  $\lambda^{(4)}$  adeo infra nihilum deprimitur.

Coroll.

## Coroll. 4.

156. Maiorem ergo valorem numeris  $\lambda, \lambda', \lambda'', \lambda'''$  tribuendo effici poterit, vt valor ipsius  $\lambda^{(t)}$  praecise nihilo aequalis prodeat; quippe hoc fiet fumendo  $\lambda = 5 \nu = 1, 163460$ . Hinc habebitur  $\lambda - 1 = 0, 163460$  et  $\tau V(\lambda - 1) = 0, 365947$ , vnde singulae lentes simplices duplici modo per formulas exhibitas construi poterunt.

## Scholion I.

157. Si inter se comparemus hos duos casus, quibus est vel  $\lambda^{(t)} = -0, 010216$  vel  $\lambda^{(t)} = 0$ , videmus, etiamsi discrimen vix partem centesimam vnitatis superet, in constructione tamen lentium simplicium satis magnum discrimen deprehendi, cum denominatores formularum, (154) siue augeri siue diminui debeant quantitate:  $0, 365947 (a + \delta)$ ; quae differentia maior est, quam errores, qui forte ab artifice non nimis rudi committi queant. Ex quo vicissim colligimus, etiamsi in constructione harum lentium quadruplicatarum ab artifice leues errores committantur, inde vix perceptibilem effectum in spatio diffusionis vel valore ipsius  $\lambda^{(t)}$  esse metuendum, quam ob causam hae lentes imprimis ad praxin accommodatae videntur. Si scilicet opus fuerit lente, pro qua valor ipsius  $\lambda$  in nihilum abeat, multo magis his lentibus quadruplicatis Coroll. 4. descriptis erit vtendum, quam triplicatis, quas supra definiui. Quin etiam eas lentes quadruplicatas adhibere licebit,



in quibus est  $\lambda^{(4)} = -0,010216$ , quia hic numerus vix nihilo est minor; ac si in constructione a praescriptis mensuris aberretur, ille valor adhuc propius ad nihilum perducatur, ita ut hoc casu adeo errores commissi scopo magis attingendo inferuiant. Infra autem videbimus, plurimos dari casus, quibus eiusmodi lentibus uti conueniat, pro quibus valor ipsius  $\lambda$  non solum sit nihilo aequalis, sed etiam tantillum infra nihilum deprimatur; tunc optimo cum successu huiusmodi lentes quadruplicatas in usum vocabimus. Sin autem eiusmodi lentes sufficiant, pro quibus valor ipsius  $\lambda$  sit  $0,042165$  vel aliquantillum excedat, lentes triplicatae erunt commendandae, quarum constructionem supra (§. 136) dedimus; quemadmodum duplicatae negotium conficient si non opus fuerit minore valore ipsius  $\lambda$  quam  $0,191827$ . Tota autem instrumentorum dioptricum perfectio in hoc maxime consistit, ut lentes habeantur, pro quibus valor ipsius  $\lambda$  sit quam minimus, cum eae sint aptissimae ad confusionem penitus tollendam; ex quo lentium quadruplicatarum hic descriptarum usus erit amplissimus.

### Scholion 2.

158. Si haec, quae de lentibus quadruplicatis hic tradidimus, attente considerentur, facile patebit, quomodo lentes quintuplicatae magisque multiplicatae ad usum sint accommodandae. Eiusmodi scilicet semper constructione erit opus, quae a natura minimi pa-  
rumper

rumper recedat, quoniam hoc modo errores in praxi commissi effectum propositum minime perturbant. Cum autem vix unquam vsu veniat, ut lentibus opus sit, pro quibus valor numeri  $\lambda$  magis ultra nihilum imminuat, superfluum foret, constructionem lentium quintuplicatarum, magisque multiplicatarum ulterius prosequi. Interim tamen iuvabit, si pro huiusmodi lentibus numeri  $\lambda$  lentium numero convenienter signis  $\lambda^{(5)}$ ,  $\lambda^{(6)}$  etc. indicentur, eorum valores, quos ex natura minimi recipiunt, exposuisse. Sumamus omnes numeros  $\lambda$ ,  $\lambda'$ ,  $\lambda''$ ,  $\lambda'''$ ,  $\lambda^{(1)}$ ,  $\lambda^{(2)}$  etc. lentibus simplicibus respondentes unitati aequales, et eorum qui lentibus multiplicatis conveniunt ex natura minimi ita se habebunt:

Pro Lente

$$\text{Solitaria} \dots \lambda^{(1)} = \frac{3 - 0v}{3.1} = 1,000000$$

$$\text{Duplicata} \dots \lambda^{(2)} = \frac{3 - 3v}{3.4} = 0,191827$$

$$\text{Triplicata} \dots \lambda^{(3)} = \frac{3 - 8v}{3.9} = 0,042165$$

$$\text{Quadruplicata} \quad \lambda^{(4)} = \frac{3 - 15v}{3.16} = -0,010216$$

$$\text{Quintuplicata} \quad \lambda^{(5)} = \frac{3 - 24v}{3.25} = -0,034461$$

$$\text{Sextuplicata} \quad \lambda^{(6)} = \frac{3 - 35v}{3.36} = -0,047632$$

$$\text{Septuplicata} \quad \lambda^{(7)} = \frac{3 - 48v}{3.49} = -0,055573$$

$$\text{Octuplicata} \quad \lambda^{(8)} = \frac{3 - 63v}{3.64} = -0,060727$$

$$\text{Noncuplicata} \quad \lambda^{(9)} = \frac{3 - 80v}{3.81} = -0,064261$$

$$\text{Decuplicata} \quad \lambda^{(10)} = \frac{3 - 99v}{3.100} = -0,066788$$



ac si huiusmodi lentes in infinitum multiplicentur, valor numeri respondentis  $\lambda^{(\infty)}$  erit  $= -\frac{v}{3} = -0,077564$  ita vt nunquam infra hunc numerum deprimi possit: ex quo patet vix vnquam casum existere posse, quo lente saltem quintuplicata opus esset. Ceterum etiam in genere, quicunque alii valores praeter vnitatem, numeris  $\lambda$ ,  $\lambda'$ ,  $\lambda''$ ,  $\lambda'''$  etc. tribuantur, ex formulis superioribus numeri  $\lambda^{(5)}$ ,  $\lambda^{(6)}$  etc. facile deriuabuntur; quin etiam distantiae determinatrices singularum lentium simplicium indidem sine difficultate definientur, cum lex progressionis satis sit manifesta. Verum per totum hoc caput probe tenendum est, vbique crassitiam lentium tanquam euanescentem esse consideratam, in sequente autem capite iterum lentes in genere non neglecta crassitie sumus contemplaturi.



C A P V T IV.  
DE  
CONFVSIONE VISIONIS  
NEC NON DE MAGNITVDINE  
APPARENTE ET CLARITATE.

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Definitio I.

159.

Visio est distincta, si omnes radii, qui ex quolibet obiecti puncto in oculum ingrediuntur in fundo oculi super retina iterum in vnum punctum congregantur.

Scholion.

160. Ad visionem distinctam requiritur, vt obiecta in certa quadam ab oculo distantia reperiantur, quæ distantia pro varia oculorum indole maxime solet esse diuersa, dum Myopes exiguam, ii qui oculis valent ingentem, ac presbytes non solum infinitam sed quandoque etiam negatiuam exigunt; cuiusmodi distantia cum in veris obiectis locum habere nequeat, ope perspicillorum sibi satisfacere solent. Quilibet ergo oculus ad certam quandam distantiam obiectorum est instructus quam eius distantiam iustam appellabo; vbi quidem insignis latitudo locum habet,

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propterea quod structura oculi ita est artificiosa, vt contractione ac elongatione quadam se ad distantias aliquanto maiores et minores accommodare possit. Quando ergo obiecta in distantia ab oculo iusta reperiuntur, visio est distincta, dum singula obiectorum puncta super retina singulis punctis exprimuntur.

### Definitio 2.

161. Visio est confusa, si radii ex quolibet obiecti puncto in oculum immissi non in vno retinae puncto congregantur, sed per aliquod spatium retinam afficiunt.

### Corollarium 1.

162. Eo maior ergo erit confusio, quo maius fuerit hoc spatium in retina, per quod radii ex eodem obiecti puncto emissi dissipantur. Ex quo huius spatii magnitudo veram confusiois mensuram suppeditabit.

### Coroll. 2.

163. Visio ergo erit confusa, cum obiecti visi distantia multum fuerit diuersa ab oculi distantia iusta. Paruum enim discrimen vel per se nullam confusioem parit vel oculus se ad obiecti distantiam, sua qua pollet, volubilitate accommodare valet.

### Scholion.

164. Si obiecta per lentes distincte repraesentarentur, visio imaginum eadem lege teneretur atque ipso-

ipforum obiectorum; iusta scilicet earum ab oculo distantia visionem distinctam, admodum autem diuersa confusam produceret. Verum si imago per aliquod spatium fuerit diffusa etiamsi ab oculo ad distantiam iustam sit remota, inde tamen in visione confusio oriatur necesse est. Quod si nempe non obiecta ipsa, sed earum imagines per vnā pluresue lentes repraesentatas intueamur ob duplicem causam visio poterit esse confusa; altera, si distantia imaginis ab oculo multum fuerit diuersa a distantia iusta, altera vero si ipsa imago per aliquod spatium diffundatur. Priorem quidem causam tollere in nostra est potestate siquidem lentes ita disponere licet, vt imaginem in iusta ab oculo distantia exhibeant; quam lentium dispositionem propterea hic perpetuo assumamus. Quamobrem in hoc capite inuestigare constitui quanta confusio in visione imaginum a spatio diffusionis oriri debeat, eiusque quantitate determinata deinceps in hoc erit elaborandum quemadmodum lentes formatas ac dispositas esse oporteat, vt confusio inde in visione nata datum limitem, quo adhuc est tolerabilis, non excedat. In genere quidem certum est, quo maius fuerit spatium diffusionis, eo maiorem inde in visionem induci debere confusionem, verum tamen mox videbimus, confusionem visionis non esse spatio diffusionis proportionalem, sed aliam omnino legem sequi, quam accurate determinasse maximi erit momenti, cum ex hoc fonte constructio omnium instrumentorum Dioptricum ad visionem accommodatorum sit repetenda.



## Problema I.

Tab. II.  
Fig. 10.

165. Si oculus aspiciat imaginem cuiuspiam obiecti ab vna pluribusue lentibus per spatium  $Ll$  diffusam, atque imago principalis  $L$  in iusta ab oculo distantia reperiatur, definire confusionem, qua visio huius imaginis afficietur.

## Solutio.

In spatio diffusionis  $Ll$ , in quo punctum  $L$  a radiis per lentium medium, punctum  $l$  vero a radiis per lentium extremitates transmissis exhibeatur, considerari oportet primo eius magnitudinem  $Ll$  deinde radiorum in  $l$  concurrentium inclinationem ad axem. Supra autem vidimus has res a primae lentis apertura, cuius semidiameter sit  $= x$ , ita pendere ut sit ipsum spatium  $Ll = V x x$ , et angulus  $O/T = \mathfrak{B} x$  pro quocumque scilicet lentium numero valores horum coefficientium  $V$  et  $\mathfrak{B}$  determinauimus. Repraesentet iam circulus  $OV$  oculum, quem hic ut exiguam cameram obscuram spectare licet, sed ita perfectam, ut radios ex vno puncto emissos iterum in vno puncto colligat, etsi enim in oculo plures fiunt refractiones, tamen conditione illa seruata vnica lens earum loco considerari potest, quae sit in  $TO T$ , a qua retina remota sit intervallo  $OV = u$ . Cum nunc imago principalis  $L$  in debita ab oculo distantia, quae vocetur  $OL = l$  existat, puncti  $L$  imago in oculo in ipsam retinam incidet, et in  $V$  distin-

cte

ete depingetur: puncti vero  $l$  imago non in  $V$ , sed ante retinam in  $v$  referetur, quod interuallum  $Vv$ , si lentis in  $TOT$  conceptae crassities euanescat, secundum §. 62 ita exprimitur, vt sit  $Vv = \frac{OV^2}{OL^2}$ .  $Ll = \frac{uu}{ll} Vxx$ . Cum autem radii punctum  $l$  formantes ad axem inclinati sint angulo  $OT = \mathfrak{B}x$ , ii per lentis oculi puncta  $TT$  ita intrabunt, vt neglecto interuallo  $Ll$  prae distantia  $OL = l$ , sit  $OT = l \mathfrak{B}x$ , ex quo ii in puncto  $v$  concurrentes cum axe angulum facient  $OvT = \frac{OT}{Ov} = \frac{OT}{OV} = \frac{l}{u} \mathfrak{B}x$ . Hinc ergo ultra ad retinam pergentes super ea circulum  $UU$  effingent cuius radius erit  $VU = \frac{l}{u} \mathfrak{B}x$ .  $Vv = \frac{l}{u} \mathfrak{B}x \cdot \frac{uu}{ll} Vxx$  et a punctis inter  $L$  et  $l$  mediis hoc spatium circulare super retina replebitur. Quare imago per spatium  $Ll$  diffusa super retina circello repraesentabitur, cuius radius  $VU = \frac{u}{l} \mathfrak{B} Vx^3$  qui circellus veram confusionis menuram supeditat. Quodlibet scilicet obiecti punctum, quod lentibus per spatium  $Ll$  diffusum exhibetur, in oculo super retina non puncto sed circello exprimetur cuius radius erit  $= \frac{u}{l} \mathfrak{B} Vx^3$ .

## Coroll. I.

166. Videmus ergo confusionem, qua visio afficitur non solum a quantitate spatii diffusionis  $Ll = Vxx$ , sed insuper ab angulo, quo radii in  $l$  concurrentes ad axem inclinantur, qui est  $= \mathfrak{B}x$  pendere; radiumque circelli confusionem metientis producto illius spatii per hunc angulum esse proportionalem.



## Coroll. 2.

167. Cum igitur posito semidiametro aperture primae lentis  $=x$ , spatium diffusionis sit ut eius quadratum  $xx$ ; radius circelli confusionem metientis est ut eius cubus  $x^3$ . Ac si confusio ipsa areae huius circelli proportionalis aestimetur erit ea ut  $x^6$ , seu ut cubus aperture primae lentis.

## Coroll. 3.

168 Hinc ergo intelligitur quanti intersit per idoneam lentium dispositionem spatium diffusionis  $Vxx$  diminuisse non solum enim in eadem ratione qua spatium diffusionis  $Vxx$  seu quantitas  $V$  diminuitur sed adeo in ratione duplicata, ipsa confusio visa minor redditur.

## Scholion.

169. Assumsi hic pupillam tam late patere, ut radios ab  $l$  diuergentes recipiat; at si apertura pupillae minor esset, radios a puncto  $l$  venientes nequidem caperet: hoc ergo casu res eodem rediret ac si contracta primae lentis apertura spatium diffusionis  $Ll$  eo usque diminueretur quoad pupilla omnes radios ab imagine emissos recipere posset hocque casu manifestum est, confusionem minorem esse prodituram. Verum in sequentibus ostendetur, tam in Telescopiis quam Microscopiis hunc casum vix unquam locum inuenire cum plerumque conus radiosus ex puncto  $l$  emissus circa ingressum in oculum mul-

to

to tenuior sit quam pupillae apertura; quamobrem ne opus quidem est hunc casum etsi per se facile expediretur, hic expendere. Ceterum tametsi oculus non commode cum lente simplici cuius crassities evanescat, conferri possit hincque expressio spatii  $Vv$  secundum §. 84. aliquantum diuersa prodire potuisset hic ad istam circumstantiam non attendimus propterea quod spatium  $Vv$  secundum certam quandam rationem auctum vel diminutum prodiiisset; Hic autem non adeo necesse est ipsam quantitatem absolutam confusionis cognouisse, dummodo rationem quam sequitur accurate definiuerimus. Cum enim nostrum calculum cum experientia contulerimus, terminum cognoscemus, quem si confusio nostro more expressa superauerit, intolerabilis euadat; hincque perpetuo sufficiet confusionem simili modo expressam infra hunc terminum reduxisse. Interim tamen notasse iuvabit, per oculi conformationem confusionem adhuc multo minorem effici posse.

### Problema 2.

170. Positis iisdem, quae in praecedente problemate sunt assumpta, eam definire oculi conformationem, qua minima confusio percipiatur.

Tab. II.  
Fig. II.

### Solutio.

Oculus omnis facultate praeditus est sese aliquantillum aliter conformandi, ut etiam obiecta, quorum distantia a iusta non nimis differt, distincte videat:



videat: quod quomodocunque perficiatur, id ita fieri concipere licet, ac si retina ad pupillam seu potius eam lentem, quam oculi loco consideramus, propius admoueretur seu longius ab ea detorqueretur. Cum ergo ante retinam in  $V$  sitam sumus contemplati, hic primo sumamus retinam per ipsum punctum  $v$  transire, ac manifestum est super ea punctum  $l$  distincte expressum iri. Verum quia punctum  $L$  alios radios nisi axi proximos non emittit etiam punctum  $L$  sine vlla confusione in  $v$  repraesentabitur; ita vt tantum puncta media inter  $L$  et  $l$  confusionem sint paritura. Consideretur ergo quoduis punctum intermedium  $\lambda$ , quod esset extremum in spatio confusionis, si semidiameter aperturae primae lentis minor foret quam  $x$ , qui ergo ponatur  $=z$ : eritque  $L\lambda = Vz$ , et inclinatio radiorum in  $\lambda$  concurrentium ad axem  $=\mathfrak{B}z$ : hinc istius puncti imago intra oculum referretur in  $s$ , vt sit  $Vs = \frac{uu}{ll} Vz$ , ideoque ob  $Vv = \frac{uu}{ll} Vx$  erit interuallum  $vs = \frac{uu}{ll} V(xx - zz)$ : et radiorum in  $s$  conuergentium inclinatio ad axem  $= \frac{l}{u} \mathfrak{B}z$ ; ex quo circelli super retina in  $v$  existentis radiis puncti  $\lambda$  expressi semidiameter erit  $= \frac{u}{l} \mathfrak{B} Vz(xx - zz)$  qui euanescit, vti iam monuimus, siue sit  $z=0$ , siue  $z=x$ . Quaeratur iam ille valor ipsius  $z$ , quo ille circellus fiat maximus; id quod eueniet, si  $xx - 3zz = 0$  seu  $z = \frac{x}{\sqrt{3}}$ ; sicque circelli confusionem hoc casu mentientis semidiameter erit  $= \frac{2u}{3l\sqrt{3}} \mathfrak{B}Vx^3$ , quod est multo minus quam casu praecedente, quo retina erat in  $V$ . At

At si retina aliquantillum a  $v$  versus  $V$  remoueatur, hic circellus adhuc minor effici poterit. Ponatur enim spatium  $VS=s$ , vt fit  $vS=\frac{uu}{ll}Vxx-s$ , et  $Ss=s-\frac{uu}{ll}Vzz$ . Punctum ergo  $l$ , cuius effigies in  $u$  exhibetur, super retina  $S$  circello referetur, cuius radius est  $=\frac{l}{u}\mathfrak{B}x(\frac{uu}{ll}Vxx-s)$ , punctum vero  $\lambda$ , cuius effigies est in  $s$ , super retina  $S$  circello, cuius radius erit  $=\frac{l}{u}\mathfrak{B}z(s-\frac{uu}{ll}Vzz)$ ; nunc igitur id punctum  $\lambda$  inuestigemus, vnde iste circellus minimus euadat: quod fit si fit  $s=\frac{3uu}{ll}Vzz$ ; eritque huius circuli radius  $=\frac{2u}{l}\mathfrak{B}Vz^3$ , qui simul confusio- nem exhiberet, si modo a puncto  $l$ , non maior cir- culus oriretur: sed substituto pro  $s$  valore inuento huius circuli radius erit,  $\frac{u}{l}\mathfrak{B}Vx(xx-3zz)$ , qui ergo illi  $\frac{2u}{l}\mathfrak{B}Vz^3$  aequalis statuatur, vnde nascitur haec aequatio:

$$x^3-3xzz=2z^3, \text{ hincque } z=\frac{1}{2}x.$$

Quare circelli, quo minima confusio mensuratur, ra- dius erit  $=\frac{u}{4l}\mathfrak{B}Vx^3$ , quadruplo minor, quam si re- tina esset in  $V$ ; indeque ipsa confusio sedecies minor. Etsi autem illa aequatio cubica tres habet radices, praeter  $z=\frac{1}{2}x$  binae reliquae sunt aequales et  $z=-x$ , sicque  $zz=xx$ , vnde non minima sed quasi maxima confusio nasceretur.

## C O R O L L. I.

171. Cum igitur, vt oculus minimam confu- sionem sentiat, debeat esse  $z=\frac{1}{2}x$ , ideoque  $s=\frac{3uu}{4ll}Vxx$ ,

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patet retinam in eum locum S cogi debere, vt fit  $VS = \frac{3}{4} Vv$  et  $vS = \frac{1}{4} Vv$ .

### Coroll. 2.

172. Cum ergo oculus non solum praeditus sit facultate sese parumper immutandi sed etiam hac facultate vti soleat ad confusionem euitandam, nullum est dubium quin circellorum confusionem producentium radius sit  $= \frac{u}{+1} \mathfrak{B} V x^3$ , ideoque quadruplo minor, quam problemate praecedente inueneramus dummodo imago spectanda L l propemodum in distantia iusta reperiatur.

### Coroll. 3.

173. Haec igitur expressio  $\frac{u}{+1} \mathfrak{B} V x^3$  iustam nobis exhibet mensuram confusionis, cum exprimat radium circellorum, quibus singula obiecti puncta, quatenus id per lentes spectatur, super retina repraesentantur. Si modo imago in distantia iusta ab oculo reperiatur.

### Scholion.

174. Si hi circelli euanescerent, visio plane esset distincta hoc autem fieri nequit, nisi ipsum interuallum diffusionis  $Vxx$  euanescat: vnde patet si apertura primae lentis ad nihilum reduceretur, nullam confusionem sentiri debere. Verum visio non ita est delicata, vt prorsus nullam confusionem pati possit sed dummodo confusio certum quendam terminum

num non superet, quasi esset nulla considerari potest. Iste terminus, seu valor quem expressio  $\frac{n}{4l} 3Vx^3$  excedere non debet, ex experientia potius peti debet, quam ex Theoria isque idcirco insigni adhuc latitudine continetur: unde fit vt pro diuerso scopo modo maiore gradu confusionis contenti esse soleamus, modo autem minorem gradum exigamus: quas circumstantias cum ad praxin propius accedemus, accuratius sumus examinaturi. Ceterum notandum est pro interuallo  $u$ , quo quasi profunditas oculi exhibetur, vnum circiter pollicem assumi posse, quae quidem mensura fere erit arbitraria, cum deinceps limites confusionis per experientiam constituemus.

### Problema. 3.

175. Si oculus per vnicam lentem obiectum Tab. I. aspiciat ita vt imago visa ab oculo in distantia iu- Fig. 3. sta reperiatur, definire confusionem, qua visio afficietur.

### Solutio.

Sit distantia obiecti  $E\varepsilon$  ante lentem  $AE = a$ , imaginis vero principalis  $F\zeta$  post lentem  $BF = \alpha$ , lentisque crassities  $AB = v$ , quantitas vero arbitraria qua cum binis distantis determinatricibus  $a$  et  $\alpha$  lentis facies determinatur, sit  $= k$ , ponatur autem breuitatis gratia  $\frac{k-v}{k+v} = i$ . Tum autem posita ratione refractionis  $= n$ , debet esse lentis  $PP$ .

$$\text{radius faciei anterioris} = \frac{(n-1)a(k+v)}{k+v+2na}$$

$$\text{radius faciei posterioris} = \frac{(n-1)\alpha(k-v)}{k-v-2n\alpha}$$

T 2

AC



ac si semidiameter aperturæ faciei anterioris sit  $=x$ , posterioris vero  $>ix$  spatium diffusionis  $Ff$  erit  $=Pa\alpha x x$  existente

$$P = \frac{n}{2(n-1)} \left( \frac{1}{iz} \left( \frac{n}{a} + \frac{2}{k+v} \right) \left( \frac{1}{a} + \frac{2}{k+v} \right)^2 + ii \left( \frac{n}{a} - \frac{2}{k-v} \right) \left( \frac{1}{a} - \frac{2}{k-v} \right)^2 \right)$$

radiorum vero in  $f$  concurrentium inclinatio ad axem  $=i. \frac{x}{a}$ . His, quæ §. 86. sunt stabilita, præmissis, sit distantia oculi post lentem  $BO=O$ , et quia imago  $F\zeta$  ante oculum in distantia iusta existere assumitur, erit  $O=\alpha+l$ , ideoque  $\alpha=O-l$ , si locus oculi ut datus consideretur. Hinc ergo habebimus  $V=Pa\alpha$  et  $\mathfrak{B}=\frac{i}{a}$  ideoque  $\mathfrak{B}V=i\alpha P$ . Consequenter radius circellorum in oculo confusionem metientium, seu ut in posterum loquemur mensura confusionis erit

$$=\frac{u}{4l}. i\alpha x^3. P = \frac{i u}{4l} (O-l)x^3 P.$$

Si lentis crassities  $v$  euanescat, et pro determinatione lentis numerus arbitrarius  $\lambda$  loco  $k$  introducatur, erit ex §. 91.  $P=\mu \left( \frac{1}{a} + \frac{1}{a} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{a} \right)^2 + \frac{v}{a\alpha} \right)$ , et  $i=1$  ex quo loco citato etiam ipsa lentis constructio est petenda.

### COROLL. I.

176. Si igitur lens fuerit data, locus oculi post lentem ita definitur, ut debeat esse distantia  $BO=O=\alpha+l$ , existente  $l$  distantia oculi iusta: Sin autem locus oculi detur, pro lentis constructione distantia determinatrix  $\alpha$  ita capi debet, ut sit  $\alpha=O-l$ .

Coroll.

## Coroll. 2.

177. Quare si oculus ita fuerit comparatus, ut exigit distantiam iustam  $l = \infty$ , fiet  $\alpha = -\infty$ , et mensura confusiois erit  $= -\frac{1}{4} i u x^3$ . P seu  $= \frac{1}{4} i u x^3$ . P, quia signum  $-$  nihil mutat in magnitudine circellorum confusioem producentium.

## Coroll. 3.

178. Etsi ergo hoc casu quo  $\alpha = \infty$  spatium diffusionis  $Ff$  est infinitum, tamen inde confusio in visione orta est finita, quia hoc non obstante valor ipsius P manet finitus: erit enim

$$P = \frac{n}{2(n-1)^2} \left( \frac{1}{ii} \left( \frac{n}{a} + \frac{2}{k+v} \right) \left( \frac{1}{a} + \frac{2}{k+v} \right)^2 - \frac{8ii}{(k-v)^2} \right)$$

Radius autem faciei posterioris fit  $= -\frac{(n-1)}{2n} (k-v)$ .

## Coroll. 4.

179. Quia est  $i = \frac{k-v}{k+v}$ , valor ipsius P etiam in genere ita exprimi potest, ut sit:

$$P = \frac{n}{2(n-1)^2} \left( \left( \frac{n}{a} + \frac{2}{k+v} \right) \left( \frac{1}{ia} + \frac{2}{k-v} \right)^2 + \left( \frac{n}{a} - \frac{2}{k-v} \right) \left( \frac{i}{a} - \frac{2}{k+v} \right)^2 \right)$$

vbi est  $\frac{n}{2(n-1)^2} = \frac{310}{121} = 2,561983$  ob  $n = \frac{31}{20}$ . Sicque etiam valores reliquarum litterarum Q, R, S, etc. in §. 86 transformari poterunt.

## Problema 4.

180. Si oculus per duas lentes obiectum  $Ee$  aspiciat, ita ut imago per eas repraesentata  $G\eta$  in distantia iusta  $OG = l$  ob oculo sit remota, definire confusioem, qua visio afficietur.

Tab. I.  
Fig. 5.

T 3

Solutio



## Solutio.

Sint ut hactenus pro lente PP distantiae determinatrices  $AE = a$ ,  $aF = \alpha$ , crassities  $Aa = v$ , et distantia arbitraria  $= k$ , ita ut sit

$$\text{radius faciei anterioris} = \frac{(n-1)a(k+v)}{k+v+2na}$$

$$\text{radius faciei posterioris} = \frac{(n-1)\alpha(k-v)}{k-v-2n\alpha}$$

tum vero posito  $\frac{k-v}{k+v} = i$  ponatur

$$P = \frac{n}{2(n-1)^2} \left( \left( \frac{n}{a} + \frac{2}{k+v} \right) \left( \frac{1}{i a} + \frac{2}{k-v} \right)^2 + \left( \frac{n}{\alpha} - \frac{2}{k-v} \right) \left( \frac{i}{\alpha} - \frac{2}{k+v} \right)^2 \right)$$

Deinde pro lente posteriori QQ sint distantiae determinatrices  $BF = b$ ,  $bG = \mathfrak{E}$ , crassities lentis  $Bb = v'$  et distantia arbitraria  $= k'$  ut sit

$$\text{radius faciei anterioris} = \frac{(n-1)b(k'+v')}{k'+v'+2nb}$$

$$\text{radius faciei posterioris} = \frac{(n-1)\mathfrak{E}(k'-v')}{k'-v'-2n\mathfrak{E}}$$

tum vero posito  $\frac{k'-v'}{k'+v'} = i'$  ponatur

$$Q = \frac{n}{2(n-1)^2} \left( \left( \frac{n}{b} + \frac{2}{k'+v'} \right) \left( \frac{1}{i' b} + \frac{2}{k'-v'} \right)^2 + \left( \frac{n}{\mathfrak{E}} - \frac{2}{k'-v'} \right) \left( \frac{i'}{\mathfrak{E}} - \frac{2}{k'+v'} \right)^2 \right)$$

Atque iam erit spatium diffusionis

$$Gg = \mathfrak{E}\mathfrak{E}xx \left( \frac{1}{vi} \cdot \frac{a\alpha}{bb} P + ii \cdot \frac{bb}{\alpha\alpha} Q \right) = Vxx$$

et radiorum in  $g$  concurrentium inclinatio ad axem

$$ii' \cdot \frac{bx}{\alpha\mathfrak{E}} = \mathfrak{B}x.$$

Sit iam oculus in  $O$  existente  $OG = l$ , ac ponatur eius post lentem QQ distantia  $BO = O$  erit  $O = \mathfrak{E} + l$ , ideoque

ideoque  $\xi = 0 - l$ ; quibus positis cum sit mensura confusionis  $= \frac{u}{4l} \mathfrak{B}x$ . Vix erit ea pro nostro casu:

$$\frac{ii'u}{4l} \cdot \frac{bb}{\alpha} x^3 \left( \frac{1}{i'i'} \cdot \frac{\alpha\alpha}{bb} P + ii \cdot \frac{bb}{\alpha\alpha} Q \right)$$

vel pro  $\xi$  posito valore  $0 - l$  et signo mutato

$$\frac{1}{4} i i' u \left( 1 - \frac{0}{l} \right) \cdot \frac{bb}{\alpha} x^3 \left( \frac{1}{i'i'} \cdot \frac{\alpha\alpha}{bb} P + ii \cdot \frac{bb}{\alpha\alpha} Q \right).$$

### COROLL. I.

181. Si ergo pro oculo fuerit  $l = \infty$ , ideoque et  $\xi = \infty$ , etsi spatium diffusionis  $Gg$  sit infinitum tamen ob  $\frac{0}{l} = 0$  confusio visionem afficiens nihilominus erit finita: neque enim  $P$  neque  $Q$  ob  $\xi = \infty$  fit infinita.

### COROLL. 2.

182. Si crassities lentium evanescat, ut sit  $v = 0$  et  $v' = 0$  erit  $i = 1$  et  $i' = 1$ . ideoque hoc casu mensura confusionis

$$\frac{1}{4} u \left( 1 - \frac{0}{l} \right) \cdot \frac{bb}{\alpha} x^3 \left( \frac{\alpha\alpha}{bb} P + \frac{bb}{\alpha\alpha} Q \right)$$

At si loco  $k$  et  $k'$  introducantur numeri  $\lambda$  et  $\lambda'$  erit

$$P = \mu \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{v}{a\alpha} \right) \text{ et}$$

$$Q = \mu \left( \frac{1}{b} + \frac{1}{\beta} \right) \left( \lambda' \left( \frac{1}{b} + \frac{1}{\beta} \right)^2 + \frac{v}{b\beta} \right)$$

existente

$$\mu = 0,938191 \text{ et } v = 0,232692$$

Coroll.



## Coroll. 3.

183. Eodem autem hoc casu constructio binarum lentium ita se habebit :

$$\begin{array}{l} \text{radius faciei} \\ \text{Pro lente PP} \left\{ \begin{array}{l} \text{anterioris} = \frac{a\alpha}{\varrho\alpha + \sigma\alpha \pm \tau(a+\alpha)\sqrt{(\lambda-1)}} \\ \text{posterioris} = \frac{a\alpha}{\varrho\alpha + \sigma\alpha \mp \tau(a+\alpha)\sqrt{(\lambda-1)}} \end{array} \right. \\ \text{Pro lente QQ} \left\{ \begin{array}{l} \text{anterioris} = \frac{b\beta}{\varrho\beta + \sigma\beta \pm \tau(b+\beta)\sqrt{(\lambda'-1)}} \\ \text{posterioris} = \frac{b\beta}{\varrho\beta + \sigma\beta \mp \tau(b+\beta)\sqrt{(\lambda'-1)}} \end{array} \right. \end{array}$$

existente  $\varrho=0,190781$ ;  $\sigma=1,627401$  et  $\tau=0,905133$ .

## Problema 5.

Tab. II.  
Fig. 6.

184. Si oculus per tres lentes PP, QQ et RR obiectum  $E\epsilon$  aspiciat, ita vt imago per eas repraesentata  $H\theta$  in distantia iusta  $OH=l$  ab oculo O fit remota, definire confusionem, qua visio afficietur.

## Solutio.

Positis duabus prioribus lentibus PP et QQ vt in problemate praecedente, indeque determinatis valoribus P et Q cadat imago per has duas lentes repraesentata principalis in  $G\eta$ , post quam tertia lens RR ita collocata fit, vt sint eius distantiae determinatrices  $CG=c$ ,  $cH=\gamma$ , crassities  $Cc=v''$  et quantitas arbitraria  $=k''$  vt fit:

$$\text{radius faciei anterioris} = \frac{(n-1)c(k''+v'')}{k''+v''+2nc}$$

$$\text{radius faciei posterioris} = \frac{(n-1)\gamma(k''-v'')}{k''-v''-2n\gamma}$$

Tum

Tum vero posito  $\frac{k''-v''}{k''+v''} = i''$  ponatur

$$R = \frac{n}{2(n-1)} \left( \left( \frac{n}{c} + \frac{2}{k''+v''} \right) \left( \frac{1}{i''c} + \frac{2}{k''-v''} \right)^2 + \left( \frac{n}{\gamma} - \frac{2}{k''-v''} \right) \left( \frac{i''}{\gamma} - \frac{2}{k''+v''} \right)^2 \right)$$

atque iam spatium diffusionis erit

$$Hb = \gamma \gamma x x \left( \frac{1}{i' i'' i'''} \cdot \frac{\alpha \alpha \epsilon \epsilon}{b b c c} P + \frac{i i}{i'' i'''} \cdot \frac{b b \epsilon \epsilon}{\alpha \alpha c c} Q + i i \cdot i' i'' \cdot \frac{b b c c}{\alpha \alpha \epsilon \epsilon} R \right)$$

quod est valor ipsius  $Vxx$ . Radiorum vero in  $b$  concurrentium inclinatio ad axem est

$$i i' i'' \cdot \frac{b c x}{\alpha \epsilon \gamma} = \mathfrak{B} x$$

fit iam oculus in  $O$ , ac ponatur eius distantia post lentem  $RR = O$  erit  $O = \gamma + l$ , ideoque  $\gamma = O - l$ .

Hinc mensura confusiois in oculo ortae colligitur

$$\frac{1}{4} i i' i'' u \left( 1 - \frac{O}{l} \right) \frac{b c}{\alpha \epsilon} x^3 \left( \frac{1}{i' i'' i'''} \cdot \frac{\alpha \alpha \epsilon \epsilon}{b b c c} P + \frac{i i}{i'' i'''} \cdot \frac{b b \epsilon \epsilon}{\alpha \alpha c c} Q + i i \cdot i' i'' \cdot \frac{b b c c}{\alpha \alpha \epsilon \epsilon} R \right)$$

### COROLL. I.

185. Hic iterum vt ante patet, si fuerit  $l = \infty$  ideoque et  $\gamma = \infty$ , quo casu diffusionis spatium in infinitum extenditur, mensuram confusiois terminis finitis contineri, quod etiam locum habet pro quouis lentium numero.

### COROLL. 2.

186. Si lentium crassities pro euanescente habeatur, ob  $i = 1$   $i' = 1$ ,  $i'' = 1$ , mensura confusiois ita simplicius exprimetur, vt fit

$$= \frac{1}{4} u \left( 1 - \frac{O}{l} \right) \frac{b c}{\alpha \epsilon} x^3 \left( \frac{\alpha \alpha \epsilon \epsilon}{b b c c} P + \frac{b b \epsilon \epsilon}{\alpha \alpha c c} Q + \frac{b b c c}{\alpha \alpha \epsilon \epsilon} R \right)$$

Tom. I.

V

At



At hoc casu loco  $k, k', k''$  introductis numeris  $\lambda, \lambda', \lambda''$  erit

$$P = \mu \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{\nu}{a\alpha} \right)$$

$$Q = \mu \left( \frac{1}{b} + \frac{1}{\beta} \right) \left( \lambda' \left( \frac{1}{b} + \frac{1}{\beta} \right)^2 + \frac{\nu}{b\beta} \right)$$

$$R = \mu \left( \frac{1}{c} + \frac{1}{\gamma} \right) \left( \lambda'' \left( \frac{1}{c} + \frac{1}{\gamma} \right)^2 + \frac{\nu}{c\gamma} \right)$$

### Coroll. 3.

187. Eodem vero casu constructio lentium per istos numeros  $\lambda, \lambda', \lambda''$  ita erit dirigenda ut sit

Radius faciei

$$\text{Pro lente PP} \begin{cases} \text{anterioris} = \frac{a\alpha}{\rho\alpha + \sigma\alpha \pm \tau(a+\alpha)\sqrt{(\lambda-1)}} \\ \text{posterioris} = \frac{a\alpha}{\rho\alpha + \sigma\alpha \mp \tau(a+\alpha)\sqrt{(\lambda-1)}} \end{cases}$$

$$\text{Pro lente QQ} \begin{cases} \text{anterioris} = \frac{b\beta}{\rho\beta + \sigma\beta \pm \tau(b+\beta)\sqrt{(\lambda'-1)}} \\ \text{posterioris} = \frac{b\beta}{\rho\beta + \sigma\beta \mp \tau(b+\beta)\sqrt{(\lambda'-1)}} \end{cases}$$

$$\text{Pro lente RR} \begin{cases} \text{anterioris} = \frac{c\gamma}{\rho\gamma + \sigma\gamma \pm \tau(c+\gamma)\sqrt{(\lambda''-1)}} \\ \text{posterioris} = \frac{c\gamma}{\rho\gamma + \sigma\gamma \mp \tau(c+\gamma)\sqrt{(\lambda''-1)}} \end{cases}$$

### Scholion.

188. Hinc satis manifestum est, quemadmodum hae formulae pro pluribus lentibus progrediantur: verum antequam eas exponam, conueniet alias quoque circumstantias, quae hinc facillime deducuntur, perpendi, scilicet magnitudinem obiecti visam, et copiam radiorum a singulis eius punctis in oculum transmissorum ut hae simul cum confusione visionis de-

deinceps coniunctim pro quouis lentium numero exhiberi queant; quo pacto plures taediosas repetitiones evitabimus. Duae autem res ex haftenus allatis facile definiri possunt, quarum altera est quantitas, qua imago obiecti per lentes repraesentata ab oculo cernitur, quae quantitas aestimanda est ex angulo, sub quo imago videtur, ut is deinceps comparari possit, cum eo angulo sub quo ipsum obiectum in data distantia a nudo oculo spectaretur, unde qua ratione magnitudo per lentes visa augeatur, intelligetur. Altera res in copia radiorum, a singulis obiecti punctis in oculum transmissorum versatur, qua claritas visione percepta continetur, a quolibet scilicet puncto conus seu cylindrus radiosus in oculum ingreditur, qui si pupillam penitus expleat, claritas ad summum gradum erit euecta, nisi forte maiori illustratione ipsi obiecto maius lumen concilietur. At si sectio illius coni aut cylindri, qua in oculum intrat, minor fuerit pupilla, in eadem ratione claritas decrescet; quod cum in omnibus instrumentis dioptricis, quibus magnitudinem visam vehementer augere propositum est, usu venire soleat, plurimum intererit amplitudinem illius coni seu cylindri, qua in oculum penetrat, accurate determinasse.

### Problema 6.

189. Definire quantitatem, sub qua quaevis obiecti portio per lentes quocunque ab oculo in distantia iusta ab imagine ultima remoto cernetur.

V 2

Solutio.



## Solutio.

Sit  $z$  linea in obiecto concepta, quae quanta per lentes oculo fit apparitura, definiri oporteat. Ostensum autem est in praecedentibus, quotcumque fuerint lentes, imaginem principalem huius lineae iterum esse lineam, cuius longitudo ad  $z$  certam teneat rationem a distantis determinatricibus lentium et numeris  $i, i', i'', i'''$  etc. pendentem (86). Sit ergo haec longitudo imaginis  $= Mz$ , quae cum ab oculo in distantia  $l$  remoto aspiciatur, apparebit sub angulo  $= \frac{Mz}{l}$ ; vel cuius tangens potius sit  $= \frac{Mz}{l}$ ; sed quia hic angulus rarissime ultra aliquot gradus assurgere solet, tangens tuto pro ipso arcu assumitur. Effigies autem quae ab hac linea  $Mz$  in oculo exprimitur, erit  $= \frac{Mu}{l}z$ ; hic enim cogitationes a confusione abstraho, qua utique fit, ut effigies maior imprimatur, propterea quod singula puncta circellis exhibentur. Iam positis iisdem lentium determinationibus, quibus supra §. 86 sum usus, pro vario lentium numero angulus, sub quo linea  $z$  in obiecto sumpta cernetur ita se habebit:

|                                | Angulus visionis                                                                             | situ    |
|--------------------------------|----------------------------------------------------------------------------------------------|---------|
| Pro vnica lente . . . . .      | $\frac{1}{i} \cdot \frac{\alpha}{a} \cdot \frac{z}{l}$                                       | inuerso |
| Pro duabus lentibus . . . . .  | $\frac{1}{i i'} \cdot \frac{\alpha \beta}{a b} \cdot \frac{z}{l}$                            | erecto  |
| Pro tribus lentibus . . . . .  | $\frac{1}{i i' i''} \cdot \frac{\alpha \beta \gamma}{a b c} \cdot \frac{z}{l}$               | inuerso |
| Pro quatuor lentibus . . . . . | $\frac{1}{i i' i'' i'''} \cdot \frac{\alpha \beta \gamma \delta}{a b c d} \cdot \frac{z}{l}$ | erecto  |
|                                | etc.                                                                                         |         |

Scilicet

Scilicet si hae formulae sint positivae, oculus lineam  $z$  situ siue erecto siue inuerso videbit, prout est notatum sin autem fuerint negativae, situm indicatum in contrarium verti oportet.

### COROLL. I.

190. Si eadem obiecti linea  $z$  in distantia  $=b$  ab oculo nudo cerneretur, ea apparitura esset sub angulo  $=\frac{z}{b}$ , vnde perspicitur, quanto ea vel maior vel minor per lentes videatur.

### COROLL. 2.

191. Si distantia oculi a postrema lente ponatur  $=O$ , erit casus vnus lentis  $\alpha=O-l$ , ideoque  $\frac{\alpha}{l}=-\left(1-\frac{O}{l}\right)$  vnde angulus opticus lineae obiecti  $z$  respondens erit  $=\frac{1}{l}\left(1-\frac{O}{l}\right)\frac{z}{\alpha}$  pro situ erecto, quia signum mutauimus.

### COROLL. 3.

192. Simili modo casu duarum lentium ob  $\xi=O-l$  erit iste angulus  $=\frac{1}{l'l'}\left(1-\frac{O}{l}\right)\frac{\alpha z}{a'b}$  pro situ inuerso.

Casu vero trium lentium ob  $\gamma=O-l$  erit

iste angulus  $=\frac{1}{l'l'v''}\left(1-\frac{O}{l}\right)\frac{\alpha\xi z}{a'b'c}$  pro situ erecto

Casu quatuor lentium ob  $\delta=O-l$  erit

iste angulus  $=\frac{1}{l'l'v''v'''}\left(1-\frac{O}{l}\right)\frac{\alpha\xi\gamma z}{a'b'c'd}$  pro situ inuerso, et ita porro pro pluribus lentibus.



## Scholion.

193. Hinc etiam modus se offert confusione[m] ob lentium aperturam in oculo natam distinctius aestimandi. Scilicet cum singula obiecti puncta in oculo exprimantur circulis, quorum radius est  $= \frac{u}{4l} \mathfrak{B} V x^3$ , verus autem circulus cuius radius  $= z$ , oculo in distantia  $l$  expositus in oculo referatur circulo cuius radius  $= \frac{u}{l} z$  singula puncta illius obiecti per lentes spectati, aequae magna apparebunt, ac orbes circulares radii  $z = \frac{1}{4} \mathfrak{B} V x^3$ , si in distantia ab oculo  $= l$  spectarentur. Vel cum horum orbium semidiameter apparens sit  $= \frac{z}{l}$ , ob confusione[m] singula obiecti puncta instar circulorum videbuntur, quorum semidiameter apparens esset  $= \frac{1}{4l} \mathfrak{B} V x^3$ . In expressionibus igitur ante pro confusione inuentis deleatur quantitas  $u$ , et habebitur semidiameter apparens circulorum confusione[m] exprimentium. Hinc iudicari poterit quam parua esse debeat confusio, ut non amplius sentiatur, scilicet si oculus non amplius percipere valeat spatium circulare, cuius semidiameter esset  $1''$ , seu  $\frac{1}{603}$  pars radii circiter, euidens est, si fuerit nostra expressio  $\frac{1}{4l} \mathfrak{B} V x^3 = \frac{1}{603}$ , confusione[m] fore imperceptibilem. Ac experientiam consulentes deprehendimus multo maiores angulos non amplius percipi posse, ita ut confusio non sit metuenda etiam si expressio  $\frac{1}{4l} \mathfrak{B} V x^3$  notabiliter maior fuerit, quam  $\frac{1}{603}$ ; ne autem hic temere quicquam statuamus, ponamus limitem, quem formula  $\frac{1}{4l} \mathfrak{B} V x^3$  excedere non

non debeat, esse  $= \frac{1}{x^3}$ , ita vt esse oporteat  $\frac{1}{2} V x^3 < \frac{1}{x^3}$ .  
 Postea igitur ad praxin descendentes poterimus pro  $x$   
 numerum vel 40 vel minorem assumere prout expe-  
 rientia quouis casu postulauerit. Quod si ergo hoc  
 modo confusionis rationem habeamus, profunditas oculi  
 $u$  non amplius in computum ingreditur.

### Definitio 3.

194. Semidiameter confusionis, est semidiamete-  
 ter apparens circuli, qui ab oculo aequae magnus vi-  
 detur, ac singula obiecti puncta ipsi ob confusionem  
 apparent.

### Corollarium.

195. Inueniemus igitur facile semidiametrum  
 confusionis si formulas supra pro confusione repertas  
 per profunditatem oculi  $u$  diuidamus quo pacto eae  
 formulae ad numeros absolutos reducentur.

### Definitio 4.

196. Multiplicatio per lentes producta ex ra-  
 tione quantitatis, qua obiecta per lentes spectantur,  
 ad quantitatem, qua eadem obiecta in data distantia  
 ab oculo nudo cernerentur aestimatur. Exponens au-  
 tem multiplicationis inuenitur, si magnitudo, qua  
 linea quaecunque in obiecto concepta per lentes vide-  
 tur, diuidatur per magnitudinem, qua eadem linea  
 in data distantia ab oculo nudo spectata esset apparitura.

Coroll.



## Coroll. 1.

197. Inuoluit ergo diiudicatio multiplicationis distantiam quandam fixam, in qua eadem obiecta a nudo oculo aspici assumimus quae prout diuersa assumatur, multiplicatio alio atque alio modo exprimitur.

## Coroll. 2.

198. Si haec distantia fixa, ex qua multiplicatio diiudicatur ponatur  $=b$  et exponens multiplicationis  $=m$ , fit linea quaequam in obiecto concepta  $=z$ , quae ergo nudo oculo in distantia  $b$  appareret sub angulo  $=\frac{z}{b}$ ; eadem autem linea per lentes spectetur sub angulo  $=\frac{Mz}{l}$  (88) ex quo erit exponens multiplicationis  $m = \frac{Mb}{l}$ .

## Coroll. 3.

199. In ratione ergo  $m:1$  dimensiones lineares per lentes augeri sunt censendae; vnde superficies auctae apparebunt in ratione  $mm:1$ , et ipsa corpora in ratione  $m^3:1$ . Cum exponente autem multiplicationis coniungi debet situs, quo obiecta apparent siue is sit erectus siue inuersus.

## Scholion.

200. Distantia haec fixa  $b$  ad quam multiplicatio refertur non eodem modo perpetuo assumi solet, quippe quod etiam pro diuersitate obiectorum omnino fieri non posset. Nam si per lentes obiecta  
valde

valde remota veluti coelestia contuemur, quoniam ea nunquam in distantia modica spectare solemus, conuenit utique magnitudinem per lentes visam cum ea comparare, qua in ea ipsa a nobis distantia nudis oculis cernerentur: ideoque his casibus distantia fixa  $b$  ipsi distantiae  $a$ , qua a lentibus obiecta sunt remota, aequalis constitui solet. Scilicet si distantia  $a$  fuerit valde magna, qui est casus Telescopiorum, statuitur  $b = a$ , et magnitudo per haec instrumenta visa, cum magnitudine per nudos oculos visa in eadem distantia commodissime comparatur. Sic de Telescopiis dicitur, quoties diametri corporum coelestium multiplicentur, eoque casu littera  $m$  exponentem huius rationis indicabit. Sin autem obiecta propiora contemplamur, qui est vsus Microscopiorum ea plerumque ita prope ad instrumentum admouentur, ut in tam exigua distantia nudis oculis nunquam distincte cerni possent: neque ergo his casibus statui  $b = a$  conueniret. Aliam ergo rationem ineundo pro  $b$  sumi solet eiusmodi distantia modica, in qua obiecta commode ac distincte cernere liceat, quae etsi utique pro diuersa oculorum indole diuersa sumi deberet; tamen ut aliquid fixi statuatur, pro  $b$  distantia 8 pollicum utpote maximae oculorum parti conueniens accipi solet, ita ut his casibus definiamus, quoties huiusmodi obiecta maiora *appareant*, quam si eadem nudis oculis in distantia 8 digitorum aspicerentur. Interim tamen si multiplicatio ad hanc distantiam fuerit relata, non difficile erit eam ad quam-



libet aliam referre, ita vt haec hypothesis naturam horum instrumentorum afficere non sit censenda.

### Problema 7.

Tab. II.  
Fig. 10.

201. Si obiectum per lentes quotcunque aspi-  
ciatur, definire amplitudinem coni seu cylindri  
luminosi, qui a singulis obiecti punctis in oculum  
transmittitur.

### Solutio.

Reperiatur vt supra oculus in distantia iusta  $l$   
post imaginem postremam per lentes repraesentatam  
quae etsi est per spatium aliquod  $Ll$  diffusa, hic  
tamen a confusione inde nata mentem abstrahimus,  
quoniam confusionem iam seorsim determinauimus.  
Ex puncto igitur  $l$  oculum versus diffunditur conus  
luminosus, cuius radii extremi ad axem inclinati sunt  
angulo  $O/T$  quem posuimus supra  $= Bx$ . Huius igitur  
coni sectio circa ingressum in oculum consideretur,  
cuius semidiameter erit  $= B/x$ , vnde pro vario lentium  
numero hic semidiameter sequenti modo definietur.

Pro vnica lente . . . . .  $i/l \frac{x}{\alpha}$

Pro duabus lentibus . . . . .  $i i' / l \frac{b x}{\alpha \beta}$

Pro tribus lentibus . . . . .  $i i' i'' / l \frac{\delta c x}{\alpha \beta \gamma}$

Pro quatuor lentibus . . . . .  $i i' i'' i''' / l \frac{h c d x}{\alpha \beta \gamma \delta}$

etc.

hicque perinde est siue hae formulae sint positivae  
siue negatiuae; quia circulus siue radio positiuo siue  
negatiuo describatur eiusdem prodit magnitudinis.

Coroll.

## Coroll. 1.

202. Si iste semidiameter  $\mathfrak{B}/x$  maior fuerit semidiametro pupillae, tota pupillae apertura radiis impletur neque propterea vis clarior procurari poterit, nisi forte ipsum obiectum fortiori illuminatione splendidius reddatur.

## Coroll. 2.

203. Sin autem haec quantitas  $\mathfrak{B}/x$  minor fuerit semidiametro pupillae claritatis mensura existet, quae eo maior erit quo maior fuerit ista quantitas: dum contra minuta hac quantitate claritas tam exigua euadere potest, ut non amplius sensui visus excitando sufficiat.

## Definitio 5.

204. Gradus claritatis per lentes perceptae commodissime definietur semidiametro coni luminosi, qui a quouis obiecti puncto in oculum transmittitur.

## Coroll. 1.

205. Gradus ergo claritatis definitur quantitate supra inuenta  $\mathfrak{B}/x$ , ita ut si gradum claritatis ponamus  $=y$  habeamus  $y = \mathfrak{B}/x$ , quo cognito facillime iudicabimus, quonam gradu visio sit clara habenda.

## Coroll. 2.

206. Scilicet si semidiametrum pupillae ponamus  $=\omega$ , quamdiu fuerit  $y > \omega$ , claritate plena  
X 2
fruemur,



fruemur, quae nullius augmenti est capax nisi forte ipsam pupillam magis dilatare valeamus.

### C O R O L L. 3.

207. At si fuerit  $y < \omega$ , claritatem utique minorem percipiemus: ac si claritatem plenam unitate designemus, claritas ex casu  $y < \omega$  resultans erit  $= \frac{y}{\omega}$ ; propterea quod copia radiorum in oculum immissorum est ut quadratum semidiametri  $y$ .

### C O R O L L. 4.

208. Quod si gradus claritatis  $y$  eousque decrescat, ut copia radiorum nimis sit parva, quam ut sensum visus excitare possit, nihil ob summam caliginem percipi poterit, unde manifestum est ad visionem requiri, ut gradus claritatis certum quempiam limitem superet.

### S c h o l i o n.

209. Tam in Telescopiis quam Microscopiis maxime necesse est; ut obiecta certo claritatis gradu exhibeantur ne repraesentatio nimis fiat obscura. Hic autem gradus plurimum a lumine proprio obiectorum pendet, quae quo fuerint illustriora minor claritatis gradus iis satis clare videndis sufficit; ideoque stellas illustriores contemplantes minori gradu claritatis contenti esse possumus: terrestria vero obiecta multo maiorem claritatis gradum postulant. Quo igitur haec  
ad

ad omnes casus accommodare valeamus gradum claritatis hic littera indefinita  $y$  contentum in computum ducturus. Quamobrem his de multiplicatione et claritate praemissis, haec duo elementa simul cum confusione pro quouis lentium numero exhibebo: ac primo quidem non neglecta lentium crassitie, tum vero eadem scorsum lentium crassitie neglecta exponi conueniet.

### Problema 8.

210. Si oculus per quocunque lentes PP, QQ, RR, SS etc. obiectum  $E\varepsilon$  aspiciat; ita vt imago postrema per eas repraesentata ante oculum in iusta distantia  $= l$  reperiatur; determinare tam multiplicationem et claritatem, quam confusionem, qua visio perturbabitur.

### Solutio

Quocunque fuerint lentes, sint pro singulis distantiae determinatrices, vt et crassities cum quantitate arbitraria, vt sequitur.

| Pro Lente  | Diff: determinatr: | Crassities | quant: arb: |
|------------|--------------------|------------|-------------|
| Prima PP   | $EA=a; aF=\alpha;$ | $Aa=v;$    | $k$         |
| Secunda QQ | $FB=b; bG=\beta;$  | $Bb=v';$   | $k'$        |
| Tertia RR  | $GC=c; cH=\gamma;$ | $Cc=v'';$  | $k''$       |
| Quarta SS  | $HD=d; dI=\delta;$ | $Dd=v''';$ | $k'''$      |
|            | etc.               |            |             |



atque hinc posita ratione refractionis  $\frac{z'}{z''} = n$  constructio lentium ita se habebit;

| Pro Lente  | Radius Faciei                                   |                                                           |
|------------|-------------------------------------------------|-----------------------------------------------------------|
|            | anterioris                                      | posterioris                                               |
| Prima P P  | $\frac{(n-1)a(k+v)}{k+v+2na}$                   | $\frac{(n-1)\alpha(k-v)}{k-v-2n\alpha}$                   |
| Secunda QQ | $\frac{(n-1)b(k'+v')}{k'+v'+2nb}$               | $\frac{(n-1)\beta(k'-v')}{k'-v'-2n\beta}$                 |
| Tertia RR  | $\frac{(n-1)c(k''+v'')}{k''+v''+2nc}$           | $\frac{(n-1)\gamma(k''-v'')}{k''-v''-2n\gamma}$           |
| Quarta SS  | $\frac{(n-1)d(k''' + v''')}{k''' + v''' + 2nd}$ | $\frac{(n-1)\delta(k''' - v''')}{k''' - v''' - 2n\delta}$ |
|            | etc.                                            |                                                           |

Tam posito breuitatis gratia:

$$\frac{k-v}{k+v} = i; \frac{k'-v'}{k'+v'} = i'; \frac{k''-v''}{k''+v''} = i''; \frac{k'''-v'''}{k''' + v'''} = i''' \text{ etc.}$$

si aperturæ primæ lentis in facie anteriori semidiameter fuerit  $= x$ , tam pro facie posteriori quam pro vtraque facie singularum lentium sequentium aperturæ maiores esse debent vel saltem non minores, quam sequens tabula ostendit:

| Pro Lente   | Semidiameter aperturæ in facie        |                                            |
|-------------|---------------------------------------|--------------------------------------------|
|             | anteriori                             | posteriori                                 |
| Prima P P   | $x$                                   | $i x$                                      |
| Secunda Q Q | $i. \frac{bx}{a}$                     | $i i'. \frac{bx}{a}$                       |
| Tertia R R  | $i i'. \frac{bcx}{a\beta}$            | $i i' i''. \frac{bcx}{a\beta}$             |
| Quarta S S  | $i i' i''. \frac{bcdx}{a\beta\gamma}$ | $i i' i'' i'''. \frac{bcdx}{a\beta\gamma}$ |
|             | etc.                                  |                                            |

Denique

Denique ad abbreviandum ponatur :

$$P = \frac{n}{2(n-1)^2} \left( \left( \frac{n}{a} + \frac{2}{k+v} \right) \left( \frac{1}{ia} + \frac{2}{k-v} \right)^2 + \left( \frac{n}{a} - \frac{2}{k-v} \right) \left( \frac{1}{a} - \frac{2}{k+v} \right)^2 \right)$$

$$Q = \frac{n}{2(n-1)^2} \left( \left( \frac{n}{b} + \frac{2}{k'+v'} \right) \left( \frac{1}{ib'} + \frac{2}{k'-v'} \right)^2 + \left( \frac{n}{b} - \frac{2}{k'-v'} \right) \left( \frac{1}{b} - \frac{2}{k'+v'} \right)^2 \right)$$

$$R = \frac{n}{2(n-1)^2} \left( \left( \frac{n}{c} + \frac{2}{k''+v''} \right) \left( \frac{1}{ic''} + \frac{2}{k''-v''} \right)^2 + \left( \frac{n}{c} - \frac{2}{k''-v''} \right) \left( \frac{1}{c} - \frac{2}{k''+v''} \right)^2 \right)$$

$$S = \frac{n}{2(n-1)^2} \left( \left( \frac{n}{d} + \frac{2}{k''' + v'''} \right) \left( \frac{1}{id} + \frac{2}{k''' - v'''} \right)^2 + \left( \frac{n}{d} - \frac{2}{k''' - v'''} \right) \left( \frac{1}{d} - \frac{2}{k''' + v'''} \right)^2 \right)$$

etc.

His positis ponamus oculum in distantia  $= O$  ab ultima lente locari ita ut post imaginem ultimam reperiatur in distantia  $= l$ : magnitudinem autem visam comparari cum magnitudine qua idem obiectum nudo oculo in distantia fixa  $= b$  cerneretur; ac ponatur exponens multiplicationis  $= m$ .

Deinde pro claritate sit gradus claritatis  $= y$ , ita ut  $y$  indicet semidiametrum coni luminosi in oculum intrantis.

Confusio autem aestimetur per semidiametrum confusionis supra (194) definitum.

Iam pro quovis lentium numero hae tres res ita se habebunt.

I. Pro unica lente  $O = \alpha + l$

1. Exponens multiplicationis  $m = \frac{1}{i} \cdot \frac{\alpha b}{a i}$  situ inverso

2. Gradus claritatis  $y = i l \cdot \frac{\alpha}{a}$

3. Semid: Confusionis  $= \frac{1}{4} i \frac{\alpha}{l} x^3 \cdot P$

II.



II. Pro duabus Lentibus  $O = \epsilon + l$ 

1. Exp. mult:  $m = \frac{1}{ii} \cdot \frac{\alpha \epsilon b}{a b l}$  situ erecto
2. Gradus claritatis  $y = ii' l \cdot \frac{b x}{\alpha \epsilon}$
3. Semid. Confusionis  $= \frac{1}{i' i'} \cdot \frac{\epsilon}{l} \cdot \frac{b}{\alpha} x^3 \left( \frac{1}{i' i'} \cdot \frac{\alpha \alpha}{b b} P + ii \cdot \frac{b b}{\alpha \alpha} Q \right).$

III. Pro tribus Lentibus  $O = \gamma + l$ 

1. Exp. mult:  $m = \frac{1}{i' i' i'} \cdot \frac{\alpha \epsilon \gamma b}{a b c l}$  situ inuerso:
  2. Gradus claritatis  $y = ii' i'' l \cdot \frac{b c x}{\alpha \epsilon \gamma}.$
  3. Semidiameter confusionis;
- $$\frac{1}{i' i' i''} \cdot \frac{\gamma}{l} \cdot \frac{b c}{\alpha \epsilon} x^3 \left( \frac{1}{i' i' i'' i''} \cdot \frac{\alpha \alpha \epsilon \epsilon}{b b c c} P + \frac{i i}{i'' i''} \cdot \frac{b b \epsilon \epsilon}{\alpha \alpha c c} Q + ii \cdot i' i' \cdot \frac{b b c c}{\alpha \alpha \epsilon \epsilon} R \right).$$

IV. Pro quatuor Lentibus  $O = \delta + l$ 

1. Exp. multipl:  $m = \frac{1}{i'' i'' i'' i''} \cdot \frac{\alpha \epsilon \gamma \delta b}{a b c d l}$  situ erecto
  2. Gradus Claritatis  $y = ii' i'' i''' l \cdot \frac{b c d x}{\alpha \epsilon \gamma \delta}$
  3. Semidiameter Confusionis;
- $$\frac{1}{i' i' i'' i''} \cdot \frac{\delta}{l} \cdot \frac{b c d}{\alpha \epsilon \gamma} x^3 \left\{ \frac{1}{i' i' i'' i'' i'' i''} \cdot \frac{\alpha \alpha \epsilon \epsilon \gamma \gamma}{b b c c d d} P + \frac{i i}{i'' i'' i'' i''} \cdot \frac{b b \epsilon \epsilon \gamma \gamma}{\alpha \alpha c c d d} Q \right. \\ \left. + \frac{i i \cdot i' i'}{i'' i''} \cdot \frac{b b c c \gamma \gamma}{\alpha \alpha \epsilon \epsilon d d} R + ii \cdot i' i' \cdot i'' i'' \cdot \frac{b b c c d d}{\alpha \alpha \epsilon \epsilon \gamma \gamma} S \right\}$$

V. Pro quinque Lentibus  $O = \varepsilon + l$

1. Exp. multipl:  $m = \frac{1}{i' i'' i''' i''''} \cdot \frac{\alpha \varepsilon \gamma \delta \varepsilon b}{a b c d e l}$  situ inuerſo

2. Gradus claritatis  $y = i' i'' i''' i'''' l \cdot \frac{b c d e x}{\alpha \varepsilon \gamma \delta \varepsilon}$

3. Semidiameter confuſionis

$$\frac{1}{4} i' i'' i''' i'''' \cdot \frac{\varepsilon}{l} \cdot \frac{b c d e}{\alpha \varepsilon \gamma \delta} x^3 \left\{ \begin{array}{l} + \frac{1}{i' i'' i''' i''''} \cdot \frac{\alpha \alpha \varepsilon \varepsilon \gamma \gamma \delta \delta}{b b c c d d e e} P \\ + \frac{i i}{i'' i''' i''''} \cdot \frac{b b \varepsilon \varepsilon \gamma \gamma \delta \delta}{\alpha \alpha c c d d e e} Q \\ + \frac{i i \cdot i' i'}{i''' i''''} \cdot \frac{b b c c \gamma \gamma \delta \delta}{\alpha \alpha \varepsilon \varepsilon d d e e} R \\ + \frac{i i \cdot i' i' \cdot i'' i''}{i''''} \cdot \frac{b b c c d d \delta \delta}{\alpha \alpha \varepsilon \varepsilon \gamma \gamma e e} S \\ + i i \cdot i' i' \cdot i'' i'' \cdot i''' i''' \cdot \frac{b b c c d d e e}{\alpha \alpha \varepsilon \varepsilon \gamma \gamma \delta \delta} T \end{array} \right.$$

atque hinc etiam progressus ad plures lentes est manifestus.

### Coroll. 1.

211. His omnibus casibus euidentis est fore generatim

$$m y = \frac{b x}{\alpha}.$$

Datis scilicet multiplicatione  $m$  cum claritate  $y$  statim definitur apertura primae lentis nempe  $x = m y \cdot \frac{\alpha}{b}$ . Quo maior scilicet tam multiplicatio quam claritas desideratur, eo maiorem esse oportet aperturam lentis primae.

### Coroll. 2.

212. Quum autem  $x$  maius accipere non liceat, quam vt confusio infra certum limitem contineatur, dato  $x$  cum exponente multiplicationis  $m$

Tom. I.

Y

defi-



definitur claritatis gradus  $y = \frac{bx}{ma}$ , vnde patet reliquis paribus, quo maior multiplicatio exigatur, eo minori claritate contentos nos esse oportere.

### C o r o l l. 3.

213. Imprimis autem hic obseruandum est, has formulas aequè negotium conficere, quamcunque crassitiem lentes habuerint. *Euadent autem tractabiliores*, si lentium crassities negligatur, qui casus seorsim tractari meretur.

### P r o b l e m a 9.

214. Iisdem positis, quae in problemate praecedente, si lentium crassities vt euanesceat consideretur, determinare tam multiplicationem et claritatem, quam confusionem, qua visio perturbabitur.

### S o l u t i o.

Haec tractatio a praecedenti in isto differt, quod lentium crassities  $v, v', v''$  etc. euanesceant, et loco quantitatum arbitrariarum  $k, k', k''$  etc. numeri arbitrarii  $\lambda, \lambda', \lambda''$  etc. in calculum introducantur. Ponatur ergo

| Pro Lente |     | Diff. determinatr.    | num. arb.    |
|-----------|-----|-----------------------|--------------|
| Prima     | P P | $EA = a; aF = a$      | $\lambda$    |
| Secunda   | Q Q | $FB = b; bG = \beta$  | $\lambda'$   |
| Tertia    | R R | $GC = c; cH = \gamma$ | $\lambda''$  |
| Quarta    | S S | $HD = d; dI = \delta$ | $\lambda'''$ |
|           |     | etc.                  |              |

Hinc

Hinc si sit breuitatis gratia!

$\varrho = 0, 190781$ ;  $\sigma = 1, 627401$ ; et  $\tau = 0, 905133$   
lentium constructio ita est instituenda

Radius Faciei

| Pro Lente  | anterioris                                                                        | posterioris                                                                       |
|------------|-----------------------------------------------------------------------------------|-----------------------------------------------------------------------------------|
| Prima PP   | $\frac{a \alpha}{\varrho a + \sigma a + \tau(a + \alpha)\sqrt{(\lambda - 1)}}$    | $\frac{a \alpha}{\varrho a + \sigma a + \tau(a + \alpha)\sqrt{(\lambda - 1)}}$    |
| Secunda QQ | $\frac{b \beta}{\varrho b + \sigma b + \tau(b + \beta)\sqrt{(\lambda' - 1)}}$     | $\frac{b \beta}{\varrho b + \sigma b + \tau(b + \beta)\sqrt{(\lambda' - 1)}}$     |
| Tertia RR  | $\frac{c \gamma}{\varrho c + \sigma c + \tau(c + \gamma)\sqrt{(\lambda'' - 1)}}$  | $\frac{c \gamma}{\varrho c + \sigma c + \tau(c + \gamma)\sqrt{(\lambda'' - 1)}}$  |
| Quarta SS  | $\frac{d \delta}{\varrho d + \sigma d + \tau(d + \delta)\sqrt{(\lambda''' - 1)}}$ | $\frac{d \delta}{\varrho d + \sigma d + \tau(d + \delta)\sqrt{(\lambda''' - 1)}}$ |

Cum iam lentis primae PP semidiameter aperturæ sit  $= x$ , et in qualibet lente vtriusque faciei eadem sit ratio, vt omnes radii per primam ingressi simul per reliquas transmittantur, apertura reliquarum sequentes limites superare debet:

Semid. aperturæ

Lentis secundae QQ  $> \frac{b}{a} x$

Lentis tertiæ RR  $> \frac{bc}{a\beta} x$

Lentis quartæ SS  $> \frac{bcd}{a\beta\gamma} x$

etc.

Tum vero posito breuitatis ergo  $\mu = 0, 938191$  et  $\nu = 0, 232692$  statuatur:

$$P = \mu \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{\nu}{a\alpha} \right)$$

$$Q = \mu \left( \frac{1}{b} + \frac{1}{\beta} \right) \left( \lambda' \left( \frac{1}{b} + \frac{1}{\beta} \right)^2 + \frac{\nu}{b\beta} \right)$$

$$R = \mu \left( \frac{1}{c} + \frac{1}{\gamma} \right) \left( \lambda'' \left( \frac{1}{c} + \frac{1}{\gamma} \right)^2 + \frac{\nu}{c\gamma} \right)$$

$$S = \mu \left( \frac{1}{d} + \frac{1}{\delta} \right) \left( \lambda''' \left( \frac{1}{d} + \frac{1}{\delta} \right)^2 + \frac{\nu}{d\delta} \right)$$

etc.

Y 2

Sit



Sit iam  $O$  distantia oculi post ultimam lentem, ita ut ab imagine postrema distet intervallo  $= l$ , comparatur magnitudo visa cum ea, qua idem obiectum in distantia fixa  $b$  nudo oculo cerneretur, sitque exponens multiplicationis  $= m$ ; et gradus claritatis  $= y$ , quibus positus erit pro quouis lentium numero ut sequitur.

I. Pro vnica lente  $O = \alpha + l$

1. Exponens multiplicationis  $m = \frac{\alpha b}{\alpha l}$  situ inuerso
2. Gradus claritatis  $y = l \cdot \frac{x}{\alpha}$  hinc  $my = \frac{bx}{\alpha}$
3. Semid : confusionis  $= \frac{\alpha}{\alpha l} \cdot x^3 P$

II. Pro duabus lentibus  $O = \epsilon + l$

1. Exponens multiplicationis  $m = \frac{\alpha \epsilon b}{\alpha b l}$  situ erecto
2. Gradus claritatis  $y = l \cdot \frac{bx}{\alpha \epsilon}$ , hinc  $my = \frac{bx}{\alpha}$
3. Semid : confusionis  $= \frac{\epsilon}{\alpha l} \cdot \frac{b}{\alpha} x^3 (\frac{\alpha \alpha}{b b} P + \frac{b b}{\alpha \alpha} Q)$

III. Pro tribus lentibus  $O = \gamma + l$

1. Exponens multiplicationis  $m = \frac{\alpha \epsilon \gamma b}{\alpha b c l}$  situ inuerso
2. Gradus claritatis  $y = l \cdot \frac{bcx}{\alpha \epsilon \gamma}$  hinc  $my = \frac{bx}{\alpha}$
3. Semidiameter confusionis

$$\frac{\gamma}{\alpha l} \cdot \frac{bc}{\alpha \epsilon} x^3 (\frac{\alpha \alpha \epsilon \epsilon}{b b c c} P + \frac{b b \epsilon \epsilon}{\alpha \alpha c c} Q + \frac{b b c c}{\alpha \alpha \epsilon \epsilon} R).$$

IV. Pro quatuor lentibus  $O = \delta + l$

1. Exponens multiplicationis  $m = \frac{\alpha \epsilon \gamma \delta b}{\alpha b c d l}$  situ erecto
2. Gradus claritatis  $y = l \cdot \frac{bcdx}{\alpha \epsilon \gamma \delta}$  hinc  $my = \frac{bx}{\alpha}$
3. Semidiameter confusionis :

$$\frac{\delta}{\alpha l} \cdot \frac{bcd}{\alpha \epsilon \gamma} x^3 (\frac{\alpha \alpha \epsilon \epsilon \gamma \gamma}{b b c c d d} P + \frac{b b \epsilon \epsilon \gamma \gamma}{\alpha \alpha c c d d} Q + \frac{b b c c \gamma \gamma}{\alpha \alpha \epsilon \epsilon d d} R + \frac{b b c c d d}{\alpha \alpha \epsilon \epsilon \gamma \gamma} S)$$

V.

V. Pro quinque lentibus  $O = \varepsilon + l$

1. Exponens multipl.  $m = \frac{\alpha \varepsilon \gamma \delta \varepsilon b}{a b c d e l}$  situ inuerso.
2. Gradus claritatis  $y = l \cdot \frac{b c d e x}{\alpha \varepsilon \gamma \delta \varepsilon}$  hinc  $my = \frac{b x}{a}$
3. Semidiameter confusionis

$$\frac{\varepsilon}{a l} \cdot \frac{b c d e}{\alpha \varepsilon \gamma \delta} x^3 \left\{ \begin{array}{l} + \frac{\alpha \alpha \varepsilon \varepsilon \gamma \gamma \delta \delta}{b b c c d d e e} P + \frac{b b \varepsilon \varepsilon \gamma \gamma \delta \delta}{\alpha \alpha c c d d e e} Q \\ + \frac{b b c c \gamma \gamma \delta \delta}{\alpha \alpha \varepsilon \varepsilon a d e e} R \\ + \frac{b b c c d d \delta \delta}{\alpha \alpha \varepsilon \varepsilon \gamma \gamma e e} S + \frac{b b c c d d e e}{\alpha \alpha \varepsilon \varepsilon \gamma \gamma \delta \delta} T \end{array} \right.$$

vnde non difficile erit has formulas ad lentes etiam plures continuare.

### COROLL. 1.

215. Lentes simplices adhibendo numeros  $\lambda, \lambda', \lambda'',$  etc. vnitatem minores accipi non possunt. Verum hic nihil obstat quo minus loco lentium simplicium lentes duplicatae, triplicatae vel etiam quadruplicatae in usum vocentur, quo pacto in his formulis numeri  $\lambda, \lambda', \lambda'',$  etc. non solum infra vnitatem diminui sed ad nihilum usque perducere poterunt. Tum autem constructio lentium harum multiplicatarum ex superiori capite peti et ad distantias determinatrices hic positas accommodari debet.

### COROLL. 2.

216. Veluti si pro lente prima PP debeat esse

Y 3

$\lambda =$



$\lambda = 0, 191827$ , hanc lentem ita ex duabus componi oportet, vt fit

Radius faciei

| Pro lente  | anterioris                                           | posterioris                                         |
|------------|------------------------------------------------------|-----------------------------------------------------|
| Priori . . | $\frac{2a\alpha}{(2\rho - \sigma)\alpha + \sigma a}$ | $\frac{2a\alpha}{(2\sigma - \rho)\alpha + \rho a}$  |
| Posteriori | $\frac{2a\alpha}{\rho\alpha + (2\sigma - \rho)a}$    | $\frac{2a\alpha}{\sigma\alpha + (2\rho - \sigma)a}$ |

### Coroll. 3.

217. Sin autem velimus, vt pro lente prima fit  $\lambda = 0, 042165$ , ea erit triplicanda, hoc modo

Radius faciei

| Pro lente  | anterioris                                                    | posterioris                                                   |
|------------|---------------------------------------------------------------|---------------------------------------------------------------|
| Priori     | $\frac{3a\alpha}{(3\rho - 2\sigma)\alpha + \sigma a}$         | $\frac{3a\alpha}{(3\sigma - 2\rho)\alpha + \rho a}$           |
| Media      | $\frac{3a\alpha}{(2\rho - \sigma)\alpha + (2\sigma - \rho)a}$ | $\frac{3a\alpha}{(2\sigma - \rho)\alpha + (2\rho - \sigma)a}$ |
| Posteriori | $\frac{3a\alpha}{\rho\alpha + (3\sigma - 2\rho)a}$            | $\frac{3a\alpha}{\sigma\alpha + (3\rho - 2\sigma)a}$          |

### Coroll. 4.

218. At si pro lente prima requiratur  $\lambda = -0, 010216$ , quadruplicata erit vtendum ita construenda:

Radius faciei

| Pro lente | anterioris                                                     | posterioris                                                    |
|-----------|----------------------------------------------------------------|----------------------------------------------------------------|
| Prima     | $\frac{4a\alpha}{(4\rho - 3\sigma)\alpha + \sigma a}$          | $\frac{4a\alpha}{(4\sigma - 3\rho)\alpha + \rho a}$            |
| Secunda   | $\frac{4a\alpha}{(\rho - 2\sigma)\alpha + (2\sigma - \rho)a}$  | $\frac{4a\alpha}{(3\sigma - 2\rho)\alpha + (2\rho - \sigma)a}$ |
| Tertia    | $\frac{4a\alpha}{(2\rho - \sigma)\alpha + (3\sigma - 2\rho)a}$ | $\frac{4a\alpha}{(2\sigma - \rho)\alpha + (3\rho - 2\sigma)a}$ |
| Quarta    | $\frac{4a\alpha}{\rho\alpha + (4\sigma - 3\rho)a}$             | $\frac{4a\alpha}{\sigma\alpha + (4\rho - 3\sigma)a}$           |

Coroll.

## Coroll. 5.

219. Simili autem modo lens secunda  $QQ$  ex suis distantibus determinatricibus  $b$  et  $\delta$  per multiplicationem erit construenda, numerus que ei respondens  $\lambda'$  debet esse vel  $0,191827$ , vel  $0,042165$  vel  $-0,010216$ ; quod idem de reliquis lentibus est intelligendum.

## Scholion I.

220. Alias lentium species hic nollem in praxi adhiberi, cum hae folae sine metu enormis erroris confici queant; tum vero etsi aliae ab his non admodum discrepantes fere aequo successu in praxin introduci possent, tamen quia discrimen non admodum est notabile, iis facile carere poterimus. Cum igitur sit  $\varrho = 0,190781$  et  $\sigma = 1,627401$ , ideoque

$$\varrho = 0,190781; \quad \sigma = 1,627401$$

$$2\varrho - \sigma = -1,245839; \quad 2\sigma - \varrho = 3,064021$$

$$3\varrho - 2\sigma = -2,682459; \quad 3\sigma - 2\varrho = 4,500641$$

$$4\varrho - 3\sigma = -4,119079; \quad 4\sigma - 3\varrho = 5,937261$$

easdem constructiones in numeris euolutis exhiberi conueniet.

I. Si igitur pro lente  $PP$  debeat esse  $\lambda = 1$ ,

ea erit simplex hoc modo construenda.

Radius faciei            anterioris            posterioris

$$\frac{a\alpha}{+0,190781\alpha + 1,627401\alpha}; \quad \frac{a\alpha}{+1,627401\alpha + 0,190781\alpha}$$

II.



II. Si pro lente PP debeat esse  $\lambda = 0,191827$

ea erit duplicata hoc modo construenda

Radius Faciei

| Pro lente  | anterioris                                    | posterioris                                   |
|------------|-----------------------------------------------|-----------------------------------------------|
| Priori     | $\frac{a\alpha}{-0,62251\alpha + 0,813700a}$  | $\frac{a\alpha}{+1,532510\alpha + 0,095390a}$ |
| Posteriori | $\frac{a\alpha}{+0,095310\alpha + 0,532010a}$ | $\frac{a\alpha}{+0,813700\alpha - 0,622519a}$ |

III. Si pro lente PP debeat esse  $\lambda = 0,042165$

ea erit triplicata hoc modo construenda

Radius Faciei

| Pro lente  | anterioris                                    | posterioris                                   |
|------------|-----------------------------------------------|-----------------------------------------------|
| Priori     | $\frac{a\alpha}{-0,894153\alpha + 0,542467a}$ | $\frac{a\alpha}{+1,500214\alpha + 0,063594a}$ |
| Media      | $\frac{a\alpha}{-0,415200\alpha + 1,021340a}$ | $\frac{a\alpha}{+1,021340\alpha - 0,415280a}$ |
| Posteriori | $\frac{a\alpha}{+0,063594\alpha + 1,500214a}$ | $\frac{a\alpha}{+0,542467\alpha - 0,894153a}$ |

IV. Si pro lente PP debeat esse  $\lambda = -0,010216$

ea erit quadruplicata hoc modo construenda

Radius Faciei

| Pro lente | anterioris                                    | posterioris                                   |
|-----------|-----------------------------------------------|-----------------------------------------------|
| Prima     | $\frac{a\alpha}{-1,029770\alpha + 0,406850a}$ | $\frac{a\alpha}{+1,484315\alpha + 0,047695a}$ |
| Secunda   | $\frac{a\alpha}{-0,670615\alpha + 0,766005a}$ | $\frac{a\alpha}{+1,125160\alpha - 0,311460a}$ |
| Tertia    | $\frac{a\alpha}{-0,311460\alpha + 1,125160a}$ | $\frac{a\alpha}{+0,766005\alpha - 0,670615a}$ |
| Quarta    | $\frac{a\alpha}{+0,047695\alpha + 1,484315a}$ | $\frac{a\alpha}{+0,406850\alpha - 1,029770a}$ |

Scholion

## Scholion 2.

221. Interim tamen si lentes desideremus, in quibus valor ipsius  $\lambda$  maior sit, quam hic est assumptus, constructio earum leui additione ex his ipsis formulis concinnari poterit. In binis scilicet fractionibus, quibus radii facierum cuiusque lentis designantur, alter denominator augeri alter vero diminui debet eadem quantitate, quae quantitas semper  $= \tau(a+\alpha)\sqrt{v}$  denotante  $v$  excessum valoris ipsius  $\lambda$  supra ante assumptum. Ita si esse debeat

I.  $\lambda = 1+v$ ; II.  $\lambda = 0,191827+v$ ; III.  $\lambda = 0,042165+v$   
vel IV.  $\lambda = -0,010216+v$

denominatores fractionum in Scholio praecedente traditarum pro quavis lente simplici alternatim sunt augendi et minuendi quantitate  $0,905133(a+\alpha)\sqrt{v}$ . Vnde si lens quadruplicata desideretur, pro qua sit praecise  $\lambda = 0$ , erit  $v = 0,010216$  et  $\tau\sqrt{v} = 0,091487$ ; hincque pro quavis lente simplici denominatorum alter augeri alter vero diminui debet hac quantitate

$0,091487\alpha + 0,091487a$

ex quo talis nascitur constructio huiusmodi lentis quadruplicatae pro qua est  $\lambda = 0$

Radius faciei

| Pro lente | anterioris                                    | posterioris                                   |
|-----------|-----------------------------------------------|-----------------------------------------------|
| Prima     | $\frac{a\alpha}{-1,121257\alpha + 0,315363a}$ | $\frac{a\alpha}{+1,575802\alpha + 0,139182a}$ |
| Secunda   | $\frac{a\alpha}{-0,762102\alpha + 0,674518a}$ | $\frac{a\alpha}{+1,216647\alpha - 0,219973a}$ |
| Tertia    | $\frac{a\alpha}{-0,402547\alpha + 1,033673a}$ | $\frac{a\alpha}{+0,857452\alpha - 0,579128a}$ |
| Quarta    | $\frac{a\alpha}{-0,043792\alpha + 1,352828a}$ | $\frac{a\alpha}{+0,458337\alpha - 0,938283a}$ |
| Tom. I.   | Z                                             | Sed                                           |



Sed hoc casu quaelibet lens adhuc alio modo construi potest: sic quarta ordine retrogrado exposita permutatis  $a$  et  $\alpha$  dabit alteram formam lentis primae.

### Supplementum III.

#### Ad Problem. 8.

Si lentes ratione refractionis discrepent, vt sit ratio refractionis pro prima lente  $=n$ , pro secunda  $=n'$ , pro tertia  $=n''$ ; inde neque in multiplicatione  $m$  neque in gradu claritatis quicquam mutatur; at vero in semidiametro confusionis valores litterarum P, Q, R sequenti modo immutari debent.

$$\begin{aligned}
 P &= \frac{n}{2(n-1)^2} \left\{ \left( \frac{n}{\alpha} + \frac{2}{k+v} \right) \left( \frac{1}{ia} + \frac{2}{k-v} \right)^2 \right. \\
 &\quad \left. + \left( \frac{n}{\alpha} - \frac{2}{k-v} \right) \left( \frac{i}{\alpha} - \frac{2}{k+v} \right)^2 \right\} \\
 Q &= \frac{n'}{2(n'-1)^2} \left\{ \left( \frac{n'}{b} + \frac{2}{k'+v'} \right) \left( \frac{1}{i'b} + \frac{2}{k'-v'} \right)^2 \right. \\
 &\quad \left. + \left( \frac{n'}{b} - \frac{2}{k'-v'} \right) \left( \frac{i'}{b} - \frac{2}{k'+v'} \right)^2 \right\} \\
 R &= \frac{n''}{2(n''-1)^2} \left\{ \left( \frac{n''}{c} + \frac{2}{k''+v''} \right) \left( \frac{1}{i''c} + \frac{2}{k''-v''} \right)^2 \right. \\
 &\quad \left. + \left( \frac{n''}{c} - \frac{2}{k''-v''} \right) \left( \frac{i''}{c} - \frac{2}{k''+v''} \right)^2 \right\} \\
 S &= \frac{n'''}{2(n'''-1)^2} \left\{ \left( \frac{n'''}{d} + \frac{2}{k''' + v'''} \right) \left( \frac{1}{i'''d} + \frac{2}{k''' - v'''} \right)^2 \right. \\
 &\quad \left. + \left( \frac{n'''}{d} - \frac{2}{k''' - v'''} \right) \left( \frac{i'''}{d} - \frac{2}{k''' + v'''} \right)^2 \right\} \\
 &\quad \text{etc.}
 \end{aligned}$$

#### Ad Problem. 9.

Si lentes ratione refractionis discrepent, loco litterarum  $\varrho$ ,  $\sigma$ ,  $\tau$  in radiis facierum scribi oportet  
pro

pro secunda lente  $\varrho'$ ,  $\sigma'$ ,  $\tau'$ ; pro tertia  $\varrho''$ ,  $\sigma''$ ,  $\tau''$ , etc.  
tum vero expressiones pro multiplicatione et gradu  
claritatis nulla mutatione egent; pro confusione au-  
tem notari oportet, fore

$$P = \mu \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( \lambda \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{\nu}{a\alpha} \right)$$

$$Q = \mu' \left( \frac{1}{b} + \frac{1}{\beta} \right) \left( \lambda' \left( \frac{1}{b} + \frac{1}{\beta} \right)^2 + \frac{\nu'}{b\beta} \right)$$

$$R = \mu'' \left( \frac{1}{c} + \frac{1}{\gamma} \right) \left( \lambda'' \left( \frac{1}{c} + \frac{1}{\gamma} \right)^2 + \frac{\nu''}{c\gamma} \right)$$

etc.





# CAPUT V.

## DE

### CAMPO APPARENTE OBLIQUE LOCO MAXIME IDONEO.

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#### Problema I.

222.

Tab. II.  
Fig. 12.

**S**i ex obiecti puncto extra axem sumto  $\varepsilon$  radius quicunque  $\varepsilon M$  per lentem  $PP$  transmittatur, definire eius concursum cum axe  $O$ .

#### Solutio.

Sit  $F\zeta$  imago principalis per lentem proiecta, a diffusionem enim hic mentem abstrahimus, ac ponamus pro lente eius distantias determinatrices  $AE = a$ ,  $aF = \alpha$ , eius crassitiem  $Aa = v$ , et quantitatem arbitrariam  $= k$ , sitque breuitatis ergo  $\frac{k-v}{k+v} = i$ . His positis si puncti  $\varepsilon$  imago cadat in  $\zeta$ , voceturque  $E\varepsilon = z$ , erit  $F\zeta = \frac{1}{i} \cdot \frac{\alpha z}{a}$  at pro puncto lentis  $M$  statuatur eius distantia ab axe  $AM = x$ ; ac supra ostendimus si radius a puncto  $E$  per punctum  $M$  transmitteretur, eum ita per  $m$  ad punctum  $F$  progressurum esse, ut foret  $am = ix$ . Radius igitur  $\varepsilon M$  per punctum  $N$  ad  $\zeta$  feretur, ita ut, cum obliquitas radio-

rum

rum in facies lentis incidentium perpetuo valde parva statuatur, fit proxime angulus  $EM\epsilon$  ad  $NMm$  vt  $n$  ad  $1$  posito  $n = \frac{31}{20}$ . Hinc erit  $E\epsilon$  ad  $mN$  in ratione composita istorum angulorum, et distantiarum  $AE$  ad  $Aa$ , seu  $E\epsilon : mN = na : v$ , vnde fit  $mN = \frac{vz}{na}$ , ideoque  $aN = ix - \frac{vz}{na}$ . Iam vero radius a puncto  $N$  recta ad  $\zeta$  pergit, et propterea axem ita in  $O$  secabit, vt fit  $aN + F\zeta : aF = aN : aO$  sicque

$$aO = \left( i\alpha x - \frac{\alpha vz}{na} \right) : \left( ix - \frac{vz}{na} + \frac{\alpha z}{ia} \right) = \frac{n i i \alpha \alpha x - i \alpha v z}{n i i \alpha x - i v z + n \alpha z}$$

$$\text{et } FO = \frac{n \alpha \alpha z}{n i i \alpha x - i v z + n \alpha z}$$

Si  $x$  fit semidiameter aperturæ lentis in facie anteriori, hæc intersectio  $O$  respondet casui, quo radius ab  $\epsilon$  per lentis terminum summum transeat: sin autem is per terminum imum transmittatur, sumto  $x$  negatiuo fiet

$$FO = \frac{n \alpha \alpha z}{-n i i \alpha x - i v z + n \alpha z}$$

At si radius ex  $\epsilon$  per centrum lentis  $A$  transeat, punctum intersectionis ita cadet in  $O$  vt fit

$$FO = \frac{n \alpha \alpha}{n \alpha - i v}$$

### Coroll. I.

223. Si igitur  $x$  denotet semidiametrum aperturæ lentis in facie anteriori  $PMA P$ , vt omnes radii a puncto  $\epsilon$  in hanc faciem incidentes per lentem transmittantur, necesse est, vt faciei posterioris

Z 3

PNaP



$PNaP$  semidiameter sit maior quam  $\pm ix - \frac{vz}{na}$ , sumto  $x$  tam negatiuo quam positiuo. Vnde hic semidiameter minor esse nequit, quam  $ix + \frac{vz}{na}$ .

### Coroll. 2.

224. Si apertura in facie anteriori euanescat, vt sit  $x=0$  radii a puncto  $\varepsilon$ , cuius ab axe distantia  $E\varepsilon=z$ , per lentem non transmittentur, nisi in facie posteriori semidiameter aperturae sit  $=\frac{vz}{na}$  vel maior. Vnde patet quo maior fuerit lentis crassities  $v$  eo maiori apertura in facie posteriori esse opus.

### Coroll. 3.

225. Vicissim ergo si detur lentis apertura in facie posteriori, cuius semidiameter sit  $=\frac{vz}{na}$ ; inde simul in obiecto extremum punctum  $\varepsilon$  determinatur, a quo radius in centrum lentis  $A$  incidens per eam transmittatur.

### Coroll. 4.

226. Hic autem radius per  $A$  immissus post transitum cum axe in  $O$  occurret, vt ob  $x=0$  sit interuallum  $FO = \frac{n\alpha\alpha}{n\alpha - iv}$ . seu  $aO = \frac{-i\alpha v}{n\alpha - iv}$ . Nisi ergo oculus in hoc axis loco teneatur, radius transmissus non in oculum ingredietur; si quidem apertura pupillae vt infinite parua spectetur.

Coroll.

## Coroll. 5.

227. At si semidiameter pupillae ponatur  $= \omega$ , oculus etiam in  $o$  positus illum radium excipiet, si fuerit  $o \omega = \omega$ . Cum autem casu  $x = o$  fit

$$a N \left( \frac{vz}{na} \right) : a O \left( \frac{i\alpha v}{n\alpha - iv} \right) = o \omega (\omega) : O o \text{ erit}$$

$$O o = \frac{n i a \alpha \omega}{z (n \alpha - i v)}.$$

Quod interuallum cum aeque posituum ac negativum accipi possit, pro loco oculi  $o$  habebimus

$$a o = - \frac{i \alpha v z + n i a \alpha \omega}{z (n \alpha - i v)} = - \frac{i \alpha (v z + n a \omega)}{z (n \alpha - i v)}.$$

## Scholion.

228. In hac tractatione, vbi campum apparentem et locum oculi idoneum inuestigamus, tam aperturam lentis obiectivae in facie anteriori quam pupillae amplitudinem pro nihilo habebimus, vt quaestiones obtineamus determinatas. Quare horum elementorum, quae in instrumentis dioptricis ad visionem accommodatis maximi sunt momenti, sequentes definitiones constituemus.

## Definitio I.

229. Campus apparens est spatium in obiecto, ex cuius singulis punctis radii in centrum lentis obiectivae incidentes per reliquas lentes omnes transmittuntur. Quod spatium cum sit circulare, eius radius vocatur semidiameter campi apparentis.

Coroll.



## Coroll. 1.

230. Si ergo  $E\varepsilon = z$  fuerit semidiameter campi apparentis, erit  $\varepsilon$  punctum obiecti extremum, sed ab axe maxime remotum, ex quo adhuc radii in centrum A lentis obiectivae incidentes per omnes lentes transmittuntur.

## Coroll. 2.

231. Determinatur igitur magnitudo campi apparentis per aperturam sequentium facierum refringentium, ac fortasse per aperturam vnius, si scilicet radii a puncto quodam magis remoto quam  $\varepsilon$  venientes nullum transitum per eam inuenirent, etiam si per reliquas facies transmitterentur.

## Scholion.

232. Si radii ex solo puncto E, quod est centrum campi apparentis, considerentur, ii quidem qui in A incidunt, quoniam perpetuo secundum axem progrediuntur, per omnes reliquas facies refringentes, quantumvis parua fuerit earum apertura, certe transmittentur; hincque campus apparens nunquam penitus evanescere potest. At quo magis punctum  $\varepsilon$  ab axe distans accipitur, ut radii ab eo per omnes facies transmittantur, eo maior earum apertura requiritur, quae cum ab earum curvatura pendeat, neque certum limitem superare debeat, hinc ultimum punctum  $\varepsilon$ , unde radii etiam nunc transmittuntur, ac propterea semidiameter campi apparentis determinatur.

natur. In sequentibus quidem propositionibus campum apparentem seu eius semidiametrum  $E\varepsilon=z$  vt datum assumam, et quanta esse debeat cuiusque faciei apertura, inuestigabo: hinc enim facile vicissim, si quaeque apertura fuerit cognita, campum apparentem ipsum definire licebit. Ceterum in hac definitione assumpsi aperturam primae faciei esse euanescentem: ex quo facile intelligitur ea aucta etiam campum aliquantum extendi oportere; verum hoc augmentum postea in lucrum cedit, quod cum nunquam soleat esse notabile, eius rationem hic non habendam censui; quemadmodum etiam in sequente definitione aperturae pupillae rationem non habebo.

### Definitio 2.

233. Locus oculi idoneus est id punctum in axe, in quo radii ab extremitate campi apparentis per lentes transmissi axem intersecant. Oculus scilicet in hoc loco constitutus totum campum apparentem conspiciet.

### Coroll. I.

234. Hinc igitur idoneus locus oculo assignabitur, si illa intersectio radiorum extremorum cum axe post lentem vltimam cadat, sin autem haec intersectio ante lentem vltimam reperiatur, fieri nequit vt oculus in eo loco teneatur neque propterea totum campum contueri poterit.



## Coroll. 2.

235. At si ista intersectio pone lentem ultimam cadat, oculus totum campum apparentem perspiciet, etiamsi pupilla maxime esset contracta: neque tamen ob maiorem pupillae amplitudinem maiorem campum percipere valet.

## Coroll. 3.

236. Verum ob amplitudinem pupillae hoc commodi assequimur, ut oculus etiamsi extra locum idoneum constituitur, dummodo distantia non sit nimis magna, tamen totum campum apparentem conspicere possit: id quod egregie usu veniet iis casibus, quibus locus idoneus oculi ante faciem refringentem extremam cadit. Tum enim fieri poterit, ut oculus huic faciei immediate applicatus tamen totum campum percipiat.

## Scholion.

237. Quando hic de visione loquor, id ita in genere est interpretandum, ut a puncto viso radius in oculum ingrediatur, neque hic curo, utrum visio sit distincta nec ne? in sequentibus enim docebitur, quomodo lentes disponi conveniat, ut oculus in loco idoneo positus etiam in iusta ab ultima imagine distantia reperiatur, quo visio distincta reddatur. Hic igitur sine ullo respectu ad visionem distinctam habito, eum oculo locum assigno, ubi ab omnibus punctis

ctis in campo apparente contentis radios recipiat, et quoniam singula momenta, quae ad visionem pertinent, seorsim expediri conuenit; hic etiam non ad spatium diffusionis respicio, quod quidem semper per se euanescit, si apertura faciei primae euanescebat statuatur.

### Problema 2.

238. Si obiectum per unicam lentem aspiciatur, determinare tam campum apparentem, quam locum idoneum oculi.

Tab. III.  
Fig. 12.

### Solutio.

Sint ut in problemate superiori distantiae determinatrices huius lentis, scilicet distantia obiecti ante lentem  $AE = a$ , et imaginis post lentem  $aF = a$ , tum vero lentis crassities  $Aa = v$ , et distantia arbitraria constructionem lentis plene determinans  $= k$ . Deinde ponamus semidiametrum campi apparentis  $E\varepsilon = z$ , ita ut posita faciei anterioris  $PAP$  apertura infinite parua etiamnum a puncto  $\varepsilon$  radii per lentem transmittantur. Sit porro breuitatis gratia  $\frac{k-v}{k+v} = i$ , atque supra demonstrauius campum apparentem ad  $\varepsilon$  usque extendi, si pro facie posteriori  $PaP$  fuerit semidiameter aperturae  $aN = \frac{vz}{na}$  existente  $n = \frac{31}{25}$ ; perinde enim est, siue hic semidiameter affirmatiue accipiat siue negatiue. Hinc ergo vicissim si semidiameter huius aperturae ponatur  $= a$ , erit semidiameter campi apparentis  $E\varepsilon = z = \frac{na}{v} a$ .

Aa 2

Quod



Quod ad locum idoneum oculi attinet, qui fit in O, quoniam inuenimus  $FO = \frac{n\alpha\alpha}{n\alpha - iv}$ , erit distantia  $aO = \frac{-i\alpha v}{n\alpha - iv}$ ; ideoque negatiua, nisi sit vel  $i$  numerus negatiuus, vel  $iv > n\alpha$ . Sin autem haec distantia  $aO$  fuerit positiua, oculus in O positus totum campum apparentem perspiciet.

### C o r o l l. 1.

239. Si crassities lentis  $v$  euanescat, ob  $z = \frac{n\alpha}{v}$ , campus apparens euadet infinitus, seu potius indeterminatus; distantia vero  $aO$  euanescet. Oculus igitur lenti immediate applicatus tantum spatium conspiciet, quantum per propriam indolem complecti valebit.

### C o r o l l. 2.

240. Sin autem ob lentis crassitiem  $Aa = v$  distantia  $aO = \frac{-i\alpha v}{n\alpha - iv}$  prodeat positiua, oculus in O positus totum campum apparentem aspicere valebit, seu in obiecto spatium circulare spectabit, cuius semidiameter  $E\varepsilon = z = \frac{n\alpha}{v} \alpha = \frac{3}{2} \frac{\alpha^2}{v}$ , quod ergo eo erit maius, quo tenuior fuerit lens.

### C o r o l l. 3.

241. At si distantia  $aO$  resultet negatiua, oculum in loco idoneo constitui non licet, quoniam is necessario post lentem teneri debet. Vbicunque autem is post lentem collocetur, non vniuersum campum apparentem contuebitur, sed tantum eius partem, et  
quidem

quidem eo minorem, quo magis post lentem removeatur, propterea quod hoc modo magis a loco idoneo recedit.

#### Coroll. 4.

242. Hoc igitur casu conveniet oculum immediate ad faciem lentis posteriorem applicare, quo situ eatenus tantum radios accipiet, quatenus pupilla patet; sicque campus visus ab apertura pupillae pende-  
bit; quae si esset nulla etiam campus apparens evanesceret.

#### Coroll. 5.

243. Hinc patet si pupilla excedat aperturam faciei  $PaP$  seu si sit  $\omega > a$ , denotante  $\omega$  semidiameter pupillae, quia tum oculus huic faciei applicatus omnes radios transmissos recipit, eum totum campum apparentem esse visurum. Sin autem sit  $\omega < a$ , partem tantum totius campi perspiciet, cuius semidiameter erit  $= \frac{n a}{v} \omega = \frac{3 + a \omega}{2 + v}$ , scilicet non maiorem quam si apertura faciei  $PaP$  aequalis esset pupillae.

#### Scholion.

244. Interim tamen si hoc casu postremo apertura faciei lentis  $PaP$ , cui oculus est applicatus maior sit quam pupilla, nihil obstat quominus ea successive totam aperturam peragret, sicque pedetentim totum campum apparentem con-



spicere poterit, etiamsi eum simul contueri non valeat. Ceterum notandum est, si etiam faciei anteriori apertura tribuatur, inde campum apparentem aliquantum augeri, sed partes obiecti vltiores, quia non per medium lentis radios transmittunt, obscuriores apparebunt, vnde merito a campo apparente excluduntur. At si faciei anteriori tribuatur apertura, cuius semidiameter  $= x$ , vt facies posterior omnes radios in illam incidentes transmittat, eius apertura tanto maior esse debet, secundum regulas supra traditas; scilicet eius semidiametrum esse oportet  $= ix + a$  seu  $= ix + \frac{vz}{na}$ .

### Problema. 3.

245. Si instrumentum dioptricum duabus constet lentibus, definire campum apparentem, et locum oculi idoneum.

### Solutio.

Tab. III.

Fig. 13.

Sit pro his lentibus vt haecenus:

pro PP:  $AE = a, aF = \alpha, Aa = v$ ; dist. arb:  $= k$ , et  $\frac{k-v}{k+v} = i$

pro QQ:  $BF = b, bG = \beta, Bb = v'$ ; dist. arb:  $= k'$  et  $\frac{k'-v'}{k'+v'} = i'$

ac posito semidiametro campi apparentis  $E\varepsilon = z$ , in imaginibus erit  $F\zeta = \frac{1}{i} \cdot \frac{\alpha}{a} z$  et  $G\eta = \frac{1}{i'} \cdot \frac{\alpha\beta}{a'b} z$ . Consideretur iam radius a puncto  $\varepsilon$  in centrum lentis primae A incidens et per lentes transmissus qui vtique per extremitates imaginum  $\zeta$  et  $\eta$  transibit.

Iam

Iam ex problemate praecedente patet esse  $aN = \frac{v}{na}$ ,  
vnde in altera lente punctum  $M'$  ita definietur,  
vt fit

$$F\zeta - aN : aF = BM' - F\zeta : BF \text{ siue}$$

$$BM' = F\zeta + \frac{BF(F\zeta - aN)}{aF} = \frac{aBF\zeta - BF \cdot aN}{aF}$$

vnde fit

$$BM' = \frac{1}{i} \cdot \frac{\alpha + b}{a} z - \frac{bvz}{na\alpha}$$

Nunc punctum  $N'$  hinc perinde definietur, atque  
ex problemate primo ex puncto  $M$  determinabatur  
punctum  $N$ ; erit quippe

$$bN' = i' \cdot BM' - \frac{v'}{n'b} \cdot F\zeta \text{ ideoque}$$

$$bN' = \frac{i'}{i} \cdot \frac{\alpha + b}{a} z - \frac{i'bv}{na\alpha} z - \frac{1}{i} \cdot \frac{\alpha v'}{n'ab} z$$

Hinc autem punctum  $O$ , vbi est locus oculi idoneus  
facile assignabitur erit enim  $bN' + G\eta : bG = bN' : bO$ ,  
indeque

$$bO = \frac{bG \cdot bN'}{bN' + G\eta} = \frac{\frac{i'}{i} \cdot \frac{\alpha + b}{a} z - \frac{i'bv}{na\alpha} z - \frac{1}{i} \cdot \frac{\alpha v'}{n'ab} z}{\frac{i'}{i} \cdot \frac{\alpha + b}{a} z - \frac{i'bv}{na\alpha} z - \frac{1}{i} \cdot \frac{\alpha v'}{n'ab} z + \frac{1}{n'} \cdot \frac{\alpha}{ab} z} \cdot bG$$

Quod si ergo ponamus semidiametrum aperturæ

$$\text{pro lente PP} \begin{cases} \text{faciei anterioris} = \mathcal{A} = x \\ \text{faciei posterioris} = a \end{cases}$$

$$\text{pro lente QQ} \begin{cases} \text{faciei anterioris} = \mathcal{B} \\ \text{faciei posterioris} = b \end{cases}$$

habe-



habebimus

$$\mathfrak{A} = 0$$

$$a = \frac{v}{n} z$$

$$\mathfrak{B} = \left( \frac{1}{i} \cdot \frac{\alpha+b}{a} - \frac{b}{na\alpha} \right) z$$

$$b = \left( \frac{i'}{i} \cdot \frac{\alpha+b}{a} - \frac{i'bv}{na\alpha} - \frac{1}{i} \cdot \frac{\alpha v'}{na b} \right) z$$

ac si distantia oculi post lentem  $Q Q$  ponatur  
 $bO = 0$  erit

$$O = \frac{b}{b + \frac{1}{i'} \cdot \frac{\alpha b}{a}} z \cdot \xi$$

### Coroll. 1.

246. Si ambarum lentium crassities evanescat  
 erit  $v = 0$ ,  $v' = 0$ , et  $i = i' = 1$ ; quo ergo casu nostrae  
 formulae in sequentes abibunt:

$$\mathfrak{A} = 0; a = 0; \mathfrak{B} = \frac{\alpha+b}{a} z \text{ et } b = \frac{\alpha+b}{a} z$$

### Coroll. 2.

247. Datis ergo vicissim aperturis lentium ex  
 aequationum traditarum ea, pro qua quantitas  $z$   
 minimum valorem adipiscitur, definietur campus  
 adparens.

### Coroll. 3.

248. Si igitur crassities lentium evanescat,  
 campus apparens ex apertura lentis posterioris facilli-  
 me determinatur. Erit enim  $z = \frac{a \mathfrak{B}}{\alpha+b}$ ; id quod intel-  
 ligendum est, si distantia  $bO = 0$  fuerit positiva ocu-  
 lusque in  $O$  collocetur.

Coroll.

## Coroll. 4.

249. Sin autem distantia  $bO=O$  prodeat negatiua oculusque vltimae lenti immediate applicetur, tum eius apertura plus non praestat quam si amplitudini pupillae esset aequalis. Quare si  $b$  maior fuerit semidiametro pupillae  $\omega$  loco  $b$  scribatur  $\omega$  et ex vltima aequatione verus valor ipsius  $z$  elicietur nisi forte ex aliqua reliquarum aequationum adhuc minor valor pro  $z$  esset proditurus.

## Problema 4.

250. Si instrumentum dioptricum ex tribus constet lēntibus determinare cum campum apparentem, tum locum oculi idoneum. Tab. III.  
Fig. 14.

## Solutio.

Existentibus imaginibus per has lentes successiue repraesentatis in  $F\zeta$ ,  $G\eta$  et  $H\theta$ , obiecto vero ipso in  $E\epsilon$ , ponamus vt haecenus:

Pro Lente Distantias crassitiem dist. arb: et

Prima PP ..  $AE=a$ ;  $aF=\alpha$ ;  $Aa=v$ ;  $k$ ;  $\frac{k-v}{k+v}=i$

Secunda QQ ..  $BF=b$ ;  $bG=\beta$ ;  $Bb=v'$ ;  $k'$ ;  $\frac{k'-v'}{k'+v'}=i'$

Tertia RR ..  $CG=c$ ;  $cH=\gamma$ ;  $Cc=v''$ ;  $k''$ ;  $\frac{k''-v''}{k''+v''}=i''$

Semidiametros vero aperturarum

Pro Lente

PP faciei  $\left\{ \begin{array}{l} \text{anterioris} = \mathfrak{A} = o \\ \text{posterioris} = a \end{array} \right.$

Tom. I.

B b

QQ



$$QQ \text{ faciei} \begin{cases} \text{anterioris} = \mathfrak{B} \\ \text{posterioris} = \mathfrak{b} \end{cases}$$

$$RR \text{ faciei} \begin{cases} \text{anterioris} = \mathfrak{C} \\ \text{posterioris} = \mathfrak{c} \end{cases}$$

Tum vero sit semidiameter campi apparentis  $E\varepsilon = z$ ,  
supraque ostendimus fore :

$$F\zeta = \frac{1}{i} \frac{\alpha}{a} z; \quad G\eta = \frac{1}{i'} \frac{\alpha\varepsilon}{ab} z; \quad H\theta = \frac{1}{i'i'} \frac{\alpha\varepsilon\gamma}{abc} z.$$

His positis habebimus  $aN = a = \frac{vz}{na}$ ; vnde si per F  
recta ipsi  $NM'$  parallela ducta intelligatur erit

$$F\zeta - aN : aF = BM' - F\zeta : BF \text{ siue } BM' = \frac{aB \cdot F\zeta - BF \cdot aN}{a \cdot F}$$

ideoque

$$BM' = \frac{\alpha + b}{\alpha} \cdot \frac{1}{i} \cdot \frac{\alpha}{a} z - \frac{bvz}{na\alpha} \text{ seu}$$

$$\mathfrak{B} = \frac{1}{i} \cdot \frac{\alpha + b}{a} z - \frac{bv}{na\alpha} z.$$

Porro vero est ex problemate primo  $bN' = i' \cdot BM' - \frac{v'}{nb} \cdot F\zeta$ ,

$$\text{hincque } \mathfrak{b} = \frac{i'}{i} \cdot \frac{\alpha + b}{a} z - \frac{i'bv}{na\alpha} z - \frac{1}{i} \cdot \frac{\alpha v'}{nab} z;$$

simili modo per G ducta intelligatur recta ipsi  $N'M''$   
parallela, eritque  $bG : bN' + G\eta = CG : CM'' - G\eta$   
siue

$$CM'' = \mathfrak{C} = G\eta + \frac{c}{\varepsilon} (\mathfrak{b} + G\eta) = \frac{\varepsilon + c}{\varepsilon} \cdot G\eta + \frac{c}{\varepsilon} \mathfrak{b},$$

vnde fit

$$\mathfrak{C} = \frac{1}{i'} \cdot \frac{\alpha(\varepsilon + c)}{ab} z + \frac{i'}{i} \cdot \frac{c(\alpha + b)}{a\varepsilon} z - \frac{i'bcv}{na\alpha\varepsilon} z - \frac{1}{i} \cdot \frac{\alpha c v'}{nab\varepsilon} z.$$

Deinde

Deinde ex  $CM''$  ita definitur  $cN''$  per problema primum, ut fit

$$cN'' = i'' \cdot CM'' - \frac{v''}{nc} \cdot G\eta \text{ seu } c = i'' \cdot \mathfrak{C} - \frac{1}{i i'} \cdot \frac{\alpha \mathfrak{E} v''}{nabc} z$$

hincque

$$c = \frac{i''}{i i'} \cdot \frac{\alpha(\mathfrak{E} + c)}{ab} z + \frac{i' i''}{i} \cdot \frac{c(\alpha + b)}{a\mathfrak{E}} z - i' i'' \cdot \frac{bcv'}{na\alpha\mathfrak{E}} z - \frac{i''}{i} \cdot \frac{\alpha cv'}{nab\mathfrak{E}} z - \frac{1}{i i'} \cdot \frac{\alpha \mathfrak{E} v''}{nabc} z$$

Isti ergo valores sequenti modo determinantur:

$$\mathfrak{A} = 0$$

$$\mathfrak{a} = \frac{v}{na} \cdot E\epsilon$$

$$\mathfrak{B} = \frac{\alpha + b}{\alpha} \cdot F\zeta - \frac{b}{\alpha} \cdot \mathfrak{a}$$

$$\mathfrak{b} = i' \cdot \mathfrak{B} - \frac{v'}{nb} \cdot F\zeta$$

$$\mathfrak{C} = \frac{\mathfrak{E} + c}{\mathfrak{E}} \cdot G\eta + \frac{c}{\mathfrak{E}} \cdot \mathfrak{b}$$

$$c = i'' \cdot \mathfrak{C} - \frac{v''}{nc} \cdot G\eta$$

Cum iam punctum  $O$  praebeat locum oculi iustum; si ponamus  $cO = O$ , erit  $cN'' + H\theta : cH = cN''$ ;  $cO$  unde reperitur

$$O = \frac{\gamma c}{c + H\theta}, \text{ vel } \frac{O}{c} = \frac{1}{\gamma} + \frac{1}{i i' i''} \cdot \frac{\alpha \mathfrak{E}}{ab} \cdot \frac{z}{c}$$

### COROLL. I.

251. Ex his aequationibus sequitur fore

$$\frac{\mathfrak{B}}{b} + \frac{\mathfrak{a}}{\alpha} = \frac{\alpha + b}{\alpha b} \cdot F\zeta = \frac{1}{i} \left( \frac{\alpha}{b} + 1 \right) \frac{z}{a} \text{ et}$$

$$\frac{\mathfrak{C}}{c} - \frac{\mathfrak{b}}{\mathfrak{E}} = \frac{\mathfrak{E} + c}{\mathfrak{E} c} \cdot G\eta = \frac{1}{i i'} \left( \frac{\alpha \mathfrak{E}}{bc} + \frac{\alpha}{b} \right) \frac{z}{a}$$

hincque porro:

$$\frac{i\mathfrak{a}}{\alpha} + \frac{i\mathfrak{B}}{b} + \frac{i i' \mathfrak{b}}{\mathfrak{E}} - \frac{i i' \mathfrak{C}}{c} = \left( 1 - \frac{\alpha \mathfrak{E}}{bc} \right) \frac{z}{a}$$

B b 2

Coroll.



## Coroll. 2.

252. Si ergo crassities lentium euanescat, cum sit  $a=0$ ,  $b=B$  et  $c=C$ , definitio campi apparentis reducitur ad has duas aequationes.

$$\text{I. } B \cdot \frac{1}{b} = \left(1 + \frac{a}{b}\right) \frac{z}{a}$$

$$\text{II. } B \left(\frac{1}{b} + \frac{1}{c}\right) - C \cdot \frac{1}{c} = \left(1 - \frac{ac}{bc}\right) \frac{z}{a}$$

vnde minor valor ipsius  $z$  praebet semidiametrum campi apparentis.

## Coroll. 3.

253. Deinde si distantia  $cO=O$  prodeat negativa, vt oculus cogamur lenti vltimae immediate applicare, pro  $c$  scribi oportet semidiametrum pupillae  $\omega$ , et ex vltima aequatione definietur semidiameter campi  $z$ , nisi forte ex alia aequatione adhuc minor valor pro  $z$  prodeat.

## Problema 5.

254. Si instrumentum Dioptricum ex quatuor lentibus super eodem axe dispositis constet, determinare cum campum apparentem, tum locum oculi idoneum.

## Solutio.

Tab. III.  
Fig. 15.

Existente obiecto  $E\varepsilon=z$ , sint imagines per lentes successiue repraesentatae  $F\zeta$ ,  $G\eta$ ,  $H\theta$ , et  $I\iota$ , ponamusque

musque pro lentium singularum determinatione vt  
hactenus :

Pro lente distantias crassitiem dist. arb. et

|         |    |         |              |            |               |                                                                                  |
|---------|----|---------|--------------|------------|---------------|----------------------------------------------------------------------------------|
| Prima   | PP | $AE=a;$ | $aF=\alpha;$ | $Aa=v;$    | $\dots k;$    | $\left  \begin{array}{l} \frac{k-v}{k+v}=i \end{array} \right $                  |
| Secunda | QQ | $BF=b;$ | $bG=\xi;$    | $Bb=v';$   | $\dots k';$   | $\left  \begin{array}{l} \frac{k'-v'}{k'+v'}=i' \end{array} \right $             |
| Tertia  | RR | $CG=c;$ | $cH=\gamma;$ | $Cc=v'';$  | $\dots k'';$  | $\left  \begin{array}{l} \frac{k''-v''}{k''+v''}=i'' \end{array} \right $        |
| Quarta  | SS | $DH=d;$ | $dI=\delta;$ | $Dd=v''';$ | $\dots k''';$ | $\left  \begin{array}{l} \frac{k'''-v'''}{k''' + v'''}=i''' \end{array} \right $ |

Semidiametri vero aperturarum sint :

Pro Lente

Prima PP faciei  $\left\{ \begin{array}{l} \text{anterioris} = \mathcal{A} = \circ \\ \text{posterioris} = a \end{array} \right.$

Secunda QQ faciei  $\left\{ \begin{array}{l} \text{anterioris} = \mathcal{B} \\ \text{posterioris} = b \end{array} \right.$

Tertia RR faciei  $\left\{ \begin{array}{l} \text{anterioris} = \mathcal{C} \\ \text{posterioris} = c \end{array} \right.$

Quarta SS faciei  $\left\{ \begin{array}{l} \text{anterioris} = \mathcal{D} \\ \text{posterioris} = d \end{array} \right.$

Si iam  $E\varepsilon=z$  exhibeat semidiametrum campi appa-  
rentis erit, vt iam supra ostendimus :

$$F\zeta = \frac{1}{i} \cdot \frac{\alpha}{a} z; \quad G\eta = \frac{1}{i'} \cdot \frac{\alpha\xi}{ab} z; \quad H\theta = \frac{1}{i''} \cdot \frac{\alpha\xi\gamma}{abc} z;$$

$$\text{et } I\iota = \frac{1}{i'''} \cdot \frac{\alpha\xi\gamma\delta}{abcd} z.$$

Bb 3

Quod



Quod si ratiocinium nunc ut ante instituamus, obtinebimus sequentes aequationes:

$$A=0;$$

$$a = \frac{v}{na} \cdot E \varepsilon$$

$$B = (1 + \frac{b}{\alpha}) F \zeta - \frac{b}{\alpha} a:$$

$$b = i' B - \frac{v'}{nb} \cdot F \zeta$$

$$C = (1 + \frac{c}{\beta}) G \eta + \frac{c}{\beta} b:$$

$$c = i'' C - \frac{v''}{nc} \cdot G \eta$$

$$D = (1 + \frac{d}{\gamma}) H \theta + \frac{d}{\gamma} c:$$

$$d = i''' D - \frac{v'''}{nd} \cdot H \theta$$

Ex quarum ordine priori consequimur:

$$\frac{B}{b} + \frac{a}{\alpha} = (\frac{1}{\alpha} + \frac{1}{b}) F \zeta = \frac{1}{i} (1 + \frac{\alpha}{b}) \frac{z}{a}$$

$$\frac{c}{c} - \frac{b}{\beta} = (\frac{1}{\beta} + \frac{1}{c}) G \eta = \frac{1}{i'} (\frac{\alpha}{b} + \frac{\alpha \beta}{b c}) \frac{z}{a}$$

$$\frac{D}{d} - \frac{c}{\gamma} = (\frac{1}{\gamma} + \frac{1}{d}) H \theta = \frac{1}{i'' i'''} (\frac{\alpha \beta}{b c} + \frac{\alpha \beta \gamma}{b c d}) \frac{z}{a}$$

ex ordine vero posteriori

$$a = \frac{v z}{n a}$$

$$b = i' B - \frac{1}{i} \cdot \frac{\alpha}{b} \cdot \frac{v z}{n a}$$

$$c = i'' C - \frac{1}{i' i''} \cdot \frac{\alpha \beta}{b c} \cdot \frac{v z}{n a}$$

$$d = i''' D - \frac{1}{i' i'' i'''} \cdot \frac{\alpha \beta \gamma}{b c d} \cdot \frac{v z}{n a}$$

Si denique pro loco oculi idoneo ponamus  $dO=0$  erit

$$O = \frac{\delta d}{d+1} i \text{ seu } \frac{1}{O} = \frac{1}{\delta} + \frac{1}{i' i'' i'''} \cdot \frac{\alpha \beta \gamma}{b c d} \cdot \frac{z}{a d}.$$

Coroll.

## Coroll. I.

255. Ex aequationibus prioribus deducimus sequentes :

$$\frac{ia}{a} + \frac{i\mathfrak{B}}{b} = \left(1 + \frac{\alpha}{b}\right) \frac{z}{a}$$

$$\frac{ia}{a} + i\frac{\mathfrak{B}}{b} + \frac{ii'\mathfrak{C}}{\mathfrak{C}} - \frac{ii'\mathfrak{C}}{c} = \left(1 - \frac{\alpha\mathfrak{C}}{bc}\right) \frac{z}{a}$$

$$\frac{ia}{a} + \frac{i\mathfrak{B}}{b} + \frac{ii'\mathfrak{C}}{\mathfrak{C}} - \frac{ii'\mathfrak{C}}{c} - \frac{ii'i''\gamma}{\gamma} + \frac{ii'i''\mathfrak{D}}{d} = \left(1 + \frac{\alpha\mathfrak{C}\gamma}{bcd}\right) \frac{z}{a}$$

quae quomodo ad plures lentes sint continuandae, facile perspicitur.

## Coroll. 2.

256. Si lentium crassities evanescat, fiet  $\alpha=0$ ;  $b=\mathfrak{B}$ ;  $c=\mathfrak{C}$  et  $d=\mathfrak{D}$ , porro  $i=i'=i''=i'''=1$ , unde hae aequationes in sequentes formas abibunt :

$$\mathfrak{B} \cdot \frac{1}{b} = \left(1 + \frac{\alpha}{b}\right) \frac{z}{a}$$

$$\mathfrak{B} \left(\frac{1}{b} + \frac{1}{\mathfrak{C}}\right) - \mathfrak{C} \frac{1}{c} = \left(1 - \frac{\alpha\mathfrak{C}}{bc}\right) \frac{z}{a}$$

$$\mathfrak{B} \left(\frac{1}{b} + \frac{1}{\mathfrak{C}}\right) - \mathfrak{C} \left(\frac{1}{c} + \frac{1}{\gamma}\right) + \mathfrak{D} \frac{1}{d} = \left(1 + \frac{\alpha\mathfrak{C}\gamma}{bcd}\right) \frac{z}{a}$$

ex quibus tribus aequationibus, uti in genere, valor ipsius  $z$ , qui prodierit minimus verum semidiametrum campi apparentis praebebit.

## Coroll. 3.

257. Totus iste campus apparens reuera spectabitur ab oculo in puncto O constituto, dummodo distantia  $dO=O$  fuerit positiva. Sed si ea sit negativa, oculusque lenti ultimo SS immediate applicetur,

pona-



ponatur  $b = \omega$ , scilicet semidiametro pupillae, et ex vltima aequatione elicietur semidiameter spatii in obiecto reuera conspicui.

### Scholion I.

258. Hinc igitur perspicitur, quomodo campus apparens a singularum lentium apertura pendeat; simulque patet quanta esse debeat cuiusque lentis apertura, vt campus apparens datae magnitudinis obtineatur. Si enim quantitas  $z$  cum quantitatibus ad lentium determinationem pertinentibus pro data assumatur, per nostras formulas successiue semidiametri aperturarum pro singulis lentibus definiuntur: vbi quidem deinceps est dispiciendum, num lentes tantae aperturae sint capaces. Hinc scilicet campo apparenti limites praefiniuntur, quos transgredi non liceat; vnde sequitur campum apparentem maiorem assumi non posse, quam vt aperturae inde pro singulis lentibus oriundae admitti queant. His autem definitis perinde est siue cuique lenti ea ipsa, quae fuerit inuenta apertura tribuatur, siue maior, dum ne sit minor quandoquidem hic aperturam lentis obiectivae euanescentem assumimus. Verum si insuper claritatis ratio habeatur, necesse est vt vera apertura cuiusque lentis eam, quam hic assignauimus, aliquantum superet, et quidem ea quantitate, quam supra pro limitibus ob claritatem requisitis exhibuimus, nisi enim hoc augmentum accesserit, extremitas in campo apparente minore lumine praedita erit quam

quam medium. Tum autem campus apparens latius patebit oramque obscuriorem complectetur; quamobrem si circa extremitates minori lumine contenti esse velimus, ne opus quidem est, ut lentibus maior apertura, quam quidem per formulas nostras definitur, tribuatur; superfluumque foret aperturas ultra hos limites augere, ita ut hinc cuique lenti conueniens apertura constituatur.

### Scholion 2.

259. Etsi pro casu, quo lentium crassities negligitur formulae nostrae non multo simpliciores euadunt, tamen in iis percommode vsu venit, ut litterarum  $\mathfrak{B}$ ,  $\mathfrak{C}$ ,  $\mathfrak{D}$ , coefficientes, scilicet  $\frac{1}{b} + \frac{1}{e}$ ;  $\frac{1}{c} + \frac{1}{\gamma}$ ;  $\frac{1}{d} + \frac{1}{\delta}$ ; ipsam distantiam focalem, inuoluant cuiusque lentis; in praxi autem apertura satis tuto ex distantia focali colligi solet. Nam si lentis QQ distantia focalis ponatur  $= q$ , erit  $\frac{1}{b} + \frac{1}{e} = \frac{1}{q}$ , sicque  $\mathfrak{B}(\frac{1}{b} + \frac{1}{e}) = \frac{\pi}{q}$ ; ac ne arcus nimis magni in apertura comprehendantur, necesse est ut sit  $\mathfrak{B} < \frac{1}{2}q$ ; et pro varia lentis forma valor fractionis  $\frac{\pi}{q}$  vsque ad  $\frac{1}{4}$ , vel  $\frac{1}{6}$  diminui debet. Quare si ponamus

$$\mathfrak{B}(\frac{1}{b} + \frac{1}{e}) = \pi; \mathfrak{C}(\frac{1}{c} + \frac{1}{\gamma}) = \pi'; \mathfrak{D}(\frac{1}{d} + \frac{1}{\delta}) = \pi''$$

hae litterae  $\pi$ ,  $\pi'$ ,  $\pi''$  eiusmodi denotabunt fractiones, quarum valor ut plurimum erit vel  $\frac{1}{3}$  vel  $\frac{1}{4}$  vel  $\frac{1}{5}$ , id scilicet tantum cauendum est, ne his litteris nimis magnus valor tribuatur. Quo obseruato cum arbitrio nostro relinquuntur, imprimis conueniet has ipsas lit-



teras in calculum introduci, ex iisque reliquas determinari; earum enim beneficio campus apparens facillime definitur. Quin etiam ipse campus apparens statim quoque in calculum induci poterit, quippe cuius determinatio deinceps per formulam simplicissimam expedietur. Hunc in finem ut tota inuestigatio ad meros numeros redigatur, ponam  $\frac{z}{a} = \Phi$  ita ut  $\Phi$  sit angulus, sub quo semidiameter campi apparentis ab oculo ad lentem obiectiuam collocato spectaretur. Videamus ergo quomodo per hos numeros  $\pi, \pi', \pi'',$  etc. et  $\Phi$  reliquae quantitates definiantur.

### Definitio 3.

260. Ratio aperturae cuiusque lentis mihi vocabitur quotus, qui oritur, si semidiameter aperturae diuidatur per distantiam focalem lentis; eius crassitie pro nihilo habita.

### Coroll. 1.

261. Ita si  $b$  et  $\varepsilon$  sint distantiae determinatrices lentis, et  $\mathfrak{B}$  semidiameter aperturae eiusdem, quia distantia focalis est  $= \frac{b\varepsilon}{b+\varepsilon}$ , ratio aperturae erit  $= \mathfrak{B}(\frac{1}{b} + \frac{1}{\varepsilon})$

### Coroll. 2.

262. Ratio igitur aperturae cuiusvis lentis est fractio minor quam  $\frac{1}{2}$ , quandoquidem hanc legem fanciuimus ut neutrius faciei arcus 60 gradibus maior in apertura contineatur.

Coroll.

## Coroll. 3.

263. Si scilicet ambae facies fuerint aequae curuac ratio aperturae per hanc legem vsque ad  $\frac{1}{2}$  augeri poterit; sin autem altera facies fuerit plana, ratio aperturae  $\frac{1}{4}$  superare vix poterit: ac si lens sit meniscus, ea adhuc minor statui debet.

## Coroll. 4.

264. Cum autem nihil sit, quo apertura lentium accuratius definiatur, ea fere arbitrio nostro relinquitur, et quouis casu commodissime per experientiam determinatur, sufficietque notasse, eam fractioni siue  $\frac{1}{3}$  siue  $\frac{1}{4}$  siue etiam  $\frac{1}{5}$  pro forma lentis aequalem statui debere.

## Scholion.

265. Vt scilicet quouis casu ratio aperturae recte definiatur, radios vtriusque faciei lentis contemplari oportet, qui si fuerint  $f$  et  $g$ , erit distantia focalis  $= \frac{fg}{(n-1)(f+g)} = \frac{20}{11} \cdot \frac{fg}{f+g}$ . Iam semidiameter aperturae minor esse debet quam  $\frac{1}{2}f$  vel quam  $\frac{1}{2}g$ , prout vel  $f$  vel  $g$  fuerit minor. Sit  $g < f$ , et cum semidiameter aperturae minor esse debeat quam  $\frac{1}{2}g$ , ratio aperturae minor accipienda est quam  $\frac{11}{40} (1 + \frac{g}{f})$ . Vnde patet, si lens sit vtrinque aequae conuexa, seu  $g=f$ , rationem aperturae capi debere infra  $\frac{11}{20}$ ; sin autem sit altera facies plana seu  $f=\infty$ , illum limitem esse  $\frac{11}{40}$ , qui adhuc minor fiet, si lens sit



meniscus, seu  $\frac{g}{f}$  numerus negatiuus. Ceterum si ratio aperturæ sit  $= \pi$ , eique hoc modo idoneus valor tribuatur, perinde est siue is negatiue siue positiue accipiat: semper autem conducet rationi aperturæ minorem valorem tribui, quam secundum hanc regulam; partim vt obliquitas radiorum incidentium diminuatur, partim vero potissimum, vt ob claritatem aperturas lentium adhuc ultra augere liceat.

### Problema 6.

266. Si instrumentum dioptricum ex quocunque lentibus, quarum crassitiem vt nullam spectare liceat, sit compositum, dataque sit ratio aperturæ pro singulis lentibus vna cum campo apparente, definire distantias determinatrices singularum lentium.

### Solutio.

Sit distantia obiecti ante lentem primam  $AE = a$ , imaginisque per eam repræsentatae  $aF = \alpha$ , ac prosequentibus lentibus ponatur.

Pro Lente Dist.: determinatrices ratio aperturæ

Secunda. . .  $BF = b$ ;  $bG = g$ ;  $\pi$

Tertia . . .  $CG = c$ ;  $cH = \gamma$ ;  $\pi^I$

Quarta . . .  $DH = d$ ;  $dI = \delta$ ;  $\pi^{II}$

Quinta . . .  $EI = e$ ;  $eK = \varepsilon$ ;  $\pi^{III}$

etc.

Hinc

Hinc ergo si semidiametri aperturarum harum lentium vt ante indicentur literis  $\mathfrak{B}$ ,  $\mathfrak{C}$ ,  $\mathfrak{D}$ ,  $\mathfrak{E}$ , etc. erit  $\pi = \mathfrak{B}(\frac{1}{b} + \frac{1}{\mathfrak{e}})$ ;  $\pi' = \mathfrak{C}(\frac{1}{c} + \frac{1}{\gamma})$ ;  $\pi'' = \mathfrak{D}(\frac{1}{d} + \frac{1}{\delta})$ ;  $\pi''' = \mathfrak{E}(\frac{1}{e} + \frac{1}{\varepsilon})$  etc. Tum vero si sit semidiameter campi apparentis  $= z$  ponatur etiam  $\frac{z}{a} = \Phi$ . Cum igitur hinc sit:

$$\mathfrak{B} = \frac{\pi \mathfrak{e}}{b + \mathfrak{e}}; \mathfrak{C} = \frac{\pi' \gamma}{c + \gamma}; \mathfrak{D} = \frac{\pi'' d \delta}{d + \delta}; \mathfrak{E} = \frac{\pi''' e \varepsilon}{e + \varepsilon} \text{ etc.}$$

habebimus ex §. 256 sequentes aequationes.:

$$\frac{\pi \mathfrak{e}}{b + \mathfrak{e}} = (1 + \frac{\alpha}{b}) \Phi$$

$$\pi - \frac{\pi' \gamma}{c + \gamma} = (1 - \frac{\alpha \mathfrak{e}}{b c}) \Phi$$

$$\pi - \pi' + \frac{\pi'' d \delta}{d + \delta} = (1 + \frac{\alpha \mathfrak{e} \gamma}{b c d}) \Phi$$

$$\pi - \pi' + \pi'' - \frac{\pi''' e \varepsilon}{e + \varepsilon} = (1 - \frac{\alpha \mathfrak{e} \gamma \delta}{b c d e}) \Phi$$

etc.

Quo iam facilius hinc per  $\pi$ ,  $\pi'$ ,  $\pi''$ ,  $\pi'''$  etc. et  $\Phi$  distantiae determinatrices lentium definiri queant, ponatur  $\alpha = A a$ ;  $\mathfrak{e} = B b$ ;  $\gamma = C c$ ;  $\delta = D d$ ;  $\varepsilon = E e$  etc.

Ita vt litterae  $A, B, C, D, E$  etc. denotent numeros absolutos, ac nostrae aequationes induent has formas.:

$$\frac{B \pi}{B + 1} = (1 + \frac{A a}{b}) \Phi$$

$$\pi - \frac{C \pi'}{C + 1} = (1 - \frac{A B a}{c}) \Phi$$

$$\pi - \pi' + \frac{D \pi''}{D + 1} = (1 + \frac{A B C a}{d}) \Phi$$

$$\pi - \pi' + \pi'' - \frac{E \pi'''}{E + 1} = (1 - \frac{A B C D a}{e}) \Phi$$

etc.



vnde eliciuntur sequentes determinationes:

$$b = \frac{A(B+1)a\Phi}{B\pi - (B+1)\Phi}$$

$$c = \frac{AB(C+1)a\Phi}{C\pi' - (C+1)(\pi - \Phi)}$$

$$d = \frac{ABC(D+1)a\Phi}{D\pi'' - (D+1)(\pi' - \pi + \Phi)}$$

$$e = \frac{ABCD(E+1)a\Phi}{E\pi''' - (E+1)(\pi'' - \pi' + \pi - \Phi)}$$

etc.

Datis ergo praeter numeros  $\Phi, \pi, \pi', \pi'', \pi'''$  etc. numeris  $A, B, C, D, E$  etc. cum distantia obiecti  $AE = a$ , per has formulas distantiae  $b, c, d, e$  etc. determinantur indeque insuper alterae  $\alpha, \beta, \gamma, \delta, \varepsilon$  etc. hoc modo:

$$\alpha = Aa$$

$$\beta = \frac{AB(B+1)a\Phi}{E\pi - (B+1)\Phi}$$

$$\gamma = \frac{ABC(C+1)a\Phi}{C\pi' - (C+1)(\pi - \Phi)}$$

$$\delta = \frac{ABCD(D+1)a\Phi}{D\pi'' - (D+1)(\pi' - \pi + \Phi)}$$

$$\varepsilon = \frac{ABCDE(E+1)a\Phi}{E\pi''' - (E+1)(\pi'' - \pi' + \pi - \Phi)}$$

etc.

Hinc nanciscimur distantias focales lentium

$$\text{Primae } PP = \frac{Aa}{A+1}$$

$$\text{Secundae } QQ = \frac{ABa\Phi}{B\pi - (B+1)\Phi}$$

$$\text{Tertiae } RR = \frac{ABCa\Phi}{C\pi' - (C+1)(\pi - \Phi)}$$

$$\text{Quartae } SS = \frac{ABCDa\Phi}{D\pi'' - (D+1)(\pi' - \pi + \Phi)}$$

$$\text{Quintae } TT = \frac{ABCDEa\Phi}{E\pi''' - (E+1)(\pi'' - \pi' + \pi - \Phi)}$$

etc.

Coroll.

## Coroll. 1.

267. Ex angulo  $\Phi$  cum distantia obiecti ante lentem primam  $AE = a$ , ita definitur semidiameter campi apparentis  $z$ , ut sit  $z = a\Phi$ : neque tamen campus apparens pro lubitu assumi potest, sed is per multiplicationem determinabitur, ut mox videbimus.

## Coroll. 2.

268. Cum omnes numeri hic in calculum introducti aequae negative ac positive accipi queant, observandum est eos perpetuo ita assumi debere, ut intervalla lentium quae sunt  $\alpha + b$ ;  $\beta + c$ ;  $\gamma + d$ ;  $\delta + e$ ; etc. omnia prodeant positiva.

## Coroll. 3.

269. Quod ad aperturam cuiusque lentis attinet eius semidiameter habebitur si eius distantia focalis multiplicetur per rationem aperturae littera  $\pi$  insignitam.

## Scholion.

270. Quo formulas hic inventas simpliciores reddamus quoniam litteris  $\mathcal{A}$ ,  $\mathcal{B}$ ,  $\mathcal{C}$ ,  $\mathcal{D}$ ; etc. non amplius indigebimus, ponamus ad abbreviandum;

$$\frac{A}{A+1} = \mathcal{A}; \frac{B}{B+1} = \mathcal{B}; \frac{C}{C+1} = \mathcal{C}; \frac{D}{D+1} = \mathcal{D}; \frac{E}{E+1} = \mathcal{E} \text{ etc.}$$

ut sit

$$A = \frac{\mathcal{A}}{1-\mathcal{A}}; B = \frac{\mathcal{B}}{1-\mathcal{B}}; C = \frac{\mathcal{C}}{1-\mathcal{C}}; D = \frac{\mathcal{D}}{1-\mathcal{D}}; E = \frac{\mathcal{E}}{1-\mathcal{E}} \text{ etc.}$$

at-



atque habebimus:

$$a = Aa;$$

$$b = \frac{ABa\Phi}{B\pi - \Phi}$$

$$c = \frac{ABCa\Phi}{E\pi' - \pi + \Phi}$$

$$d = \frac{ABCDa\Phi}{D\pi'' - \pi' + \pi - \Phi}$$

$$e = \frac{ABCDEa\Phi}{E\pi''' - \pi'' + \pi' - \pi + \Phi}$$

$$b = \frac{Aa\Phi}{B\pi - \Phi}$$

$$c = \frac{ABa\Phi}{E\pi' - \pi + \Phi}$$

$$d = \frac{ABCa\Phi}{D\pi'' - \pi' + \pi - \Phi}$$

$$e = \frac{ABCDa\Phi}{E\pi''' - \pi'' + \pi' - \pi + \Phi}$$

etc.

Hincque porro definientur distantiae focales lentium:

$$\text{Primae } PP = Aa$$

$$\text{Secundae } QQ = \frac{ABa\Phi}{B\pi - \Phi}$$

$$\text{Tertiae } RR = \frac{ABCa\Phi}{E\pi' - \pi + \Phi}$$

$$\text{Quartae } SS = \frac{ABCDa\Phi}{D\pi'' - \pi' + \pi - \Phi}$$

$$\text{Quintae } TT = \frac{ABCDEa\Phi}{E\pi''' - \pi'' + \pi' - \pi + \Phi}$$

etc.

et lentium intervalla:

$$\text{I et II} = \frac{ABa\pi}{B\pi - \Phi}$$

$$\text{II et III} = \frac{ABa\Phi(E\pi' - (1-B)\pi)}{(B\pi - \Phi)(E\pi' - \pi + \Phi)}$$

$$\text{III et IV} = \frac{ABCa\Phi(D\pi'' - (1-E)\pi')}{(E\pi' - \pi + \Phi)(D\pi'' - \pi' + \pi - \Phi)}$$

$$\text{IV et V} = \frac{ABCDa\Phi(E\pi''' - (1-D)\pi'')}{(D\pi'' - \pi' + \pi - \Phi)(E\pi''' - \pi'' + \pi' - \pi + \Phi)}$$

Quae intervalla debent esse positiua.

### Problema 7.

271. Positis iisdem, quae in problemate praecedente sunt assumpta, definire locum idoneum oculi, unde totus campus apparens conspici queat.

So-

## Solutio.

Maneant omnes denominationes vt ante, et quia apertura lentis PP vt nulla spectatur, pro reliquis lentibus ex data aperturæ ratione, semidiameter aperturæ cuiusque ita se habebit

Lentis            semidiameter aperturæ

$$\text{Secundæ } QQ \dots \frac{A \mathfrak{B} a \pi}{\mathfrak{B} \pi - \Phi} \Phi$$

$$\text{Tertiæ } RR \dots \frac{A B \mathfrak{C} a \pi'}{\mathfrak{C} \pi' - \pi' + \Phi} \Phi$$

$$\text{Quartæ } SS \dots \frac{A B C \mathfrak{D} a \pi''}{\mathfrak{D} \pi'' - \pi' + \pi - \Phi} \Phi$$

$$\text{Quintæ } TT \dots \frac{A B C D \mathfrak{E} a \pi'''}{\mathfrak{E} \pi''' - \pi'' + \pi' - \pi + \Phi} \Phi$$

vbi litteræ  $\mathfrak{B}$ ,  $\mathfrak{C}$ ,  $\mathfrak{D}$ ,  $\mathfrak{E}$ , etc. valores in præcedente Scholio assignatos obtinent.

Deinde magnitudines singularum imaginum considerari conuenit, quæ ob  $E \varepsilon = z = a \Phi$  et

$$\alpha = A a, \varepsilon = B b; \gamma = C c; \delta = D d \text{ etc.}$$

erunt

$$F \zeta = \alpha \Phi = A a \Phi$$

$$G \eta = A B a \Phi$$

$$H \theta = A B C a \Phi$$

$$I \iota = A B C D a \Phi$$

etc.

Iam pro quolibet lentium numero locus oculi idoneus seorsim definiri debet; denotante ergo O distantiam oculi post vltimam lentem

Tom. I.

Dd

I.



## I. Pro vnica lente

Quia crassities lentis vt nulla spectatur, eu-  
dens est pro loco oculi idoneo fore  $O=0$ .

## II. Pro duabus Lentibus

Tab. III.

Fig. 13.

Cum hic fit  $bN' + G\eta$ :  $bG = bN'$ :  $bO$  erit  
 $bO = 0 = \frac{bN'}{bN' + G\eta} \xi$ . Sed est  $bN' = \frac{A\mathfrak{B}\pi}{\mathfrak{B}\pi - \Phi} a\Phi$  et  $G\eta = ABa\Phi$   
vnde fit

$$bN' + G\eta = Aa\Phi \cdot \frac{(B+1)\mathfrak{B}\pi - B\Phi}{\mathfrak{B}\pi - \Phi} = \frac{ABa\Phi(\pi - \Phi)}{\mathfrak{B}\pi - \Phi}$$

$$\text{ob } (B+1)\mathfrak{B} = B. \text{ Erit ergo } \frac{bN'}{bN' + G\eta} = \frac{\mathfrak{B}\pi}{B(\pi - \Phi)},$$

quae fractio per  $\xi = \frac{ABa\Phi}{(\mathfrak{B}\pi - \Phi)}$  multiplicata dat locum  
oculi idoneum

$$O = \frac{A\mathfrak{B}a\pi\Phi}{(\pi - \Phi)(\mathfrak{B}\pi - \Phi)}$$

## III. Pro tribus Lentibus

Fig. 14.

Cum hic fit  $cN'' + H\theta$ :  $cH = cN''$ :  $cO$ , erit  
 $cO = 0 = \frac{cN''}{cN'' + H\theta} \gamma$ . Sed est

$$cN'' = \frac{AB\mathfrak{C}a\pi'\Phi}{\mathfrak{C}\pi' - \pi + \Phi} \text{ et } H\theta = ABCa\Phi$$

hincque ob  $(C+1)\mathfrak{C} = C$  fiet

$$cN'' + H\theta = \frac{ABCa\Phi(\pi' - \pi + \Phi)}{\mathfrak{C}\pi' - \pi + \Phi}$$

et  $\frac{cN''}{cN'' + H\theta} = \frac{\mathfrak{C}\pi'}{C(\pi' - \pi + \Phi)}$ . Nunc igitur ob  $\gamma = \frac{ABCa\Phi}{\mathfrak{C}\pi' - \pi + \Phi}$  ha-  
bebimus distantiam oculi idoneam:

$$O = \frac{AB\mathfrak{C}a\pi'\Phi}{(\pi' - \pi + \Phi)(\mathfrak{C}\pi' - \pi + \Phi)}$$

## IV.

## IV. Pro quatuor Lentibus

Cum fit  $dN''' + I : dI = dN''' : dO$ , erit Tab. III.  
 $dO = O = \frac{dN'''}{dN''' + I} \cdot \delta$ . Sed est Fig. 15.

$$dN''' = \frac{ABCD a \pi'' \Phi}{2\pi'' - \pi' + \pi - \Phi} \text{ et } I = ABCD a \Phi$$

hincque ob  $(D + 1)\Phi = D$  fiet

$$dN''' + I = \frac{ABCD a \Phi (\pi'' - \pi' + \pi - \Phi)}{2\pi'' - \pi' + \pi - \Phi} \text{ et}$$

$$\frac{dN'''}{dN''' + I} = \frac{2\pi''}{D(\pi'' - \pi' + \pi - \Phi)}. \text{ Ergo ob } \delta = \frac{ABCD a \Phi}{2\pi'' - \pi' + \pi - \Phi}$$

prodit distantia oculi idonea

$$O = \frac{ABCD a \pi'' \Phi}{(\pi'' - \pi' + \pi - \Phi)(2\pi'' - \pi' + \pi - \Phi)}.$$

## V. Pro quinque Lentibus

Si ratiocinium simili modo ad casum quinque lentium extendatur, reperiemus distantiam oculi idoneam

$$O = \frac{ABCDE a \pi''' \Phi}{(\pi''' - \pi'' + \pi' - \pi + \Phi)(2\pi''' - \pi'' + \pi' - \pi + \Phi)}.$$

## VI. Pro sex Lentibus

Eodemque modo progrediendo colligitur fore pro casu sex lentium distantiam oculi idoneam

$$O = \frac{ABCDE \delta a \pi'''' \Phi}{(\pi'''' - \pi''' + \pi'' - \pi' + \pi - \Phi)(2\pi'''' - \pi''' + \pi'' - \pi' + \pi - \Phi)}$$

ficque ulterius, quousque libuerit progredi licet.



## Coroll. 1.

272. Quouis ergo casu necesse est, vt distantia oculi idonea prodeat positiua: si enim fieret negatiua totus campus apparens nusquam conspici posset.

## Coroll. 2.

273. Iis autem casibus, quibus distantia  $O$  fit negatiua oculum immediate vltimae lenti applicari conueniet; Tum vero oculus plus non cernet, quam si vltimae lentis apertura aequalis esset amplitudini pupillae.

## Coroll. 3.

274. Hoc ergo casu statuatur semidiameter aperturae vltimae lenti  $= \omega$  semidiametro pupillae, ex eaque aequatione eliciatur valor ipsius  $\Phi$ , quo invento erit  $a\Phi$  semidiameter campi apparentis, qui in obiecto reuera conspicietur.

## Problema 8.

275. Positis iisdem atque in problematibus praecedentibus, eam conditionem in lentium dispositione definire, vt oculus in loco idoneo positus obiectum simul distincte videat.

## Solutio.

Quia aperturam lentis obiectiuae euanescentem assumimus, in visione alia confusio locum habere nequit nisi quatenus oculus non in distantia iusta ab  
ultima

ultima imagine, quam intuetur existit, quae ergo tolletur, si lentes ita disponantur, ut imago ultima ante oculum in O situm in distantia iusta, quam littera  $l$  designauimus, versetur. Cum igitur in figuris locus oculi ante imaginem ultimam cadat haec distantia negative sumpta ipsi  $l$  aequalis est ponenda; unde pro quouis lentium numero sequentes habebimus determinationes.

### I. Pro vnica lente

Cum hic sit  $O = 0$ , et  $OF = a = Aa$ , oportet esse  $Aa = -l$ , ideoque  $A = -\frac{l}{a}$ : et  $a = -l$ , unde indoles huius lentis determinatur, ita ut eius distantia focalis esse debeat  $= \frac{al}{l-a}$ .

Tab. III.

Fig. 12.

### II. Pro duabus lentibus

Ex inuenta distantia  $bO = 0$ , erit  $OG = \frac{0}{b_N}$ . Fig. 13.  
 Gn. Est vero  $\frac{0}{b_N} = \frac{1}{\pi - \phi}$ ; unde fit  $\frac{ABa\phi}{\pi - \phi} = -l$ , hincque pro secunda lente  $B = -\frac{(\pi - \phi)l}{Aa\phi}$  et  $\mathfrak{B} = \frac{B}{B+1}$ . Vel cum sit  $Aa\phi = -\frac{(\pi - \phi)l}{B}$ , erit pro loco oculi

$$O = \frac{-\mathfrak{B}\pi}{B(\mathfrak{B}\pi - \phi)} l$$

### III. Pro tribus lentibus

Hic est  $OH = \frac{0}{c_N}$ .  $H\theta = \frac{H\theta}{\pi' - \pi + \phi}$ , unde obtinetur: Fig. 14.

$$OH = \frac{ABCa\phi}{\pi' - \pi + \phi} = -l$$

sicque pro vltima lente habebitur:

$$C = -\frac{(\pi' - \pi + \phi)l}{ABa\phi}$$

Dd 3

At



At si pro determinatione primae lentis capiatur  
 $Aa\Phi = \frac{-(\pi' - \pi + \Phi)l}{BC}$ , erit distantia oculi

$$O = \frac{-\epsilon\pi'}{C(\epsilon\pi' - \pi + \Phi)} l$$

#### IV. Pro quatuor lentibus

Tab. III.  
Fig. 15.

Cum fit  $OI = \frac{O}{a_{N''}}$ ,  $I = \frac{I'}{\pi'' - \pi' + \pi - \Phi}$ , habebitur

$$OI = \frac{ABCDa\Phi}{\pi'' - \pi' + \pi - \Phi} = -l$$

Vnde pro vltima lente

$$D = -\frac{(\pi'' - \pi' + \pi - \Phi)l}{ABCa\Phi}$$

Cum autem fit  $ABCa\Phi = -\frac{(\pi'' - \pi' + \pi - \Phi)l}{D}$  erit in loco oculi hoc valore furrogando

$$O = \frac{-\mathfrak{D}\pi''}{D(\mathfrak{D}\pi'' - \pi' + \pi - \Phi)} l$$

#### V. Pro quinque Lentibus

Simili modo pro quinque lentibus vltima ita comparata esse debet, vt fit

$$E = \frac{-(\pi''' - \pi'' + \pi' - \pi + \Phi)l}{ABCDEa\Phi}$$

Prima autem inde definita fit

$$O = \frac{-\epsilon\pi'''}{E(\epsilon\pi''' - \pi'' + \pi' - \pi + \Phi)} l$$

#### VI. Pro sex Lentibus

Eodem modo patet pro sex lentibus fore

$$F = \frac{-(\pi'''' - \pi''' + \pi'' - \pi' + \pi - \Phi)l}{ABCDEa\Phi}$$

atque

atque si hinc  $Aa$  definiatur :

$$O = \frac{-8\pi''''}{F(8\pi'''' - \pi'''' + \pi'' - \pi' + \pi - \Phi)} l$$

quas formulas quousque libuerit, continuare licet.

### C o r o l l. I.

276. Si distantia oculi iusta  $l$  fuerit infinita erit vt sequitur :

- I. Pro vna lente  $A = \infty$  et  $\mathcal{A} = 1$
  - II. Pro duabus lentibus  $B = \infty$  et  $\mathcal{B} = 1$
  - III. Pro tribus lentibus  $C = \infty$  et  $\mathcal{C} = 1$
  - IV. Pro quatuor lentibus  $D = \infty$  et  $\mathcal{D} = 1$
- etc.

### C o r o l l. 2.

277. Casu ergo, quo distantia oculi iusta  $l$  est infinita, distantia oculi post lentem vltimam erit pro quouis lentium numero :

- I. Pro vnica lente  $O = 0$
  - II. Pro duabus lentibus  $O = \frac{Aa\pi\Phi}{(\pi - \Phi)}$
  - III. Pro tribus lentibus  $O = \frac{ABa\pi'\Phi}{(\pi' - \pi + \Phi)^2}$
  - IV. Pro quatuor lentibus  $O = \frac{ABCa\pi''\Phi}{(\pi'' - \pi' + \pi - \pi)^2}$
- etc.

### S c h o l i o n.

278. Hactenus campum apparentem  $\Phi$  vt datum consideraui, ex eoque tam lentium indolem quam



quam earum dispositionem determinavi, ut campus datae amplitudinis appareat, nihilque ob stare deprehendimus, quominus huic conditioni satisfiat; cum numeri  $A, B, C, D$ , etc. penitus arbitrio relinquantur, aperturarum vero rationes  $\pi, \pi', \pi''$ , etc. infra  $\frac{1}{3}$  vel  $\frac{1}{4}$  accipi debeant. Verum hic multiplicationis ratio nondum in computum est ducta qua simul campus apparens ita adstringitur, ut certum limitem excedere nequeat. Quoniam igitur in omnibus instrumentis dioptricis multiplicatio imprimis proposita esse solet, quemadmodum per eam campus apparens definiatur in sequente problemate exponamus.

### Problema 9.

279. Si instrumentum dioptricum ex quocunque lentibus fuerit compositum, quarum quidem crassities ut nulla spectetur, simul vero multiplicationis ratio sit proposita, determinare campum apparentem.

### Solutio.

Manentibus omnibus denominationibus, quibus hactenus sumus usi, ita ut  $a\Phi$  semidiametrum campi apparentis denotet, sit  $b$  distantia, ad quam multiplicationem referamus. Magnitudo igitur  $a\Phi$  in distantia hac  $= b$  nudo oculo cerneretur sub angulo, cuius tangens est  $= \frac{a\Phi}{b}$ . Quare si multiplicationis ratio statuatur  $= m$ , necesse est ut eadem magnitudo  $a\Phi$  per lentes spectetur sub angulo, cuius tangens sit  $= \frac{m a \Phi}{b}$ .

Iam

Iam vero ex iis, quae in problematibus praecedentibus sunt tradita iste angulus facile assignatur, sicque obtinebitur angulus  $\Phi$ , indeque semidiameter campi apparentis  $a\Phi$ . Cum autem haec multiplicatio non ad ipsos angulos sed eorum tangentes referatur evidens est tantum partes obiecti minimas circa centrum E fitas in ratione proposita multiplicari, remotiores vero in ratione minore. Quo notato hanc multiplicationis rationem  $m$  pro quouis lentium numero contemplemur.

### I. Pro vnica Lente

Tangens anguli, quo imago  $F\zeta$  ab oculo in O constituto conspicitur, est  $\frac{F\zeta}{OF} = \frac{F\zeta}{aF}$  ob  $aO = 0$ , Erit ergo  $\frac{ma\Phi}{b} = \Phi$ , seu  $ma = b$ . Hoc ergo casu campus apparens non determinatur sed multiplicationis ratio est  $m = \frac{b}{a}$ . Verum vt visio sit distincta, per superius problema debet esse  $A = -\frac{l}{a}$  et distantia oculi post lentem  $O = 0$ . At obiectum situ erecto cernetur:

Tab. III.  
Fig. 12.

### II. Pro duabus Lentibus

Tangens anguli, quo imago  $G\eta$  ab oculo in O constituto cernetur, est  $= \frac{bN'}{bO} = \pi - \Phi = \frac{ma\Phi}{b}$ ; vnde sequitur semidiameter campi apparentis

Fig. 13.

$$\Phi = \frac{\pi b}{ma + b} \text{ pro situ inuerso}$$

Quo inuento, vt visio sit distincta oportet esse

$$B = \frac{-ml}{Ab}, \text{ et } \mathfrak{B} = \frac{-ml}{Ab - ml}$$

Tom. I.

Ee

hinc-



hincque prodit distantia oculi post lentem ocularem

$$O = \frac{A b l (m a + b)}{m m a l + A b b}$$

### III. Pro tribus Lentibus

Tab. III.

Fig. 14.

Tangens anguli, quo imago  $H\theta$  ab oculo in  $O$  constituto cernitur est  $\frac{c N''}{c O} = \pi' - \pi + \Phi = \frac{m a \Phi}{b}$ , unde fit semidiameter campi apparentis:

$$\Phi = \frac{(\pi' - \pi) b}{m a - b} \text{ pro situ erecto}$$

Deinde vt visio fit distincta, oportet esse

$$C = \frac{-m l}{A B b} \text{ et } \mathfrak{C} = \frac{-m l}{A B b - m l}$$

Cum igitur sit  $\pi' - \pi + \Phi = \frac{m a (\pi' - \pi)}{m a - b}$  erit

$$\mathfrak{C} \pi' - \pi + \Phi = \pi' - \pi + \Phi - \frac{A B b \pi'}{A B b - m l} = \frac{m a (\pi' - \pi)}{m a - b} - \frac{A B b \pi'}{A B b - m l}$$

hincque pro loco oculi

$$O = \frac{-A B b l \pi'}{(A B b - m l)(\mathfrak{C} \pi' - \pi + \Phi)} = \frac{A B b l (m a - b) \pi'}{(m m a l - A B b b) \pi' + m a (A B b - m l) \pi'}$$

### IV. Pro quatuor lentibus

Fig. 15. Tangens anguli, quo imago  $I\iota$  ab oculo in  $O$  constituto cernitur est  $= \frac{d N'''}{d O} = \pi'' - \pi' + \pi - \Phi = \frac{m a \Phi}{b}$ ;

unde elicitur:

$$\Phi = \frac{(\pi'' - \pi' + \pi) b}{m a + b} \text{ pro situ inuerso}$$

Hinc porro pro visione distincta esse debet

$$D = \frac{-m l}{A B C b} \text{ et } \mathfrak{D} = \frac{-m l}{A B C b - m l}$$

unde fit

$$\mathfrak{D} \pi'' - \pi' + \pi - \Phi = \frac{m a (\pi'' - \pi' + \pi)}{m a + b} - \frac{A B C b \pi''}{A B C b - m l}$$

et

et pro loco oculi

$$O = \frac{ABCbl(ma+b)\pi''}{(mmal+ABChb)\pi''+ma(ABCh-mL)(\pi'-\pi)}.$$

### V. Pro quinque lentibus

Eodem modo progrediendo pro campo apparente reperitur :

$$\Phi = \frac{(\pi''' - \pi'' + \pi' - \pi)b}{ma - b} \text{ pro situ erecto}$$

et vt visio euadat distincta

$$E = \frac{-ml}{ABCD b} \text{ et } \mathcal{E} = \frac{-ml}{ABCD b - ml}$$

vnde pro loco oculi idoneo concluditur

$$O = \frac{ABCDbl(ma-b)\pi'''}{(mmal-ABCDhb)\pi''' + ma(ABCDh-mL)(\pi'' - \pi' + \pi)}.$$

### VI. Pro sex lentibus

Hic campus apparens ita definitur , vt fit :

$$\Phi = \frac{(\pi'''' - \pi''' + \pi'' - \pi' + \pi)b}{ma + b} \text{ pro situ inuerso}$$

visio vero distincta exigit

$$F = \frac{-ml}{ABCDEb} ; \mathcal{F} = \frac{-ml}{ABCDEb - ml}.$$

vnde pro loco oculi idoneo

$$O = \frac{ABCDEbl(ma+b)\pi''''}{(mmal+ABCDEhb)\pi'''' + ma(ABCDEh-mL)(\pi'''' - \pi''' + \pi'' - \pi)}$$

ficque progressio ad plures lentes est manifesta.

### Coroll. I.

280. Datis ergo rationibus aperturarum singularum lentium  $\pi$ ,  $\pi'$ ,  $\pi''$ , etc. vna cum ratione multiplicationis  $m$ , distantia  $b$ , ad quam multiplicatio

Ee 2

refer-



refertur, et distantia obiecti ante instrumentum  $a$ , determinatur campus apparens.

### Coroll. 2.

281. Vt ergo campus apparens pro data multiplicatione maximus obtineatur litteris  $\pi$ ,  $\pi'$ ,  $\pi''$ , etc. ita valores maximos tribui conueniet, vt alternatim sint positiui et negatiui.

### Coroll. 3.

282. Si igitur valores  $\pi$ ,  $\pi'$ ,  $\pi''$ , etc. vsque ad  $\frac{1}{3}$  augeri liceat, maximus valor ipsius  $\Phi$  pro quouis lentium numero erit vt sequitur:

$$\text{Pro casu duarum lentium } \Phi = \frac{b}{3(ma+b)}$$

$$\text{Pro casu trium lentium } \Phi = \frac{2b}{3(ma-b)}$$

$$\text{Pro casu quatuor lentium } \Phi = \frac{3b}{3(ma+b)}$$

$$\text{Pro casu quinque lentium } \Phi = \frac{4b}{3(ma-b)}$$

etc.

### Coroll. 4.

283. Quo plures ergo lentes adhibentur eo magis campus apparens augeri potest simul vero patet, quo maior multiplicatio desideretur, eo minorem fieri campum apparentem.

Coroll.

## Coroll. 5.

284. Ratio multiplicationis  $m$  tam positue quam negatiue capi potest. Si positue accipitur pro lentium numero pari situm inuersum, pro impari autem situm erectum declarat. Contrarium vero euenit, si  $m$  fuerit numerus negatiuus.

## Scholion I.

285. Hic autem imprimis notandum est, valorem ipsius  $\Phi$  tum solum angulum  $EA\varepsilon$  praeberere, quando fuerit tam exiguus, vt aliquot gradus non superet; si enim valor ipsius  $\Phi$  prodeat multo maior, tum tangentem huius anguli  $EA\varepsilon$  exprimit. Plerumque autem, si quidem multiplicatio sit modica, iste valor ipsius  $\Phi$  tam paruus reperitur, vt sine errore pro ipso angulo  $EA\varepsilon$  accipi possit. Hic igitur ob aliam causam amplitudo campi apparentis restringitur, vt certum limitem superare nequeat; cum enim angulus, quo radii in oculum incidentes in  $O$  ad axem inclinantur, nunquam possit esse rectus, neque fortasse vix  $60^\circ$  superare queat, quandoquidem ne nudo quidem oculo spatium in coelo maius quam  $120^\circ$  conspiciere valeamus; si illum angulum maximum, quem oculus capere valeat, circiter  $63^\circ$  statuamus, vt eius tangens sit  $= 2$  pro quouis lentium numero habebimus  $\frac{m a \Phi}{b} = 2$ , vnde fit  $\Phi = \frac{2 b}{m a}$ , et  $a \Phi = \frac{2 b}{m}$ . Data ergo multiplicatione  $m$  et distantia  $b$  ad quam refertur, semidiameter spatii in obiecto



conspicui nunquam maior existere potest quam  $\frac{2b}{m}$  quotcunque etiam adhibeantur lentes, eaeque ita disponantur ut maximum campum patefaciant. In Telescopiis ergo, ubi sumitur  $b = a$ , et semidiameter campi ex ipso angulo  $\Phi$  aestimatur, eius tangens nunquam maior esse potest quam  $\frac{2}{m}$ : unde sequentem tabellam adiungo, quae pro quavis multiplicatione semidiametrum campi apparentis maximi ostendit, quem nunquam superare liceat.

| Multipli-<br>catio<br>$m$ | Semidiam :<br>campi app:<br>maximi | Multipli-<br>catio<br>$m$ | Semidiam :<br>campi app:<br>maximi |
|---------------------------|------------------------------------|---------------------------|------------------------------------|
| 5                         | $21^{\circ}, 48'$                  | 60                        | $1^{\circ}, 54', 33''$             |
| 10                        | $11, 18$                           | 70                        | $1, 38, 13$                        |
| 15                        | $7, 35$                            | 80                        | $1, 25, 57$                        |
| 20                        | $5, 42$                            | 90                        | $1, 16, 24$                        |
| 25                        | $4, 34$                            | 100                       | $1, 8, 45$                         |
| 30                        | $3, 49$                            | 150                       | $0, 45, 50$                        |
| 35                        | $3, 16\frac{1}{2}$                 | 200                       | $0, 34, 22$                        |
| 40                        | $2, 52$                            | 250                       | $0, 27, 30$                        |
| 45                        | $2, 33$                            | 300                       | $0, 22, 55$                        |
| 50                        | $2, 17\frac{1}{2}$                 | 400                       | $0, 17, 11$                        |
|                           |                                    | 500                       | $0, 13, 45$                        |

Quod si ergo numerum lentium multiplicando iam fere ad tantum campum apparentem pertigerimus, is ulterius nullo modo augeri poterit.

Scho-

## Scholion 2.

286. Quod ad locum oculi idoneum attinet, eum ideo in O constituimus, vt omnes radios per lentes transmissos accipiat, etiamsi pupilla maxime esset constricta: ex quo patet ob amplitudinem pupillae oculum de hoc loco sine vlllo detrimento aliquantillum remoueri posse, ita vt superfluum foret hunc locum nimis sollicite obseruare, nisi forte apertura vltimae lentis fuerit admodum magna. Sin ea autem pupillam non superet, eaque adeo sit minor, manifestum est oculum ei immediate applicatum aequae omnes radios excipere, et eundem campum contueri, ac si in loco idoneo esset constitutus. His igitur casibus, si forte distantia O pro loco oculi prodeat negatiua, nihil de campo apparente perit, dummodo oculus lenti vltimae immediate applicetur. His itaque, quae ad visionem per instrumenta dioptrica in genere pertinent, expeditis, superest, vt inuestigemus, quantum visio ob diuersam radiorum refrangibilitatem turbetur, et quemadmodum hanc perturbationem euitare queamus.

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## CAPVT VI.

DE

## CONFUSIONE

## A DIVERSA RADIORVM INDOLE

## ORIVNDA.

## Problema I.

287.

Si a puncto dato E radii per lentem PP transmittantur definire variationem in loco imaginis F, quae a diuersa radiorum refrangibilitate oritur.

## Solutio.

Sit distantia puncti E ante lentem  $AE = a$ , facierum autem lentis radius anterioris  $= f$  posterioris  $= g$  et crassities  $Aa = v$ , quae quantitates sunt constantes. Posita nunc refractionis ratione ex aere in vitrum  $= n:1$ , ob diuersam radiorum naturam numerus  $n$  erit variabilis, ideoque etiam locus imaginis F post lentem expressae, cuius distantia si ponatur  $aF = \alpha$ , erit ex supra inuentis

$$\frac{n-1}{f} = \frac{1}{a} + \frac{2n}{k+v} \quad \text{et} \quad \frac{n-1}{g} = \frac{1}{\alpha} - \frac{2n}{k-v}$$

vbi

vbi quantitas  $k$  etiam pro variabili est habenda, quia tantum  $a$ ,  $f$ ,  $g$  et  $v$  sunt constantes. Quaestio ergo huc redit, vt si numerus  $n$  differentiali suo  $dn$  crescere sumatur, definiatur differentiale distantiae  $\alpha$ . Quare differentientur ambae aequationes illae;

$$\frac{dn}{f} = \frac{2dn}{k+v} - \frac{2ndk}{(k+v)^2}$$

$$\frac{dn}{g} = -\frac{d\alpha}{\alpha\alpha} - \frac{2dn}{k-v} + \frac{2ndk}{(k-v)^2}.$$

indeque eliminato  $dk$  habebitur

$$\frac{dn(k+v)^2}{f} + \frac{dn(k-v)^2}{g} = 2dn(k+v) - 2dn(k-v) - \frac{d\alpha(k-v)^2}{\alpha\alpha}$$

Restituantur pro  $f$  et  $g$  valores initio positi, ac peruenietur ad hanc aequationem:

$$\frac{dn(k+v)^2}{a} + \frac{dn(k-v)^2}{\alpha} + 4vdn + \frac{(n-1)d\alpha}{\alpha\alpha}(k-v)^2 = 0$$

vnde reperitur:

$$\frac{d\alpha}{\alpha\alpha} = -\frac{dn}{(n-1)\alpha} \left( \frac{k+v}{k-v} \right)^2 - \frac{dn}{(n-1)\alpha} - \frac{4vdn}{(n-1)(k-v)^2} \text{ seu}$$

$$d\alpha = -\frac{\alpha dn}{n-1} \left( 1 + \frac{\alpha}{a} \left( \frac{k+v}{k-v} \right)^2 + \frac{4\alpha v}{(k-v)^2} \right)$$

Tum vero cum etiam  $k$  sit quantitas variabilis erit

$$dk = -\frac{(k+v)dn}{n(n-1)} \left( 1 + \frac{k+v}{2a} \right)$$

Cum ergo posuerimus  $\frac{k-v}{k+v} = i$ , ob  $di = \frac{2vdk}{(k+v)^2}$  erit

$$di = -\frac{vdn}{n(n-1)} \left( \frac{1}{a} + \frac{2}{k+v} \right).$$

### COROLL. I.

288. Si  $n:1$  denotet rationem refractionis radiorum mediae naturae, vt fit  $n = \frac{31}{20} = 1,55$ . erit

Tom. I.

Ff

pro



pro radiis rubris seu minime refractis  $n=1,54$ , et violaceis  $n=1,56$ . quorum valorum discrimen a medio, cum sit  $=\frac{1}{105}$ , pro differentiali  $dn$  haberi poterit.

### COROLL. 2.

289. Quare si  $\alpha$  denotet distantiam imaginis a radiis mediis formatae, pro ea quae a rubris formatur, erit  $dn=-\frac{1}{105}$  et  $\frac{dn}{n-1}=-\frac{1}{55}$ . Hinc distantia imaginis rubrae post lentem erit  $=\alpha+\frac{\alpha}{55}\left(1+\frac{\alpha}{a}\left(\frac{k+v}{k-v}\right)^2+\frac{4\alpha v}{(k-v)^2}\right)$   
Distantia autem imaginis violaceae post lentem erit

$$\alpha - \frac{\alpha}{55}\left(1+\frac{\alpha}{a}\left(\frac{k+v}{k-v}\right)^2+\frac{4\alpha v}{(k-v)^2}\right).$$

### COROLL. 3.

290. Si crassities lentis euanescat, vt fit  $v=0$ , ob variabilitatem numeri  $n$  erit

$$d\alpha = \frac{-\alpha dn}{n-1}\left(1+\frac{\alpha}{a}\right) = \frac{-\alpha \alpha dn}{n-1}\left(\frac{1}{a}+\frac{1}{\alpha}\right)$$

Ac si distantia focalis lentis ponatur  $=p$ , cum sit

$$\frac{1}{a}+\frac{1}{\alpha}=\frac{1}{p}, \text{ erit } d\alpha = \frac{-\alpha \alpha dn}{(n-1)p} = \frac{-\alpha \alpha dn}{11p}.$$

### SCHOLIUM.

291. Hoc ergo modo ob diuersam radiorum naturam valor distantiae  $\alpha$  immutatur, vnde si ab ea tanquam ab obiecto radii porro ad lentem secundam emittantur, fiet etiam respectu huius lentis distantia obiecti variabilis. Quam ob causam in loco imaginis ab ea formatae duplex variatio orietur: id quod deinceps etiam in lentibus sequentibus multo magis

magis eueniet. Hanc igitur variationem, quae pro quauis lente in loco imaginis nascitur, in sequente problemate determinemus.

### Problema 2.

292. Si locus imaginis F, quae respectu lentis QQ vicem obiecti gerit, ob diuersam radiorum naturam ipse fit variabilis, determinare variationem, quam ob eandem causam imago sequens in G patietur.

### Solutio.

Sit pro radiis mediae naturae, quibus respondet numerus  $n$  distantia obiecti F ante lentem  $BF = b$ ; imaginisque inde proiectae distantia post lentem  $bG = g$ ; dum autem  $n$  abit in  $n + dn$ , hae distantiae ambae  $b$  et  $g$  capiant sua incrementa differentialia  $db$  et  $dg$ . Ad quae inuenienda sit lentis QQ radius faciei anterioris  $= f$ , posterioris  $= g$ , et crassities  $Bb = v$ , eritque vt ante:

$$\frac{n-1}{f} = \frac{1}{b} + \frac{2n}{k+v} \text{ et } \frac{n-1}{g} = \frac{1}{g} - \frac{2n}{k-v}$$

vbi  $k$  cum  $b$  et  $g$  pro variabili est habenda. Differentiatione ergo instituta habebitur.

$$\frac{dn}{f} = -\frac{db}{bb} + \frac{2dn}{k+v} - \frac{2ndk}{(k+v)^2}$$

$$\frac{dn}{g} = -\frac{dg}{gg} - \frac{2dn}{k-v} + \frac{2ndk}{(k-v)^2}$$

vnde eliminato  $dk$ , fit

$$\frac{dn(k+v)^2}{f} + \frac{dn(k-v)^2}{g} = -\frac{db}{bb}(k+v)^2 - \frac{dg}{gg}(k-v)^2 + 4vdn$$

Ff 2

quae



quae multiplicata per  $n-1$ , si pro  $f$  et  $g$  valores dati substituantur prodit

$$\frac{dn(k+v)^2}{b} + 2n dn(k+v) + \frac{dn(k-v)^2}{g} - 2n dn(k-v) \\ = -\frac{(n-1)db}{bb}(k+v)^2 - \frac{(n-1)dg}{gg}(k-v)^2 + 4(n-1)v dn$$

feu

$$\frac{dn(k+v)^2}{(n-1)b} + \frac{dn(k-v)^2}{(n-1)g} + \frac{4v dn}{n-1} + \frac{db}{bb}(k+v)^2 + \frac{dg}{gg}(k-v)^2 = 0$$

Atque hinc elicitur:

$$dg = -\frac{ggdb}{bb}\left(\frac{k+v}{k-v}\right)^2 - \frac{gdn}{n-1}\left(1 + \frac{g}{b}\left(\frac{k+v}{k-v}\right)^2 + \frac{4gv}{(k-v)^2}\right)$$

Pro variabilitate autem ipsius  $k$  reperietur

$$\frac{dk}{(k+v)^2} = -\frac{db}{2nb} - \frac{dn}{2n(n-1)b} - \frac{dn}{n(n-1)(k+v)}$$

Quare si ponatur  $\frac{k-v}{k+v} = i$ , ob  $di = \frac{2v dk}{(k+v)^2}$  erit

$$di = \frac{vdb}{nb} - \frac{v dn}{n(n-1)b} - \frac{2v dn}{n(n-1)(k+v)}$$

ac si loco  $k$  numerus  $i$  introducatur, erit

$$dg = -\frac{ggdb}{iibb} - \frac{bdn}{n-1}\left(1 + \frac{g}{iib} + \frac{(1-i)^2 g}{iiv}\right) \text{ et}$$

$$di = \frac{vdb}{nb} - \frac{v dn}{n(n-1)}\left(\frac{1}{b} + \frac{1-i}{v}\right).$$

### COROLL. I.

293. Inuenta aequatio differentialis etiam hac forma repraesentari potest ut sit

$$\frac{idg}{gg} + \frac{db}{ib} = \frac{-dn}{n-1}\left(\frac{i}{g} + \frac{1}{ib} + \frac{(1-i)^2}{iv}\right)$$

siue

siue restituendo  $k$

$$\frac{d\varepsilon}{\varepsilon\varepsilon}\left(\frac{k-v}{k+v}\right) + \frac{db}{bb}\left(\frac{k+v}{k-v}\right) = \frac{-dn}{n-1}\left(\frac{1}{\varepsilon}\left(\frac{k-v}{k+v}\right) + \frac{1}{b}\left(\frac{k+v}{k-v}\right) + \frac{v}{kk-vv}\right)$$

vbi haec obseruanda est analogia, vt quemadmodum ad  $b$  refertur  $\frac{k+v}{k-v}$ , ita ad  $\varepsilon$  referatur  $\frac{k-v}{k+v}$ .

### Coroll. 2.

294. Si lentis huius crassities euanescat, fit  $v=0$ , et  $i=1$  vbi figura lentis non amplius in computum ingreditur sed sola distantia focalis, vnde variatio in loco imaginis  $G$  ita erit comparata, vt fit

$$\frac{d\varepsilon}{\varepsilon\varepsilon} + \frac{db}{bb} = \frac{-dn}{n-1}\left(\frac{1}{\varepsilon} + \frac{1}{b}\right) \text{ ideoque}$$

$$d\varepsilon = -\frac{\varepsilon\varepsilon}{b}db - \frac{\varepsilon\varepsilon dn}{n-1}\left(\frac{1}{\varepsilon} + \frac{1}{b}\right).$$

### Coroll. 3.

295. Si casum a radiis mediae naturae, ad quos formulae haecenus traditae sunt accommodatae, ad radios rubros transferre velimus, poni oportet  $dn = -\frac{1}{100}$ , sin autem ad radios violaceos  $dn = +\frac{1}{100}$ .

### Problema 3.

296. Si radii ab obiecto  $E$  per lentes quotcunque transmittantur, determinare variationem in locis singularum imaginum, quae a diuersa radiorum refrangibilitate proficiscitur.

F f 3

Solutio.



## Solutio.

Retineantur omnes denominationes, quibus in superioribus capitibus sumus vfi, ac sint  $a, \alpha, b, \beta, c, \gamma$  etc. distantiae determinatrices lentium pro radiis mediae naturae. Variata ergo ratione refractionis etiam hae distantiae variabuntur, quarum variationes differentialibus indicemus. Cum autem distantiae inter binas lentes maneant constantes, illae variationes ita erunt comparatae ut fit

$$d\alpha + db = 0, d\beta + dc = 0, d\gamma + dd = 0 \text{ etc.}$$

Cum iam distantia obiecti  $AE = a$  fit invariabilis erit ex problemate primo si ponatur  $\frac{k-v}{k+v} = i$

$$d\alpha = -db = \frac{-\alpha \alpha dn}{i(n-1)} \left( \frac{1}{\alpha} + \frac{1}{ia} + \frac{(1-i)^2}{i v} \right) = \frac{-\alpha \alpha dn}{n-1} \cdot \frac{k+v}{k-v} \left( \frac{k-v}{\alpha(k+v)} + \frac{k+v}{\alpha(k-v)} + \frac{1}{kk-vv} \right)$$

$$\text{et } di = \frac{-v dn}{n(n-1)} \left( \frac{1}{a} + \frac{1-i}{v} \right)$$

Deinde pro secunda lente, ad quam referuntur distantiae determinatrices  $b$ , et  $\beta$  cum arbitraria  $k'$  et crassitie  $v'$ , vnde fecimus  $\frac{k'-v'}{k'+v'} = i'$ , habebimus

$$d\beta = -dc = \frac{-\beta \beta db}{i' b b} - \frac{\beta \beta dn}{i'(n-1)} \left( \frac{i'}{\beta} + \frac{1}{i' b} + \frac{(1-i')^2}{i' v'} \right)$$

$$\text{et } di' = \frac{-v' db}{n b b} - \frac{v' dn}{n(n-1)} \left( \frac{1}{b} + \frac{1-i'}{v'} \right)$$

Simili modo pro tertia lente, ad quam referuntur distantiae determinatrices  $c$  et  $\gamma$  cum arbitraria  $k''$  et crassitie  $v''$  posito  $\frac{k''-v''}{k''+v''} = i''$  adipiscemur:

$$d\gamma = -dd = \frac{-\gamma \gamma dc}{i'' c c} - \frac{\gamma \gamma dn}{i''(n-1)} \left( \frac{i''}{\gamma} + \frac{1}{i'' c} + \frac{(1-i'')^2}{i'' v''} \right)$$

$$\text{et } di'' = \frac{-v'' dc}{n c c} - \frac{v'' dn}{n(n-1)} \left( \frac{1}{c} + \frac{1-i''}{v''} \right)$$

atque

atque ulterius progrediendo obtinebimus sequentes formulas :

$$d\delta = -de = -\frac{\delta\delta d\delta}{i''i''d\delta} - \frac{\delta\delta dn}{i''(n-1)} \left( \frac{i'''}{\delta} + \frac{1}{i''\delta} + \frac{(1-i''')^2}{i''i''v''} \right)$$

$$\text{et } di''' = -\frac{v''d\delta}{n\delta} - \frac{v''dn}{n(n-1)} \left( \frac{1}{\delta} + \frac{1-i'''}{v''} \right)$$

vnde haec differentialia facile ad quotcunque lentes extenduntur. Atque si hic successiue valores differentialium  $dd$ ,  $dc$ ,  $db$ , iam ante definiti substituantur omnia haec differentialia tam distantiarum determinatricium, quam numerorum  $i$ ,  $i'$ ,  $i''$ ,  $i'''$  etc. per differentiale  $dn$  exprimentur.

Si ratio refractionis pro singulis lentibus sit diuersa pro iisque ordine exprimatur numeris,  $n$ ,  $n'$ ,  $n''$  etc. perspicuum est, differentialia hic inuenta sequenti modo expressum iri :

$$\text{I. } d\alpha = -db = -\frac{\alpha\alpha dn}{i(n-1)} \left( \frac{i}{\alpha} + \frac{1}{i\alpha} + \frac{(1-i)^2}{i v} \right)$$

$$di = -\frac{v dn}{n(n-1)} \left( \frac{1}{\alpha} + \frac{1-i}{v} \right)$$

$$\text{II. } d\beta = -dc = -\frac{\beta\beta db}{i' i' b b} - \frac{\beta\beta dn'}{i'(n'-1)} \left( \frac{i'}{\beta} + \frac{1}{i'\beta} + \frac{(1-i')^2}{i' v'} \right)$$

$$di' = -\frac{v' db}{n' b b} - \frac{v' dn'}{n'(n'-1)} \left( \frac{1}{\beta} + \frac{1-i'}{v'} \right)$$

$$\text{III. } d\gamma = -dc = -\frac{\gamma\gamma dc}{i'' i'' c c} - \frac{\gamma\gamma dn''}{i''(n''-1)} \left( \frac{i''}{\gamma} + \frac{1}{i''\gamma} + \frac{(1-i'')^2}{i'' v''} \right)$$

$$di'' = -\frac{v'' dc}{n'' c c} - \frac{v'' dn''}{n''(n''-1)} \left( \frac{1}{\gamma} + \frac{1-i''}{v''} \right)$$

etc.

ex quibus formulis etiam mutationes singularum imaginum ac proinde etiam tandem vltimae imaginis facile



facile definiri poterunt, seu potius loco imaginum angulos, sub quibus eae oculo ad iustam distantiam  $l$  posita sint adpariturae, consideremus.

Pro vna lente  $\frac{1}{i} \cdot \frac{\alpha}{a} \cdot \frac{z}{l}$

Pro duabus lentibus  $\frac{1}{i \cdot i'} \cdot \frac{\alpha \beta}{a b} \cdot \frac{z}{l}$

Pro tribus lentibus  $\frac{1}{i \cdot i' \cdot i''} \cdot \frac{\alpha \beta \gamma}{a b c} \cdot \frac{z}{l}$

etc.

quatenus scilicet praeter distantias  $a, \alpha; b, \beta; c, \gamma$  etc. etiam litterae  $i, i', i'', i'''$  etc. sunt variables.

### C o r o l l. I.

297. Hinc igitur per differentiationem definire licet, quanta mutatio in loco vltimae imaginis, quae obiectum visionis constituit, ob diuersam radiorum refrangibilitatem oriri debeat.

### C o r o l l. 2.

298. Deinde cum etiam magnitudinem cuiusque imaginis supra per distantias determinatrices et numeros  $i, i', i''$  etc. definiuerimus, pro magnitudine imaginis habetur  $\frac{1}{i \cdot i' \cdot i''} \cdot \frac{\alpha \beta \gamma \delta}{a b c d} \cdot \frac{z}{l}$  (189). simili modo mutatio assignari potest, quam magnitudo vltimae imaginis ob diuersam refrangibilitatem radiorum patietur.

### C o r o l l. 3.

299. Cognita autem vtraque mutatione, quam vltima imago tam respectu loci, quam magnitudinis subit,

subit, non difficulter colligetur, quanta confusione ipsa visio ob diuersam radiorum refrangibilitatem perturbetur.

## Coroll. 4.

300. Si crassities lentium euanescat fiet

$$d\alpha = -db = -\frac{\alpha \alpha dn}{n-1} \left( \frac{1}{a} + \frac{1}{\alpha} \right)$$

$$d\beta = -dc = -\frac{\beta \beta db}{b\beta} - \frac{\beta \beta dn}{n-1} \left( \frac{1}{b} + \frac{1}{\beta} \right)$$

$$d\gamma = -dd = -\frac{\gamma \gamma dc}{c\gamma} - \frac{\gamma \gamma dn}{n-1} \left( \frac{1}{c} + \frac{1}{\gamma} \right)$$

$$d\delta = -de = -\frac{\delta \delta dd}{d\delta} - \frac{\delta \delta dn}{n-1} \left( \frac{1}{d} + \frac{1}{\delta} \right)$$

etc.

numeri autem  $i$ ,  $i'$ ,  $i''$  etc. abeunt in vnitatem nulloque mutationi amplius sunt obnoxii.

Si ergo crassities lentium euanescat, pro diuersa refractione singularum lentium formulae superiores abibunt in sequentes:

$$\text{I. } d\alpha = -db = -\frac{\alpha \alpha dn}{n-1} \left( \frac{1}{a} + \frac{1}{\alpha} \right)$$

$$\text{II. } d\beta = -dc = -\frac{\beta \beta db}{b\beta} - \frac{\beta \beta dn'}{n'-1} \left( \frac{1}{b} + \frac{1}{\beta} \right)$$

$$\text{III. } d\gamma = -dd = -\frac{\gamma \gamma dc}{c\gamma} - \frac{\gamma \gamma dn''}{n''-1} \left( \frac{1}{c} + \frac{1}{\gamma} \right)$$

etc.

Ceterum per se manifestum est, quando §. 288  $dn$  vel  $+\frac{1}{100}$  vel  $-\frac{1}{100}$  significare dicitur, id tantum de illa vitri specie, pro qua est refractione radiorum

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mediorum  $n = \frac{31}{30}$  esse intelligendum; et pro aliis vitri speciebus differentialia  $dn'$ ,  $dn''$ ,  $dn'''$ , etc. haud mediocriter ab  $\frac{1}{100}$  discrepare posse. Quanta autem futura sit haec diuersitas, optandum esset, vt ea potius experimentis, quam ex theoria quapiam definiretur.

### Scholion

301. Cum igitur ob diuersam radiorum refrangibilitatem cuique imagini duplex alteratio inducatur, quarum altera eius magnitudinem, altera vero eius locum afficit, duplex inde confusio in visionem infertur. Si enim formulae in superioribus capitibus exhibitae ad radios mediae naturae restringantur, pro quibus est  $n = \frac{31}{30}$ , posito  $dn = -\frac{1}{100}$ , ex formulis hic traditis differentialibus locus et magnitudo imaginis a radiis rubris formatae definietur, posito autem  $dn = +\frac{1}{100}$ , locus et magnitudo imaginis violaceae declarabitur. Scilicet si Ii fuerit imago vltima visioni obiecta, quae a radiis mediae naturae formatur, per formulas modo inuentas, prout vel  $dn = -\frac{1}{100}$  vel  $dn = +\frac{1}{100}$  ponatur, definietur tam imago rubra Rg quam violacea Vv; atque ex natura differentialium manifestum est cum interualla IR et IV inter se aequalia esse debere, tum etiam differentias Ii—Rg et Vv—Ii, ita vt oculo series innumerabilium imaginum inter extremas Rg et Vv sitatum simul cernenda offeratur, vnde eo maior confusio oriatur necesse est, quo maior fuerit differentia tam ratione loci quam magnitudinis. Quare haec confusio penitus tolleretur, si eiusmodi lentium dispositio definiri posset, vt tam interuallum RV, quam

Tab. III.  
Fig. 16.

quam differentia inter imagines  $Rg$  et  $Vv$  ad nihilum redigeretur, quod vtrumque nisi simul praestari queat, confusionem perfecte tollere non licet Verumtamen etiamsi neutri harum conditionum satisfieri possit, tamen dabitur pro oculo eiusmodi locus  $O$ , ubi confusio minime sensibilis percipiat, qui erit in concursu rectae  $vg$  productae cum axe: ibi enim omnes extremitates  $giv$  communibus radiis cernentur, neque propterea extremitas obiecti colore tincta apparebit. Quare si simul punctum  $O$  conueniat cum loco oculi idoneo alia confusio non percipietur, nisi quae inde originem trahit, quod forte imagines extremae  $Rg$  et  $Vv$  nimis a distantia iusta discrepent, siquidem media  $I$  ad distantiam iustam ob oculo fuerit remota. Neque tamen hinc ora obiecti coloribus iridis cincta apparebit, cui confusionis speciei maxime est occurrendum; ideoque ea quae adhuc adfuerit confusio facile tolerari poterit, quae vero etiam omnino tolleretur si modo interuallum  $RV$  vel in nihilum redigi vel saltem satis paruum reddi posset. Hinc ergo intelligimus vitium illud, quo obiecta coloribus iridis circumdata saepe repraesentantur, non tam necessario cum instrumentis dioptricis esse coniunctum, vt nullo pacto ab iis separari queat, quamobrem eo magis operae erit pretium, vt investigemus quomodo haec instrumenta ab isto vitio liberari possint. Quae tota investigatio huc redit, vt determinetur punctum  $O$ , ubi recta per terminos imaginum  $v$ ,  $i$ ,  $g$  ducta cum axe concurrat hocque punctum cum loco oculi iam



supra definito conueniens reddatur, si quidem fieri potest: vnde perspicitur locum oculi O hac proprietate praeditum esse oportere, vt angulus, sub quo vltima imago cernitur ob variabilitatem numero  $n$  tributam nullam mutationem patiatur. Tum vero insuper videndum erit, num interualla IR et IV vel ad nihilum reduci, vel minima reddi queant.

### Problema 4.

Tab. III.  
Fig. 12.

302. Proposita vnica lente definire locum oculi, vnde obiectum sine margine colorato cernatur.

### Solutio.

Sit obiecti E $\epsilon$  ante lentem distantia EA =  $a$ , imago vero per radios mediae naturae in F $\zeta$  repraesentetur, ponaturque aF =  $\alpha$ . Pro lente vero sit eius crassities Aa =  $v$ , et quantitas arbitraria =  $k$ , vnde capiatur  $\frac{k-v}{k+v} = i$ . Hinc posito E $\epsilon$  =  $z$  erit F $\zeta$  =  $\frac{1}{i} \cdot \frac{\alpha}{a} z$  (86): quare si pro loco oculi statuatur distantia aO =  $O$ , quae est fixa, erit OF =  $\alpha - O$ , et anguli FO $\zeta$  tangens =  $\frac{1}{i} \cdot \frac{\alpha}{a} \cdot \frac{z}{\alpha - O}$ , quae formula ob diuersam radiorum refrangibilitatem nullam mutationem subire debet. Inde autem quantitates  $\alpha$  et  $i$  tantum variantur, cum reliquae manent constantes. Quare istius formulae differentiale logarithmicum nihilo aequale positum praebet hanc aequationem

$$-\frac{di}{i} + \frac{d\alpha}{\alpha} - \frac{d\alpha}{\alpha - O} = 0 \text{ seu } \frac{-di}{i} - \frac{O d\alpha}{\alpha(\alpha - O)} = 0$$

vbi

vbi si valores supra inuenti substituantur, prodit

$$\frac{v d n}{i n (n-1)} \left( \frac{1}{a} + \frac{1-i}{v} \right) + \frac{O \alpha d n}{i (n-1) (\alpha-O)} \left( \frac{i}{\alpha} + \frac{1}{i a} + \frac{(1-i)^2}{i v} \right) = 0$$

quae aequatio per  $\frac{d n}{i (n-1)}$  diuisa praebet:

$$\frac{v}{n a} + \frac{1-i}{n} + \frac{O \alpha}{\alpha-O} \left( \frac{i}{\alpha} + \frac{1}{i a} + \frac{(1-i)^2}{i v} \right) = 0$$

vnde locus oculi definiri poterit: qui si debeat conuenire cum supra inuento (238), vbi inuenimus

$$O = \frac{-i \alpha v}{n \alpha - i v}, \text{ erit}$$

$\alpha - O = \frac{n \alpha \alpha}{n \alpha - i v}$  et  $\frac{O \alpha}{\alpha - O} = \frac{-i v}{n}$ , hincque nostra aequatio per  $n$  multiplicata abit in:

$$\frac{v}{a} + 1 - i - \frac{i v}{\alpha} - \frac{v}{a} - (1-i)^2 = 0$$

seu  $i - i i - \frac{i v}{\alpha} = 0$ , ideoque  $i = \frac{\alpha}{\alpha + v} = \frac{k-v}{k+v}$ . Quamobrem quantitatem arbitrariam  $k$  ita definiri conueniet vt fit  $k = 2\alpha + v$ : et cum sit  $i = \frac{\alpha}{\alpha + v}$  pro loco oculi habebimus  $O = \frac{-\alpha v}{n \alpha + (n-1)v} = \frac{-20 \alpha v}{31 \alpha + 11 v}$  ob  $n = \frac{31}{20}$ .

Quod si porro hinc variationem in loco imaginis desideremus definiri oportet differentiale  $d\alpha$ , quod fiet:

$$d\alpha = \frac{-\alpha \alpha d n}{i (n-1)} \left( \frac{i}{\alpha} + \frac{1}{i a} + \frac{(1-i)^2}{i v} \right)$$

et pro  $i$  posito valore  $\frac{\alpha}{\alpha + v}$

$$d\alpha = \frac{-(\alpha + v) d n}{n-1} \left( 1 + \frac{\alpha + v}{a} \right)$$

qui valor si ad nihilum redigi posset, confusio omnis a diuersa radiorum refrangibilitate oriunda perfecte tolleretur.

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Coroll.



## Coroll. 1.

303. Pro lentis ergo constructione quantitas arbitraria  $k$  ita accipi debet, vt sit  $k = 2\alpha + v$ ; atque tum oculus in eo loco constitutus, vbi totum campum apparentem percipiat, simul nullam confusionem a diuersa radiorum indole sentiet.

## Coroll. 2.

304. Vt autem oculus simul imaginem in distantia iusta aspiciat, oportet sit  $\alpha - O = -l$ , ideoque  $l = \frac{-\frac{1}{2}\alpha + \frac{1}{2}v}{\frac{1}{2}\alpha + \frac{1}{2}v}$ . Vnde colligitur:  $\alpha = \frac{-l}{2} - \frac{v}{2} - V(\frac{1}{4}ll + \frac{9}{64}vl + \frac{1}{4}vv)$ , hincque  $O = \frac{1}{2}l - \frac{1}{2}v - V(\frac{1}{4}ll + \frac{9}{64}vl + \frac{1}{4}vv)$  et  $k = -l - 2V(\frac{1}{4}ll + \frac{9}{64}vl + \frac{1}{4}vv)$

## Coroll. 3.

305. Potest vero insuper effici, vt etiam  $d\alpha$  euanescat, quod euenit si  $\alpha + \alpha + v = 0$ , hoc est  $\alpha = \frac{1}{2}l - \frac{1}{2}v + V(\frac{1}{4}ll + \frac{9}{64}vl + \frac{1}{4}vv)$ . Verum cum hoc casu ob  $\alpha = -\alpha - v$  imago in ipsum obiectum cadet, ita, vt radii nullam refractionem pati sint censendi, visio per lentem perinde erit comparata atque nudis oculis.

## Coroll. 4.

306. Si crassities lentis  $v$  plane euanescat, tam ob campum apparentem, quam diuersam radiorum refrangibilitatem sit  $O = 0$  hoc ergo casu oculus lenti immediate applicatus nullam confusionem ob diuersam radiorum naturam percipiet. Dum ergo fuerit  $\alpha = -l$  visio erit distincta. Scho-

## Scholion.

307. Hic scilicet penitus mentem abstrahimus a confusione iam supra determinata, quae a lentium apertura oritur, ideoque aperturam primae lentis ut euanescentem spectamus. Eam igitur hic tantum confusionis speciem contemplamur, quae a diuersa radiorum refrangibilitate originem ducit; quam plerumque tolli obseruauimus, si angulus ad O inuariabilis reddatur; tum enim ora obiecti satis bene terminata conspicietur, neque coloribus iridis cincta. Interim tamen adhuc aliqua confusio sentiri poterit inde oriunda, quod si imago media iustam ab oculo distantiam teneat, imaginum extremarum altera sit nimis propinqua alter animis remota, verum si earum interuallum non sit admodum magnum, confusio haec parum erit sensibilis. Ita hic inuenimus, quod experientia satis comprobatur, si obiecta per vnicam lentem spectemus, ea margine colorato destituta apparere dummodo oculus immediate applicetur quod si quando secus euenire videatur, causa aperturae lentis sine dubio erit tribuenda, cui conditioni rationes hic allegatae refragantur.

## Problema 5.

308. Si instrumentum dioptricum duabus instructum sit lentibus, definire locum oculi, unde obiectum sine margine colorato videatur.

Tab. III.

Fig. 13.

Solutio



## Solutio.

Positi obiecti distantia  $AE = a$ , sint pro radiis mediae naturae reliquae distantiae determinatrices  $aF = \alpha$ ,  $BF = b$  et  $bG = \xi$  crassities vero lentium  $Aa = v$ ,  $Bb = v'$ , et distantiae arbitrariae  $k$  et  $k'$ , ponaturque  $\frac{k-v}{k+v} = i$  et  $\frac{k'-v'}{k'+v'} = i'$ . His positis si magnitudo obiecti  $E\varepsilon$  vocetur  $= z$ , erit imago  $G\eta = \frac{1}{i'} \cdot \frac{\alpha \xi}{ab} z$ , unde si oculi distantia ponatur  $bO = O$ ,

ob  $GO = \xi - O$ , erit anguli  $GO\eta$  tangens  $\frac{1}{i'} \cdot \frac{\alpha \xi}{ab} \cdot \frac{z}{\xi - O}$  cuius differentiale logarithmicum nihilo aequatum praebet,

$$-\frac{di}{i} - \frac{di'}{i'} + \frac{d\alpha}{\alpha} - \frac{db}{b} + \frac{d\xi}{\xi} - \frac{d\xi}{\xi - O} = 0$$

quae valoribus supra (296) inuentis substitutis abit in:

$$\frac{v}{in(1-i)} \left( \frac{1}{a} + \frac{1-i}{v} \right) + \frac{v'}{n'i'bb} + \frac{v'}{i'n'(n'-1)} \left( \frac{1}{b} + \frac{1-i'}{v'} \right) - \frac{db}{\alpha} - \frac{db}{b} + \frac{O}{\xi - O} \left( \frac{\xi db}{i'bb} + \frac{\xi dn'}{i'(n'-1)} \left( \frac{i'}{\xi} + \frac{1}{i'b} + \frac{(1-i')^2}{i'v'} \right) \right) = 0$$

Verum conditio campi exigebat  $O = \frac{\xi \xi}{\xi + \frac{1}{i'} \cdot \frac{\alpha \xi}{ab} z}$ , existente

$$\xi = \left( \frac{i'}{i} \cdot \frac{\alpha + b}{a} - \frac{i'bv}{na\alpha} - \frac{1}{i} \cdot \frac{\alpha v'}{n'ab} \right) z, \quad (245), \text{ unde fit}$$

$$\frac{O}{\xi - O} = \frac{\xi \xi}{\frac{1}{i'} \cdot \frac{\alpha \xi \xi}{ab} z} = \frac{i'ab}{\alpha \xi} \left( \frac{i'}{i} \cdot \frac{\alpha + b}{a} - \frac{i'bv}{na\alpha} - \frac{1}{i} \cdot \frac{\alpha v'}{n'ab} \right)$$

Nunc

Nunc vero est  $db = \frac{\alpha \alpha dn}{i(n-1)} \left( \frac{i}{\alpha} + \frac{1}{ia} + \frac{(1-i)^2}{iv} \right)$ , quem valorem antequam substituamus, transformemus aequationem nostram in hanc formam

$$\frac{dn}{n-1} \left( \frac{v}{ina} + \frac{1-i}{in} + \frac{v'}{i'nb} + \frac{1-i'}{i'n} + \frac{\epsilon \circ}{i'(\epsilon-0)} \left( \frac{i'}{\epsilon} + \frac{1}{i'b} + \frac{(1-i')^2}{i'v'} \right) \right. \\ \left. + db \left( \frac{v'}{ni'bb} - \frac{1}{a} - \frac{1}{b} + \frac{\epsilon \circ}{i'i'bb(\epsilon-0)} \right) \right) = 0$$

vbi posterius membrum abit in  $-\frac{iv}{n\alpha\alpha} db$ , tum vero erit

$$0 = \frac{dn}{n-1} \left\{ 1 + \frac{b}{\alpha} + \frac{i'ib}{\epsilon} \left( 1 + \frac{b}{\alpha} \right) + \frac{(1-i')^2 b(\alpha+b)}{\alpha v'} + \frac{2-i-i'}{n} \right. \\ \left. - \frac{iv}{n\alpha} - \frac{i'v'}{n\epsilon} - \frac{ibv}{n\alpha\alpha} - \frac{ii'i'bbv}{n\alpha\alpha\epsilon} - \frac{i(1-i')^2 bbv}{n\alpha\alpha v'} \right\}$$

distinguendo  $n'$  ab  $n$  erit

$$\frac{dn}{n-1} \left( \frac{iv}{\alpha n} - \frac{(1-i)}{n} \right) = \frac{dn'}{n'-1} \left\{ 1 + \frac{b}{\alpha} + \frac{i'ib}{\epsilon} \left( 1 + \frac{b}{\alpha} \right) - \frac{ii'i'bbv}{n\alpha\alpha\epsilon} - \frac{i'v'}{n'\epsilon'} + \frac{1-i'}{n'} \right. \\ \left. - \frac{ibv}{n\alpha\alpha} + \frac{i(1-i')^2}{v'} \left( 1 + \frac{b}{\alpha} \right) - \frac{i(1-i')^2 bbv}{n\alpha\alpha v'} \right\}$$

cuius aequationis complicatio obstat, quominus quicquam commode inde concludi possit.

### COROLL. I.

309. Si ambae lentes crassitie careant, vt sit  $v=0$ ,  $v'=0$ , et  $i=i'=1$ , aequatio differentialis prima est

$$\frac{d\alpha}{\alpha} - \frac{db}{b} - \frac{0 d\epsilon}{\epsilon(\epsilon-0)} = 0$$

tum vero:

$$d\alpha = -db = -\frac{\alpha \alpha dn}{n-1} \left( \frac{1}{\alpha} + \frac{1}{a} \right) \text{ et}$$

$$d\epsilon = -\frac{\epsilon \epsilon db}{bb} - \frac{\epsilon \epsilon dn}{n-1} \left( \frac{1}{\epsilon} + \frac{1}{b} \right) = -\frac{dn}{n-1} \left( \frac{\alpha \epsilon \epsilon}{bb} \left( \frac{1}{\alpha} + \frac{1}{a} \right) + \epsilon \epsilon \left( \frac{1}{\epsilon} + \frac{1}{b} \right) \right)$$

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quibus valoribus substitutis et per  $\frac{dn}{n-1}$  diuisione facta fit

$$-a\alpha\left(\frac{1}{\alpha}+\frac{1}{a}\right)\left(\frac{1}{\alpha}+\frac{1}{b}\right)+\frac{0}{\varepsilon-0}\left(\frac{\alpha\alpha}{b\bar{b}}\left(\frac{1}{\alpha}+\frac{1}{a}\right)+\frac{1}{\varepsilon}+\frac{1}{b}\right)=0$$

Vnde si oculus lenti posteriori immediate applicaretur  
vt esset

$$0=0, \text{ deberet esse } \left(\frac{1}{\alpha}+\frac{1}{a}\right)\left(\frac{1}{\alpha}+\frac{1}{b}\right)=0.$$

### Coroll. 2.

310. Verum in eadem hypothesi, vt locus  
oculi congruat cum eo, quem visio campi exigit,  
debet esse  $O=\frac{b\varepsilon(\alpha+b)}{b(\alpha+b)+\alpha\varepsilon}$ , vnde fit  $\varepsilon-O=\frac{\alpha\varepsilon\varepsilon}{b(\alpha+0)+\alpha\varepsilon}$ ,

$$\text{ideoque } \frac{0}{\varepsilon-0}=\frac{b(\alpha+b)}{\alpha\varepsilon}=\frac{bb}{\varepsilon}\left(\frac{1}{\alpha}+\frac{1}{b}\right)$$

quo valore substituto nostra aequatio erit

$$0=\left(\frac{1}{\alpha}+\frac{1}{b}\right)\left(-a\alpha\left(\frac{1}{\alpha}+\frac{1}{a}\right)+bb\left(\frac{\alpha\alpha}{b\bar{b}}\left(\frac{1}{\alpha}+\frac{1}{a}\right)+\frac{1}{\varepsilon}+\frac{1}{b}\right)\right)$$

quae reducitur ad hanc formam

$$0=bb\left(\frac{1}{\alpha}+\frac{1}{b}\right)\left(\frac{1}{\varepsilon}+\frac{1}{b}\right)$$

distinguendo  $n'$  ab  $n$ , membra a  $dn$  pendentia se de-  
struunt et oritur

$$0=\frac{dn'}{n'-1}bb\left(\frac{1}{\alpha}+\frac{1}{b}\right)\left(\frac{1}{\varepsilon}+\frac{1}{b}\right)$$

quod ita ostenditur:

Cum aequatio prima differentialis praebeat:

$$\frac{d\alpha}{\alpha}-\frac{db}{b}-\frac{0d\varepsilon}{\varepsilon(\varepsilon-0)}=0 \text{ siue}$$

$$d\alpha\left(\frac{1}{\alpha}+\frac{1}{b}\right)-\frac{0d\varepsilon}{\varepsilon(\varepsilon-0)}=0$$

Cum

Cum igitur ex conditione campi apparentis sit

$$\frac{0}{e-0} = \frac{bb}{e} \left( \frac{1}{a} + \frac{1}{b} \right) \text{ erit}$$

$$d\alpha \left( \frac{1}{a} + \frac{1}{b} \right) - \frac{bb}{e^2} \left( \frac{1}{a} + \frac{1}{b} \right) d\epsilon = 0$$

ideoque loco  $d\epsilon$  suum valorem substituendo

$$d\alpha \left( \frac{1}{a} + \frac{1}{b} \right) - \frac{bb}{e^2} \left( \frac{1}{a} + \frac{1}{b} \right) \left( \frac{ee d\alpha}{b\alpha} - \frac{e^2 dn'}{n'-1} \left( \frac{1}{b} + \frac{1}{e} \right) \right) = 0$$

ubi membra, quae  $d\alpha$  continent, manifesto se destruunt, et tota quaestio ad hanc aequationem perducitur

$$\frac{dn'}{n'-1} \cdot bb \left( \frac{1}{a} + \frac{1}{b} \right) \left( \frac{1}{e} + \frac{1}{b} \right) = 0.$$

### Coroll. 3.

311. Quod si ergo huic conditioni satisfieri possit, obiectum sine margine colorato apparebit: praeterea vero confusio penitus tolleretur, si reddi liceret  $d\epsilon = 0$ , quod fit per hanc aequationem

$$\frac{\alpha\alpha}{bb} \left( \frac{1}{a} + \frac{1}{a} \right) + \frac{1}{e} + \frac{1}{b} = 0 \text{ siue}$$

$$+ \frac{dn}{n-1} \cdot \frac{\alpha\alpha}{bb} \left( \frac{1}{a} + \frac{1}{a} \right) = 0$$

$$+ \frac{dn'}{n'-1} \left( \frac{1}{e} + \frac{1}{b} \right) = 0$$

$$\alpha\alpha \left( \frac{1}{a} + \frac{1}{a} \right) + bb \left( \frac{1}{e} + \frac{1}{b} \right) = 0.$$

### Coroll. 4.

312. Priori autem aequationi satisfieri nequit nisi fuerit vel  $a+b=0$  vel  $\frac{1}{e} + \frac{1}{b} = 0$ . Illo casu ambae lentes coniungerentur, ut unicam constituerent, hoc vero posterioris distantia focalis fieret infinita, qui



casus iterum ad casum vnicae lentis rediret; foret enim  $O = -\alpha - b$ , ob  $\xi = -b$ , et oculus priori lenti immediate applicari deberet.

### Scholion.

313. Si simili modo has inuestigationes ad plures lentes extendere vellemus, non neglecta earum crassitie in formulas plane inextricabiles delaberemur vnde vix quicquam concludi posset. Verum quia in omnibus fere instrumentis dioptricis, praecipue quae pluribus lentibus constant, iis tam exigua crassities tribui solet, vt sine notabili errore pro nihilo haberi possit, tam taediosae indagationi facile supersedere poterimus. Ad quod accedit, quod hic non de summo rigore geometrico agatur, sed contenti esse queamus, dummodo hanc confusionem satis prope cognouerimus: ex quo sufficiet in consideratione plurium lentium earum crassitiem prorsus neglexisse.

### Supplementum IV.

Si ratio refractionis in singulis lentibus sit diuersa solutio sequenti modo absoluetur:

I. Prima aequatio differentialis prorsus se habebit, vt in problemate, ita, vt sit

$$d\alpha\left(\frac{1}{\alpha} + \frac{1}{b}\right) + d\xi\left(\frac{1}{\xi} + \frac{1}{c}\right) - \frac{O d\gamma}{\gamma(\gamma - O)} = 0$$

et quia etiam, vt ante, est

$$\frac{O}{\gamma - O} = \frac{b b c c}{\alpha \xi^2 \gamma} \left(1 + \frac{\alpha}{b}\right) + \frac{c c}{\xi \gamma} \left(1 + \frac{\xi}{c}\right)$$

erit

erit nostra aequatio

$$\left(\frac{1}{a} + \frac{1}{b}\right)(d\alpha - \frac{bbcc}{e^2\gamma^2} d\gamma) + \left(\frac{1}{e} + \frac{1}{c}\right)(d\epsilon - \frac{cc}{\gamma\gamma} d\gamma) = 0.$$

II. Nunc autem ratio diuersae refractionis est habenda; vnde in superioribus additamentis inuenimus esse

$$d\alpha = -\frac{\alpha dn}{n-1} \left(\frac{1}{a} + \frac{1}{\alpha}\right)$$

$$d\epsilon = \frac{ee d\alpha}{b b} - \frac{ee dn'}{n'-1} \left(\frac{1}{b} + \frac{1}{e}\right)$$

$$d\gamma = \frac{\gamma\gamma d\epsilon}{cc} - \frac{\gamma\gamma dn''}{n''-1} \left(\frac{1}{c} + \frac{1}{\gamma}\right)$$

Hincque ergo fiet

$$d\alpha - \frac{bbcc}{e^2\gamma^2} d\gamma = d\alpha - \frac{bbd\epsilon}{e^2} + \frac{bbcc dn''}{e^2(n''-1)} \left(\frac{1}{c} + \frac{1}{\gamma}\right)$$

$$= \frac{bb \cdot d\alpha'}{n'-1} \left(\frac{1}{b} + \frac{1}{e}\right) + \frac{bbcc dn''}{e^2(n''-1)} \left(\frac{1}{c} + \frac{1}{\gamma}\right)$$

Deinde

$$d\epsilon - \frac{cc}{\gamma\gamma} d\gamma = \frac{cc dn''}{n''-1} \left(\frac{1}{c} + \frac{1}{\gamma}\right)$$

III. Ex his ergo nostra aequatio differentialis abibit in hanc formam:

$$\left(\frac{1}{a} + \frac{1}{b}\right) \left( \frac{bb dn'}{n'-1} \left[\frac{1}{b} + \frac{1}{e}\right] + \frac{bbcc^2 dn''}{e^2(n''-1)} \left[\frac{1}{c} + \frac{1}{\gamma}\right] \right)$$

$$+ \left(\frac{1}{e} + \frac{1}{c}\right) \left( \frac{cc dn''}{n''-1} \left[\frac{1}{c} + \frac{1}{\gamma}\right] \right)$$

siue

$$\frac{dn'}{n'-1} \cdot bb \cdot \left(\frac{1}{a} + \frac{1}{b}\right) \left(\frac{1}{e} + \frac{1}{b}\right)$$

$$+ \frac{dn''}{n''-1} \cdot cc \left(\frac{1}{c} + \frac{1}{\gamma}\right) \left( \frac{bb}{e^2} \left(\frac{1}{a} + \frac{1}{b}\right) + \frac{1}{e} + \frac{1}{c} \right) = 0$$

IV. Pro casu autem illo singulari, quo oculus lenti vltimae immediate, debet adplicari, ob  $O=0$ , habebitur simpliciter haec aequatio

$$d\alpha \left(\frac{1}{a} + \frac{1}{b}\right) + d\beta \left(\frac{1}{e} + \frac{1}{c}\right) = 0$$

Hh 3

siue



siue, substituto valore  $d\mathcal{E}$

$$\begin{aligned} d\alpha\left(\frac{1}{\alpha} + \frac{1}{b}\right) + \frac{\mathcal{E}\mathcal{E}d\alpha}{bb}\left(\frac{1}{\mathcal{E}} + \frac{1}{c}\right) \\ - \frac{\mathcal{E}\mathcal{E}dn'}{n'-1}\left(\frac{1}{b} + \frac{1}{\mathcal{E}}\right) = 0. \\ d\alpha\left(\frac{1}{\alpha} + \frac{1}{b} + \frac{\mathcal{E}\mathcal{E}}{bb}\left[\frac{1}{\mathcal{E}} + \frac{1}{c}\right]\right) \\ - \frac{\mathcal{E}^2dn'}{n'-1}\left(\frac{1}{b} + \frac{1}{\mathcal{E}}\right) = 0, \end{aligned}$$

seu tandem

$$\begin{aligned} + \frac{dn}{n-1}\alpha\alpha\left(\frac{1}{b} + \frac{1}{\alpha}\right)\left(\frac{1}{\alpha} + \frac{1}{b} + \frac{\mathcal{E}^2}{b^2}\left[\frac{1}{\mathcal{E}} + \frac{1}{c}\right]\right) \\ + \frac{dn'}{n'-1}\mathcal{E}\mathcal{E}\left(\frac{1}{b} + \frac{1}{\mathcal{E}}\right) = 0. \end{aligned}$$

V. Hoc modo tantum margo coloratus tollitur; ut autem tota confusio tollatur, quod fit si  $d\gamma = 0$ , insuper satisfieri debet huic aequationi

$$\begin{aligned} 0 = \frac{dn}{n-1} \cdot \frac{\alpha\alpha\mathcal{E}\mathcal{E}\gamma\gamma}{bbcc}\left(\frac{1}{a} + \frac{1}{\alpha}\right) \\ + \frac{dn'}{n'-1} \cdot \frac{\mathcal{E}\mathcal{E}\gamma\gamma}{cc}\left(\frac{1}{b} + \frac{1}{\mathcal{E}}\right) \\ + \frac{dn''}{n''-1} \cdot \gamma\gamma \cdot \left(\frac{1}{c} + \frac{1}{\gamma}\right). \end{aligned}$$

### Problema 6.

Tab. III.

Fig. 14.

314. Si instrumentum dioptricum tribus constet lentibus, quarum crassities euanescat, eam definire dispositionem, ut oculus in eo loco, quem campus postulat, constitutus obiectum sine margine colorato conspiciat.

### Solutio.

Posita ergo distantia obiecti ante lentem obiectiuam  $AE = a$ , eiusque magnitudine  $E\varepsilon = z$ , vo-

centur

centur distantiae imaginum a radiis mediae naturae formatarum, vt haecenus

$aF = \alpha$ ,  $BF = b$ ;  $bG = \mathfrak{e}$ ,  $CG = c$ , et  $cH = \gamma$ .  
eritque imago vltima  $H\theta = \frac{\alpha\mathfrak{e}\gamma}{abc}z$ , et posita oculi post lentem vltimam distantia  $cO = O$ , erit  $OH = \gamma - O$  et anguli  $HO\theta$  tangens  $= \frac{\alpha\mathfrak{e}\gamma}{abc} \cdot \frac{z}{\gamma - O}$ , quae debet esse invariabilis. Posito ergo eius differentiali logarithmico  $= 0$  habebimus hanc aequationem:

$$\frac{d\alpha}{\alpha} - \frac{db}{b} + \frac{d\mathfrak{e}}{\mathfrak{e}} - \frac{dc}{c} + \frac{d\gamma}{\gamma} - \frac{d\gamma}{\gamma - O} = 0$$

At et formulis supra (300) erutis habemus

$$d\alpha = -db = -\frac{dn}{n-1} \cdot \alpha \alpha \left( \frac{1}{a} + \frac{1}{a} \right)$$

$$d\mathfrak{e} = -dc = -\frac{\mathfrak{e}\mathfrak{e}dn}{n-1} \left( \frac{\alpha\alpha}{b\mathfrak{e}} \left( \frac{1}{a} + \frac{1}{a} \right) + \frac{1}{b} + \frac{1}{\mathfrak{e}} \right)$$

$$d\gamma = -\frac{\gamma\gamma dn}{n-1} \left( \frac{\alpha\alpha\mathfrak{e}\mathfrak{e}}{b\mathfrak{e}c\gamma} \left( \frac{1}{a} + \frac{1}{a} \right) + \frac{\mathfrak{e}\mathfrak{e}}{c\gamma} \left( \frac{1}{b} + \frac{1}{\mathfrak{e}} \right) + \frac{1}{c} + \frac{1}{\gamma} \right)$$

vnde nostra aequatio erit

$$d\alpha \left( \frac{1}{a} + \frac{1}{b} \right) + d\mathfrak{e} \left( \frac{1}{\mathfrak{e}} + \frac{1}{c} \right) - \frac{O d\gamma}{\gamma(\gamma - O)} = 0$$

Sed ob campum apparentem supra 254 inuenimus:

$$\mathfrak{B} = \mathfrak{e} = b \left( 1 + \frac{\alpha}{b} \right) \frac{z}{a}$$

$$\mathfrak{C} = c = \frac{bc}{\mathfrak{e}} \left( 1 + \frac{\alpha}{b} \right) \frac{z}{a} + \frac{ac}{b} \left( 1 + \frac{\mathfrak{e}}{c} \right) \frac{z}{a}$$

hincque

$$O = \frac{\gamma c}{c + H\theta} \text{ et } \frac{O}{\gamma - O} = \frac{c}{H\theta} \text{ vnde fit ob}$$

$$H\theta = \frac{\alpha\mathfrak{e}\gamma}{bc} \cdot \frac{z}{a}$$

$$\frac{O}{\gamma - O} = \frac{b\mathfrak{e}c\gamma}{\alpha\mathfrak{e}\mathfrak{e}\gamma} \left( 1 + \frac{\alpha}{b} \right) + \frac{cc}{\mathfrak{e}\gamma} \left( 1 + \frac{\mathfrak{e}}{c} \right) \text{ seu}$$

$$\frac{O}{\gamma(\gamma - O)} = \frac{b\mathfrak{e}c\gamma}{\mathfrak{e}\mathfrak{e}\gamma\gamma} \left( \frac{1}{\alpha} + \frac{1}{b} \right) + \frac{cc}{\gamma\gamma} \left( \frac{1}{\mathfrak{e}} + \frac{1}{c} \right)$$

Valo-



Valoribus iam his substitutis habebimus:

$$\left(\frac{1}{a} + \frac{1}{b}\right) \left(d\alpha - \frac{bbcc}{ee\gamma\gamma} d\gamma\right) + \left(\frac{1}{e} + \frac{1}{c}\right) \left(d\epsilon - \frac{cc}{\gamma\gamma} d\gamma\right) = 0$$

At est

$$d\alpha - \frac{bbcc}{ee\gamma\gamma} d\gamma = \frac{dn}{n-1} \left(bb\left(\frac{1}{b} + \frac{1}{e}\right) + \frac{bbcc}{ee} \left(\frac{1}{c} + \frac{1}{\gamma}\right)\right)$$

$$d\epsilon - \frac{cc}{\gamma\gamma} d\gamma = \frac{dn}{n-1} (cc\left(\frac{1}{c} + \frac{1}{\gamma}\right))$$

Quare facta diuisione per  $\frac{dn}{n-1}$  nanciscemur hanc aequationem:

$$bb\left(\frac{1}{a} + \frac{1}{b}\right)\left(\frac{1}{b} + \frac{1}{e} + \frac{cc}{ee}\left(\frac{1}{c} + \frac{1}{\gamma}\right)\right) + cc\left(\frac{1}{e} + \frac{1}{c}\right)\left(\frac{1}{c} + \frac{1}{\gamma}\right) = 0$$

Quodsi vero oculus lenti postremae immediate applicetur, seu sit  $O=0$ , conditio praescripta hanc postulat aequationem:

$$aa\left(\frac{1}{a} + \frac{1}{b}\right)\left(\frac{1}{a} + \frac{1}{e}\right) + ee\left(\frac{1}{e} + \frac{1}{c}\right)\left(\frac{aa}{bb}\left(\frac{1}{a} + \frac{1}{e}\right) + \frac{1}{b} + \frac{1}{e}\right) = 0$$

Confusio vero a diuersa radiorum refrangibilitate oriunda perfecte tolletur si praeterea fuerit  $d\gamma=0$  seu

$$\frac{aaee}{bbcc}\left(\frac{1}{a} + \frac{1}{e}\right) + \frac{ee}{cc}\left(\frac{1}{b} + \frac{1}{e}\right) + \frac{1}{c} + \frac{1}{\gamma} = 0.$$

### COROLL. I.

315. Si rationes aperturarum lentium in computum ducantur, eaque pro lente secunda ponatur  $=\pi$ , pro tertia  $=\pi'$  erit  $bN' = \frac{\pi b e}{b+e}$  et  $CM'' = cN'' = \frac{\pi' c \gamma}{c+\gamma}$ . Tum vero posito  $\frac{z}{a} = \Phi$ , erit  $G\eta = \frac{ae}{ab} a\Phi$ , et  $H\theta = \frac{ae\gamma}{ab c} \cdot a\Phi$ . Hinc fiet  $\frac{O}{\gamma-0} = \frac{cN''}{H\theta} = \frac{\pi' a b c c}{ae(c+\gamma)a\Phi}$  et  $\frac{O}{\gamma(\gamma-0)} = \frac{\pi' a b c c}{ae\gamma(c+\gamma)} \cdot \frac{1}{a\Phi}$ .

Coroll.

## Coroll. 2.

316. Quodsi porro vt supra ponatur

$$a = Aa, \text{ } \mathfrak{e} = Bb, \gamma = Cc \text{ erit } \frac{0}{\gamma(\gamma-1)} = \frac{\pi'}{ABC(C+1)a\Phi};$$

$$\text{tum vero } bN' = \frac{\pi Bb}{B+1} \text{ et } CM'' = \frac{\pi' Cc}{C+1} : \text{ ac } G\eta = ABa\Phi;$$

Cum iam fit

$$bN' + G\eta : bG = CM'' - G\eta : CG \text{ erit } \frac{bN' + G\eta}{bG} = \frac{CM'' - G\eta}{CG}$$

ideoque

$$\frac{1}{bG} + \frac{1}{CG} = \frac{CM''}{CG \cdot G\eta} - \frac{bN'}{bG \cdot G\eta} \text{ hoc est } \frac{1}{\mathfrak{e}} + \frac{1}{c} = \frac{\pi' C}{AB(C+1)a\Phi} - \frac{\pi}{AB(B+1)a\Phi}$$

$$\text{simili vero modo est } \frac{1}{a} + \frac{1}{b} = \frac{\pi B}{A(B+1)a\Phi}.$$

## Coroll. 3.

317. Per easdem substitutiones fit

$$\frac{1}{a} + \frac{1}{\alpha} = \frac{A+1}{Aa}; \quad \frac{1}{b} + \frac{1}{\mathfrak{e}} = \frac{B+1}{Bb}; \quad \frac{1}{c} + \frac{1}{\gamma} = \frac{C+1}{Cc} \text{ etc.}$$

hincque :

$$a\alpha \left( \frac{1}{a} + \frac{1}{\alpha} \right) = A(A+1)a$$

$$\mathfrak{e}\mathfrak{e} \left( \frac{1}{b} + \frac{1}{\mathfrak{e}} \right) = B(B+1)b$$

$$\gamma\gamma \left( \frac{1}{c} + \frac{1}{\gamma} \right) = C(C+1)c \text{ sicque porro.}$$

## Coroll. 4.

318. His ergo nouis denominationibus introductis differentialia ex §. 300 ita exprimentur

$$da = -db = -\frac{dn}{n-1} \cdot A(A+1)a$$

$$d\mathfrak{e} = -dc = -\frac{dn}{n-1} \cdot B(B+1)b$$

$$d\gamma = -d.c = -\frac{dn}{n-1} \cdot C(C+1)c$$

quae formulae commodius in calculum introducentur.



Quemadmodum hic nouae formae adhibentur in sequentibus vsurpandae; ita et pro casu diuersae refractionis sequentibus formulis in posterum vti licebit:

$$\begin{aligned}\frac{0}{\gamma(\gamma-0)} &= \frac{\pi'}{ABC(C+1)a\Phi} \\ \frac{\frac{1}{a} + \frac{1}{b}}{\frac{1}{a} + \frac{1}{b}} &= \frac{\pi B}{A(B+1)a\Phi} \\ \frac{\frac{1}{b} + \frac{1}{c}}{\frac{1}{b} + \frac{1}{c}} &= \frac{\pi' C}{AB(C+1)a\Phi} - \frac{\pi}{AB(B+1)a\Phi} \\ \frac{\frac{1}{c} + \frac{1}{d}}{\frac{1}{c} + \frac{1}{d}} &= \frac{\pi'' D}{ABC(D+1)a\Phi} - \frac{\pi'}{ABC(C+1)a\Phi} \\ &\text{etc.}\end{aligned}$$

tum vero formulae differentiales erunt

$$\begin{aligned}d\alpha &= -db = -\frac{dn}{n-1} \cdot A(A+1)a \\ d\beta &= -dc = B^2 d\alpha \frac{-dn'}{n'-1} \cdot B(B+1)b \\ d\gamma &= -dd = C^2 d\beta \frac{-dn''}{n''-1} \cdot C(C+1)c \\ d\delta &= -de = D^2 d\gamma \frac{-dn'''}{n'''-1} \cdot D(D+1)d \\ &\text{etc.}\end{aligned}$$

Atque ex his formulis vt margo coloratus euanescat, satisfieri debet huic aequationi

$$0 = \frac{dn'}{n'-1} \cdot \frac{\pi b}{Aa\Phi} + \frac{dn''}{n''-1} \cdot \frac{\pi' c}{ABa\Phi}$$

Vt autem haec confusio penitus tollatur, fieri debet

$$\begin{aligned}d\gamma &= -\frac{dn}{n-1} \cdot A(A+1)B^2 C^2 a \\ &\quad - \frac{dn'}{n'-1} \cdot B(B+1)C^2 \cdot b \\ &\quad - \frac{dn''}{n''-1} \cdot C(C+1)c = 0.\end{aligned}$$

fiue

siue

$$\begin{aligned} & \frac{dn}{n-1} \cdot \frac{(A+1)z}{A} \\ & + \frac{dn'}{n'-1} \cdot \frac{(B+1)b}{AAB} = 0 \\ & + \frac{dn''}{n''-1} \cdot \frac{(C+1)c}{A^2 B^2 C} \end{aligned}$$

## Problema 7.

319. Si instrumentum dioptricum quatuor constet lentibus quarum crassities negligi queat, eam definire dispositionem ut oculus in eo loco quem campus apparens postulat, constitutus obiectum sine confusione a diuersa radiorum indole oriunda conspiciat.

Tab. III.  
Fig. 15.

## Solutio.

Posita distantia obiecti ante instrumentum  $AE=a$ , eiusque magnitudine  $E\varepsilon=z$ , vocentur distantiae imaginum a radiis mediae naturae formatarum ut haecenus.

$aF=\alpha$ ,  $BF=b$ ;  $bG=\varepsilon$ ,  $CG=c$ ;  $cH=\gamma$ ;  $DG=d$ ;  $dI=\delta$

Tum vero ponamus praeterea

$$\alpha=Aa, \varepsilon=Bb; \gamma=Cc; \delta=Dd$$

atque introducantur rationes aperturarum pro singulis lentibus post primam, quae sint  $\pi$  pro secunda  $QQ$ ,  $\pi'$  pro tertia  $RR$  et  $\pi''$  pro quarta  $SS$ , sumto pro campo  $\frac{z}{a}=\Phi$ . His positae erunt imagines  $F\zeta=Aa\Phi$ ;  $G\eta=ABa\Phi$ ;  $H\theta=ABCa\Phi$  et  $I_1=ABCDa\Phi=\frac{\alpha\varepsilon\gamma\delta}{a b c d}z$ .

I i 2

Iam



Iam posita oculi distantia post instrumentum  $dO=O$ ,  
 ut sit  $OI=\delta-O$  erit anguli  $IO$  tangens  $\frac{a\gamma\delta}{abc a} \frac{z}{\delta-O}$ ,  
 quae cum ob diuersam radiorum refrangibilitatem  
 immutata manere debeat, differentiata dabit hanc  
 aequationem

$$\frac{d\alpha}{a} - \frac{db}{b} + \frac{d\epsilon}{c} - \frac{dc}{c} + \frac{d\gamma}{\gamma} - \frac{dd}{d} - \frac{O d\delta}{\delta(\delta-O)} = 0$$

quae ob  $db=-d\alpha$ ,  $dc=-d\epsilon$  et  $dd=-d\gamma$  abit in hanc

$$d\alpha\left(\frac{1}{a} + \frac{1}{b}\right) + d\epsilon\left(\frac{1}{c} + \frac{1}{d}\right) + d\gamma\left(\frac{1}{\gamma} + \frac{1}{d}\right) - \frac{O d\delta}{\delta(\delta-O)} = 0$$

Verum modo ante notauimus fore (318)

$$d\alpha = \frac{-dn}{n-1} \cdot A(A+1)a$$

$$d\epsilon = BB d\alpha - \frac{dn}{n-1} \cdot B(B+1)b$$

$$d\gamma = CC d\epsilon - \frac{dn}{n-1} \cdot C(C+1)c$$

$$d\delta = DD d\gamma - \frac{dn}{n-1} \cdot D(D+1)d$$

vbi pro  $b, c, d$ , valores §. 266 assignati substitui  
 debent. Porro vero iam animaduertimus esse (316)

$$\frac{1}{a} + \frac{1}{b} = \frac{1}{Aa\Phi} \cdot \frac{\pi B}{B+1}$$

$$\frac{1}{c} + \frac{1}{d} = \frac{1}{ABa\Phi} \left( \frac{\pi' C}{C+1} - \frac{\pi}{B+1} \right)$$

$$\frac{1}{\gamma} + \frac{1}{d} = \frac{1}{ABCa\Phi} \left( \frac{\pi'' D}{D+1} - \frac{\pi'}{C+1} \right)$$

atque

$$\frac{O}{\delta(\delta-O)} = \frac{\pi'''}{ABCD(D+1)a\Phi}$$

Quod

Quodsi iam priores valores in posterioribus successiue substituantur, habebimus

$$d\alpha = \frac{-dn}{n-1} \cdot A(A+1)a$$

$$d\beta = \frac{-dn}{n-1} \cdot B(AB(A+1)a + (B+1)b)$$

$$d\gamma = \frac{-dn}{n-1} \cdot C(AB^2C(A+1)a + BC(B+1)b + (C+1)c)$$

$$d\delta = \frac{-dn}{n-1} \cdot D(AB^2C^2D(A+1)a + BC^2D(B+1)b + CD(C+1)c + (D+1)d)$$

His igitur valoribus in aequatione differentiali substitutis, et diuisione per  $\frac{dn}{(n-1)^2\Phi}$  facta aequatio nostra secundum singula membra distributa erit

$$-\frac{\pi B}{(B+1)} \cdot (A+1)a$$

$$-\frac{1}{A} \left( \frac{\pi' C}{C+1} - \frac{\pi}{B+1} \right) (AB(A+1)a + (B+1)b)$$

$$-\frac{1}{AB} \left( \frac{\pi'' D}{D+1} - \frac{\pi'}{C+1} \right) (AB^2C(A+1)a + BC(B+1)b + (C+1)c)$$

$$+\frac{1}{ABC \cdot D+1} (AB^2C^2D(A+1)a + BC^2D(B+1)b + CD(C+1)c + (D+1)d) = 0$$

Terna autem priora membra negatiua sola collecta praebent

$$-\frac{BCD(A+1)\pi''}{D+1} a - \frac{CD(B+1)\pi''}{A(D+1)} b - \frac{D(C+1)\pi''}{AB(D+1)} c$$

$$+\frac{\pi}{A} b + \frac{\pi'}{AB} c$$

quibus si quantum addatur prodit

$$\frac{\pi}{A} b + \frac{\pi'}{AB} c + \frac{\pi''}{ABC} d$$

Iam duo hic casus considerari oportet, alterum quo punctum O post lentem vltimam cadit, alterum



vero, quo ob distantiam  $O$  negatiuam oculus lenti vltimae immediate applicatur. Pro priori casu, quo distantia  $dO = O$  prodit positiua habetur ista aequatio si quidem in aequatione modo inuenta pro  $b, c, d$  valores §. 266 inuenti substituantur

$$\frac{(B+1)\pi}{B\pi-(B+1)\Phi} + \frac{(C+1)\pi'}{C\pi'-(C+1)(\pi-\Phi)} + \frac{(D+1)\pi''}{D\pi''-(D+1)(\pi'-\pi+\Phi)} = 0$$

Pro casu posteriori, quo distantia  $O$  euanescens assumitur facta multiplicatione per  $\frac{D+1}{a\Phi}$ .

$$\frac{BCD(A+1)\pi''}{\Phi} + \frac{CD(B+1)^2\pi''}{E\pi-(B+1)\Phi} + \frac{D(C+1)^2\pi''}{C\pi'-(C+1)(\pi-\Phi)} \\ = \frac{(B+1)(\Gamma+1)\pi}{B\pi-(B+1)\Phi} + \frac{(C+1)(\Gamma+1)\pi'}{C\pi'-(C+1)(\pi-\Phi)}$$

Hoc modo efficitur, vt obiectum sine margine colorato appareat at omnis confusio tolletur si praeterea fuerit

$$AB'C^2D(A+1) + \frac{ABC^2D(B+1)^2\Phi}{B\pi-(B+1)\Phi} + \frac{ABCD(C+1)^2\Phi}{C\pi'-(C+1)(\pi-\Phi)} + \frac{ABC(D+1)^2\Phi}{D\pi''-(D+1)(\pi'-\pi+\Phi)} = 0$$

seu per  $ABC$  diuidendo:

$$BCD(A+1) + \frac{C\Gamma(B+1)^2\Phi}{B\pi-(B+1)\Phi} + \frac{D(C+1)^2\Phi}{C\pi'-(C+1)(\pi-\Phi)} + \frac{(\Gamma+1)^2\Phi}{D\pi''-(D+1)(\pi'-\pi+\Phi)} = 0 \\ BCD(A+1) + \frac{CD(B+1)b}{Aa} + \frac{D(C+1)c}{ABa} + \frac{(\Gamma+1)d}{ABCa} = 0$$

### Coroll. I.

320. Si vtrique conditioni satisfieri potest, vt nec locus nec magnitudo imaginis vllam mutationem patiat locus oculi non amplius in computum ingreditur sed imago, vndecunque cernatur, ab omni confusione prorsus libera erit.

Coroll.

## Coroll. 2.

321. Vt ergo hunc summum perfectionis gradum consequamur, binis his aequationibus satisfieri oportet.

$$\frac{(B+1)\pi}{B\pi-(B+1)\Phi} + \frac{(C+1)\pi'}{C\pi'-(C+1)(\pi-\Phi)} + \frac{(D+1)\pi''}{D\pi''-(D+1)(\pi'-\pi+\Phi)} = 0$$

et

$$A+1 + \frac{(B+1)^2\Phi}{B(B\pi-(B+1)\Phi)} + \frac{(C+1)^2\Phi}{C(C\pi'-(C+1)(\pi-\Phi))} + \frac{(D+1)^2\Phi}{D(D\pi''-(D+1)(\pi'-\pi+\Phi))} = 0.$$

## Coroll. 3.

322. Si porro vt supra fecimus ponamus

$$\frac{A}{A+1} = \mathfrak{A}, \quad \frac{B}{B+1} = \mathfrak{B}, \quad \frac{C}{C+1} = \mathfrak{C} \text{ et } \frac{D}{D+1} = \mathfrak{D}$$

istae aequationes sequenti modo simplicius exprimentur:

$$+ \frac{\pi}{\mathfrak{B}\pi-\Phi} + \frac{\pi}{\mathfrak{C}\pi'-\pi+\Phi} + \frac{\pi''}{\mathfrak{D}\pi''-\pi'+\pi-\Phi} = 0 \text{ et}$$

$$\mathfrak{A} + \frac{\Phi}{\mathfrak{A}\mathfrak{B}(\mathfrak{B}\pi-\Phi)} + \frac{\Phi}{\mathfrak{A}\mathfrak{B}\mathfrak{C}(\mathfrak{C}\pi'-\pi+\Phi)} + \frac{\Phi}{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D}(\mathfrak{D}\pi''-\pi'+\pi-\Phi)} = 0$$

hic prima membra multiplicari debent per  $\frac{d n}{n-1}$

secunda per  $\frac{d n'}{n'-1}$

tertia per  $\frac{d n''}{n''-1}$

etc.

## Coroll. 4.

323. Sin autem non liceat has ambas aequationes adimplere, curandum est, vt saltem priori, qua magnitudo apparens a confusione liberatur, satisfiat



fiat hoc enim modo obiectum finè margine colorato apparebit vbi quidem ad duos casus respici conuenit prout distantia oculi  $dO$  prodierit positua vel negatiua.

### Scholion.

324. Introducendis ergo rationibus aperturarum quibus supra iam commodè sumus vfi ad campum apparentem definiendum aequationes etiam istae confusionem a diuersa radiorum refrangibilitate oriundam tollentes satis fiunt simplices, vt sine molestia tractari queant; si quidem crassities lentium negligatur. Hoc ergo modo problema generale, quicumque fuerit lentium numerus, expediri conueniet.

### Supplementum V.

Si ratio refractionis in singulis lentibus discrepet, prodit primo quidem eadem aequatio differentialis

$$d\alpha\left(\frac{1}{\alpha} + \frac{1}{b}\right) + d\epsilon\left(\frac{1}{\epsilon} + \frac{1}{c}\right) + d\gamma\left(\frac{1}{\gamma} + \frac{1}{d}\right) \frac{-\infty d\delta}{\delta(\delta - O)} = 0$$

in qua ergo casu  $O = 0$  vltimum membrum abiici debet.

I) Si autem  $O$  habeat valorem posituum, erit, vt ante,

$$\frac{O}{\delta(\delta - O)} = \frac{\pi''}{ABCD(D+1)\alpha\Phi}$$

atque etiam valores  $\frac{1}{\alpha} + \frac{1}{b}$ ;  $\frac{1}{\epsilon} + \frac{1}{c}$ ;  $\frac{1}{\gamma} + \frac{1}{d}$  manent iidem, vt ante.

At

At vero ob diuersitatem refractionis habebimus

$$d\alpha = \frac{-dn}{n-1} \cdot A(A+1)a$$

$$d\beta = BBd\alpha - \frac{dn'}{n'-1} \cdot B(B+1)b$$

$$d\gamma = CCd\beta - \frac{dn''}{n''-1} \cdot C(C+1)c$$

$$d\delta = DDd\gamma - \frac{dn'''}{n'''-1} \cdot D(D+1)d$$

Hinc ergo aequatio nostra successiue ita formetur:

$$\frac{-0d\delta}{\delta(\delta-0)} = \frac{-\pi''d\delta}{ABCD(D+1)a\Phi}$$

$$= \frac{-\pi''Dd\gamma}{ABC(D+1)a\Phi} + \frac{dn'''}{n'''-1} \cdot \frac{\pi''d}{ABC \cdot a\Phi}$$

$$\text{Addatur } d\gamma\left(\frac{1}{\gamma} + \frac{1}{d}\right) = \frac{d\gamma \cdot \pi''D}{ABC(D+1)a\Phi} - \frac{\pi' d\gamma}{ABC(C+1)a\Phi}$$

$$\text{et prodibit } \frac{-\pi' d\gamma}{ABC(C+1)a\Phi} + \frac{dn'''}{n'''-1} \cdot \frac{\pi''d}{ABC \cdot a\Phi}$$

et pro  $d\gamma$  substituto valore

$$\frac{-\pi' C d\beta}{AB(C+1)a\Phi} + \frac{dn''}{n''-1} \cdot \frac{\pi' c}{AB \cdot a\Phi} + \frac{dn'''}{n'''-1} \cdot \frac{\pi''d}{ABC \cdot a\Phi}$$

Iam addatur

$$d\beta\left(\frac{1}{\beta} + \frac{1}{c}\right) = \frac{\pi' C d\beta}{AB(C+1)a\Phi} - \frac{\pi d\beta}{AB(B+1)a\Phi}$$

proditque

$$\frac{-\pi d\beta}{AB(B+1)a\Phi} + \frac{dn''}{n''-1} \cdot \frac{\pi' c}{AB \cdot a\Phi} + \frac{dn'''}{n'''-1} \cdot \frac{\pi''d}{ABC \cdot a\Phi}$$

et substituto valore ipsius  $d\beta$

$$\frac{-\pi B d\alpha}{A(B+1)a\Phi} + \frac{dn'}{n'-1} \cdot \frac{\pi \cdot b}{A \cdot a\Phi} + \frac{dn''}{n''-1} \cdot \frac{\pi' c}{AB \cdot a\Phi} + \frac{dn'''}{n'''-1} \cdot \frac{\pi''d}{ABC \cdot a\Phi}$$

$$\text{denique addatur } d\alpha\left(\frac{1}{\alpha} + \frac{1}{b}\right) = \frac{\pi B d\alpha}{A(B+1)a\Phi}$$

ac aequatio quacsita, qua margo coloratus euanesceat, erit

$$\frac{dn'}{n'-1} \cdot \frac{\pi b}{A \cdot a\Phi} + \frac{dn''}{n''-1} \cdot \frac{\pi' c}{AB \cdot a\Phi} + \frac{dn'''}{n'''-1} \cdot \frac{\pi''d}{ABC \cdot a\Phi} = 0$$

Tom. I.

K k

II)



II) Sin autem  $O$  habeat valorem negativum :  
tunc sumi debet  $O=0$  et pro eodem scopo aequatio  
ita formabitur :

Cum fit

$$d\gamma\left(\frac{1}{\gamma} + \frac{1}{d}\right) = \frac{\pi'' D d \gamma}{ABC(D+1)a\Phi} - \frac{\pi' d \gamma}{ABC(C+1)a\Phi}$$

substituto valore  $d\gamma$  fiet

$$CC d\xi \left( \frac{\pi'' D}{ABC(D+1)a\Phi} - \frac{\pi'}{ABC(C+1)a\Phi} \right) \\ - \frac{dn''}{n''-1} \cdot C(C+1)c \left( \frac{\pi'' D}{ABC(D+1)a\Phi} - \frac{\pi'}{ABC(C+1)a\Phi} \right)$$

Addatur

$$d\xi\left(\frac{1}{\xi} + \frac{1}{c}\right) = \frac{\pi' Cd\xi}{AB(C+1)a\Phi} - \frac{\pi d\xi}{AB(B+1)a\Phi}$$

eritque

$$d\xi \left( \frac{CD\pi''}{AB(D+1)a\Phi} - \frac{\pi}{AB(B+1)a\Phi} \right) - \frac{dn''}{n''-1} (C+1)c \left( \frac{\pi'' D}{AB(D+1)a\Phi} - \frac{\pi'}{AB(C+1)a\Phi} \right)$$

et substituto valore ipsius  $d\xi$

$$d\alpha \left( \frac{BCD\pi''}{A(D+1)a\Phi} - \frac{B\pi}{A(B+1)a\Phi} \right) \\ - \frac{dn'}{n'-1} \left( \frac{(B+1)CD\pi''}{A(D+1)a\Phi} - \frac{\pi b}{A \cdot a\Phi} \right) - \frac{dn''}{n''-1} (C+1)c \left( \frac{\pi'' D}{AB(D+1)a\Phi} - \frac{\pi'}{AB(C+1)a\Phi} \right)$$

Addatur

$$d\alpha\left(\frac{1}{\alpha} + \frac{1}{b}\right) = \frac{\pi Bd\alpha}{A(B+1)a\Phi}$$

eritque

$$- \frac{dn}{n-1} \cdot \frac{(A+1)BCD\pi''}{(D+1)a\Phi} \\ - \frac{dn'}{n'-1} \left( \frac{(B+1)CD\pi'' - (D+1)c\pi}{A(D+1)a\Phi} \right) \\ - \frac{dn''}{n''-1} \left( \frac{(C+1)Dc\pi'' - (D+1)c\pi'}{AB(D+1)a\Phi} \right)$$

Vnde

Vnde pro casu  $O=0$  aequatio, qua margo coloratus destruitur, erit

$$0 = \frac{a dn}{n-1} \cdot (A+1) B C D \pi'' \\ + \frac{b dn'}{n'-1} \cdot \left( \frac{(B+1) C D \pi'' - (D+1) \pi}{A} \right) \\ + \frac{c d n''}{n''-1} \cdot \left( \frac{(C+1) D \pi'' - (D+1) \pi'}{A B} \right)$$

III) Vt autem praeterea omnis confusio tollatur, insuper reddi oportet  $d\delta=0$ ; vnde haec aequatio nascitur

$$0 = \begin{cases} \frac{a d n}{n-1} \cdot A (A+1) B^2 C^2 D^2 \\ + \frac{b d n'}{n'-1} \cdot B (B+1) C^2 D^2 \\ + \frac{c d n''}{n''-1} \cdot C (C+1) D^2 \\ + \frac{d d n'''}{n'''-1} \cdot D (D+1) \end{cases}$$

quae per  $A^2 B^2 C^2 D^2$  diuisa dat

$$0 = \begin{cases} \frac{a d n}{n-1} \cdot \frac{A+1}{A} \\ + \frac{b d n'}{n'-1} \cdot \frac{B+1}{A^2 B} \\ + \frac{c d n''}{n''-1} \cdot \frac{C+1}{A^2 B^2 C} \\ + \frac{d d n'''}{n'''-1} \cdot \frac{D+1}{A^2 B^2 C^2 D} \end{cases}$$

Circa hanc autem aequationem inprimis, notandum est, si omnes lentes pari facultate refringente sint praeditae, ei satisfieri haud posse; ex quo haec aequatio proprie pertinet ad casum, quo diuersae refractiones locum habent.



## Problema 8.

325. Si instrumentum dioptricum ex quocunque lentibus sit compositum, quarum crassitiem negligere liceat, eam determinare dispositionem, ut oculus in eo loco, quem campus apparens postulat positus nullam confusionem sentiat.

## Solutio.

Sit obiecti ante lentem primam distantia  $AE=a$ , eiusque magnitudo  $E\varepsilon=z$ , quae quidem conspici queat, et statuatur  $\frac{z}{a}=\Phi$ . Deinde sint distantiae imaginum a radiis mediae naturae formatarum ut supra:

$$\begin{aligned} AE=a; BF=b; CG=c; DH=d; EI=e \\ aF=a; bG=b; cH=c; dI=d; eK=e \\ \text{etc.} \end{aligned}$$

ac vocemus breuitatis gratia

$$a=Aa, b=Bb, c=Cc, d=Dd, e=Ee \text{ etc.}$$

tum vero etiam

$$\frac{A}{A+1}=\mathfrak{A}; \frac{B}{B+1}=\mathfrak{B}; \frac{C}{C+1}=\mathfrak{C}; \frac{D}{D+1}=\mathfrak{D}, \frac{E}{E+1}=\mathfrak{E} \text{ etc.}$$

Iam apertura primae lentis PP ut evanescente considerata sit ratio aperturae pro reliquis lentibus

$$QQ=\pi; RR=\pi'; SS=\pi''; TT=\pi''' \text{ etc.}$$

His

His positis supra vidimus ( 266 ) fore praeter  $\alpha = Aa$

$$b = \frac{Aa\Phi}{2\pi - \Phi}$$

$$\xi = \frac{ABa\Phi}{2\pi - \Phi}$$

$$c = \frac{ABa\Phi}{\epsilon\pi' - \pi + \Phi}$$

$$\gamma = \frac{ABCa\Phi}{\epsilon\pi' - \pi + \Phi}$$

$$d = \frac{ABCa\Phi}{2\pi'' - \pi' + \pi - \Phi}$$

$$\delta = \frac{ABCDa\Phi}{2\pi'' - \pi' + \pi - \Phi}$$

$$e = \frac{ABCDa\Phi}{\epsilon\pi''' - \pi'' + \pi' - \pi + \Phi}$$

$$\varepsilon = \frac{ABCDEa\Phi}{\epsilon\pi''' - \pi'' + \pi' - \pi + \Phi}$$

etc.

etc.

Deinde singularum imaginum magnitudo ita se habebit

$$F\zeta = Aa\Phi; G\eta = ABa\Phi; H\theta = ABCa\Phi; I\iota = ABCDa\Phi$$

etc.

et ipsi aperturarum semidiametri :

$$\text{Lentis secundae } QQ = \frac{Aa\Phi}{2\pi - \Phi} \pi$$

$$\text{Lentis tertiae } RR = \frac{ABa\Phi}{\epsilon\pi' - \pi + \Phi} \pi'$$

$$\text{Lentis quartae } SS = \frac{ABCa\Phi}{2\pi'' - \pi' + \pi - \Phi} \pi''$$

$$\text{Lentis quintae } TT = \frac{ABCDEa\Phi}{\epsilon\pi''' - \pi'' + \pi' - \pi + \Phi} \pi'''$$

Mutata iam refractionis lege  $n : 1$  infinite parum, distantiae  $\alpha, b, \xi, c; \gamma, d, \delta$  etc. tales mutationes recipiunt.

$$d\alpha = -\frac{dn}{n-1} \cdot Aa(A+1)$$

$$d\xi = \frac{-dn}{n-1} \cdot ABa((A+1)B + \frac{(B+1)\Phi}{2\pi - \Phi})$$

$$d\gamma = \frac{-dn}{n-1} \cdot ABCa((A+1)BC + \frac{(B+1)C\Phi}{2\pi - \Phi} + \frac{(C+1)\Phi}{\epsilon\pi' - \pi + \Phi})$$

$$d\delta = \frac{-dn}{n-1} \cdot ABCDa((A+1)BCD + \frac{(B+1)CD\Phi}{2\pi - \Phi} + \frac{(C+1)D\Phi}{\epsilon\pi' - \pi + \Phi} + \frac{(D+1)\Phi}{2\pi'' - \pi' + \pi - \Phi})$$

etc.

Kk 3

Porro



Porro vero habetur :

$$\frac{1}{a} + \frac{1}{b} = \frac{1}{Aa\Phi} \mathfrak{B} \pi$$

$$\frac{1}{c} + \frac{1}{e} = \frac{1}{ABa\Phi} (\mathfrak{C} \pi^I - \frac{\mathfrak{B} \pi}{B})$$

$$\frac{1}{d} + \frac{1}{f} = \frac{1}{ABCa\Phi} (\mathfrak{D} \pi^{II} - \frac{\mathfrak{C} \pi^I}{C})$$

$$\frac{1}{g} + \frac{1}{h} = \frac{1}{ABCDa\Phi} (\mathfrak{E} \pi^{III} - \frac{\mathfrak{D} \pi^{II}}{D})$$

etc.

His positis pro quolibet lentium numero seorsim formulas quaesito satisfaciennes expediamus: posita distantia oculi post lentem ultimam  $= 0$ .

### I. Pro vnica Lente

Habetur haec aequatio differentialis  $\frac{-0 d \alpha}{\alpha(\alpha-0)} = 0$ , pro quo casu tam magnitudo imaginis quam eius locus manebit inuariatus si fuerit  $d\alpha = 0$ , hoc est  $A(A+1) = 0$ , vnde deberet esse vel  $A = 0$  vel  $A = -1$ : quorum prius visio non admittit posterius autem lentem tollit. Tum vero ob campum esse debet  $0 = 0$ .

### II. Pro duabus lentibus

Habetur haec aequatio differentialis, qua margo coloratus tollitur:

$$d\alpha \left( \frac{1}{\alpha} + \frac{1}{b} \right) - \frac{0 d e}{e(e-0)} = 0$$

at ob campum apparentem est  $\frac{0}{e(e-0)} = \frac{1}{Bb} \left( \frac{1}{\alpha} + \frac{1}{b} \right)$ , ita vt habeamus:

$$d\alpha \left( \frac{1}{\alpha} + \frac{1}{b} \right) - \frac{d e}{Bb} \left( \frac{1}{\alpha} + \frac{1}{b} \right) = 0$$

Verum

Verum est  $d\mathcal{E} = BBd\alpha - \frac{dn}{n-1} \cdot ABa \cdot \frac{(B+1)\Phi}{\mathfrak{B}\pi - \Phi}$

Quare si O habeat valorem positium erit ob

$$\mathfrak{B}(B+1) = B, \frac{\pi}{\mathfrak{B}\pi - \Phi} = 0$$

Sin autem valor O prodeat negatiuus, quo casu capi-  
tur  $O = 0$  erit  $\frac{(A+1)\mathfrak{B}\pi}{\Phi} = 0$ .

Omnis autem confusio penitus tolletur, si insuper  
fuerit

$$\frac{1}{\mathfrak{A}} + \frac{\Phi}{A\mathfrak{B}(\mathfrak{B}\pi - \Phi)} = 0$$

### III. Pro tribus Lentibus

Si calculum eodem modo prosequamur, obiectum  
sine margine colorato conspicietur:

1. Si ex campo apparente distantia O prodeat po-  
sitiua hanc aequationem adimplendo:

$$\frac{\pi}{\mathfrak{B}\pi - \Phi} + \frac{\pi'}{\mathfrak{E}\pi' - \pi + \Phi} = 0.$$

2. Si ob distantiam O prodeuntem negatiuam ca-  
piatur  $O = 0$  huic aequationi erit satisfaciendum:

$$\frac{\pi'}{\mathfrak{A}\Phi} + \frac{-\pi'}{A\mathfrak{B}(\mathfrak{B}\pi - \Phi)} = \frac{\pi}{AB\mathfrak{E}(\mathfrak{B}\pi - \Phi)}$$

Omnis autem confusio penitus tolletur si fuerit insuper

$$\frac{1}{\mathfrak{A}} + \frac{\Phi}{A\mathfrak{B}(\mathfrak{B}\pi - \Phi)} + \frac{\Phi}{AB\mathfrak{E}(\mathfrak{E}\pi' - \pi + \Phi)} = 0.$$

### IV. Pro quatuor Lentibus

Vt obiectum sine margine colorato spectetur:

1. Si



1. Si ex campo apparente distantia  $O$  prodeat positiua huic aequationi erit satisfaciendum :

$$\frac{\pi}{\mathfrak{B}\pi - \Phi} + \frac{\pi'}{\mathfrak{E}\pi' - \pi + \Phi} + \frac{\pi''}{\mathfrak{D}\pi'' - \pi' + \pi - \Phi} = 0$$

2. Sin autem capiatur  $O=0$ , huic

$$\frac{\pi''}{\mathfrak{A}\Phi} + \frac{\pi''}{\mathfrak{A}\mathfrak{B}(\mathfrak{B}\pi - \Phi)} + \frac{\pi''}{\mathfrak{A}\mathfrak{B}\mathfrak{E}(\mathfrak{E}\pi' - \pi + \Phi)} = \frac{\pi}{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D}(\mathfrak{B}\pi - \Phi)} + \frac{\pi'}{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D}(\mathfrak{E}\pi' - \pi + \Phi)}$$

Omnis vero confusio penitus tolletur si fuerit praeterea

$$\frac{1}{\mathfrak{A}} + \frac{\Phi}{\mathfrak{A}\mathfrak{B}(\mathfrak{B}\pi - \Phi)} + \frac{\Phi}{\mathfrak{A}\mathfrak{B}\mathfrak{E}(\mathfrak{E}\pi' - \pi + \Phi)} + \frac{\Phi}{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D}(\mathfrak{D}\pi'' - \pi' + \pi - \Phi)} = 0$$

### V. Pro quinque Lentibus

Vt obiectum tantum sine margine colorato conspiciatur

1. Si ex campo apparente distantia  $O$  prodeat positiua, huic aequationi erit satisfaciendum :

$$\frac{\pi}{\mathfrak{B}\pi - \Phi} + \frac{\pi'}{\mathfrak{E}\pi' - \pi + \Phi} + \frac{\pi''}{\mathfrak{D}\pi'' - \pi' + \pi - \Phi} + \frac{\pi'''}{\mathfrak{E}\pi''' - \pi'' + \pi' - \pi + \Phi} = 0$$

2. Sin autem capiatur  $O=0$  huic

$$\frac{\pi'''}{\mathfrak{A}\Phi} + \frac{\pi'''}{\mathfrak{A}\mathfrak{B}(\mathfrak{B}\pi - \Phi)} + \frac{\pi'''}{\mathfrak{A}\mathfrak{B}\mathfrak{E}(\mathfrak{E}\pi' - \pi + \Phi)} + \frac{\pi'''}{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D}(\mathfrak{D}\pi'' - \pi' + \pi - \Phi)}$$

$$\frac{1}{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D}\mathfrak{E}} \left( \frac{\pi}{\mathfrak{B}\pi - \Phi} + \frac{\pi'}{\mathfrak{E}\pi' - \pi + \Phi} + \frac{\pi''}{\mathfrak{D}\pi'' - \pi' + \pi - \Phi} \right)$$

Omnis autem confusio penitus tolletur, si praeterea satisfiat huic aequationi:

$$\frac{1}{\mathfrak{A}} + \frac{\Phi}{\mathfrak{A}\mathfrak{B}(\mathfrak{B}\pi - \Phi)} + \frac{\Phi}{\mathfrak{A}\mathfrak{B}\mathfrak{E}(\mathfrak{E}\pi' - \pi + \Phi)} + \frac{\Phi}{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D}(\mathfrak{D}\pi'' - \pi' + \pi - \Phi)} + \frac{\Phi}{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D}\mathfrak{E}(\mathfrak{E}\pi''' - \pi'' + \pi' - \pi + \Phi)} = 0$$

Atque hinc manifesta est progressio ad maiorem lentium numerum.

Coroll.

## Coroll. I.

326. In casu ergo vnicae lentis licet quidem obiectum a margine colorato liberare neque vero confusionem penitus tollere. Casu autem duarum lentium ne margo quidem coloratus tolli potest, si quidem oculus in eo loco, quem campi apparentis conditio postulat, teneatur.

## Coroll. 2.

327. Quodsi vero plures duabus habeantur lentes sufficiens quantitatum numerus adest, quarum determinatione non solum margo coloratus deleri, sed etiam forte omnis confusio penitus auferri posse videtur praecipue si lentium numerus ternarium superet.

## Scholion.

328. Quod ergo incommodum a diuersa radiorum natura oriundum adeo graue vel summo Newtono est visum, vt instrumenta dioptrica nullo modo ab eo liberari posse sit arbitratus, id quidem saltem quod ad marginem coloratum attinet ad quem Newtonus inprimis spectabat, iam satis feliciter tolli posse certum est; ita vt saltem ob hanc causam non opus sit ad Telescopia Catoptrica confugere. Hic autem vitio sublato si praeterea alterum confusionis fontem obstruamus lentes scilicet nullam confusionem parientes adhibendo, nullum est dubium quin instrumenta dioptrica ad summum perfectionis gradum euehi queant.

Tom. I.

LI

Quae



Quae igitur hactenus particulatim circa singulas horum instrumentorum affectiones proposuimus, ea colligi conueniet, vnde in capite sequente praecepta generalia pro omnium instrumentorum dioptricum constructione tradere est visum.

## Supplementum VI.

Ex iis, quae ante sunt adiecta, poterimus etiam problematis solutionem pro casu exhibere, quo singulae lentes peculiari refractione sunt praeditae, vbi quidem tantum postremae aequationes pro confusione vitanda mutationem quandam postulant; interim tamen etiam priores formulas, quibus locus oculi, quem campus apparens requirit, distinctius repraesentemus.

I. *Distantia Oculi post ultimam lentem* pro quouis lentium numero se habebit, vt sequitur.

| Num.<br>lentium | O id est, distantia oculi post lentem ultimam                                                                                                                                            |
|-----------------|------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|
| I.              | o                                                                                                                                                                                        |
| II.             | $\frac{A B a \pi \Phi}{(\pi - \Phi)(B \pi - \Phi)}$ seu $\frac{B b \pi}{\pi - \Phi}$                                                                                                     |
| III.            | $\frac{A P . \epsilon a \pi' \Phi}{(\pi' - \pi + \Phi)(\epsilon \pi' - \pi + \Phi)}$ seu $\frac{\epsilon c \pi'}{\pi' - \pi + \Phi}$                                                     |
| IV.             | $\frac{A B C D . a \pi'' \Phi}{(\pi'' - \pi' + \pi - \Phi)(D \pi'' - \pi' + \pi - \Phi)}$ seu $\frac{D d \pi''}{\pi'' - \pi' + \pi - \Phi}$                                              |
| V.              | $\frac{A B C D E . a \pi''' \Phi}{(\pi''' - \pi'' + \pi' - \pi + \Phi)(\epsilon \pi''' - \pi'' + \pi' - \pi + \Phi)}$ seu $\frac{\epsilon e \pi'''}{\pi''' - \pi'' + \pi' - \pi + \Phi}$ |
|                 | etc.                                                                                                                                                                                     |

II. Si valor ipsius  $O$  sit positivus, ad marginem coloratum tollendum sequentes aequationes sunt adimplendae:

Num.  
lentium

I.  $O = 0$

II.  $O = \frac{bdn'}{n'-1} \cdot \frac{\pi}{Aa\Phi}$

III.  $O = \frac{bdn'}{n'-1} \cdot \frac{\pi}{Aa\Phi} + \frac{cdn''}{n''-1} \cdot \frac{\pi'}{ABa\Phi}$

IV.  $O = \frac{bdn'}{n'-1} \cdot \frac{\pi}{Aa\Phi} + \frac{cdn''}{n''-1} \cdot \frac{\pi'}{ABa\Phi} + \frac{ddn'''}{n'''-1} \cdot \frac{\pi''}{ABCa\Phi}$

etc.

III. Si valor ipsius  $O$  prodeat negativus, quo casu capi debet  $O = 0$ , ad marginem coloratum tollendum sequentes aequationes sunt adimplendae.

Num.  
lentium

I.  $O = 0$

II.  $O = \frac{adn}{n-1} \cdot (A+1) B \pi$

III.  $O = \frac{adn}{n-1} (A+1) BC \pi' + \frac{bdn'}{n'-1} \cdot \frac{(B+1)C\pi' - (C+1)\pi}{A}$

IV.  $O = \frac{adn}{n-1} \cdot (A+1) BCD \pi'' + \frac{bdn'}{n'-1} \left( \frac{(B+1)CD\pi'' - (D+1)\pi}{A} \right) + \frac{cdn''}{n''-1} \left( \frac{(C+1)D\pi'' - (D+1)\pi'}{AB} \right)$

etc.

L 1 2

IV.



IV. Vt autem insuper omnis confusio huius generis tollatur, sequentes aequationes sunt adimplendae:

Num.  
lentium

$$\begin{aligned}
 \text{I.} \quad & 0 = \frac{a d n}{n-1} \cdot \frac{A+1}{A} \\
 \text{II.} \quad & 0 = \frac{a d n}{n-1} \cdot \frac{A+1}{A} + \frac{b d n'}{n'-1} \cdot \frac{B+1}{A^2 B} \\
 \text{III.} \quad & 0 = \frac{a d n}{n-1} \cdot \frac{A+1}{A} + \frac{b d n'}{n'-1} \cdot \frac{B+1}{A^2 B} + \frac{c d n''}{n''-1} \cdot \frac{C+1}{A^2 B^2 C} \\
 \text{IV.} \quad & 0 = \frac{a d n}{n-1} \cdot \frac{A+1}{A} + \frac{b d n'}{n'-1} \cdot \frac{B+1}{A^2 B} + \frac{c d n''}{n''-1} \cdot \frac{C+1}{A^2 B^2 C} + \frac{d d n'''}{n'''-1} \cdot \frac{D+1}{A^2 B^2 C^2 D} \\
 & \text{etc.}
 \end{aligned}$$

Quarum formularum ordo hinc distinctius perspicitur, quam in problemate.



## CAPUT VII.

DE

## CONSTRUCTIONE

## INSTRUMENTORVM DIOPTRICORVM

## IN GENERE.

## Problema I.

329.

Si instrumentum dioptricum vnica constet lente Tab. III.  
 crassitie cuiuscunque, PP, definire omnia momenta Fig. 12.  
 ad visionem pertinentia.

## Solutio.

Quod primo ad ipsius lentis structuram attinet  
 ponatur obiecti  $E\epsilon$  ante eam distantia  $AE = a$ , ima-  
 ginisque  $F\zeta$  post eam proiectae  $aF = \alpha$ ; tum vero  
 lentis crassities  $Aa = v$  et quantitas arbitraria  $= k$ ,  
 vnde capiatur  $\frac{k-v}{k+v} = i$ . Hinc facies lentis ita erunt  
 formatae, vt sit existente  $n = \frac{31}{25}$ . Conf. §. 68.

$$\text{radius faciei anterioris} = \frac{(n-1)a(k+v)}{k+v+2na}$$

$$\text{radius faciei posterioris} = \frac{(n-1)\alpha(k-v)}{k-v-2n\alpha}.$$

Ponatur porro semidiameter obiecti conspicui  $E\epsilon = z$ ,  
 et semidiameter aperturæ in facie anteriori  $AM = x$ ,  
 in facie autem posteriori semidiameter aperturæ non

L 1 3

fit



fit minor quam  $ix$ . Denique oculi post lentem distantia vocetur  $aO=O$ , cuius distantia iusta fit  $=l$ . His positis sequentia momenta perpendi oportet.

1. Debet esse  $O=\alpha+l$  vt distantia imaginis visae naturae oculi conueniat.

2. Consideranda venit multiplicatio, quae ex ratione definitur, quam diameter obiecti visus tenet ad eiusdem diametrum visum si nudo oculo in distantia data  $=b$  spectaretur. Quodsi ergo hic exponens multiplicationis ponatur  $=m$  erit

$$m = \frac{1}{i} \cdot \frac{\alpha b}{a l} \text{ pro situ inuerso}$$

3. Gradus claritatis determinatur semidiametro coni luminosi, qui a quouis obiecti puncto in oculum immittitur, qui si ponatur  $=y$  erit

$$y = i l \cdot \frac{x}{\alpha} = \frac{b x}{m a}$$

4. Confusio inquinans visionem mensuratur semidiametro apparente circuli, qui nudo oculo aequae magnus cernitur, ac singula obiecti puncta per lentem spectata. Hanc mensuram vocavi (194) semidiametrum confusionis, ad quem definiendum si ponatur:

$$P = \frac{n}{2(n-1)^2} \left( \left( \frac{n}{a} + \frac{2}{k+v} \right) \left( \frac{i}{ia} + \frac{2}{k-v} \right)^2 + \left( \frac{n}{\alpha} - \frac{2}{k-v} \right) \left( \frac{i}{\alpha} - \frac{2}{k+v} \right)^2 \right)$$

erit semidiameter confusionis  $= \frac{1}{4} i \cdot \frac{\alpha}{l} x^3 P = \frac{1}{4} i^2 \cdot \frac{m a}{b} x^3 \cdot P$ .

5. Vt oculus maximum campum apparentem percipiat, debet esse  $O = \frac{i \alpha v}{n \alpha - i v}$

6. Quae

6. Quae distantia si fuerit positua, semidiameter campi seu obiecti conspicui  $E\varepsilon = z$ , ita pendet ab apertura faciei posterioris, vt posito huius semidiametro  $= a$  sit  $z = \frac{na}{v} \cdot a$ .

7. Sin autem distantia illa pro  $O$  assignata prodierit negatiua, oculum lenti immediate applicari conueniet, vt sit  $O = 0$ ; tum vero pro  $a$  posito semidiametro pupillae  $\omega$  erit pro campo apparente  $z = \frac{na}{v} \omega$ .

8. Vt obiectum sine margine colorato appareat, existente distantia  $O$  ante inuenta positua, debet esse  $i = \frac{a}{a+v}$  seu  $k = 2a + v$ .

9. Sin autem ob illam distantiam negatiuam prodeuntem capiatur  $O = 0$ , vt margo coloratus euitetur, oportet esse  $\frac{v}{a} + 1 - i = 0$ , vnde fit  $k = -2a - v$ .

10. Omnis denique confusio a diuersa radiorum refrangibilitate oriunda prorsus tolletur, si insuper fieri posset  $(a+v)(a+a+v) = 0$

### C o r o l l. I.

330. Si exponens multiplicationis  $m$  cum gradu claritatis  $y$  proponatur, erit  $my = \frac{bx}{a}$ . quae proprietas ad lentium numerum quantumuis magnum patet. Cum ergo  $\frac{b}{a}$  sit quantitas data, erit  $x$  vt  $my$ ; tum vero  $y$  vt  $\frac{x}{m}$ , ac  $m$  vt  $\frac{x}{y}$

### C o r o l l. 2.

331. Tam maior ergo multiplicatio quam maior claritatis gradus postulat maiorem aperturam.

Verum



Verum aucto  $x$  confusio augetur in ratione triplicata, siquidem quantitas  $P$  maneat eadem: quare si  $x$  ex confusione etiamnum tolerabili determinetur, simul quantitas  $my$  determinatur.

### Coroll. 3.

332. Vt igitur tam multiplicationem  $m$  quam claritatem  $y$  salua confusione maxime augere liceat, quantitatem arbitrariam  $k$  ita definiri conueniet, vt litterae  $P$  minimus valor concilietur.

Cum autem posuerimus  $\frac{k-v}{k+v} = i$ , valor ipsius  $P$  fiet minimus, huic aequationi satisfaciendo:

$$2i^4 \left( \frac{n}{a} + \frac{1}{v} \right) \left( \frac{1}{a} + \frac{1}{v} \right)^2 - \frac{2ni^3}{a v} \left( \frac{1}{a} + \frac{1}{v} \right) - \frac{i^3}{v} \left( \frac{1}{a} + \frac{2}{v} \right)^2 \\ + \frac{2ni}{a v} \left( \frac{1}{a} + \frac{1}{v} \right) + \frac{i}{v} \left( \frac{1}{a} + \frac{2}{v} \right)^2 - 2 \left( \frac{n}{a} + \frac{1}{v} \right) \left( \frac{1}{a} + \frac{1}{v} \right)^2 = 0$$

vti §. 45 inuenimus.

### Coroll. 4.

333. Si loco  $i$  valor substituaturs suus  $\frac{k-v}{k+v}$ , aequatio transibit in hanc formam:

$$+ \frac{2n}{a^3} (k+v)^4 + \frac{(2n+1)}{a a v} (k+v)^3 (k+3v) + \frac{4(n+2)}{a v} (k+v)^2 \\ - \frac{2n}{a^3} (k-v)^4 - \frac{(2n+1)}{a a v} (k-v)^3 (k-3v) + \frac{4(n+2)}{a v} (k-v)^2 + 32k = 0$$

quae per  $v$  multiplicata etiam ad casum accommodari potest quo est  $v=0$ : tum autem reperitur

$$k = \frac{4(n+2)a\alpha}{(2n+1)(a-\alpha)} \text{ et}$$

$$P = \frac{n}{8(n-1)^2(n+2)} \left( \frac{1}{a} + \frac{1}{\alpha} \right) \left( (4n-1) \left( \frac{1}{a} + \frac{1}{\alpha} \right)^2 + \frac{4(n-1)}{a\alpha} \right)$$

Scholion

## Scholion

334. Hic non generatim omnem visionem, quae fit per vnicam lentem, considero, sed tantum eam, qua maximus campus apparens conspicitur; quamobrem locum oculi ita definiui, vt ipsi maximus campus adferatur. Sin autem minore campo velimus esse contenti oculus quoque alibi post lentem constitutus obiecta distincte cernere poterit, quemadmodum per lentes satis amplas, quae tabularum vitrearum nomine sunt notae, fieri solet. Hunc autem casum quoniam per se facile expeditur, atque in instrumentis dioptricis magnitudo campi potissimum spectatur, hic non attingo.

## Problema 2.

335. Si instrumentum dioptricum duabus conflet lentibus cuiuscunque crassitiei definire omnia momenta ad visionem pertinentia.

Tab. III.  
Fig. 13.

## Solutio.

Obiecto constituto in  $E\varepsilon$ , eius imagines proiciantur per istas lentes in  $F\zeta$  et  $G\eta$ , ac ponantur quantitates vtramque lentem determinantes

$AE=a$ ,  $aF=a$ , crassities  $Aa=v$  et distantia arbitraria  $=k$

$BF=b$ ,  $bG=b$ , crassities  $Bb=v'$  et distantia arbitraria  $=k'$

Tom. I.

Mm

ponatur-



ponaturque breuitatis gratia  $\frac{k-v}{k+v} = i$  et  $\frac{k'-v'}{k'+v'} = i'$ .  
Hinc existente  $n = \frac{31}{20}$ , constructio vtriusque lentis  
ita se habebit,

Radius faciei anterioris posterioris

Pro Lente PP  $\cdot \frac{(n-1)a(k+v)}{k+v+2na} \cdot \frac{(n-1)\alpha(k-v)}{k-v-2n\alpha}$

Pro Lente QQ  $\cdot \frac{(n-1)b(k'+v')}{k'+v'+2nb} \cdot \frac{(n-1)\varepsilon(k'-v')}{k'-v'-2n\varepsilon}$

Sit porro semidiameter aperturae lentis primae PP  
in facie anteriori  $= x$ , in facie autem posteriori  
maior sit quam  $ix$ . Tum vero lentis secundae QQ  
semidiameter aperturae in facie anteriori maior esse  
debet quam  $i \cdot \frac{bx}{\alpha}$ , in posteriori vero maior quam  $i' \cdot \frac{bx}{\alpha}$ .  
Deinde sit semidiameter obiecti conspicui  $E\varepsilon = z$ ,  
distantia oculi  $bo = O$ , eiusque distantia iusta  $= l$ .  
His positis ad sequentia momenta erit attendendum.

1. Vt oculus imaginem  $G\eta$  in distantia iusta con-  
spiciat, oportet esse  $O = \varepsilon + l$ .

2. Posita distantia  $= b$  ad quam multiplicatio  
referatur, sit exponens multiplicationis  $= m$ , ac  
supra inuenimus esse oportere  $m = \frac{1}{i' \cdot \frac{\alpha \varepsilon b}{a b l}}$  pro situ  
erecto.

3. Denotante  $y$  gradum claritatis fiet

$$y = i i' l \cdot \frac{bx}{\alpha \varepsilon} \text{ ideoque } my = \frac{bx}{\alpha}$$

4. Pro confusione posito breuitatis ergo

$$P = \frac{n}{2(n-1)^2} \left( \left( \frac{n}{a} + \frac{2}{k+v} \right) \left( \frac{1}{ia} + \frac{2}{k-v} \right)^2 + \left( \frac{n}{a} + \frac{2}{k-v} \right) \left( \frac{i}{\alpha} - \frac{2}{k+v} \right)^2 \right)$$

$$Q = \frac{n}{2(n-1)^2} \left( \left( \frac{n}{b} + \frac{2}{k'+v'} \right) \left( \frac{1}{ib} + \frac{2}{k'-v'} \right)^2 + \left( \frac{n}{b} - \frac{2}{k'-v'} \right) \left( \frac{i'}{\varepsilon} - \frac{2}{k'+v'} \right)^2 \right)$$

erit

eritque

semid. confusionis  $= \frac{1}{4} i i' \cdot \frac{b}{l} \cdot \frac{b}{\alpha} x^3 \left( \frac{1}{i' i'} \cdot \frac{\alpha \alpha}{b b} P + i i' \cdot \frac{b b}{\alpha \alpha} Q \right)$ .

5. Pro campo apparente definiendo, qui pendet ab apertura singularum facierum lentium, si ponamus

Semidiametrum aperturæ

Pro lente PP  $\begin{cases} \text{faciei anterioris} = \mathfrak{A} = x \\ \text{faciei posterioris} = a \end{cases}$

Pro lente QQ  $\begin{cases} \text{faciei anterioris} = \mathfrak{B} \\ \text{faciei posterioris} = b \end{cases}$

habebimus sequentes aequationes

$$a = \frac{v}{n a} z$$

$$\mathfrak{B} = \left( \frac{1}{i'} \cdot \frac{\alpha + b}{a} - \frac{b v}{n a \alpha} \right) z$$

$$b = \left( \frac{i'}{i} \cdot \frac{\alpha + b}{a} - \frac{i' b v}{n a \alpha} - \frac{1}{i} \cdot \frac{\alpha v'}{n a b} \right) z$$

ex quibus tribus aequationibus valor ipsius  $z$  minimus praebet semidiametrum campi apparentis.

6. Quem campum vt oculus reuera perspiciat, eius locus ita debet sumi, vt fit

$$O = \frac{\frac{i'}{i} \frac{\alpha + b}{a} - \frac{i' b v}{n a \alpha} - \frac{1}{i} \frac{\alpha v'}{n a b}}{\frac{i'}{i} \frac{\alpha + b}{a} - \frac{i' b v}{n a \alpha} - \frac{1}{i} \frac{\alpha v'}{n a b} + \frac{1}{i' i'} \frac{\alpha \alpha}{b b}} \cdot \mathfrak{C}$$

si quidem haec distantia fuerit positiua.

7. Sin autem haec distantia prodierit negatiua, oculum lenti QQ immediate applicari conuenit, vt fit  $O = 0$ ; tum vero pro campo apparente inue-



niendo in aequationibus n°. 5 allatis loco  $\mathcal{E}$  scribatur semidiameter pupillae  $\omega$ , atque ex iisdem valor ipsius  $\mathcal{Z}$  minimus erutus dabit semidiametrum campi apparentis.

8. Obiectum porro sine margine colorato cerne-  
tur, existente distantia  $O$  n°. 6 inuenta positiua, si  
huic aequationi satisfiat:

$$0 = 1 + \frac{b}{\alpha} + \frac{i' i' b}{\mathcal{E}} \left( 1 + \frac{b}{\alpha} \right) + \frac{(1-i')^2 b (\alpha + b)}{\alpha v'} + \frac{2-i-i'}{n} \\ - \frac{i v}{n \alpha} - \frac{i' v'}{n \mathcal{E}} - \frac{i b v}{n \alpha \alpha} - \frac{i i' i' b b v}{n \alpha \alpha \mathcal{E}} - \frac{i (1-i')^2 b b v}{n \alpha \alpha v'}$$

9. Sin autem capiatur distantia  $O=0$ , obiectum  
margine colorato carebit, si fuerit

$$0 = \frac{1-i}{in} + \frac{1-i'}{i'n} + \frac{v}{ina} + \frac{v'}{inb} + \alpha \left( 1 + \frac{\alpha}{ia} + \frac{(1-i)^2 \alpha}{i.i.v} \right) \left( \frac{v'}{ni'bb} - \frac{1}{\alpha} - \frac{1}{b} \right)$$

10. Omnis autem confusio penitus tolletur, quae  
quidem a diuersa radiorum refrangibilitate proficisci-  
tur, si huic aequationi satisfacere licuerit  $\mathcal{E} = 0$ ,  
seu huic

$$0 = \frac{\alpha \alpha}{i i' b b} \left( \frac{i}{\alpha} + \frac{1}{ia} + \frac{(1-i)^2}{iv} \right) + \frac{i'}{\mathcal{E}} + \frac{1}{i'b} + \frac{(1-i')^2}{i' v'}$$

### Coroll. 1.

336. Quia ambarum lentium distantia  $aB$  ne-  
cessario est positiua, oportet sit  $\alpha + b > 0$ . Tum vero  
aperturae ita debent esse comparatae vt sit  $\alpha > ix$ ;  
 $\mathcal{B} > i. \frac{bx}{\alpha}$  et  $\mathcal{E} > i i'. \frac{bx}{\alpha}$ .

### Coroll. 2.

337. Casu igitur, quo ob valorem ipsius  $O$   
prodeuntem negatiuum distantia  $O=0$  assumitur,  
quia

quia tum  $b = \omega$  statuitur, etiam esse debet  $\omega > ii^{\frac{bx}{a}}$  seu  $x < \frac{\alpha\omega}{u^b}$ . Scilicet hoc casu inutile esset primam aperturam maiorem sumere, quia radii non in oculum ingrederentur.

### Scholion

338. Si simili modo instrumenta dioptrica pluribus lentibus instructa prosequi vellemus, in calculos tantopere intricatos delaberemur, ut ex iis nihil fere concludere liceret. Oritur autem haec calculi complicatio a crassitie lentium, qua neglecta omnia fiunt multo concinniora. Quare si pluribus lentibus uti velimus, crassitiem tam parvam assumamus, ut sine errore pro nihilo haberi possit, id quod in praxi etiam sedulo observari solet. Praeterea etiam campi apparentis locique oculi determinatio non capit rigorem geometricum, parumque refert si in ea aliquantum aberretur. Similis quoque ratio est conditionum, quibus effectus a diversa radiorum refrangibilitate oriundus tollitur; sufficiet enim iis proxime satisfecisse, cum perfecta huius confusionis destructione sperari quidem possit. Quocirca in sequentibus, ubi instrumenta pluribus lentibus instructa evoluemus, crassitiem lentium in calculo penitus praetermittamus: in quo negotio ne eadem toties repetere opus habeamus, problema generale praemittamus in quo omnia momenta generatim tantum exponamus eaque deinceps pro quovis lentium numero accurate describamus.



## Problema 5.

339. Si instrumentum dioptricum compositum sit ex lentibus quotcunque quarum crassitiem negligere liceat, elementa eius constructionem continentia exponere, ex quibus deinceps regulæ dirigentes constructionem ipsam stabiliri possint.

## Solutio.

Obiecto in E constituto ponantur distantiae determinatrices singularum lentium una cum numeris arbitrariis ad singulas pertinentibus ut sequitur.

Pro Lente prima  $EA = a$ ;  $aF = \alpha$ ; num: arb:  $= \lambda$

Pro Lente secunda  $FB = b$ ;  $bG = \beta$ ; num: arb:  $= \lambda'$

Pro Lente tertia  $GC = c$ ;  $cH = \gamma$ ; num: arb:  $= \lambda''$

Pro Lente quarta  $HD = d$ ;  $dI = \delta$ ; num: arb:  $= \lambda'''$

Pro Lente quinta  $IE = e$ ;  $eK = \epsilon$ ; num: arb:  $= \lambda''''$

Pro Lente sexta  $KF = f$ ;  $fL = \zeta$ ; num: arb:  $= \lambda'''''$

etc.

ex quibus elementis quomodo singulae lentes debeant fermari supra est expositum.

Nunc autem porro ponatur breuitatis gratia:

$\alpha = Aa$ ;  $\beta = Bb$ ;  $\gamma = Cc$ ;  $\delta = Dd$ ;  $\epsilon = Ee$ ;  $\zeta = Ff$

etc.

tum

tum vero statuatur etiam

$$\frac{A}{A+1} = \mathcal{A}; \frac{B}{B+1} = \mathcal{B}; \frac{C}{C+1} = \mathcal{C}; \frac{D}{D+1} = \mathcal{D}; \frac{E}{E+1} = \mathcal{E}; \frac{F}{F+1} = \mathcal{F}$$

etc.

ita ut sint distantiae focales lentium :

$$\mathcal{A}a; \mathcal{B}b; \mathcal{C}c; \mathcal{D}d; \mathcal{E}e; \mathcal{F}f \quad \text{etc.}$$

Nunc sit semidiameter aperturæ primæ lentis obiectivæ  $= x$ , pro reliquis vero lentibus ratio aperturarum litteris  $\pi, \pi', \pi'', \pi''', \pi''''$  etc. exponatur; ita ut semidiameter aperturæ, cuiusque maior accipi debeat quam secundum has rationes : scilicet capi oportebit :

Semid : apert :

$$\text{Lentis Secundæ} > \pi \mathcal{B}b$$

$$\text{Lentis tertiæ} > \pi' \mathcal{C}c$$

$$\text{Lentis quartæ} > \pi'' \mathcal{D}d$$

$$\text{Lentis quintæ} > \pi''' \mathcal{E}e$$

$$\text{Lentis sextæ} > \pi'''' \mathcal{F}f \quad \text{etc.}$$

Quod si iam semidiameter spatii in obiecto conspiciui vocetur  $= z$ , fiatque  $\frac{z}{a} = \Phi$ , ut sit  $z = a\Phi$ , atque hinc distantiae determinatrices ita exprimuntur, ut sit primo  $a = Aa$  tum vero :

$$b = \frac{Aa\Phi}{2\pi - \Phi}; \quad \mathcal{B} = \frac{ABa\Phi}{2\pi - \Phi}$$

$$c = \frac{ABa\Phi}{\pi' - \pi + \Phi}; \quad \mathcal{C} = \frac{ABCa\Phi}{\pi' - \pi + \Phi}$$

$$d = \frac{ABCa\Phi}{2\pi'' - \pi' + \pi - \Phi}; \quad \mathcal{D} = \frac{ABCDa\Phi}{2\pi'' - \pi' + \pi - \Phi}$$

$$e = \frac{ABCDa\Phi}{\pi''' - \pi'' + \pi' - \pi + \Phi}; \quad \mathcal{E} = \frac{ABCDEa\Phi}{\pi''' - \pi'' + \pi' - \pi + \Phi}$$

$$f = \frac{ABCDEa\Phi}{2\pi'''' - \pi''' + \pi'' - \pi' + \pi - \Phi}; \quad \mathcal{F} = \frac{ABCDEFa\Phi}{2\pi'''' - \pi''' + \pi'' - \pi' + \pi - \Phi} \quad \text{etc.}$$

Circa



Circa has expressiones primum notandum est, aggregata  $\alpha + b$ ;  $\xi + c$ ;  $\gamma + d$ ;  $\delta + e$ ;  $\varepsilon + f$  etc. esse oportere positiua quippe quibus lentium interualla exprimuntur. Erit nempe

Interuallum  
Lentium

$$\text{I et II} = \frac{A B a \pi}{2 \pi - \Phi} > 0$$

$$\text{II et III} = \frac{A B a \Phi (\varepsilon \pi' - (1 - 2) \pi)}{(2 \pi - \Phi) (\varepsilon \pi' - \pi + \Phi)} > 0$$

$$\text{III et IV} = \frac{A B C a \Phi (2 \pi'' - (1 - \varepsilon) \pi')}{(\varepsilon \pi' - \pi + \Phi) (2 \pi'' - \pi' + \pi - \Phi)} > 0$$

$$\text{IV et V} = \frac{A B C D a \Phi (\varepsilon \pi''' - (1 - 2) \pi'')}{(2 \pi'' - \pi' + \pi - \Phi) (\varepsilon \pi''' - \pi'' + \pi' - \pi + \Phi)} > 0$$

$$\text{V et VI} = \frac{A B C D E a \Phi (2 \pi'''' - (1 - \varepsilon) \pi''')}{(\varepsilon \pi''' - \pi'' + \pi' - \pi + \Phi) (2 \pi'''' - \pi''' + \pi'' - \pi' + \pi - \Phi)} > 0$$

etc.

Praeterea vt lentes sequentes omnes radios a prima exceptos transmittant, debet esse

Semid: aperturae

$$\text{Lentis secundae} > \frac{\Phi}{2 \pi - \Phi} x$$

$$\text{Lentis tertiae} > \frac{\Phi}{\varepsilon \pi' - \pi + \Phi} x$$

$$\text{Lentis quartae} > \frac{\Phi}{2 \pi'' - \pi' + \pi - \Phi} x$$

$$\text{Lentis quintae} > \frac{\Phi}{\varepsilon \pi''' - \pi'' + \pi' - \pi + \Phi} x$$

$$\text{Lentis sextae} > \frac{\Phi}{2 \pi'''' - \pi''' + \pi'' - \pi' + \pi - \Phi} x$$

etc.

Denique

Denique ex §. 214 posito  $\mu=0,938191$  et  $\nu=0,232692$  erunt litterarum maiuscularum P, Q, R, etc. valores

$$P = \frac{\mu}{A^3 a^3} (A+1)(\lambda(A+1)^2 + \nu A)$$

$$Q = \frac{\mu}{B^3 b^3} (B+1)(\lambda'(B+1)^2 + \nu B)$$

$$R = \frac{\mu}{C^3 c^3} (C+1)(\lambda''(C+1)^2 + \nu C)$$

$$S = \frac{\mu}{D^3 d^3} (D+1)(\lambda'''(D+1)^2 + \nu D)$$

$$T = \frac{\mu}{E^3 e^3} (E+1)(\lambda''''(E+1)^2 + \nu E)$$

$$V = \frac{\mu}{F^3 f^3} (F+1)(\lambda'''''(F+1)^2 + \nu F)$$

etc.

His ita constitutis denominationibus, fit O distantia oculi post lentem ultimam, cuius distantia iuxta ponatur  $=l$ . Deinde statuatur exponens multiplicationis  $=m$ , relatus ad distantiam  $b$ , ita ut diameter obiecti per instrumentum visi  $m$  vicibus maior cernatur, quam si idem obiectum a nudo oculo in distantia  $=b$  aspiceretur: multiplicationi autem adiungi conuenit situm indicando, vtrum obiectum situm erecto an inuerso sit appariturum. Porro gradus claritatis  $\gamma$  denotet semidiametrum coni luminosi, qui a quouis obiecti puncto in oculum transmittitur posito semidiametro pupillae  $=\omega$ . Semidiametrum denique confusionis voco semidiametrum apparentem circulorum qui in nudo oculo aequae magni depinguntur, atque singula obiecti puncta per instrumentum in oculo.

Tom. I.

Nn

Coroll.



## Coroll. 1.

340. Quia pro apertura singularum lentium geminos limites inuenimus, semidiameter aperturæ cuiusque conuenientissime aggregato amborum limitum æqualis assumitur, vel saltem non minor. Vterque autem limes etsi forte alter prodeat negatiuus, affirmatiue accipi debet.

## Coroll. 2.

341. Hinc ergo sequentes consequimur formulas pro singularum lentium aperturis.

Semidiam : apert :

$$\text{Lentis secundae} = \frac{A B a \pi + x}{B \pi - \Phi} \cdot \Phi$$

$$\text{Lentis tertiæ} = \frac{A B C a \pi' + x}{C \pi' - \pi + \Phi} \cdot \Phi$$

$$\text{Lentis quartæ} = \frac{A B C D a \pi'' + x}{D \pi'' - \pi' + \pi - \Phi} \cdot \Phi$$

$$\text{Lentis quintæ} = \frac{A B C D E a \pi''' + x}{E \pi''' - \pi'' + \pi' - \pi + \Phi} \cdot \Phi$$

$$\text{Lentis sextæ} = \frac{A B C D E F a \pi'''' + x}{F \pi'''' - \pi''' + \pi'' - \pi' + \pi - \Phi} \cdot \Phi$$

etc.

## Coroll. 3.

342. Ex distantis binis determinatricibus cum numero arbitrario quaelibet lens facile construitur; id quod pro lente prima ex superioribus repetamus. Nempe si sit  $\lambda > 1$  lens simplex satisfaciet cuius constructio posito breuitatis gratia :

$$\xi = 0,190781; \sigma = 1,627401 \text{ et } \tau = 0,905130$$

ita

ita se habet :

$$\begin{array}{cc} \text{anterioris} & \text{posterioris} \\ \text{radius faciei} \dots & \frac{a\alpha}{\rho a + \sigma a \pm \tau(a+\alpha)\sqrt{\lambda-1}}; \frac{a\alpha}{\rho a + \sigma a \pm \tau(a+\alpha)\sqrt{\lambda-1}} \end{array}$$

seu numeris substitutis si fit  $\lambda = 1 + v$

$$\text{radius faciei} \begin{cases} \text{anterioris} = \frac{a\alpha}{+0,190781a+1,627401a \pm 0,905133(a+\alpha)\sqrt{v}} \\ \text{posterioris} = \frac{a\alpha}{+1,627401a+0,190781a \pm 0,905133(a+\alpha)\sqrt{v}} \end{cases}$$

### Coroll. 4.

343. Sin autem fit  $\lambda < 1$ , sed tamen  $> 0,191827$  lens est duplicanda seu ex duabus simplicibus componenda, quarum constructio, si ponatur  $\lambda = 0,191827 + v$  ita se habet

Pro Lente radius faciei

$$\text{priori} \begin{cases} \text{anterioris} = \frac{a\alpha}{-0,622911a+0,813700a \pm 0,905133(a+\alpha)\sqrt{v}} \\ \text{posterioris} = \frac{a\alpha}{+1,532010a+0,095390a \pm 0,905133(a+\alpha)\sqrt{v}} \end{cases}$$

$$\text{posteriori} \begin{cases} \text{anterioris} = \frac{a\alpha}{+0,095390a+1,532010a \pm 0,905133(a+\alpha)\sqrt{v}} \\ \text{posterioris} = \frac{a\alpha}{+0,813700a-0,622911a \pm 0,905133(a+\alpha)\sqrt{v}} \end{cases}$$

### Coroll. 5.

344. At si fit  $\lambda = 0,042165 + v$ , lente vtendum est triplicata ita construenda :

Nn 2

Pro-



Pro lente radius

$$\begin{aligned}
 \text{Priori} \quad \left\{ \begin{array}{l} \text{anterioris} = \frac{a \alpha}{-0,854153\alpha + 0,542467a + 0,05133(a+x)\sqrt{u}} \\ \text{posterioris} = \frac{a \alpha}{+1,500214\alpha + 0,063594a + 0,905133(a+\alpha)\sqrt{u}} \end{array} \right. \\
 \text{Media} \quad \left\{ \begin{array}{l} \text{anterioris} = \frac{a \alpha}{-0,415280\alpha + 1,021340a + 0,905133(a+x)\sqrt{u}} \\ \text{posterioris} = \frac{a \alpha}{+1,021340\alpha - 0,415280a + 0,905133(a+\alpha)\sqrt{u}} \end{array} \right. \\
 \text{Posteriori} \quad \left\{ \begin{array}{l} \text{anterioris} = \frac{a \alpha}{+0,063594\alpha + 1,500214a + 0,905133(a+\alpha)\sqrt{u}} \\ \text{posterioris} = \frac{a \alpha}{+0,542467\alpha - 0,854153a + 0,05133(a+\alpha)\sqrt{u}} \end{array} \right.
 \end{aligned}$$

## Coroll. 6.

345. Si denique fit  $\lambda = -0,010216 + u$  lens facienda est quadruplicata ita construenda:

Pro lente radius faciei

$$\begin{aligned}
 \text{Prima} \quad \left\{ \begin{array}{l} \text{anterioris} = \frac{a \alpha}{-1,029770\alpha + 0,906550a + 0,905133(a+x)\sqrt{u}} \\ \text{posterioris} = \frac{a \alpha}{+1,484315\alpha + 0,047615a + 0,905133(a+\alpha)\sqrt{u}} \end{array} \right. \\
 \text{Secunda} \quad \left\{ \begin{array}{l} \text{anterioris} = \frac{a \alpha}{-0,670615\alpha + 0,766005a + 0,905133(a+\alpha)\sqrt{u}} \\ \text{posterioris} = \frac{a \alpha}{+1,125160\alpha - 0,311460a + 0,905133(a+\alpha)\sqrt{u}} \end{array} \right. \\
 \text{Tertia} \quad \left\{ \begin{array}{l} \text{anterioris} = \frac{a \alpha}{-0,311460\alpha + 1,125160a + 0,905133(a+\alpha)\sqrt{u}} \\ \text{posterioris} = \frac{a \alpha}{+0,766005\alpha - 0,670615a + 0,905133(a+\alpha)\sqrt{u}} \end{array} \right. \\
 \text{Quarta} \quad \left\{ \begin{array}{l} \text{anterioris} = \frac{a \alpha}{+0,047615\alpha + 1,484315a + 0,905133(a+\alpha)\sqrt{u}} \\ \text{posterioris} = \frac{a \alpha}{+0,906550\alpha - 1,029770a + 0,905133(a+\alpha)\sqrt{u}} \end{array} \right.
 \end{aligned}$$

Scholion

## Scholion.

346. Quae in his corollariis de constructione lentis primae sunt allata, mutatis litteris  $a$  et  $\alpha$  ad reliquas lentes facile accommodari manifestum est quarum loco, si usus requirat, ut numeri  $\lambda'$ ,  $\lambda''$  etc. sint unitate minores, etiam lentes siue duplicatae siue triplicatae siue adeo quadruplicatae adhiberi debent. Ceterum hae formulae ad lentes quocunque patent ita ut proposito lentium numero quocunque tantum literae, quae ad lentes sequentes pertinerent, sint omittendae. In his autem denominationibus statim introduximus campum apparentem numero  $\Phi$  contentum, cum sit semidiameter spatii in obiecto conspicui  $z = a \Phi$ . Interim tamen campum apparentem non ad libitum augere licet, siquidem is multiplicationis ratione et numeris  $\pi$ ,  $\pi'$ ,  $\pi''$  etc. determinatur. Hi autem numeri semper infra  $\frac{1}{2}$  imo  $\frac{1}{3}$  subsistunt, ac plerumque  $\frac{1}{4}$  vel adeo  $\frac{1}{6}$  superare nequeunt; quandoque etiam minores accipi debent, id quod formulae coroll. 2. quovis casu declarabunt, quibus semidiameter aperturæ cuiusque lentis exprimitur. Quare retentis his denominationibus momenta constructionis pro quovis lentium numero determinato expendamus.

## Problema 4.

347. Si instrumentum dioptricum duabus constet lentibus, quarum crassitiem negligere liceat, definire momenta quibus constructio continetur.

N n 3.

So-



## Solutio.

Cum hic lens secunda sit vltima, momenta constructionis ita se habebunt.

1. Vt oculus imaginem in distantia iusta conspiciat, debet esse  $O = \xi + l$  ideoque

$$O = l + \frac{ABa\Phi}{\mathfrak{B}\pi - \Phi} \text{ et ex conditione campi apparentis}$$

$$O = \frac{\mathfrak{B}b\pi}{\pi - \Phi}$$

2. Exponens multiplicationis  $m$  ad distantiam  $b$  relatus praebet hanc aequationem

$$m = AB \frac{b}{l} \text{ pro situ erecto}$$

sin autem situs inuersus desideretur, numerus  $m$  negatiue est capiendus.

$$3. \text{ Pro gradu claritatis semper est } \gamma = \frac{bx}{ma}.$$

4. Pro campo apparente habebitur haec aequatio:

$$\pi - \Phi = -\frac{ma\Phi}{b} \text{ vnde fit } \Phi = \frac{-\pi b}{ma - b}.$$

5. Si distantia  $O$  hinc prodeat negatiua, capi debet  $O = 0$  seu  $\frac{ABa\Phi}{\mathfrak{B}\pi - \Phi} = -l$  tum vero  $\pi$  eiusmodi valorem induit, vt apertura lentis ocularis non superet aperturam pupillae; fiet nempe  $\pi \mathfrak{B}b = \omega$ , hincque

$$\Phi = \frac{\mathfrak{B}\pi\omega}{A\mathfrak{B}a\pi + \omega} \text{ vel } (B + 1)\omega = -\pi l.$$

6. Semidiameter confusionis sequenti formula exprimetur.

$$\frac{Bbb}{+Aal} x^3 \left( \frac{AAaa}{bb} P + \frac{bb}{AAaa} Q \right)$$

quae

quae factis substitutionibus abit in hanc :

$$\frac{\mu m x^3}{4 a a' b} \left( \frac{(A+1)(\lambda(A+1)^2 + \nu A)}{A^3} + \frac{\Phi(B+1)(\lambda'(B+1)^2 + \nu B)}{A^3 B (\mathfrak{B}\pi - \Phi)} \right)$$

7. Vt obiectum sine margine colorato appareat, si distantia O fuerit positiva, necesse est sit :

$$\frac{\pi}{\mathfrak{B}\pi - \Phi} = 0$$

8. Sin autem capiatur  $O = 0$ , fieri debet  $\frac{(A+1)\mathfrak{B}\pi}{\Phi} = 0$

$$\text{feu } \frac{\pi}{\mathfrak{B}\pi} = 0.$$

9. Omnis autem confusio a diuersa radiorum refrangibilitate oriunda penitus tollitur si insuper fuerit.

$$\frac{1}{\mathfrak{B}} + \frac{\Phi}{A\mathfrak{B}(\mathfrak{B}\pi - \Phi)} = 0$$

10. Praeterea autem erit distantia lentium

$$\frac{A\mathfrak{B}a\pi}{\mathfrak{B}\pi - \Phi} \text{ quae debet esse positiva.}$$

11. Denique Semidiameter aperturae lentis ocularis est ;

$$\frac{A\mathfrak{B}a\pi + x}{\mathfrak{B}\pi - \Phi} \cdot \Phi$$

vbi signorum  $\pm$  id capi debet, quod valorem praebet maximum siue positium siue negatiuum.

### C O R O L L. I.

348. Cum sit (n°.4)  $\Phi = \frac{-\pi b}{ma-b}$ , et ob multiplicationem  $B = \frac{ml}{Ab}$ , hincque  $\mathfrak{B} = \frac{ml}{ml + Ab}$ ; habebitur pro loco oculi  $O = \frac{-Ab l (ma-b)}{m m a l + Ab b}$ .

Tum



Tum vero est  $\mathfrak{B}\pi - \Phi = \frac{ml\pi}{ml + Ab} + \frac{\pi b}{ma - b} = \frac{\pi(mma + Abb)}{(ma - b)(ml + Ab)}$   
 unde fit lentium distantia  $\frac{A\mathfrak{B}a\pi}{\mathfrak{B}\pi - \Phi} = \frac{mAa(ma - b)}{mma + Abb} l$ .

## Coroll. 2.

349. Cum sit  $\frac{A}{A+1} = \mathfrak{A}$  et  $A = \frac{\mathfrak{A}}{1-\mathfrak{A}}$  semidia-  
 meter confusionis etiam ita exprimi potest vt sit

$$\frac{\mu m x^3}{4 a a b} \left( \frac{\lambda + v \mathfrak{A} (1 - \mathfrak{A})}{\mathfrak{A}^3} + \frac{\Phi (\lambda + v \mathfrak{B} (1 - \mathfrak{B}))}{A \mathfrak{B}^2 (\mathfrak{B}\pi - \Phi)} \right).$$

## Coroll. 3.

350. Cum sit lentium distantia  $= \frac{A\mathfrak{B}a\pi}{\mathfrak{B}\pi - \Phi}$ , obie-  
 ctum sine coloribus cerni nequit, nisi lentium di-  
 stantia euanescat; siquidem oculus in eo loco, quem  
 campus apparens postulat, teneatur. Et quoniam  
 huic conditioni satisfieri nequit, multo minus tota con-  
 fusio ex n°. 9 tolli poterit.

## Scholion.

351. Hic statim ab instrumentis duabus lenti-  
 bus instructis incepti, quoniam lens vnica facillime ex-  
 peditur. Primum enim necesse est vt sit  $O = l + a = l + Aa$ .  
 Tum vero multiplicatio praebebat  $m = \frac{b}{a}$  pro situ erecto,  
 et campus apparens non terminatur sumendo  $Aa + l = 0$ ,  
 seu  $O = 0$ , ita vt oculus lenti immediate debeat ap-  
 plicari. Semidiameter vero confusionis erit

$$= \frac{a}{4l} x^3. P = \frac{Aa x^3}{4l}. P = -\frac{1}{4} x^3. P = \frac{m a x^3}{4b}. P \text{ ob } l = -Aa \text{ et } ma = b$$

Quare semidiameter confusionis erit

$$\frac{\mu m x^3}{4 a a b} \left( \frac{(A+1)(\lambda(A+1)^2 + vA)}{A^3} \right) = \frac{\mu m x^3}{4 a a b} \cdot \frac{\lambda + v \mathfrak{A} (1 - \mathfrak{A})}{\mathfrak{A}^3}.$$

Pro

Pro gradu claritatis autem habetur vt semper  
 $y = \frac{bx}{m a}.$

Obiectum porro hoc casu sine margine colorato cernetur, tota autem confusio tolli nequit, nisi sit  $\gamma = 0$  seu  $A = -1$ , id quod scopo lentium repugnat. Quare ad considerationem plurium lentium progrediamur.

### Problema 5.

352. Si instrumentum dioptricum tribus instructum sit lentibus, quarum crassities tam sit parva, vt negligi queat, definire cuncta momenta ad constructionem dirigendam necessaria.

### Solutio.

Hic igitur lens tertia erit vltima seu ocularis, ideoque momenta sequenti modo se habebunt:

1. Vt oculo imago in distantia iusta spectanda offeratur debet esse  $O = \gamma + l$ , ac loco  $\gamma$  valore substituto

$$O = l + \frac{ABC a \Phi}{\epsilon \pi' - \pi + \Phi} \text{ et ex conditione campi apparentis}$$

$$O = \frac{\epsilon c \pi'}{\pi' - \pi + \Phi}.$$

2. Exponens multiplicationis  $m$  ad distantiam  $h$  relatus praebet hanc aequationem:

$$m = -\frac{ABC b}{l} \text{ pro situ erecto.}$$

vnde si situs inuersus desideretur, numerum  $m$  negative accipi conuenit.

Tom. I.

Oo

3. Pro



3. Pro gradu claritatis habemus, vt semper,  
 $y = \frac{b \ x}{m a}.$

4. Pro campo autem apparente definiendo habemus hanc aequationem

$$\Phi = \frac{-\pi + \pi'}{m a - b} \cdot b.$$

5. At si distantia O hinc prodeat negatiua, vt capi oporteat  $O = 0$ , ideoque  $\frac{A B C a \Phi}{\epsilon \pi' - \pi + \Phi} = -l$  campus apparens definiri debet ex hac aequatione  $\pi' \epsilon c = \omega$  seu  $\frac{A B \epsilon a \Phi \pi'}{\epsilon \pi' - \pi + \Phi} = \omega$  vel  $(C + 1) \omega = -\pi' l$ .

6. Semidiameter confusionis vero ita exprimetur:

$$\frac{c c c'}{4 A B a l} x^3 \left( \frac{A A B B a a}{c c} P + \frac{B B b b}{A A a a c c} Q + \frac{c c}{A A B B a a} R \right)$$

quae ob  $\frac{1}{l} = \frac{-m}{A B C b}$  abit signo mutato in hanc:

$$\frac{\mu m x^3}{4 a a b} \left\{ \begin{aligned} &+ \frac{(A + 1)(\lambda(A + 1)^2 + \nu A)}{A^3} \\ &+ \frac{\Phi(B + 1)(\lambda'(B + 1)^2 + \nu B)}{A^3 B^3 (\mathfrak{B} \pi - \Phi)} \\ &+ \frac{\Phi(C + 1)(\lambda''(C + 1)^2 + \nu C)}{A^3 B^3 C^3 (\epsilon \pi' - \pi + \Phi)} \end{aligned} \right.$$

7. Vt obiectum sine margine colorato appareat, si quidem distantia O prodierit positiua, esse oportet.

$$\frac{\pi}{\mathfrak{B} \pi - \Phi} + \frac{\pi'}{\epsilon \pi' - \pi + \Phi} = 0$$

8. Sin autem ob istam distantiam prodeuntem negatiuam capiatur  $O = 0$ , margo coloratus euanesceat faciendo:

$$\frac{\pi'}{2 \Phi} + \frac{\pi'}{A \mathfrak{B} (\mathfrak{B} \pi - \Phi)} = \frac{\pi}{A B \epsilon (\mathfrak{B} \pi - \Phi)}$$

9. Quod

9. Quod si marginem coloratum tollere liceat, praeterea visio ab omni confusione liberabitur, si huic aequationi satisfiat.

$$\frac{1}{2} + \frac{\Phi}{AB(\mathfrak{B}\pi - \Phi)} + \frac{\Phi}{AB\mathfrak{C}(\mathfrak{C}\pi' - \pi + \Phi)} = 0.$$

10. Denique effici debet, ut distantia lentium fiat positiua vnde habebitur:

Interuallum

Lentium

$$\text{I et II} = \frac{ABa\Phi}{\mathfrak{B}\pi - \Phi} > 0$$

$$\text{II et III} = \frac{ABa\Phi(\mathfrak{C}\pi' - (1 - \mathfrak{B})\pi)}{(\mathfrak{B}\pi - \Phi)(\mathfrak{C}\pi' - \pi + \Phi)} > 0$$

11. Tandem numeros  $\pi$  et  $\pi'$  ita accipi oportet, ut aperturae lentium non fiant nimis magnae: est vero

Semidiameter

aperturae

$$\text{Lentis secundae} = \frac{ABa\pi + x}{\mathfrak{B}\pi - \Phi} \cdot \Phi$$

$$\text{Lentis tertiae} = \frac{AB\mathfrak{C}\pi' + x}{\mathfrak{C}\pi' - \pi + \Phi} \cdot \Phi.$$

### C o r o l l. I.

353. Cum sit  $C = \frac{-ml}{ABh}$ , ideoque  $\mathfrak{C} = \frac{-ml}{ABh - ml}$ ; et  $\Phi = \frac{-\pi + \pi'}{ma - b} \cdot b$ , si hi valores substituantur erit pro loco oculi.

$$O = \frac{ABbl(ma - b)\pi'}{(mml - ABbh)\pi' + ma(ABh - ml)\pi}$$

O o 2

Coroll.



## Coroll. 2.

354. Semidiameter confusionis etiam ita expressi potest ut fit

$$\frac{\mu m x^3}{4 a a b} \left( \frac{\lambda + v \mathfrak{A} (1 - \mathfrak{A})}{\mathfrak{A}^2} + \frac{\Phi (\lambda' + v \mathfrak{B} (1 - \mathfrak{B}))}{A^2 \mathfrak{B}^2 (\mathfrak{B} \pi - \Phi)} + \frac{\Phi (\lambda'' + v \mathfrak{C} (1 - \mathfrak{C}))}{A^2 B^2 \mathfrak{C}^2 (\mathfrak{C} \pi' - \pi + \Phi)} \right)$$

cuius formae cohaerentia cum praecedentibus iam ita est manifesta, ut ad plures lentes facile extendi queat.

## Problema 6.

355. Si instrumentum dioptricum quatuor lentibus fit instructum, quarum crassities in computum duci non mereatur, definire omnia momenta ad constructionem dirigendam necessaria.

## Solutio.

Quia hic lens quarta est ultima.

1. Ut oculo imago postrema in distantia iusta offeratur debet esse  $O = \delta + l$ , ideoque

$$O = l + \frac{ABCD a \Phi}{\mathfrak{D} \pi'' - \pi' + \pi - \Phi} \text{ et ex conditione campi apparentis}$$

$$O = \frac{\mathfrak{D} d \pi''}{\pi'' - \pi' + \pi - \Phi}.$$

2. Exponens multiplicationis  $m$  ad distantiam  $b$  relatus, si obiectum ut hactenus situ erecto exhiberi assumatur, erit

$$m = +ABCD \cdot \frac{b}{l} \text{ pro situ erecto.}$$

sin autem situs desideretur inuersus numerus  $m$  negative est accipiendus.

3. Pro

3. Pro gradu claritatis perpetuo habetur  $y = \frac{bx}{ma}$ .

4. Campus autem apparens definiri debet ex hac aequatione

$$\Phi = \frac{-\pi + \pi' - \pi''}{m a - b} b$$

hinc enim  $a\Phi$  praebebit semidiametrum spatii in obiecto conspici; siquidem distantia  $O$  fuerit positiva.

5. Sin autem distantia  $O$  prodeat negativa quo casu capi oportet  $O = 0$ , seu  $\frac{ABCD a \Phi}{D \pi'' - \pi' + \pi - \Phi} = -l$  quantitas  $\Phi$  determinari debet ex hac aequatione  $\pi'' D d = \omega$  seu hac

$$\frac{ABCD a \pi'' \Phi}{D \pi'' - \pi' + \pi - \Phi} = \omega \text{ siue } (D + 1) \omega = -\pi'' l$$

6. Semidiameter confusionis ita reperietur expressus:

$$\frac{\mu m x^3}{4 a a b} \left\{ \begin{array}{l} + \frac{(A+1)(\lambda'(A+1)^2 + vA)}{A^2} \\ + \frac{\Phi(B+1)(\lambda'(B+1)^2 + vB)}{A^3 B^3 (D \pi - \Phi)} \\ + \frac{\Phi(C+1)(\lambda''(C+1)^2 + vC)}{A^2 B^2 C^3 (E \pi' - \pi + \Phi)} \\ + \frac{\Phi(D+1)(\lambda'''(D+1)^2 + vD)}{A^3 B^3 C^3 D^3 (D \pi'' - \pi' + \pi - \Phi)} \end{array} \right.$$

7. Vt obiectum sine margine colorato appareat, siquidem distantia  $O$  sit positiva esse oportet:

$$\frac{\pi}{D \pi - \Phi} + \frac{\pi'}{E \pi' - \pi + \Phi} + \frac{\pi''}{D \pi'' - \pi' + \pi - \Phi} = 0.$$

8. Verum si acceperimus  $O = 0$ , margo coloratus evanescet, huic aequationi satisfaciendo.

$$\frac{\pi''}{2\Phi} + \frac{\pi''}{A^2 B^2 D \pi - \Phi} + \frac{\pi''}{ABC(E \pi' - \pi + \Phi)} = \frac{\pi}{ABC D (D \pi - \Phi)} + \frac{\pi'}{ABC D (E \pi' - \pi + \Phi)}.$$

O O 3

9. Omnis



9. Omnis autem confusio a diuersa radiorum natura oriunda penitus tolletur, si satisfieri liceat huic aequationi:

$$\frac{1}{u} + \frac{\Phi}{AB(\mathfrak{B}\pi - \Phi)} + \frac{\Phi}{AB\mathfrak{E}(\mathfrak{E}\pi' - \pi + \Phi)} + \frac{\Phi}{ABC\mathfrak{D}(\mathfrak{D}\pi'' - \pi' + \pi - \Phi)} = 0.$$

10. Denique efficiendum est vt distantiae lentium sint positivae vnde oriuntur formulae sequentes:

Interuallum

Lentium

$$\text{I et II} = \frac{ABa\pi}{\mathfrak{B}\pi - \Phi} > 0$$

$$\text{II et III} = \frac{ABa\Phi(\mathfrak{E}\pi' - (1 - \mathfrak{B})\pi)}{(\mathfrak{B}\pi - \Phi)(\mathfrak{E}\pi' - \pi + \Phi)} > 0$$

$$\text{III et IV} = \frac{ABCa\Phi(\mathfrak{D}\pi'' - (\mathfrak{E}\pi' - \pi + \Phi))}{(\mathfrak{E}\pi' - \pi + \Phi)(\mathfrak{D}\pi'' - \pi' + \pi - \Phi)} > 0.$$

11. Aperturae autem singularum lentium ita sunt capiendae vt sit.

Semid : apert :

$$\text{Lentis secundae} = \frac{ABa\pi + x}{\mathfrak{B}\pi - \Phi} \cdot \Phi$$

$$\text{Lentis tertiae} = \frac{AB\mathfrak{E}a\pi' + x}{\mathfrak{E}\pi' - \pi + \Phi} \cdot \Phi$$

$$\text{Lentis quartae} = \frac{ABC\mathfrak{D}a\pi'' + x}{\mathfrak{D}\pi'' - \pi' + \pi - \Phi} \cdot \Phi.$$

### COROLL. I.

356. Cum sit ob multiplicationem  $D = \frac{ml}{ABCb}$  ideoque  $\mathfrak{D} = \frac{ml}{ABCb + ml}$  et  $\Phi = \frac{-\pi + \pi' - \pi''}{ma - b} \cdot b$ ; pro loco oculi habebitur haec expressio:

$$O = \frac{-ABCbl(ma - b)\pi''}{(mma + ABCbh)\pi'' - ma(ABCb + ml)(\pi' - \pi)}.$$

Coroll.

## Coroll. 2.

357. Quia hic tot occurrunt litterae determinandae, dubium nullum esse videtur, quin formulis sub n°. 7 vel 8 et 9 contentis satisfieri queat, et certum saltem est, id fieri posse, si plura adhibeantur mediorum refringentium genera.

## Problema 7.

358. Si instrumentum dioptricum quinque lentibus sit instructum, quarum crassities ob paruitatem negligi queat, definire omnia momenta constructionem dirigentia.

## Solutio.

Cum hic lens quinta vltimum locum teneat; habebimus.

1. Pro loco oculi, vt imago visa ipsi in iusta distantia offeratur debet esse  $O = \varepsilon + l$  ideoque

$$O = l + \frac{ABCDEa\Phi}{\varepsilon\pi'' - \pi'' + \pi' - \pi + \Phi} \text{ et ex conditione campi apparentis}$$

$$O = \frac{\varepsilon e \pi'''}{\pi''' - \pi'' + \pi' - \pi + \Phi}.$$

2. Exponens multiplicationis  $m$  ad distantiam  $= b$  relatus, si quidem obiectum situ erecto sit spectandum

$$m = -ABCDE. \frac{b}{l}.$$

pro situ autem inuerso numerus  $m$  negatiue capi debet.

3. Pro



3. Pro gradu claritatis est vt haecenus  $y = \frac{bx}{ma}$ .

4. Campus apparens definiri debet ex hac aequatione

$$\Phi = \frac{-\pi + \pi' - \pi'' + \pi'''}{m a - b} \cdot b$$

siquidem distantia illa O prodierit positiua.

5. Sin autem haec distantia prodeat negatiua vt capiatur  $O = 0$ , seu  $\frac{A B C D E \gamma \Phi}{\epsilon \pi''' - \pi'' + \pi' - \pi + \Phi} = -l$ , campus apparens non ex formula n. 4. sed ex hac  $\pi''' \epsilon e = \omega$  defini debet, unde si

$$\frac{A B C D E \gamma \pi''' \Phi}{\epsilon \pi''' - \pi'' + \pi' - \pi + \Phi} = \omega \text{ seu } (E + 1) \omega = -\pi''' l$$

6. Semidiameter confusionis ab apertura lentium oriundae sequenti modo exprimitur,

$$\frac{\mu m x^3}{4 a a b} \left\{ \begin{array}{l} + \frac{(A+1)(\lambda(A+1)^2 + \nu A)}{A^3} \\ + \frac{\Phi(B+1)(\lambda'(B+1)^2 + \nu B)}{A^3 B^3 (\epsilon \pi - \Phi)} \\ + \frac{\Phi(C+1)(\lambda''(C+1)^2 + \nu C)}{A^3 B^3 C^3 (\epsilon \pi' - \pi + \Phi)} \\ + \frac{\Phi(D+1)(\lambda'''(D+1)^2 + \nu D)}{A^3 B^3 C^3 D^3 (\epsilon \pi'' - \pi' + \pi - \Phi)} \\ + \frac{\Phi(E+1)(\lambda''''(E+1)^2 + \nu E)}{A^3 B^3 C^3 D^3 E^3 (\epsilon \pi''' - \pi'' + \pi' - \pi + \Phi)} \end{array} \right.$$

7. Vt obiectum sine margine colorato appareat, siquidem distantia O fuerit positiua, huic aequationi erit satisfaciendum:

$$\frac{\pi}{\epsilon \pi - \Phi} + \frac{\pi'}{\epsilon \pi' - \pi + \Phi} + \frac{\pi''}{\epsilon \pi'' - \pi' + \pi - \Phi} + \frac{\pi'''}{\epsilon \pi''' - \pi'' + \pi' - \pi + \Phi} = 0.$$

8. At

8. At si distantia  $O$  sit negatiua, capiaturque  $O=0$  pro hoc scopo obtinendo oportet sit

$$\frac{\pi'''}{2\Phi} + \frac{\pi'''}{AB(2\pi-\Phi)} + \frac{\pi'''}{ABE(\epsilon\pi'-\pi+\Phi)} + \frac{\pi'''}{ABCD(D\pi''-\pi'+\pi-\Phi)} \\ = \frac{1}{ABCD\epsilon} \left( \frac{\pi}{2\pi-\Phi} + \frac{\pi'}{\epsilon\pi'-\pi+\Phi} + \frac{\pi''}{D\pi''-\pi'+\pi-\Phi} \right).$$

9. Omnis autem confusio a diuersa radiorum natura oriunda penitus tolletur, si fieri possit

$$\frac{1}{2} + \frac{\Phi}{AB(2\pi-\Phi)} + \frac{\Phi}{ABE(\epsilon\pi'-\pi+\Phi)} + \frac{\Phi}{ABCD(D\pi''-\pi'+\pi-\Phi)} + \frac{\Phi}{ABCD\epsilon(\epsilon\pi'''-\pi''+\pi'-\pi+\Phi)} = 0$$

10. Tum vero efficiendum est, vt interualla lentium omnia sint positiua, vnde oriuntur formulae sequentes.

Interuallum

Lentium

$$\text{I. et II.} = \frac{ABa\pi}{2\pi-\Phi} > 0$$

$$\text{II. et III.} = \frac{ABa\Phi(\epsilon\pi'-(1-2)\pi)}{(2\pi-\Phi)(\epsilon\pi'-\pi+\Phi)} > 0$$

$$\text{III. et IV.} = \frac{ABCa\Phi(D\pi''-(1-\epsilon)\pi')}{(\epsilon\pi'-\pi+\Phi)(D\pi''-\pi'+\pi-\Phi)} > 0$$

$$\text{IV. et V.} = \frac{ABCDa\Phi(\epsilon\pi'''-(1-D)\pi'')}{(D\pi''-\pi'+\pi-\Phi)(\epsilon\pi'''-\pi''+\pi'-\pi+\Phi)} > 0$$

11. Aperturae denique singularum lentium ita sunt capiendae, vt sit:

Semid. apert:

$$\text{Lentis secundae} = \frac{ABa\pi+x}{2\pi-\Phi} \Phi$$

$$\text{Lentis tertiae} = \frac{ABCa\pi'+x}{\epsilon\pi'-\pi+\Phi} \Phi$$

$$\text{Lentis quartae} = \frac{ABCDa\pi''+x}{D\pi''-\pi'+\pi-\Phi} \Phi$$

$$\text{Lentis quintae} = \frac{ABCD\epsilon a\pi''' + x}{\epsilon\pi''' - \pi'' + \pi' - \pi + \Phi} \Phi.$$

Tom. I.

P p

Corol-



## Corollarium.

359. Ob multiplicationem ergo erit :

$$E = \frac{-ml}{ABCDh} \text{ et } \mathcal{E} = \frac{-ml}{ABCDh - ml}$$

Vnde oritur :

$$O = \frac{ABCDhl(ma - b)\pi'''}{(mla - ABCDhb)\pi'' + ma(ABCDh - ml)(\pi'' - \pi' + \pi)}.$$

## Scholion.

360. Hactenus supposui oculum vel in eo loco teneri, quem visio campi apparentis postulat, vel lenti postremae immediate applicari; quorum utroque casu efficiendum est, ut imago ultima oculo ad distantiam iustam offeratur. Hic quidem locum oculi, quem campus praebet statim coniunxi cum iusta imaginis spectandae distantia; sed etiam sine respectu siue ad hanc distantiam siue ad multiplicationem habito definiri potest. Ex superioribus enim §. 271 habemus

Pro casu

unius lentis  $O = 0$

duarum lentium  $O = \frac{A\mathfrak{B}a\pi\Phi}{(\pi - \Phi)(\mathfrak{B}\pi - \Phi)}$

trium lentium  $O = \frac{A\mathfrak{B}\mathcal{E}a\pi'\Phi}{(\pi' - \pi + \Phi)(\mathcal{E}\pi' - \pi + \Phi)}$

quatuor lentium  $O = \frac{ABCDa\pi''\Phi}{(\pi'' - \pi' + \pi - \Phi)(D\pi'' - \pi' + \pi - \Phi)}$

quinque lentium  $O = \frac{ABCDEa\pi'''\Phi}{(\pi''' - \pi'' + \pi' - \pi + \Phi)(\mathcal{E}\pi''' - \pi'' + \pi' - \pi + \Phi)}$   
etc.

Verum

Verum si nolumus ad campum apparentem respicere semper eiusmodi locum oculo assignare licet, unde obiectum sine margine colorato cernatur. Haud abs re autem fore arbitror hunc locum determinasse, quia ex eius distantia a loco oculi ob rationem campi assumpto facilius iudicare licebit, quanto margine cinctum apparere debeat obiectum si oculus in loco isto posteriori teneatur. Si enim aequationi n°. 7 appositae satisfieri nequeat, nil aliud intelligimus, nisi obiectum non sine margine colorato esse appariturum at si insuper constet ille locus, unde obiectum sine huiusmodi margine spectari posset, facilius quantitatem istius marginis colligere poterimus.

### Problema 8.

361. Si instrumentum dioptricum ex lentibus quocunque constet, quarum crassitiem negligere liceat, eum pro oculo locum assignare, unde obiectum sine margine colorato conspiciatur.

### Solutio.

Ponatur haec oculi post lentem ultimam distantia  $= \Omega$ , atque ex aequationibus §. 325 datis, si pro O scribatur  $\Omega$  haec ipsa distantia  $\Omega$ , quam quaerimus, elici potest. Scilicet pro quovis lentium numero erit ut sequitur.



## I. Pro vnica Lente.

Habetur haec aequatio  $\frac{\Omega d \alpha}{\alpha(\alpha - \Omega)} = 0$  vnde fit  $\frac{\Omega}{\alpha(\alpha - \Omega)} = 0$ ,  
 seu  $\frac{\Omega}{\alpha - \Omega} = 0$ . Verum etiam campus apparens exigit  
 $0 = 0$ , vnde hoc casu ambo loca oculi  $O$  et  $\Omega$  con-  
 veniunt, neque margo coloratus est pertimescendus.

## II. Pro duabus Lentibus

Hoc casu reperta est ista aequatio

$$\frac{\Omega d \epsilon}{\epsilon(\epsilon - \Omega)} = d\alpha \left( \frac{1}{a} + \frac{1}{b} \right)$$

vnde substitutis valoribus supra assignatis oritur

$$\frac{\Omega}{\epsilon(\epsilon - \Omega)} = \frac{\pi}{2\Phi} : (B + 1) AB \alpha \left( \frac{1}{a} + \frac{\Phi}{AB(2\pi - \Phi)} \right)$$

seu si ponamus

$$Y = \frac{\pi}{2\Phi} \text{ et } Z = \frac{1}{a} + \frac{\Phi}{AB(2\pi - \Phi)}$$

erit

$$\frac{\Omega}{\epsilon - \Omega} = \frac{Y}{AB(B + 1)\alpha Z} \cdot \epsilon = \frac{Y}{(B + 1)Z} \cdot \frac{\Phi}{2\pi - \Phi}$$

## III. Pro tribus Lentibus.

Eodem modo si ponamus

$$Y = \frac{\pi'}{2\Phi} + \frac{\pi'}{AB(2\pi - \Phi)} - \frac{\pi}{AB\epsilon(2\pi - \Phi)} \quad \text{et}$$

$$Z = \frac{1}{a} + \frac{\Phi}{AB(2\pi - \Phi)} + \frac{\Phi}{AB\epsilon(\epsilon\pi' - \pi + \Phi)}$$

erit pro casu trium lentium

$$\frac{\Omega}{\gamma - \Omega} = \frac{Y}{ABC(C + 1)\alpha Z} \cdot \gamma = \frac{Y}{(C + 1)Z} \cdot \frac{\Phi}{\epsilon\pi' - \pi + \Phi}$$

vbi

vbi notandum est.  $Y \neq 0$  et  $Z \neq 0$  esse eas ipsas aequationes, quae in superioribus problematibus sub numeris 8 et 9 retulimus.

#### IV. Pro quatuor lentibus

Si iam ponatur:

$$Y = \frac{\pi''}{2\phi} + \frac{\pi''}{AB(\pi - \phi)} + \frac{\pi''}{ABE(\epsilon\pi' - \pi + \phi)} - \frac{1}{ABCD} \left( \frac{\pi}{\pi - \phi} + \frac{\pi'}{\epsilon\pi' - \pi + \phi} \right)$$

$$Z = \frac{1}{2} + \frac{\phi}{AB(\pi - \phi)} + \frac{\phi}{ABE(\epsilon\pi' - \pi + \phi)} + \frac{\phi}{ABCD(\pi'' - \pi' + \pi - \phi)}$$

erit:

$$\frac{\Omega}{\delta - \Omega} = \frac{Y}{ABCD(D + 1)\alpha Z} \delta$$

$$\text{hincque } \frac{\Omega}{\delta - \Omega} = \frac{Y}{(D + 1)Z} \frac{\phi}{\pi'' - \pi' + \pi - \phi}.$$

#### V. Pro quinque Lentibus.

Posito secundum aequationes n<sup>o</sup>. 8 et 9 exhibitas

$$Y = \frac{\pi'''}{2\phi} + \frac{\pi'''}{AB(\pi - \phi)} + \frac{\pi'''}{ABE(\epsilon\pi' - \pi + \phi)} + \frac{\pi'''}{ABCD(\pi'' - \pi' + \pi - \phi)} \\ - \frac{1}{ABCDE} \left( \frac{\pi}{\pi - \phi} + \frac{\pi'}{\epsilon\pi' - \pi + \phi} + \frac{\pi''}{\pi'' - \pi' + \pi - \phi} \right)$$

$$Z = \frac{1}{2} + \frac{\phi}{AB(\pi - \phi)} + \frac{\phi}{ABE(\epsilon\pi' - \pi + \phi)} + \frac{\phi}{ABCD(\pi'' - \pi' + \pi - \phi)} \\ + \frac{\phi}{ABCDE(\epsilon\pi''' - \pi'' + \pi' - \pi + \phi)}$$

habebimus

$$\frac{\Omega}{\epsilon - \Omega} = \frac{Y}{ABCDE(E + 1)\alpha Z} \cdot \epsilon \text{ vel}$$

$$\frac{\Omega}{\epsilon - \Omega} = \frac{Y}{(E + 1)Z} \frac{\phi}{\epsilon\pi''' - \pi'' + \pi' - \pi + \phi}.$$

Lex harum formularum pro quouis maiori lentium numero hinc satis est perspicua.



## Coroll. I.

362. Cum igitur supra pro quouis lentium numero exhibuerimus aequationes n°. 8. et n°. 9. formula n°. 8. dabit valorem Y et formula n°. 9. valorem ipsius Z, quibus cognitis locus oculi, vbi margo coloratus euanescit, facile innotescit.

## Coroll. 2.

363. Si praeterea ex aequatione n°. 7. exposita ponatur

$$X = \frac{\pi}{\mathfrak{B}\pi - \Phi} + \frac{\pi'}{\mathfrak{E}\pi' - \pi + \Phi} + \frac{\pi''}{\mathfrak{D}\pi'' - \pi' + \pi - \Phi} + \frac{\pi'''}{\mathfrak{E}\pi''' - \pi'' + \pi' - \pi + \Phi}$$

erit quidem pro casu quinque lentium

$$Y = \frac{\pi'''}{\Phi} Z - \frac{1}{\text{ABCD}\mathfrak{E}} X$$

$$\text{Pro quatuor lentibus } Y = \frac{\pi''}{\Phi} Z - \frac{1}{\text{ABCD}} X$$

$$\text{Pro tribus lentibus } Y = \frac{\pi'}{\Phi} Z - \frac{1}{\text{AB}\mathfrak{E}} X$$

$$\text{et pro duabus lentibus } Y = \frac{\pi}{\Phi} Z - \frac{1}{\text{A}\mathfrak{B}} X.$$

## Scholion.

364. Quo facilius haec oculi loca littera  $\Omega$  designata cum praecedentibus littera O designatis comparare liceat, notandum est esse.

Pro casu

$$\text{Vnius Lentis } \frac{O}{\alpha - O} = 0$$

$$\text{Duarum Lentium } \frac{O}{\mathfrak{E} - O} = \frac{\pi}{(\mathfrak{E} + 1)(\mathfrak{B}\pi - \Phi)}$$

$$\text{Trium Lentium } \frac{O}{\gamma - O} = \frac{\pi'}{(\mathfrak{C} + 1)(\mathfrak{E}\pi' - \pi + \Phi)}$$

$$\text{Quatuor Lentium } \frac{O}{\delta - O} = \frac{\pi''}{(\mathfrak{D} + 1)(\mathfrak{D}\pi'' - \pi' + \pi - \Phi)}$$

$$\text{Quinque Lentium } \frac{O}{\varepsilon - O} = \frac{\pi'''}{(\mathfrak{E} + 1)(\mathfrak{E}\pi''' - \pi'' + \pi' - \pi + \Phi)} \text{ etc.}$$

Quare

Quare vt ambo loca congruant retentis pro quouis lentium numero litteris Y et Z modo adhibitis, oportet fit

Pro vna Lente  $\circ = \circ$

Pro duabus Lentibus  $\frac{Y}{Z} = \frac{\pi}{\phi}$

Pro tribus Lentibus  $\frac{Y}{Z} = \frac{\pi'}{\phi}$

Pro quatuor Lentibus  $\frac{Y}{Z} = \frac{\pi''}{\phi}$

Pro quinque Lentibus  $\frac{Y}{Z} = \frac{\pi'''}{\phi}$  etc.

vnde generatim fit  $X = 0$ , quae est ipsa aequatio n°. 7. allata. Haec igitur sunt principia generalia, ex quibus constructio instrumentorum dioptricum erit perficienda, quae quidem instrumenta hic potissimum ad visionem accommodaui. Nihilo vero minus ad repraesentationem obiectorum in camera obscura super Tabula alba adhiberi possunt, vbi praeter iam data praecepta tenendum est, has effigies etiam sine margine colorato esse apparituras, si aequationibus n°. 7. satisfiat, omnem vero confusionem a diuersa radiorum natura oriundam penitus tolli, si insuper aequationibus n°. 9 satisfiat.

## Supplementum VII.

Si ratio refractionis in singulis lentibus fuerit diuersa et secundum lentium ordinem litteris  $n, n', n'', n'''$  etc. exhibeatur, inde pro lentium constructione

I°. Litterae respondentes  $\varrho, \sigma, \tau$  et pro confusione



fusione litterae  $\mu$  et  $\nu$  conuenienter definiri debent secundum formulas supra datas §. 55.

$$\varrho = \frac{1}{2(n-1)} + \frac{1}{n+2} - 1; \sigma = \frac{1}{2(n-1)} - \frac{1}{n+2} + 1$$

$$\tau = \frac{1}{3} \left( \frac{1}{2(n-1)} + \frac{1}{n+2} \right) V(4n-1)$$

ita, vt fit

$$\varrho + \sigma = \frac{1}{n-1} \text{ et } \sigma - \varrho = \frac{2}{n+2} + 2 = \frac{2(n+1)}{n+2}.$$

Porro autem

$$\mu = \frac{1}{4(n+2)} + \frac{1}{4(n-1)} + \frac{1}{8(n-1)^2}$$

$$\nu = \frac{4(n-1)^2}{4n-1};$$

vnde intelligitur, quemadmodum ex ratione refractionis  $n'$  litterae respondentes  $\varrho', \sigma', \tau', \mu', \nu'$  definiri debeant.

II°. Quod iam ad elementa occasione campi apparentis introducta  $\pi, \pi', \pi''$  etc. attinet, distantiae determinatrices lentium inde eodem modo manent determinatae, vt supra, ita, vt non opus sit, eas formulas hic transcribere; interim tamen meminisse iuuabit; formulas primitiuas, vnde illae sunt natae, quae sunt

$$1^\circ. \frac{3\pi - \Phi}{\Phi} = \frac{Aa}{b}$$

$$2^\circ. \frac{e\pi' - \pi + \Phi}{\Phi} = \frac{ABa}{c}$$

$$3^\circ. \frac{d\pi'' - \pi' + \pi - \Phi}{\Phi} = \frac{ABCa}{d}$$

$$4^\circ. \frac{e\pi''' - \pi'' + \pi' - \pi + \Phi}{\Phi} = \frac{ABCDa}{e}$$

etc.

III.

III. Verum valores litterarum P, Q, R etc. ob diuerfas refractiones mutantur, vt sequitur.

$$P = \frac{\mu}{A^3 a^3} (A + 1) (\lambda (A + 1)^2 + \nu A)$$

$$Q = \frac{\mu'}{B^3 b^3} (B + 1) (\lambda' (B + 1)^2 + \nu' B);$$

$$R = \frac{\mu''}{C^3 c^3} (C + 1) (\lambda'' (C + 1)^2 + \nu'' C);$$

$$S = \frac{\mu'''}{D^3 d^3} (D + 1) (\lambda''' (D + 1)^2 + \nu''' D);$$

etc.

IV. Distantiae lentium etiam manent, vt ante, vna cum semidiamentris aperturarum singularum lentium; at radii binarum facierum vtriusque lentis ita sunt ad praesentem casum accommodandae, vt pro, lente cuius refractione est  $n'$ , ei respondentes litterae  $\rho'$ ,  $\sigma'$ , et  $\tau'$  vsurpari debeant, similique modo etiam pro sequentibus lentibus, quarum refractione litteris  $n''$ ,  $n'''$  etc. indicatur.

His praemissis singula momenta, quae in constructione instrumentorum dioptricorum sunt obseruanda, sequenti modo repraesentabimus.

I. Pro loco oculi seu eius distantia post vltimam lentem = O

habebimus pro singulis lentium numeris sequentes determinaciones:

Tom. I.

Q q

Num.



Num.  
lentium.

- I.  $O = 0 = \alpha + l$ ; ideoque  $\alpha = -l = Aa$
- II.  $O = \frac{Bb\pi}{\pi - \Phi} = \xi + l$ ; ideoque  $\xi = \frac{Bb\pi}{\pi - \Phi} - l = Bb$ ;  $l = \frac{-Bb(\pi - \Phi)}{\pi - \Phi}$
- III.  $O = \frac{Cc\pi'}{\pi' - \pi + \Phi} = \gamma + l$ ; hinc  $\gamma = \frac{Cc\pi'}{\pi' - \pi + \Phi} - l = Cc$ ;  
hinc  $l = -\frac{Cc(\pi' - \pi + \Phi)}{\pi' - \pi + \Phi}$
- IV.  $O = \frac{Dd\pi''}{\pi'' - \pi' + \pi - \Phi} = \delta + l$ ; hinc  $\delta = \frac{Dd\pi''}{\pi'' - \pi' + \pi - \Phi} - l$ ;  
adeoque  $l = -\frac{Dd(\pi'' - \pi' + \pi - \Phi)}{\pi'' - \pi' + \pi - \Phi}$
- V.  $O = \frac{Ee\pi'''}{\pi''' - \pi'' + \pi' - \pi + \Phi} = \varepsilon + l$ ; hinc  $\varepsilon = \frac{Ee\pi'''}{\pi''' - \pi'' + \pi' - \pi + \Phi} - l$ ;  
adeoque  $l = -\frac{Ee(\pi''' - \pi'' + \pi' - \pi + \Phi)}{\pi''' - \pi'' + \pi' - \pi + \Phi}$
- etc.

## II. Pro ratione multiplicationis ad distantiam b. relata

habebimus pro quolibet lentium numero sequentes valores litterae  $m$ .

- |      |                          |                                                                                                                                   |
|------|--------------------------|-----------------------------------------------------------------------------------------------------------------------------------|
| I.   | $m = -A \frac{b}{l}$     | } Pro situ erecto, si<br>$m$ hinc prodierit po-<br>sitivum; pro situ<br>autem inverso, si $m$<br>obtinuerit valorem<br>negativum. |
| II.  | $m = +AB \frac{b}{l}$    |                                                                                                                                   |
| III. | $m = -ABC \frac{b}{l}$   |                                                                                                                                   |
| IV.  | $m = +ABCD \frac{b}{l}$  |                                                                                                                                   |
| V.   | $m = -ABCDE \frac{b}{l}$ |                                                                                                                                   |
- etc.

III.

III. Pro gradu claritatis  $\gamma$ 

ex superioribus liquet, eum semper pari modo exprimi; quantuscunque fuerit lentium numerus; perpetuo enim erit  $\gamma = \frac{bx}{ma}$

## IV. Pro Campo apparente si quidem O habeat valorem positivum,

eius semidiameter  $\Phi$  pro quovis lentium numero sequenti modo definietur:

| Num.<br>lentium. |                                                                    |
|------------------|--------------------------------------------------------------------|
| I.               | $\Phi = \infty$ seu indefinitum                                    |
| II.              | $\Phi = \frac{-\pi b}{m a - b}$                                    |
| III.             | $\Phi = \frac{-\pi + \pi'}{m a - b} \cdot b$                       |
| IV.              | $\Phi = \frac{-\pi + \pi' - \pi''}{m a - b} \cdot b$               |
| V.               | $\Phi = \frac{-\pi + \pi' - \pi'' + \pi'''}{m a - b} \cdot b$ etc. |

## (IV.) Pro Campo apparente,

si distantia O fuerit negatiua, quo casu oculus lenti vltimae immediate adplicari debet,  $\Phi$  definietur ex aequationibus sequentibus.

| Num.<br>lentium. |                                                                             |
|------------------|-----------------------------------------------------------------------------|
| I.               | $\Phi = \infty$ indefinitum                                                 |
| II.              | $\frac{A B a \Phi}{\pi - \Phi} = \omega$                                    |
| III.             | $\frac{A B C a \Phi}{\pi' - \pi + \Phi} = \omega$                           |
| IV.              | $\frac{A B C D a \Phi}{\pi'' - \pi' + \pi - \Phi} = \omega$                 |
| V.               | $\frac{A B C D E a \Phi}{\pi''' - \pi'' + \pi' - \pi + \Phi} = \omega$ etc. |



V. Pro semidiametro confusionis  
habebuntur sequentes expressiones

$$\begin{aligned}
 \text{I.} & \frac{m x^3}{4 a a b} \cdot \frac{\mu(A+1)(\lambda(A+1)^2 + \nu A)}{A^3} \\
 \text{II.} & \frac{m x^3}{4 a a b} \left( \frac{\mu(A+1)(\lambda(A+1)^2 + \nu A)}{A^3} + \frac{\mu'(B+1)(\lambda'(B+1)^2 + \nu' B)}{A^3 B (\mathfrak{B}\pi - \Phi)} \Phi \right) \\
 \text{III.} & \frac{m x^3}{4 a a b} \left\{ \frac{\mu(A+1)(\lambda(A+1)^2 + \nu A)}{A^3} + \frac{\mu'(B+1)(\lambda'(B+1)^2 + \nu' B)}{A^3 B (\mathfrak{B}\pi - \Phi)} \Phi \right. \\
 & \quad \left. + \frac{\mu''(C+1)(\lambda''(C+1)^2 + \nu'' C)}{A^3 B^2 C^2 (\mathfrak{E}\pi' - \pi + \Phi)} \Phi \right\} \\
 \text{IV.} & \frac{m x^3}{4 a a b} \left\{ \frac{\mu(A+1)(\lambda(A+1)^2 + \nu A)}{A^3} + \frac{\mu'(B+1)(\lambda'(B+1)^2 + \nu' B)}{A^3 B (\mathfrak{B}\pi - \Phi)} \Phi \right. \\
 & \quad \left. + \frac{\mu''(C+1)(\lambda''(C+1)^2 + \nu'' C)}{A^3 B^2 C^2 (\mathfrak{E}\pi' - \pi + \Phi)} \Phi + \frac{\mu'''(D+1)(\lambda'''(D+1)^2 + \nu''' D)}{A^3 B^2 C^2 D^2 (\mathfrak{D}\pi'' - \pi' + \pi - \Phi)} \Phi \right\} \\
 \text{V.} & \frac{m x^3}{4 a a b} \left\{ \frac{\mu(A+1)(\lambda(A+1)^2 + \nu A)}{A^3} + \frac{\mu'(B+1)(\lambda'(B+1)^2 + \nu' B)}{A^3 B (\mathfrak{B}\pi - \Phi)} \Phi \right. \\
 & \quad \left. + \frac{\mu''(C+1)(\lambda''(C+1)^2 + \nu'' C)}{A^3 B^2 C^2 (\mathfrak{E}\pi' - \pi + \Phi)} \Phi + \frac{\mu'''(D+1)(\lambda'''(D+1)^2 + \nu''' D)}{A^3 B^2 C^2 D^2 (\mathfrak{D}\pi'' - \pi' + \pi - \Phi)} \Phi \right. \\
 & \quad \left. + \frac{\mu''''(E+1)(\lambda''''(E+1)^2 + \nu'''' E)}{A^3 B^2 C^2 D^2 E^2 (\mathfrak{E}\pi''' - \pi'' + \pi' - \pi + \Phi)} \Phi \right. \\
 & \quad \left. \text{etc.} \right\}
 \end{aligned}$$

VI. Pro tollendo margine colorato,

si distantia O. prodierit positiua, obiectum  
fine margine colorato apparebit, satisfaciendo sequen-  
tibus aequationibus:

|          |  |                                                                                                                                                                                              |
|----------|--|----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|
| Num.     |  |                                                                                                                                                                                              |
| lentium. |  |                                                                                                                                                                                              |
| I.       |  | $\circ = \circ$                                                                                                                                                                              |
| II.      |  | $\circ = \frac{b d n'}{n' - 1} \cdot \frac{\pi}{A a \Phi}$                                                                                                                                   |
| III.     |  | $\circ = \frac{b d n'}{n' - 1} \cdot \frac{\pi}{A a \Phi} + \frac{c d n''}{n'' - 1} \cdot \frac{\pi'}{A B a \Phi}$                                                                           |
| VI.      |  | $\circ = \frac{b d n'}{n' - 1} \cdot \frac{\pi}{A a \Phi} + \frac{c d n''}{n'' - 1} \cdot \frac{\pi'}{A B a \Phi} + \frac{d d n'''}{n''' - 1} \cdot \frac{\pi''}{A B C a \Phi} \text{ etc.}$ |

(VI.)

(VI). Pro tollendo margine colorato.

Sin autem distantia  $O$  prodierit negatiua, margo coloratus euanescet, si sequentibus aequationibus satisfiat

$$I. \quad O = 0$$

$$II. \quad O = \frac{a dn}{n-1} (A+1) B \pi$$

$$III. \quad O = \frac{a dn}{n-1} (A+1) B C \pi' + \frac{b dn' (B+1) C \pi' - (C+1) \pi}{n'-1 A}$$

$$IV. \quad O = \frac{a dn}{n-1} (A+1) B C D \pi'' + \frac{b dn' (B+1) C D \pi'' - (D+1) \pi}{n'-1 A} + \frac{c dn'' (C+1) D \pi'' - (D+1) \pi'}{n''-1 A B}$$

$$V. \quad O = \frac{a dn}{n-1} (A+1) B C D E \pi''' + \frac{b dn' (B+1) C D E \pi''' - (E+1) \pi}{n'-1 A} + \frac{c dn'' (C+1) D E \pi''' - (E+1) \pi'}{n''-1 A B} + \frac{d dn''' (D+1) E \pi''' - (E+1) \pi''}{n'''-1 A B C}$$

etc.

VII. Pro tollenda confusione omni

insuper sequentibus aequationibus est satisfaciendum

$$I. \quad O = \frac{a dn}{n-1} \cdot \frac{A+1}{A}$$

$$II. \quad O = \frac{a dn}{n-1} \cdot \frac{A+1}{A} + \frac{b dn'}{n'-1} \cdot \frac{B+1}{A^2 B}$$

$$III. \quad O = \frac{a dn}{n-1} \cdot \frac{A+1}{A} + \frac{b dn'}{n'-1} \cdot \frac{B+1}{A^2 B} + \frac{c dn''}{n''-1} \cdot \frac{C+1}{A^2 B^2 C}$$

$$IV. \quad O = \frac{a dn}{n-1} \cdot \frac{A+1}{A} + \frac{b dn'}{n'-1} \cdot \frac{B+1}{A^2 B} + \frac{c dn''}{n''-1} \cdot \frac{C+1}{A^2 B^2 C} + \frac{d dn'''}{n'''-1} \cdot \frac{D+1}{A^2 B^2 C^2 D}$$

etc.

Qq 3

Distan-



Distantiae autem lentium et semidiametri aperturarum perinde definiuntur, ac supra; tantum notetur insuper, formulas pro semidiametro confusionis exhibitas, nisi penitus ad nihilum redigi queant, aequales poni debere formulae  $\frac{1}{4k^3}$ , existente circiter  $k = 40$  (§. 193) vel adhuc minore, prout circumstantiae postulauerint.

### Supplementum VIII.

#### De Lentibus obiectiuis perfectis.

In Cap. III. nullas alias lentes obiectiuas commemorauimus, nisi quae ex eadem vitri specie et ex principio minimi sunt paratae, neque tales lentes ex diuersis vitri speciebus conficere conuenit, quoniam eae ad confusionem ex diuersa radiorum refrangibilitate oriundam, tollendam sunt ineptae. Cum igitur ista confusio aliter tolli nequeat, nisi diuersae materiae refringentes adhibeantur; hic adcuratius perpendamus, quemadmodum ex diuersis mediis diaphanis eiusmodi lentes construi queant, in quibus non solum prior confusio ex apertura lentium, sed etiam posterior ex diuersa radiorum refrangibilitate oriunda ad nihilum redigi queat, cuiusmodi lentes compositas merito *perfectas* appellare licebit.

I. Primo igitur examinemus, quomodo posteriori confusio sit occurrendum, quando quotcunque lentes inuicem iungantur, ita, vt earum distantiae quasi euanescent. Quem in finem considerentur formulae  
supra

supra datae tam pro margine colorato vitando, quam pro confusione penitus tollenda N°. VI, VII. At ex aequatione N°. VI. patet, marginem coloratum euanescere, si litterae  $\pi$ ,  $\pi'$ ,  $\pi''$  etc. sint  $=0$ , quod quidem sponte euenit, si lentes immediate coniungantur; quare in genere nullus margo coloratus est metuendus, statim atque interualla omnia lentium euanescent.

II. Vt autem aequationibus N°. VII. satisfiat, pro casu duarum lentium immediate iunctarum habemus

$$0 = \frac{a dn}{n-1} \cdot \frac{A+r}{A} + \frac{b dn'}{n'-1} \cdot \frac{B+r}{A^2 B}$$

$$0 = \frac{a dn}{n-1} \cdot \frac{a+\alpha}{\alpha} + \frac{b dn'}{n'-1} \cdot \frac{(b+\epsilon)a^2}{\alpha^2 \epsilon}$$

$$0 = \frac{dn}{n-1} \cdot \frac{a+\alpha}{a\alpha} + \frac{b dn'}{n'-1} \cdot \frac{b+\epsilon}{\alpha^2 \epsilon}; \text{ seu ob } \alpha = -b$$

$$0 = \frac{dn}{n-1} \cdot \frac{a+\alpha}{a\alpha} + \frac{dn'}{n'-1} \cdot \frac{b+\epsilon}{b\epsilon},$$

$$0 = \frac{dn}{n-1} \cdot \left(\frac{1}{a} + \frac{1}{\alpha}\right) + \frac{dn'}{n'-1} \cdot \left(\frac{1}{b} + \frac{1}{\epsilon}\right)$$

vbi duo hi coëfficientes denotant reciproca distantiarum focalium primae et secundae lentis.

III. Nunc insuper effici debet, vt prior confusio ab apertura lentium oriunda ad nihilum redigatur; vbi obseruare debemus, hoc casu esse debere  $\alpha+b=0$ ;  $\epsilon+c=0$ ;  $\gamma+d=0$  etc. pro numero lentium, quas coniungere velimus. Vnde semidiameter  
confu-



confusionis ex lentibus iunctis ortae euanesceat, si red-  
datur

$$0 = \frac{\mu(A+1)(\lambda(A+1)^2 + \nu A)}{A^3} - \frac{\mu'(B+1)(\lambda'(B+1)^2 + \nu' B)}{A^3 B^3} \\ + \frac{\mu''(C+1)(\lambda''(C+1)^2 + \nu'' C)}{A^3 B^3 C^3} \text{ etc.}$$

quoniam supra iam vidimus, fore  $\pi = 0$ ,  $\pi' = 0$  etc.  
atque hinc pro quouis numero lentium inuicem iun-  
gendarum constructionem lentium perfectarum petere  
debemus; quas inuestigationes hic suscipiamus.

De lentibus perfectis ex duabus lentibus  
compositis.

Quoniam hoc casu distantiae determinatrices  
ipsius lentis compositae sunt  $a$  et  $e$ ; statuamus

$$\frac{1}{a} + \frac{1}{a} = f\left(\frac{1}{a} + \frac{1}{e}\right) \text{ et } \frac{1}{b} + \frac{1}{e} = g\left(\frac{1}{a} + \frac{1}{e}\right) \text{ siue}$$

$$\frac{a}{a+e} = f\left(\frac{a}{a+e}\right); \quad \frac{b}{b+e} = g \cdot \frac{a}{a+e};$$

vnde pro prima conditione implenda sequitur:

$$0 = \frac{dn}{n-1} \cdot f + \frac{dn'}{n'-1} \cdot g. \text{ Porro ob } \frac{1}{a} = \frac{f-1}{a} + \frac{f}{e} \text{ et } \frac{1}{b} = \frac{g}{a} + \frac{g-1}{e} \text{ et}$$

$$a+b=0, \text{ erit } \frac{f+g-1}{a} + \frac{f+g-1}{e} = 0, \text{ seu } f+g=1.$$

Sit breuitatis gratia  $\frac{dn}{n-1} = \zeta$  et  $\frac{dn'}{n'-1} = \eta$ , vt sit  $\zeta f + \eta g = 0$ ;

ex qua aequatione cum altera  $f+g=1$  coniuncta  
vtraque littera  $f$  et  $g$  determinabitur, ita, vt sit

$$f = \frac{\eta}{\eta-\zeta} \text{ et } g = -\frac{\zeta}{\eta-\zeta}$$

deinde

deinde cum sit

$$A = \frac{\alpha}{a} = \frac{e}{e(f-1) + af} = \frac{e}{af - eg} = \frac{(\eta - \zeta)e}{a\eta + e\zeta} \text{ et}$$

$$B = \frac{e}{b} = \frac{ge + (g-1)a}{a} = \frac{ge - fa}{a} = \frac{-\zeta e - \eta a}{(1-\zeta)a};$$

$$\text{ergo } A+1 = \frac{\eta(a+e)}{a\eta + e\zeta} \text{ et } B+1 = \frac{-\zeta(a+e)}{a(\eta - \zeta)};$$

unde pro confusione priori tollenda habebimus hanc aequationem

$$0 = \frac{\mu\lambda\eta^3(a+e)^3}{(\eta a + \zeta e)^3} + \frac{\mu\nu\eta(\eta - \zeta)(a+e)e}{(\eta a + \zeta e)^2} \\ - \frac{\mu'\lambda'\zeta^3(a+e)^3}{(\eta a + \zeta e)^3} + \frac{\mu'\nu'\zeta(\eta - \zeta)(a+e)a}{(\eta a + \zeta e)^2}$$

quae reducitur ad hanc

$$0 = \frac{(a+e)^2}{\eta a + \zeta e} (\mu\lambda\eta^3 - \mu'\lambda'\zeta^3) + (\eta - \zeta)(\mu\nu\eta e + \mu'\nu'\zeta a)$$

Breuius etiam ex §. 214 deducitur haec aequatio, cui satisfieri debet  $P+Q=0$  ob  $a=-b$ , existente pro hoc casu

$$P = \mu f\left(\frac{1}{a} + \frac{1}{e}\right) \left(\lambda\left(\frac{1}{a} + \frac{1}{e}\right)^2 + \frac{1}{a}\left(\frac{f-1}{a} + \frac{f}{e}\right)\right) \text{ et}$$

$$Q = \mu'g\left(\frac{1}{a} + \frac{1}{e}\right) \left(\lambda'g^2\left(\frac{1}{a} + \frac{1}{e}\right)^2 + \frac{1}{e}\left(\frac{g}{a} + \frac{g-1}{e}\right)\right).$$

ex qua aequatione siue  $\lambda$ , siue  $\lambda'$  inueniri potest, dummodo hae litterae non minores prodeant unitate; quamobrem quosdam eiusmodi casus euoluamus et ad praxin accommodemus.

IV. Tria autem mediorum diaphanorum genera hic potissimum contemplabimur, quorum primum sit aqua pluuiæ, pro qua est  $n = \frac{4}{3} = 1,3333$ ; secundum vitrum ordinarium, pro quo est  $n = \frac{31}{20} = 1,5500$  et

Tom. I.

R. r.

tertium



tertium vitrum Crystallinum, anglice Flint-Glass, pro quo est  $n = \frac{9}{8} = 1,6000$ ; factoque calculo inuenimus, pro litteris graecis inde pendentibus sequentes valores:

|          | Aqua       | Vitrum commune | Vitrum Crystall. |
|----------|------------|----------------|------------------|
| $n$      | $= 1,3333$ | $= 1,5500$     | $= 1,6000$       |
| $\rho$   | $= 0,8000$ | $= 0,1907$     | $= 0,1111$       |
| $\sigma$ | $= 2,2000$ | $= 1,6274$     | $= 1,5555$       |
| $\tau$   | $= 1,2490$ | $= 0,9051$     | $= 0,8606$       |
| $\mu$    | $= 1,9500$ | $= 0,9381$     | $= 0,8333$       |
| $\nu$    | $= 0,1025$ | $= 0,2326$     | $= 0,2666$       |
| $\mu\nu$ | $= 0,1999$ | $= 0,2184$     | $= 0,2221$       |

V. Quod autem ad valores differentialium  $dn$  attinet, optandum esset, vt ii per experimenta accuratissime definirentur; interim tamen duae sequentes hypotheses ex theoria deductae perpendi merentur, quandoquidem *Newtoniana*, ex qua  $dn$  ipsi  $n-1$  foret proportionale, ad confusionem tollendam esset inepta. Prima autem hypothesis, quam dudum proposueram, facit  $dn$  ipsi  $n \cdot \log. n$  proportionale; altera autem ex theoria attractionis deducta dat  $dn$  ipsi  $\frac{nn-1}{n}$  proportionale; atque hinc valores formulae  $\frac{dn}{n-1}$  pro vtraque hypothesis euoluamus:

I<sup>ma</sup> Hypothesis  $dn = n \log. n$

|                  | Aqua     | Vitrum commune | Vitrum Crystall. |
|------------------|----------|----------------|------------------|
| $dn$             | $0,1665$ | $0,2950$       | $0,3265$         |
| $\frac{dn}{n-1}$ | $0,4995$ | $0,5364$       | $0,5442$         |

II<sup>da</sup>

II<sup>da</sup> Hypothesis  $dn = \frac{n n - 1}{n}$

|                  | Aqua    | Vitrum<br>commune | Vitrum<br>Cryſtall. |
|------------------|---------|-------------------|---------------------|
| $\frac{dn}{n-1}$ | 1, 7500 | 1, 6452           | 1, 6250             |

VI. Verum neutra harum hypothesium cum iis experimentis, quae Celeb. *Dollondus* circa prismata instituit, subsistere potest. Vfus nempe est duobus prisinatibus, quorum alterum ex vitro communi, erat paratum, angulo refringente existente triginta graduum; alterum autem ex vitro chryſtallino confectum continebat angulum  $19^\circ$ ; hisque duobus prismatibus inuerſe iunctis obseruauit, in spectro inde ad parietem proiecto iridis colores nullos adparere; vnde conſluit, (si  $n$  pro ratione refractionis vitri communis,  $n'$  vero pro refractione vitri chryſtallini assumatur) fore  $dn : dn' = 2 : 3$ , quam rationem vocat rationem dispersionis; quo experimento admiſſo ſequitur,  $\frac{dn}{n-1} : \frac{dn'}{n'-1} = 8 : 11$ ; ita, vt pro nostro calculo foret  $\zeta = 8$  et  $\eta = 11$ , quae ratio ab allatis hypothesibus enormiter aberrat. Prima enim dat  $\zeta : \eta = 0,5364 : 0,5442$  hoc est proxime, vt 68 : 69. Altera autem praebet  $\zeta : \eta = 1,6452 : 1,6250$  vel proxime, vti 81 : 80, vnde patet, hypothesin posteriorem non posse subsistere, quia ex ea ſequeretur  $\zeta > \eta$ ; prior vero tantum adhuc discrepat, vt cum experimento nequitiam conciliari possit; ex quo merito ancipites haeremus, quomodo nos in calculo hoc gerere debe-

R r 2

amus.



amus. Interim experimentum *Dollondi* nonnullis adhuc difficultatibus premi videtur, quas hic ob oculos ponere visum est.

### Digressio de refractione vitri chrySTALLINI secundum experimenta *Dollondi*.

Tab. III.  
Fig. 17.

I. Quo clarius theoriam, cui haec experimenta innituntur, ob oculos ponamus, rem aliquanto generalius, quam *Dollondus* fecit, consideremus. Repraesentet ergo triangulum  $ABC$  prisma triangulare ex vitro communi paratum, cuius angulus ad  $C$ , quem *Dollondus*  $30^\circ$  assumit, vocetur  $=f$ ; alterum vero prisma ex vitro chrySTALLINO factum  $DBE$  ita inuerse priori sit adplicatum, ut sit angulus  $CBD=g$ , ipseque angulus huius prismatis  $DBE=b$  ubi in experimento examinando erat  $g=0$  et  $b=19^\circ$  tum vero refractione radiorum mediorum in prismate priori sit ut  $m:1$ , in posteriori vero ut  $n:1$ , ita ut facultates radios dispergendi nobis per differentialia  $dm$  et  $dn$  repraesententur. Quibus positis *Dollondus* se observasse affirmat, si radius solis per haec duo prismata transmittatur, spectrum inde in parietem proiectum nullis coloribus inquinari. Vnde conclusit, fore  $dm:dn=2:3$ , quam conclusionem adcuratius examinemus.

II. Hic primum notari convenit, angulum, sub quo radius solis  $Op$  per faciem prioris lentis  $AC$  immittitur, non definiri, quasi perinde sit, sub quocunque

cunque angulo haec immersio fiat, id quod pleniorē expositionem postulat. Ponamus igitur angulum  $OPA = a$ , qui est complementum anguli incidentiae et anguli sequentes quavis refractione orti sint  $p, q, r, t, u, v, e$  et radius iterum emergens  $VZ$ ; ex quo statim perspicitur, fore  $q = p + f$ ,  $r = g + t$  et  $u = v + b$ ; tum vero ex ratione refractionis intelligimus I)  $\text{Cof. } a = m \cdot \text{Cof. } p$ . II)  $\text{Cof. } r = m \cdot \text{Cof. } q$ . III)  $\text{Cof. } t = n \cdot \text{Cof. } u$ . IV)  $\text{Cof. } e = n \cdot \text{Cof. } v$ .

III. Quia iam hoc transitu nulla colorum dispersio observata esse perhibetur, necesse est, ut in emersione  $V$  omnes radii colorati redditi sint inter se paralleli, ideoque angulus  $e$  constans perinde ac primus angulus  $a$ , etiamsi rationes refractionis  $m$  et  $n$  suis differentialibus  $dm$  et  $dn$  augeantur; quippe qua mutatione tantum anguli  $p, q, r, t, u, v$ , fiunt variabiles.

IV. Nunc igitur primo hanc angulorum mutationem in primo primate euoluamus, differentiando scilicet aequationes supra notatas; unde erit

$$dq = dp \text{ deinde } 0 = dm \cdot \text{cos. } p - m dp \cdot \text{sin. } p \text{ seu}$$

$$dp = \frac{dm \cdot \text{cos. } p}{m \cdot \text{sin. } p} = dq;$$

$$\text{tum vero } -dr \cdot \text{sin. } r = dm \cdot \text{cos. } q - m dq \cdot \text{sin. } q \text{ seu}$$

$$-dr \cdot \text{sin. } r = dm \cdot \text{cos. } q - \frac{dm \cdot \text{cos. } p \cdot \text{sin. } q}{\text{sin. } p}$$

$$= \frac{dm}{\text{sin. } p} (\text{cos. } q \cdot \text{sin. } p - \text{cos. } p \cdot \text{sin. } q) = \frac{dm \cdot \text{sin. } (p - q)}{\text{sin. } p}$$

$$\text{hincque tandem } dr = \frac{dm \cdot \text{sin. } f}{\text{sin. } p \cdot \text{sin. } r} \text{ ob } p - q = -f$$

R r 3

Simili



Simili modo euoluantur refractiones per alterum prisma, incipiendo ab angulo constante  $e$  vbi aequatio

$$\cos. e = n. \cos. v \text{ dat } 0 = dn. \cos. v - n dv. \sin. v \text{ siue}$$

$$dv = \frac{dn \cos. v}{n \sin. v} = du \text{ ob } u = v + b;$$

deinde vero aequatio  $\cos. t = n. \cos. u$  dabit

$$-dt. \sin. t = dn. \cos. u - n du. \sin. u = dn. \cos. u - \frac{dn \cos. v \sin. u}{\sin. v} = \frac{dn}{\sin. v}$$

$$(\cos. u. \sin. v - \cos. v. \sin. u) = \frac{dn}{\sin. v} \sin. (v - u)$$

ergo ob  $u - v = b$  habebitur  $dt = + \frac{dn \sin. b}{\sin. v. \sin. t}$ ; Cum autem sit

$$r = g + t, \text{ erit } dr = dt; \text{ ergo } \frac{dm \sin. f}{\sin. r. \sin. r} = \frac{dn \sin. b}{\sin. v. \sin. t}$$

ex qua aequatione haec relatio inter  $dm$  et  $dn$  concluditur

$$dm : dn = \frac{\sin. b}{\sin. v. \sin. t} : \frac{\sin. f}{\sin. r. \sin. r}$$

$$dm : dn = \frac{\sin. r. \sin. r}{\sin. v} : \frac{\sin. v. \sin. t}{\sin. b}$$

quae proportio generaliter locum habet, quoties angulus ad  $e$  prodit constans

V. Iam hanc conclusionem ad *Dollondi* experimentum adplicemus et quia ibi est  $g = 0$ , hincque

$r = t$  erit ratio inter vtramque dispersionem seu

$$dm : dn = \frac{\sin. p}{\sin. r} : \frac{\sin. v}{\sin. b} = \frac{CQ}{PQ} : \frac{BU}{VU};$$

ex qua aequatione iam satis clare elucet, rationem  $\frac{dm}{dn}$  ab angulo incidentiae  $a$  neutiquam esse independentem, vti *Dollondus* supposuisse videtur; ita, vt et si diffu-

diffusio colorum pro certo quodam angulo  $a$  euanes-  
cens sit deprehensa, hanc conclusionem neutiquam ad  
omnes angulos incidentiae extendere liceat. Quoniam  
igitur hic angulus a *Dollondo* non est assignatus, ex  
hoc experimento nihil certi determinari poterit.

VI. Vt autem hic nullum dubium relinquatur,  
aliquot casus praecipuos pro angulo  $a$  calculo euol-  
vamus, vt adpareat, quanta diuersitas inde in ratio-  
nem  $\frac{dm}{dn}$  ingrediatur, dum tamen plus vna vera esse  
nequit. Quare cum *Dollondus* inuenisset pro vtraque  
vitri specie a se adhibita I.  $m=1,53$  et II)  $n=1,58$ ,  
tum vero angulos  $f=30^\circ$  et  $b=19^\circ$   
sequentia exempla hinc expediamus:

### Exemplum I.

Angulum  $a$  ita constitutum concipiamus, vt  
fiat  $q$  angulus rectus, ideoque etiam  $r$ ,  $t$  et  $u$  recti;  
erit  $\sin. p = \cos. f$  et  $\sin. v = \cos. b$ , ex quo  
coligitur  $dm:dn = \cotang. f: \cotang. b$ ; siue  
 $\frac{dm}{dn} = 0,59639$ , quae iam minor est, quam  $\frac{2}{3}$

### Exemplum II.

Sit  $a=90^\circ$ , qui forte est ipse casus *Dollondi*;  
atque hinc inuenientur sequentes anguli  $p=90^\circ$ ;  
 $q=120^\circ$   $r=139^\circ, 53'$   $t=118^\circ 57'$   $v=99^\circ 57'$   
 $e=105^\circ 50'$ . Vnde sequitur  $\frac{dm}{dn} = \frac{\sin. p \cdot \sin. b}{\sin. r \cdot \sin. v} = 0,6615$ .  
ideoque proxime  $dm:dn=2:3$ , vt habet *Dollondus*.

Cete-



Ceterum hinc iam evidens est, hanc rationem pendere etiam ab angulo incidentiae, quae tamen in hoc experimento non est commemorata.

Si ergo hanc rationem cum *Dollondo* assumamus, fiet porro  $\frac{d\ m}{m-1} : \frac{d\ n}{n-1} = 116 : 159 = 7 : 10$ , ita, ut futurum sit pro lentibus obiectiuis duplicatis, quibus *Dollondus* utitur,  $\zeta = 7$ ,  $\eta = 10$ , quarum constructionem hic ex nostris principiis inuestigemus.

**Lens Obiectiua Clariss. *Dollondi* duplicata prior, anteriore lente ex vitro communi, posteriore vero ex vitro chrystallino parata.**

Cum igitur sit  $n = 1,53$  et  $n' = 1,58$ ,  $\zeta = 7$ ,  $\eta = 10$ , erit statim  $f = \frac{\eta}{\eta - \zeta} = \frac{10}{3}$  et  $g = \frac{-\zeta}{\eta - \zeta} = -\frac{7}{3}$ ;

unde distantiae determinatrices utriusque lentis erunt

Pro prima  $a$ , et  $\alpha = \frac{(\eta - \zeta) a g}{\eta a + \zeta g}$

Pro posteriore  $b$  et  $\beta = \frac{-(\eta - \zeta) a g}{\eta a + \zeta g} = -\alpha$ .

Dum ipsius lentis duplicatae distantiae determinatrices sunt  $a$  et  $g$ .

Tantum igitur restat, ut pro his binis lentibus numeri arbitrarii  $\lambda$  et  $\lambda'$  definiantur; et quoniam quaestio circa lentes obiectiuas versatur, statuamus statim  $a = \infty$ , ut adhuc resolui debeat haec aequatio

$$0 = \mu \lambda \eta^3 - \mu' \lambda' \zeta^3 + \eta(\eta - \zeta) \zeta \mu' \nu'.$$

seu  $\mu' \lambda' \zeta^3 = \mu \lambda \eta^3 + \eta \zeta (\eta - \zeta) \mu' \nu'$ . quae ob

$\zeta = 7$  et  $\eta = 10$  abit in hanc

$$343. \mu' \lambda' = 1000. \mu \lambda + 210. \mu' \nu'.$$

II.

II. Vt igitur *Dollondum* sequamur, pro his valoribus  $n = 1,53$  et  $n' = 1,58$  litteras inde derivatas quaerere debemus.

Eruitur autem ex formulis superioribus

$$\begin{array}{rcl} \mu = 0,9875 & \mu' = 0,8724 & \\ l\mu = 9.9945449. & l\mu' = 9.9407397 & \\ \hline \nu = 0,2194 & \nu' = 0,2529 & \\ l\nu = 9.3413418 & l\nu' = 9.4030044 & \\ \hline l.\mu\nu = 9.3358867. & l.\mu'\nu' = 9.3437441. & \end{array}$$

Hinc coëfficientes trium terminorum nostrae aequationis computentur.

|                       |                       |                           |
|-----------------------|-----------------------|---------------------------|
| $l. 343 = 2,5352941$  | $l. 1000 = 3,0000000$ | $l. 210 = 2,3222193$      |
| $l. \mu' = 9.9407397$ | $l. \mu = 9.9945449$  | $l. \mu'\nu' = 9.3437441$ |
| $2.4760338$           | $2.9945449$           | $1.6659634$               |
| Subtrah. $- - -$      | $2.4760338$           | $2.4760338$               |
|                       | $0.5185111$           | $9.1899296$               |
|                       | $3,3000$              | $0,1548$                  |

ita, vt fit

$$\lambda' = 3,3000 \lambda + 0,1548$$

quare sumto

$$\lambda = 1, \text{ fit } \lambda' = 3,4548; \text{ ergo}$$

$$\lambda' - 1 = 2,4548, \text{ et } \log. \sqrt{\lambda' - 1} = 0,1950080.$$

III. Iam ad radios facierum harum lentium definiendos, cum fit

$$a = \infty \text{ et } a = \frac{(\eta - \zeta)e}{\eta} \text{ et } b = -\frac{(\eta - \zeta)e}{\eta}$$

Tom. I.

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erit



erit pro lente priore radius faciei

$$\text{anterioris} = \frac{\alpha}{\sigma} = \frac{36}{10. \sigma}; \text{posterioris} = \frac{\alpha}{\rho} = \frac{36}{10. \rho}.$$

Pro lente posteriore radius faciei

$$\text{anter.} = \frac{-36}{10 \rho' - 3 \sigma' \pm 7 \tau' \sqrt{\lambda' - 1}}$$

$$\text{poster.} = \frac{-36}{10 \sigma' - 3 \rho' \pm 7 \tau' \sqrt{\lambda' - 1}}$$

quare pro duplici refractione  $n=1,53$  et  $n'=1,58$  valores litterarum  $\rho, \sigma, \tau$  quaeri oportet, qui ita se habebunt

|                   |                    |
|-------------------|--------------------|
| $n = 1,53$        | $n' = 1,58$        |
| $\rho = 0,2266$   | $\rho' = 0,1413$   |
| $\sigma = 1,6602$ | $\sigma' = 1,5827$ |
| $\tau = 0,9251$   | $\tau' = 0,8775$   |

IV. Cum igitur  $\xi$  denotet distantiam focalem ipsius lentis duplicatae, quam deinceps littera P indicabimus, pro lente priori ex vitro communi subviridi, cuius refractione est  $n=1,53$ , reperiemus vtramque faciem; erit radius faciei.

Pro prima lente

$$\text{anterioris} = 0,1807. P; \text{posterioris} = 1,3239. P.$$

Pro altera ex vitro chrySTALLINO, cuius refractione  $n=1,58$  facies ita erunt comparatae.

Pro secunda lente

$$\text{anterioris} = \frac{-3P}{-3,3351 \pm 5.6240}$$

$$\text{posterioris} = \frac{-3P}{15,4031 \pm 5.6240}$$

quia

quia nunc radios curvaturae minores euitari conuenit, signorum ambiguum ea sunt sumenda, quae denominatores minores producant; id quod fit, si signa superiora valeant; hinc ergo obtinebimus sequentes determinationes:

$$\text{radius faciei anterioris} = \frac{-3P}{6.2889} = 0.4770P$$

$$\text{posterioris} = \frac{-3P}{5.7791} = -0.5191P$$

quae lens obiectiua capax est aperturae, cuius semidiameter est  $= 0.0452P$ .

Lens obiectiua duplicata altera, anteriore lente ex vitro chrySTALLINO, posteriori ex vitro communi parata.

I. Cum igitur hic sit

$$n = 1.58 \text{ et } n' = 1.53 \text{ erit } \zeta = 10 \text{ et } \eta = 7; \text{ hinc}$$

$$f = \frac{\eta}{\eta - \zeta} = -\frac{7}{3}. \text{ et}$$

$$g = -\frac{\zeta}{\eta - \zeta} = \frac{10}{3}; \text{ Hincque}$$

$$a = \frac{-3. a g}{7a + 10g} = -\frac{3g}{7} \text{ ob}$$

$$a = \infty, \text{ et } b = \frac{3g}{7}.$$

Confusio autem primi generis vt euanescat, fieri debet

$$0 = 343 \mu \lambda - 1000 \mu' \lambda' - 210 \mu' \nu' \text{ seu}$$

$$343 \mu \lambda = 1000 \mu' \lambda' + 210 \mu' \nu'.$$

SS 2

II.



II. Huius aequationis terni coëfficientes iam quaerantur:

|                      |                       |                            |
|----------------------|-----------------------|----------------------------|
| $l. 343 = 2,5352941$ | $l. 1000 = 3,0000000$ | $l. 210 = 2,3222193$       |
| $l. \mu = 9,9407397$ | $l. \mu' = 9,9945449$ | $l. \mu' \nu' = 9,3358867$ |
| <hr/>                | <hr/>                 | <hr/>                      |
| 2.4760338            | 2.9945449             | 1.6581060                  |
|                      | 2.4760338             | 2.4760338                  |
|                      | <hr/>                 | <hr/>                      |
|                      | 0.5185111             | 9.1820722                  |
|                      | 3,3000                | 0,1520                     |
|                      | <hr/>                 | <hr/>                      |

ita, vt fit

$$\lambda = 3,3000 \lambda' + 0,1520; \text{ quare sumto}$$

$$\lambda' = 1 \text{ fiet } \lambda = 3,4520 \text{ ac propterea}$$

$$\lambda - 1 = 2,4520 \text{ et } \log. V(\lambda - 1) = 0,3895205.$$

III. Cum nunc fit

$$a = \infty \text{ et } \alpha = \frac{-3g}{7} \text{ et } b = \frac{3g}{7}$$

erunt radii facierum vtriusque lentis

Pro lente priore radius faciei

$$\text{anterioris} = \frac{-3g}{7(\sigma \pm \tau \sqrt{\lambda - 1})}; \text{ poster.} = \frac{-3g}{7g \pm 7\tau \sqrt{\lambda - 1}}$$

Pro lente posteriore radius faciei

$$\text{anterioris} = \frac{3g}{7g' + 3\sigma'}; \text{ posterioris} = \frac{3g}{7\sigma' + 3g'}$$

est autem

$$g = 0,1413; \sigma = 1,5827; \tau = 0,8775.$$

$$g' = 0,2266; \sigma' = 1,6602:$$

quibus

quibus substitutis erit pro priore lente radius faciei

$$\text{anter. } \frac{-36}{11.0789 + 5.6187}; \text{ poster. } = \frac{-36}{0.9851 + 5.6187}$$

valeant autem ob rationes supra allegatas signa inferiora, eritque radius faciei

$$\text{anterioris } = \frac{-36}{1.4602} = -2,0545 \text{ P}$$

$$\text{posterioris } = \frac{-36}{10.5078} = -0,2828 \text{ P.}$$

Pro lente autem posteriore erit radius faciei

$$\text{anterioris } = \frac{36}{6,5668} = 0,4568 \text{ P}$$

$$\text{posterioris } = \frac{36}{12,3012} = 0,2438 \text{ P.}$$

Inter quos quatuor radios minimus est 0,2438 P unde haec lens duplicata aperturam admittere potest, cuius semidiameter = 0,0609. P ita, vt haec lens maiorem admittat aperturam, quam praecedens.

## De Lentibus obiectiuis ex tribus lentibus compositis.

I. Sit pro prima lente ratio refractionis =  $n$ , pro secunda =  $n'$ , pro tertia =  $n''$ , deinde distantiae determinatrices primae  $a$  et  $\alpha$ , secundae  $b$  et  $\beta$  tertiae  $c$  et  $\gamma$  cum numeris arbitrariis  $\lambda$ ,  $\lambda'$ ,  $\lambda''$  eritque  $\alpha + b = 0$ , et  $\beta + c = 0$ , eruntque  $a$  et  $\gamma$  di-



stantiae determinatrices ipsius lentis triplicatae. Iam vero statuatur

$$\frac{1}{a} + \frac{1}{\alpha} = f \left( \frac{1}{a} + \frac{1}{\gamma} \right);$$

$$\frac{1}{b} + \frac{1}{\beta} = g \left( \frac{1}{a} + \frac{1}{\gamma} \right);$$

$$\frac{1}{c} + \frac{1}{\gamma} = h \left( \frac{1}{a} + \frac{1}{\gamma} \right);$$

fietque

$$\frac{1}{\alpha} = \frac{f-1}{a} + \frac{f}{\gamma}; \quad \alpha = \frac{a\gamma}{\gamma(f-1) + af} = -b$$

$$\frac{1}{\beta} = \frac{g}{a} + \frac{g}{\gamma} + \frac{f-1}{a} + \frac{f}{\gamma}$$

$$\frac{1}{\beta} = \frac{g+f-1}{a} + \frac{g+f}{\gamma} = -\frac{1}{c}$$

ac proinde

$$\frac{-f-g+1}{a} - \frac{f-g+1}{\gamma} = \frac{b}{a} + \frac{b}{\gamma}$$

unde sequitur fore

$$\frac{-f-g-b+1}{a} - \frac{(f+g+b-1)}{\gamma} = 0$$

$$\text{hinc } f+g+b=1.$$

II. Conditio vero confusionem colorum penitus tollens postulat, ut supra vidimus, hanc aequationem

$$0 = \frac{dn}{n-1} \left( \frac{1}{a} + \frac{1}{\alpha} \right) + \frac{dn'}{n'-1} \left( \frac{1}{b} + \frac{1}{\beta} \right) + \frac{dn''}{n''-1} \left( \frac{1}{c} + \frac{1}{\gamma} \right);$$

quae ponendo

$$\frac{dn}{n-1} = \zeta; \quad \frac{dn'}{n'-1} = \eta; \quad \frac{dn''}{n''-1} = \vartheta$$

$$\text{abit in hanc formam } 0 = \zeta f + \eta g + \vartheta h.$$

III. Ut autem confusio ab apertura pendens destruat, huic aequationi satisfieri debet:

$$0 = \frac{\mu(A+1)(\lambda(A+1)^2 + \nu A)}{A^3} - \frac{\mu'(B+1)(\lambda'(B+1)^2 + \nu' B)}{A^3 B^3} + \frac{\mu''(C+1)(\lambda''(C+1)^2 + \nu'' C)}{A^3 B^3 C^3}.$$

vbi

vbi notetur esse

$$\frac{A+1}{A} = af\left(\frac{1}{a} + \frac{1}{\gamma}\right) \text{ et}$$

$$\frac{B+1}{B} = bg\left(\frac{1}{a} + \frac{1}{\gamma}\right) \text{ et}$$

$$\frac{C+1}{C} = ch\left(\frac{1}{a} + \frac{1}{\gamma}\right);$$

vnde termini priores continentes  $\lambda$  erunt

$$a^3\left(\frac{1}{a} + \frac{1}{\gamma}\right)^3(\mu\lambda f^3 + \mu'\lambda'g^3 + \mu''\lambda''h^3)$$

termini vero tres posteriores fiunt

$$a^3\left(\frac{1}{a} + \frac{1}{\gamma}\right)\left(\frac{\mu v f}{a\alpha} + \frac{\mu' v' g}{b\beta} + \frac{\mu'' v'' h}{c\gamma}\right)$$

adeoque prodit haec aequatio

$$0 = \mu\lambda f^3 + \mu'\lambda'g^3 + \mu''\lambda''h^3$$

$$+ \frac{\text{I}}{\left(\frac{1}{a} + \frac{1}{\gamma}\right)^2} \left( \frac{\mu v f}{a\alpha} + \frac{\mu' v' g}{b\beta} + \frac{\mu'' v'' h}{c\gamma} \right)$$

vbi notandum est, ex praecedentibus esse

$$\frac{1}{a\alpha} = \frac{1}{a}\left(\frac{f-1}{a} + \frac{f}{\gamma}\right);$$

$$\frac{1}{b\beta} = -\left(\frac{f-1}{a} + \frac{f}{\gamma}\right)\left(\frac{g+f-1}{a} + \frac{f+g}{\gamma}\right)$$

$$\frac{1}{c\gamma} = -\frac{1}{\gamma}\left(\frac{f+g-1}{a} + \frac{f+g}{\gamma}\right)$$

IV. Quodsi nunc ponamus  $a=\infty$ , sequentes aequationes habebuntur resoluendae:

$$a = \frac{\gamma}{f} \text{ et } b = -\frac{\gamma}{f}; \frac{1}{\beta} = \frac{g+f}{\gamma} \text{ seu}$$

$$\beta = \frac{\gamma}{f+g} \text{ et } c = \frac{-\gamma}{f+g}.$$

Quibus



Quibus adiungi debent hae aequationes:

$f+g+b=1$ , et  $\zeta f+\eta g+\vartheta b=0$  et  
pro omni confusione tollendâ

$$\text{ob } \frac{1}{a\alpha}=0; \frac{1}{b\beta}=-\frac{f(f+g)}{\gamma^2}; \frac{1}{c\gamma}=-\frac{f+g}{\gamma^2}$$

habebitur haec aequatio

$$0=\mu\lambda f^3+\mu'\lambda'g^3+\mu''\lambda''b^3 \\ -(f+g)(\mu'\nu'fg+\mu''\nu''b).$$

V. Hic quidem assumimus, omnes tres materias, ex quibus hae lentes sunt confectae, esse diuerfas; in praxi autem nulla ratio suadet, vt tres diuerfas vitri species adhibeamus; sed potius primam et tertiam lentem ex eadem materia parari conueniet; ex quo consequimur statim

$$n''=n; \mu''=\mu; \nu''=\nu \text{ et } \vartheta=\zeta,$$

ita, vt fit

$$f+b+\frac{\eta}{\zeta}\cdot g=0;$$

quae ab altera subtracta dat

$$g-\frac{\eta}{\zeta}\cdot g=1; \text{ hincque } g=\frac{\zeta}{\zeta-\eta}$$

unde porro fit

$$f+b=-\frac{\eta}{\zeta-\eta}.$$

Ponamus hunc in finem

$$f-b=\frac{x}{\zeta-\eta} \text{ eritque}$$

$$f=\frac{x-\eta}{2(\zeta-\eta)} \text{ et } b=\frac{-x-\eta}{2(\zeta-\eta)}$$

ex quibus valoribus prodit

$$a = \frac{2\gamma(\zeta-\eta)}{x-\eta}, b = \frac{-2\gamma(\zeta-\eta)}{x-\eta};$$

$$g = \frac{2\gamma(\zeta-\eta)}{x+2\zeta-\eta}; c = \frac{-2\gamma(\zeta-\eta)}{x+2\zeta-\eta}.$$

VI. Aequatio ergo, cui adhuc satisfieri oportet, erit

$$0 = \frac{1}{4}\mu\lambda(x-\eta)^3 + 2\mu'\lambda'\zeta^3 - \frac{1}{4}\mu\lambda''(x+\eta)^3 \\ - \frac{1}{2}(x+2\zeta-\eta)(\mu'\nu'(x-\eta)\zeta - \mu\nu(x+\eta)(\zeta-\eta))$$

ex qua aequatione vel vna litterarum  $\lambda$  vel etiam numerus  $x$  definiri potest; quo posteriori casu hoc commodi consequeremur, vt singulae litterae  $\lambda = 1$  statui possent sicque singulae lentes vtpote ex principio minimi deductae facillime construi possent; hanc autem inuestigationem non aliter, nisi in exemplis particularibus suscipere licebit.

Ponamus igitur

$$\lambda = 1, \lambda' = 1 : \lambda'' = 1 \text{ fietque}$$

$$0 = -\frac{1}{2}\mu(3\eta x^2 + \eta^3) + 2\mu'\zeta^3 - \frac{1}{2}\mu'\nu'\zeta(x^2 + 2(\zeta-\eta)x - \eta(2\zeta-\eta)) \\ + \frac{1}{2}\mu\nu(\zeta-\eta)(x^2 + 2\zeta x + \eta(2\zeta-\eta))$$

quae aequatio reducitur ad hanc

$$x^2(-3\mu\eta - \mu'\nu'\zeta + \mu\nu(\zeta-\eta)) \\ + 2x(\zeta-\eta)\zeta(\mu\nu - \mu'\nu') - \mu\eta^3 + 4\mu'\zeta^3 + \mu'\nu'\eta\zeta(2\zeta-\eta) + \mu\nu.\eta(\zeta-\eta)(2\zeta-\eta) = 0.$$

VII. Adplicemus haec ad illas duas species vitri, quibus *Dollondus* vsus est, ac duo casus euolendi occurrunt. Primus igitur casus esto, quo tam prima, quam tertia lens ex vitro chrystallino con-

Tom. I.

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ficitur;



ficitur; media autem ex vitro communi subuiridi;  
ita, ut sit  $n=n''=1,58$  et  $n'=1,53$ , ideoque  $\zeta=10$   
et  $\eta=7$  tum vero  $\mu=0,8724$ ,  $\nu=0,2529$  et  
 $L\mu\nu=9.3437441$  deinde  $\mu'=0,9875$ ;  $\nu'=0,2194$   
et  $L\mu'\nu'=9.3358867$  unde superior aequatio redu-  
cetur ad hanc formam

$$+19,8255x^2=0,2340x+4177,6081, \text{ siue}$$

$$(x-3)x^2=0,0118x+210,7190$$

cuius resolutio praebet

$$x=0,0059 \pm \sqrt{210,7190}$$

$$=0,0059 \pm 14,5161$$

unde ambo valores ipsius  $x$  sunt

$$x=14,5220$$

$$x=-14,5102$$

Lens obiectiua triplicata perfecta prior  
cuius lens prima et tertia ex vitro chrystal-  
lino, pro quo  $n=1,58$ , media vero ex vitro  
communi subuiridi,  $n=1,53$ , est confecta.

I. Ob duplicem valorem ipsius  $x$  duas etiam  
eiusmodi lentes exhibere poterimus; sit igitur primo  
 $x=14,5220$  et quia est  $\zeta=10$  et  $\eta=7$ , erit  
 $f=1,2537$ ;  $g=3,3333$ ,  $h=-3,5870$

$$\frac{1}{a}=\frac{1,2537}{\gamma}; \frac{1}{b}=\frac{1,2537}{\gamma}; \frac{1}{c}=\frac{4,5870}{\gamma} \text{ et } \frac{1}{c}=\frac{1,5870}{\gamma};$$

quodsi

quodsi iam harum trium lentium radios facierum anteriorum ponamus  $F, F', F''$ , posteriorum vero

$G, G', G''$  ob  $\lambda = \lambda' = \lambda'' = 1$ , habebimus

$$\frac{1}{F} = \frac{\rho}{a} + \frac{\sigma}{\alpha} = \frac{\sigma}{\alpha} = \frac{1.2537 \cdot \sigma}{\gamma}; \quad \frac{1}{G} = \frac{\rho}{\alpha} + \frac{\sigma}{a} = \frac{\rho}{\alpha} = \frac{1.2537 \cdot \rho}{\gamma} \text{ et}$$

$$\frac{1}{F'} = \frac{\rho'}{b} + \frac{\sigma'}{\beta}; \quad \frac{1}{G'} = \frac{\rho'}{\beta} + \frac{\sigma'}{b}$$

et substitutis valoribus

$$\frac{1}{F'} = \frac{-1.2537 \cdot \rho'}{\gamma} + \frac{4.5870 \cdot \sigma'}{\gamma}$$

$$\frac{1}{G'} = \frac{4.5870 \cdot \rho'}{\gamma} - \frac{1.2537 \cdot \sigma'}{\gamma}$$

simili modo pro lente tertia

$$\frac{1}{F''} = \frac{\rho}{c} + \frac{\sigma}{\gamma}; \quad \frac{1}{G''} = \frac{\rho}{\gamma} + \frac{\sigma}{c}$$

$$\frac{1}{F''} = \frac{-4.5870 \cdot \rho + \sigma}{\gamma}; \quad \frac{1}{G''} = \frac{\rho - 4.5870 \cdot \sigma}{\gamma}$$

existente

$$\rho = 0, 1413; \quad \sigma = 1, 5827 \text{ et}$$

$$\rho' = 0, 2266; \quad \sigma' = 1, 6602$$

ac si ipsius lentis triplicatae distantia focalis ponatur  $= P$ , vt hic fit  $\gamma = P$ , reperiuntur valores sequentes

$$\frac{1}{F} = \frac{1,6843}{P}, \text{ siue } F = +0, 5039. P.$$

$$\frac{1}{G} = \frac{0,1771}{P}, \text{ hincque } G = +5, 6450. P.$$

$$\frac{1}{F'} = \frac{7,3312}{P}, \text{ siue } F' = +0, 1364. P.$$

$$\frac{1}{G'} = \frac{-1,0420}{P}, \text{ hincque } G' = -0, 9597. P.$$

$$\frac{1}{F''} = \frac{0,9346}{P}, \text{ hincque } F'' = +1, 0699. P.$$

$$\frac{1}{G''} = \frac{-7,1186}{P}, \text{ hincque } G'' = -0, 1404. P.$$



II. Inter hos sex radios cum minimus sit 0, 1364. P, eius pars quarta = 0, 0341. P dat femidiametrum maximae aperturæ, cuius haec lens triplicata est capax. Contra vero haec lens hac prerogativa gaudet, quod eius constructio ob  $\lambda=1$ ;  $\lambda'=1$ ;  $\lambda''=1$  in praxi minimae difficultati sit obnoxia.

Altera lens triplicata, pro qua pariter, ut ante, lens prima et tertia ex vitro chry-  
stallino, media vero ex vitro communi sub-  
viridi paratur.

I. Sumatur iam

$$x = -14,5102, \text{ ob } \zeta = 10, \text{ et } \eta = 7 \text{ erit}$$

$$f = -3,5850, g = 3,3333; h = +1,2517.$$

$$\frac{1}{\alpha} = -\frac{3.5850}{\gamma}; \frac{1}{b} = \frac{3.5850}{\gamma}$$

$$\frac{1}{g} = -\frac{0.2517}{\gamma}; \frac{1}{c} = +\frac{0.2517}{\gamma}$$

vnde pro radiis singularum facierum prodibunt se-  
quentes formulae:

$$\frac{1}{F} = \frac{\sigma}{\alpha} = -\frac{3.5850 \cdot \sigma}{\gamma}; \frac{1}{G} = \frac{\rho}{\alpha} = -\frac{3.5850 \cdot \rho}{\gamma}$$

$$\frac{1}{F'} = \frac{\rho'}{b} + \frac{\sigma'}{g} = \frac{3.5850 \cdot \rho' - 0.2517 \cdot \sigma'}{\gamma}$$

$$\frac{1}{G'} = \frac{\rho'}{g} + \frac{\sigma'}{b} = \frac{-0.2517 \cdot \rho' + 3.5850 \cdot \sigma'}{\gamma}$$

$$\frac{1}{F''} = \frac{\rho}{c} + \frac{\sigma}{\gamma} = \frac{0.2517 \cdot \rho + \sigma}{\gamma}$$

$$\frac{1}{G''} = \frac{\rho}{\gamma} + \frac{\sigma}{c} = \frac{\rho + 0.2517 \cdot \sigma}{\gamma}$$

vnde

vnde sequentes valores determinati pro his radiis inveniuntur :

$$\frac{1}{F} = -\frac{5,6739}{P} \text{ hincque } F = -0,1762. P.$$

$$\frac{1}{G} = -\frac{0,5065}{P} \text{ hincque } G = -1,9741. P.$$

$$\frac{1}{F'} = \frac{0,3045}{P} \text{ hincque } F' = +2,5349. P.$$

$$\frac{1}{G'} = \frac{5,8048}{P}, \text{ hincque } G' = +0,1696. P.$$

$$\frac{1}{F''} = \frac{1,6182}{P}, \text{ hincque } F'' = +0,6194. P.$$

$$\frac{1}{G''} = \frac{0,5396}{P} \text{ hincque } G'' = +1,8532. P.$$

quorum sex radiorum minimus est 0,1696 P, cuius pars quarta 0,0424. P dat semidiametrum aperturæ maximæ, cuius hæc lens triplicata est capax. Ideoque hæc lens obiectiva præcedenti est anteferenda.

VIII. Secundum §. VII. adhuc alius casus evolui debet, quo lens prima et tertia ex vitro communi subuiri  $n=n''=1,53$ ; media autem ex ChrySTALLINO  $n=1,58$  parari potest; sicque erit  $\zeta=7$ ,  $\eta=10$ . Vnde cum fiat

$$\mu=0,9875, \text{ l. } \mu\nu=9,3358867$$

$$\mu'=0,8724, \text{ l. } \mu'\nu'=9,3437441.$$

aequatio nostra pro confusione tollenda sequentem induet formam :

$$31,8198 x^2 = +0,1638 x + 245,2121 \text{ siue}$$

$$x^2 = 0,0051. x + 7,7063$$

T t 3

cuius



cuius resolutio dat

$$x = 0,0025 \pm 2,7769$$

unde bini valores ipsius  $x$  erunt

$$\text{I. } x = 2,7785$$

$$\text{II. } x = -2,7735.$$

Lens triplicata perfecta prior, cuius  
lens prima et tertia ex vitro communi,  $n=1,53$ .  
media ex chrystallino,  $n=1,58$   
est confecta.

Cum igitur sit

$$x = 2,7785 \text{ et } \zeta = 7, \eta = 10 \text{ erit}$$

$$f = +1,2036; g = -2,3333. h = 2,1297;$$

hincque

$$\frac{1}{a} = \frac{1,2036}{\gamma}, \frac{1}{b} = -\frac{1,2036}{\gamma}, \frac{1}{e} = -\frac{1,1297}{\gamma}, \frac{1}{c} = \frac{1,1297}{\gamma}$$

ex quibus colligitur

$$\frac{1}{F} = \frac{\sigma}{a} = \frac{1,2036 \cdot \sigma}{\gamma}; \quad \frac{1}{G} = \frac{\rho}{a} = \frac{1,2036 \cdot \rho}{\gamma}$$

$$\frac{1}{F'} = \frac{\rho'}{b} + \frac{\sigma'}{e} = \frac{-1,2036 \rho' - 1,1297 \sigma'}{\gamma}$$

$$\frac{1}{G'} = \frac{\rho'}{e} + \frac{\sigma'}{b} = \frac{-1,1297 \rho' - 1,2036 \sigma'}{\gamma}$$

$$\frac{1}{F''} = \frac{\rho}{c} + \frac{\sigma}{\gamma} = \frac{1,1297 \rho + \sigma}{\gamma}$$

$$\frac{1}{G''} = \frac{\sigma}{c} + \frac{\rho}{\gamma} = \frac{1,1297 \sigma + \rho}{\gamma}$$

vbi  $g = 0, 2266$ ;  $\sigma = 1, 6602$

$g' = 0, 1413$ ;  $\sigma' = 1, 5827$

Facto igitur calculo obtinebimus:

$\frac{1}{F} = \frac{1,9982}{P}$  hincque  $F = 0, 5004$ . P.

$\frac{1}{G} = \frac{0,2727}{P}$  hincque  $G = 3, 6665$ . P.

$\frac{1}{F'} = \frac{-1,9580}{P}$  hincque  $F' = -0, 5107$ . P.

$\frac{1}{G'} = \frac{-2,0645}{P}$  hincque  $G' = -0, 4843$ . P.

$\frac{1}{F''} = \frac{1,9161}{P}$  hincque  $F'' = 0, 5219$ . P.

$\frac{1}{G''} = \frac{2,1021}{P}$  hincque  $G'' = 0, 4757$ . P.

inter quos radios cum minimus sit  $0, 4757$  P eius pars quarta  $0, 1189$ . P dat semidiametrum aperturæ, quam hæc lens admittit; ideoque hæc lens duabus præcedentibus est anteferenda.

Lens triplicata perfecta altera, cuius lens prima et tertia ex vitro communi  $n = 1, 53$  media ex chrySTALLINO,  $n = 1, 58$  est parata.

Sumatur iam ex binis ipsius  $x$  valoribus alter negatiuus  $x = -2, 7735$ ; et cum sit  $\zeta = 7$ ,  $\eta = 10$ , inuenitur

$f = +2, 1289$ ;  $g = -2, 3333$   $h = +1, 2044$ ;

hincque

$\frac{1}{a} = \frac{2,1289}{\gamma}$ ;  $\frac{1}{b} = \frac{-2,1289}{\gamma}$

$\frac{1}{c} = \frac{-0,1044}{\gamma}$ ;  $\frac{1}{d} = \frac{0,1044}{\gamma}$



ex quibus consequimur

$$\frac{1}{F} = \frac{\sigma}{\alpha} = \frac{2.1289. \sigma}{\gamma}; \quad \frac{1}{G} = \frac{\rho}{\alpha} = \frac{2.280. \rho}{\gamma}$$

$$\frac{1}{F'} = \frac{\rho'}{b} + \frac{\sigma'}{b} = \frac{-2.1289. \rho' - 0.1044. \sigma'}{\gamma}$$

$$\frac{1}{G'} = \frac{\rho'}{b} + \frac{\sigma'}{b} = \frac{-0.1044. \rho' - 2.1289. \sigma'}{\gamma}$$

$$\frac{1}{F''} = \frac{\rho}{c} + \frac{\sigma}{\gamma} = \frac{0.1044. \rho + \sigma}{\gamma}$$

$$\frac{1}{G''} = \frac{\sigma}{c} + \frac{\rho}{\gamma} = \frac{0.1044. \sigma + \rho}{\gamma}$$

Vnde facto calculo pro radiis facierum singularum lentium sequentes inuenientur radii.

$$\frac{1}{F} = \frac{3.5344}{P}; \text{ seu } F = 0,2829. P.$$

$$\frac{1}{G} = \frac{0.4824}{P}; \text{ seu } G = 2,0729. P.$$

$$\frac{1}{F'} = \frac{-0.4660}{P}; \text{ seu } F' = -2,1459. P.$$

$$\frac{1}{G'} = \frac{-3.3841}{P}; \text{ seu } G' = -0,2955. P.$$

$$\frac{1}{F''} = \frac{1.5838}{P}; \text{ seu } F'' = 0,5938. P.$$

$$\frac{1}{G''} = \frac{0.3099}{P}; \text{ seu } G'' = 2,5006. P.$$

Inter quos radios minimus est 0,2829. P, cuius pars quarta 0,0707. P, dat semidiametrum aperturæ maximæ, cuius hæc lens triplicata est capax.

IX. Hæ quatuor lentes triplicatæ ideo præ ceteris, quas proponere liceret, sunt commendandæ, quod in iis assumimus  $\lambda = 1$ ;  $\lambda' = 1$ ;  $\lambda'' = 1$ . Quam ob causam in praxi facillime construï possunt, ex iisdem vero principiis etiam constructio eiusmodi lentium com-

compositarum deduci posset qua inter binas lentes vitreas aqua aliudue fluidum includeretur, ita, ut lens media ex fluido constaret: tum autem praeter conditiones ante tractatas duae novae essent implendae; scilicet ut radii facierum mediae lentis aequales essent statuendi et contrarii radiis facierum internarum primae ac tertiae lentis. Ob has igitur novas conditiones non amplius liceret numeros  $\lambda$ ,  $\lambda'$  et  $\lambda''$  unitati aequales assumere; qua positione constructioni practicae quam maxime consulitur; deinde vero etiam radii facierum tam parvi prodirent, ut tales lentes nimis exiguam aperturam essent admissurae; quam ob causam operae pretium haud videtur, in earum constructionem adcuratius inquirere.







